# Status of SIMBA

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Global Fit to  $B \rightarrow X_S \gamma$ 

Outlook on  $|V_{ub}|$ 

# Global Fit Approach to $B o X_s \gamma$ and $B o X_u \ell u$

#### Our aim: Provide global fit that combines all available information

- Simultaneously determine
  - Overall normalization:  $|V_{ub}|, \mathcal{B}(B \rightarrow X_s \gamma)$
  - Required input parameters: m<sub>b</sub>, shape function
  - $\Rightarrow$  Same strategy as for  $|V_{cb}|$  (but a bit more complicated now)
- Combine different decay modes and measurements
  - Different  $B o X_s \gamma$  spectra
  - Different  $B \rightarrow X_u \ell \nu$  partial BFs (or even better spectra)
  - Eventually also  $B \to X_s \ell^+ \ell^-$
  - External constraints on  $m_b, \mu_\pi^2$  ( $\lambda_1$ ) (from  $B \to X_c \ell \nu$  or other)

#### What we gain from a global fit

- Minimize uncertainties by making maximal use of all available data
- Consistent treatment of correlated uncertainties (experimental, theoretical, input parameters)

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## Outline

A Little Bit of Theory

2 Global Fit to  $B o X_s \gamma$ 





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# **Requirements on Theory**

#### Model-independent framework for shape function

- Small uncertainties have to be reliable uncertainties
- SF uncertainty should reflect the actual information we have
  - Perturbative constraints (RGE running and perturbative tail)
  - Constraints on moments from  $m_b$ ,  $\mu_{\pi}^2$  ( $\lambda_1$ )
  - ▶ Shape information from  $B \to X_s \gamma$  and  $B \to X_u \ell \nu$  spectra

#### Different measurements probe different phase-space regions

- SF region: large  $E_{\gamma}$ ,  $E_{\ell}$  (near peak/endpoint)
- Local OPE region: small  $E_{\gamma}$ ,  $E_{\ell}$ , large  $q^2$
- Something in between:  $m_X \sim m_D$ , moderately large  $E_\gamma, E_\ell$
- $\Rightarrow$  Combination of optimal theory descriptions for each region

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# Factorized Shape Function

[Ligeti, Stewart, FT (2008)]

$$S(\omega,\mu_\Lambda) = \int\!\mathrm{d}k\,\widehat{C}_0(\omega-k,\mu_\Lambda)\,\widehat{F}(k)$$

#### $\widehat{F}(k)$ nonperturbative part

- Determines peak region
- Fit from data

# $\widehat{C}_0(\omega,\mu_\Lambda)$ perturbative part

 Generates perturbative tail consistent with RGE



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# Factorized Shape Function

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#### $\widehat{F}(k)$ nonperturbative part

- Determines peak region
- Fit from data
- $\Rightarrow \text{ If we know } \widehat{F}(k) \text{ we can compute} \\ S(\omega, \mu_{\Lambda}) \text{ in perturbation theory}$ 
  - Vary  $\mu_{\Lambda}$  to estimate perturbative uncertainty in SF (Here:  $\mu_{\Lambda} = (1.0, 1.3, 1.8)$  GeV
    - + RGE up to  $\mu = 2.5 \, \mathrm{GeV})$

## $\widehat{C}_0(\omega,\mu_\Lambda)$ perturbative part

 Generates perturbative tail consistent with RGE



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# **Different Short Distance Schemes**

 $\widehat{C}$  and  $\widehat{F}$  defined in certain short distance scheme (can use any  $m_b$  scheme)

$$S(\omega) = \int dk C_0^{\text{pole}}(\omega - k) F^{\text{pole}}(k)$$

$$= \int dk C_0^{1S}(\omega - k) F^{1S}(k)$$

$$= \int dk C_0^{\text{kin}}(\omega - k) F^{\text{kin}}(k)$$

$$= \int dk C_0^{\text{SF}}(\omega - k) F^{\text{SF}}(k)$$

$$= \int dk C_0^{\text{SF}}(\omega - k) F^{\text{SF}}(k)$$

$$= \int dk C_0^{\text{SF}}(\omega - k) F^{\text{SF}}(k)$$

Moments of  $\widehat{F}(k)$  given by respective SD parameters  $\widehat{m}_b$ ,  $\widehat{\lambda}_1$ , etc. to all orders in  $\alpha_s$ , e.g.

$$\int \mathrm{d}k \, k^n \, F^{1S\mathrm{i}}(k) = M_n = \begin{cases} 1 & (n=0) \\ m_B - m_b^{1S} & (n=1) \\ -\lambda_1^{\mathrm{i}}/3 + (m_B - m_b^{1S})^2 & (n=2) \end{cases}$$

# Basis Expansion for $\widehat{F}(k)$

Expand  $\widehat{F}(k)$  into suitable orthonormal basis

$$\widehat{F}(\lambda x) = rac{1}{\lambda} iggl[ \sum\limits_{n=0}^{\infty} c_n f_n(x) iggr]^2 
onumber \ \int \mathrm{d} k \, \widehat{F}(k) = \sum\limits_{n=0}^{\infty} c_n^2 = 1$$

