



© Stephan Pastis/Dist. by UFS, Inc.

SUSY Predictions for the LHC

Sven Heinemeyer, IFCA (CSIC, Santander)

Kruger National Park, 12/2010

1. Introduction and motivation
2. The models and the tools
3. Prediction of the lightest Higgs boson mass M_h
4. LHC reach in the CMSSM/NUHM1
5. Conclusions

1. Introduction

The big question:

Which Lagrangian describes the world?

My guess:

It is a **supersymmetric** one

⇒ concentrate on the MSSM from now on

(other people ⇒ other guesses ⇒ other priorities . . .)

1. Introduction

The big question:

Which Lagrangian describes the world?

My guess:

It is a **supersymmetric** one

⇒ concentrate on the MSSM from now on

(other people ⇒ other guesses ⇒ other priorities . . .)

⇒ is there any possibility to know what to expect?

1. Introduction

The big question:

Which Lagrangian describes the world?

My guess:

It is a **supersymmetric** one

⇒ concentrate on the MSSM from now on

(other people ⇒ other guesses ⇒ other priorities . . .)

⇒ is there any possibility to know what to expect?

Let's see . . .

Supersymmetry (SUSY) : Symmetry between

Bosons \leftrightarrow Fermions

$$Q \text{ |Fermion} \rangle \rightarrow \text{|Boson} \rangle$$

$$Q \text{ |Boson} \rangle \rightarrow \text{|Fermion} \rangle$$

Simplified examples:

$$Q \text{ |top, } t \rangle \rightarrow \text{|scalar top, } \tilde{t} \rangle$$

$$Q \text{ |gluon, } g \rangle \rightarrow \text{|gluino, } \tilde{g} \rangle$$

\Rightarrow each SM multiplet is enlarged to its double size

Unbroken SUSY: All particles in a multiplet have the same mass

Reality: $m_