Provides model-independent description

Fit for  $\widehat{F}(k)$  by fitting basis coefficients  $c_n$ 

- Experimental uncertainties and correlations are captured in covariance matrix of fitted coefficients c<sub>n</sub>
- ⇒ Allows for *data driven*, reliable estimation of SF uncertainties

#### Basis functions 1 0.5 0 -0.5 -1 0 0.5 1 1.5 2 2.5 3 3.5 4 x



## Residual Basis Dependence from Series Truncation

$$\widehat{F}(\lambda x) = rac{1}{\lambda} iggl[ \sum\limits_{n=0}^N c_n f_n(x) iggr]^2$$

#### In practice, series must be truncated

- Induces residual basis (model) dependence
- Truncation error scales as  $1 \sum c_n^2$

#### Optimal N and $\lambda$ are determined from data

- Choose λ so series converges quickly
- Choose N so truncation error is small compared to exp. uncertainties
- Add more terms with more precise data
- ⇒ Must be careful not to "overtune"





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# Master Formula for Differential Spectra



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# Global Fit to $B o X_s \gamma$

A Little Bit of Theory

Current status of experiment to theory comparison

• HFAG extrapolation down to  $E_{\gamma}^{\rm cut} = 1.6 \, {\rm GeV}$ (adds model dependence)



 ${\cal B}(E_{\gamma}>1.6\,{
m GeV})=(3.55\pm0.24\pm0.09) imes10^{-4}$ 

• Fixed-order NNLO estimate by Misiak et al. (2006)

 ${\cal B}(E_{\gamma}>1.6\,{
m GeV})=(3.15\pm0.23) imes10^{-4}$ 

Sensitivity to new physics lies in normalization, parametrized by  $|V_{tb}V_{ts}^*C_7^{\text{incl}}|$ 

- Most sensitivity in data comes from large  $E_{\gamma}$
- Fit determines both  $|V_{tb}V_{ts}^*C_7^{\text{incl}}|$  (normalization) and  $\widehat{F}(k)$  (shape)
  - Can directly compare  $|V_{tb}V_{ts}^*C_7^{\text{incl}}|$  to its SM prediction
  - Avoids any extrapolation

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# **Theory Inputs**

$$\begin{split} \mathrm{d}\Gamma_s \propto |V_{tb}V_{ts}^*|^2 m_b^2 \bigg\{ \big| C_7^{\mathrm{incl}} \big|^2 \Big[ \big( \widehat{W}_{77}^{\mathrm{sing}} + \widehat{W}_{77}^{\mathrm{nons}} \big) \otimes \widehat{F} + \sum_n W_{77,n} F_n^{\mathrm{subl}} \Big] \\ + \sum_{i,j \neq 7} \Big[ \mathrm{Re}(C_7^{\mathrm{incl}}) 2 C_i \widehat{W}_{7i}^{\mathrm{nons}} + C_i C_j \widehat{W}_{ij}^{\mathrm{nons}} \Big] \otimes \widehat{F} + \cdots \bigg\} \end{split}$$

#### Leading $C_7^2$ contribution

- Included at full NNLL+NNLO (in 1S short-distance scheme)
- 1/mb subleading shape functions absorbed into leading one
  - Have large impact on extracted value of m<sub>b</sub>

#### Contributions from other operators $\sim C_i C_7, C_i C_j$

- Largest effects come from virtual corrections, are absorbed into C<sub>7</sub><sup>incl</sup>
  - Important charm-mass effects only enter SM prediction for C<sub>7</sub><sup>incl</sup>
- Remaining perturbative contributions included at NLO
  - Some NNLO are known, but NLO already have very small effect

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#### Belle $605 \, \mathrm{fb}^{-1}$

- Thanks to Belle, especially Antonio Limosani, for providing the covariance matrix, experimental efficiency and resolution!
- Efficiency and resolution effects folded into theory predictions

BABAR sum-over-exclusive-modes (80 fb<sup>-1</sup>), hadronic tag (210 fb<sup>-1</sup>)

- Correlations are available
- Spectra efficiency corrected, resolution not an issue
- Thanks to Francesca Di Lodovico and Henning Flächer

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# Fit Setup

 $\chi^2$  Fit

- Includes all experimental correlations
- Extensively validated with pseudo experiments
  - Just having a good  $\chi^2/\mathrm{ndf}$  is not enough

Shape function basis 
$$\widehat{F}(\lambda x) = rac{1}{\lambda} iggl[ \sum\limits_n c_n f_n(x) iggr]^2$$

- Default basis parameter:  $\lambda = 0.5 \, \text{GeV}$
- Include up to 5 basis coefficients (c<sub>0</sub> to c<sub>4</sub>)
- Fix  $\sum_n c_n^2 = 1$  to enforce correct normalization  $\int \mathrm{d}k \widehat{F}(k) = 1$

Disclaimer: What I am showing is active work in progress

- Numbers still subject to change
- Theoretical uncertainties not yet included in the fit

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Events [10<sup>3</sup>/50 MeV]

#### Fit Results



Global Fit to  $B \rightarrow X_{S} \gamma$ 00000000 Outlook on  $|V_{ub}|$ 

## Convergence of Basis Expansion



• Uncertainties underestimated with too few coefficients (c<sub>0,1</sub>)

- Would need to include additional uncertainty due to truncation
- Very little change from including 5th coefficient (c<sub>4</sub>)
  - Truncation uncertainty negligible compared to other uncertainties

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- Equivalent to fixed model with fitted 1st moment
- All with good  $\chi^2/\mathrm{ndf}$ : 37.5/40, 28.8/40, 27.8/40
- ⇒ Uncertainties underestimate model dependence



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Global Fit to  $B o X_S \gamma$ 

Outlook on  $|V_{ub}|$ 



Global Fit to  $B \rightarrow X_s \gamma$ 

Outlook on  $|V_{ub}|$ 



Global Fit to  $B \rightarrow X_{S} \gamma$ 

Outlook on  $|V_{ub}|$ 

# Effect of Perturbative and $1/m_b$ Corrections

- NNLL+NNLO corrections move result up
- Subleading shape functions cause substantial shift in m<sub>b</sub> given by their 1st moment

$$rac{-\lambda_1+3\lambda_2}{2m_b}\sim 70\,{
m MeV}$$



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# Additional Complications for $B \to X_u \ell \nu$

$$\mathrm{d}\Gamma_{u} \propto |V_{ub}|^{2} iggl\{ ig(\widehat{W}_{u}^{\mathrm{sing}} + \widehat{W}_{u}^{\mathrm{nons}}ig) \otimes \widehat{F} + \sum_{n} W_{u,n} \widehat{F}_{n}^{\mathrm{subl}} + \cdots iggr\}$$

- Combining different phase-space regions for triple differential spectrum
  - ► E.g.  $\widehat{W}_{u}^{\text{sing}}$  known to  $\mathcal{O}(\alpha_{s}^{2})$  but  $\widehat{W}_{u}^{\text{nons}}$  only to  $\mathcal{O}(\alpha_{s}, \alpha_{s}^{2}\beta_{0})$
- Subleading SFs are more tricky, cannot be absorbed anymore
  - Different  $B \to X_u \ell \nu$  spectra would help

#### Proof-of-concept fit

- BABAR  $m_X, m_X q^2, p_X^+, \, E_\ell^\Upsilon \geq 2.2 \, {
  m GeV}$
- Belle  $m_X, E_\ell^\Upsilon \geq 2.3\,{
  m GeV}$
- $B \rightarrow X_s \gamma$  spectra
- Theory: NLL+NLO, no 1/mb



Wishlist for experiments (or: if you want your measurement to be used ...)

- Correlations for spectra (or between partial branching fractions)
- Correction matrices if spectra are significantly affected by efficiency and resolution

Global Fit to  $B \rightarrow X_{s} \gamma$ 

- ⇒ Right now this unfortunately excludes a lot of valuable inputs
  - ►  $B_{ABAR}$  leptonic-tag  $B \rightarrow X_s \gamma$  spectrum (updated analysis soon)
  - BABAR  $E_{\ell}^{\Upsilon}$  partial BFs
  - Belle hadronic-tag partial BFs

Getting More Out of Existing Data

# There is much more information we can gain from $\sim 1\,{\rm ab}^{-1}$ of Belle and $B\!AB\!AR$ data

- $B \rightarrow X_u \ell \nu$  spectra will help further constrain  $m_b$  and leading (subleading) SF
- Precise  $E_{\ell}$  spectrum (maybe with cut on  $m_X$ ) would be very useful



Outlook on |V<sub>ub</sub>|

| Global Fit to B | $\rightarrow X_s \gamma$ |  |  |
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#### Summary

#### Global fit to $B \to X_s \gamma$ and $B \to X_u \ell \nu$ with a model-independent treatment of the shape function

- Minimizes uncertainties by making maximal use of available data
- SF and its uncertainties are determined by the data
  - No hidden or underestimated uncertainties from model dependence
- Eliminate extrapolation for comparison to the  $B 
  ightarrow X_s \gamma$  rate
- $\Rightarrow$  Will be key to precision  $\mathcal{B}(B \to X_s \gamma)$  and  $|V_{ub}|$  from super B factory

#### To reduce theory/parameter uncertainties with improved measurements now

- Measure (almost) the total rate, also has drawbacks:
  - Have to pay with (much) larger systematic uncertainties
  - Theory uncertainty creeps back in via signal MC model
- Measure  $B \rightarrow X_u \ell \nu$  spectra with correlations (no drawbacks)
- ⇒ Ideally, should do both. With limited manpower the second option keeps potential for future improvements open.



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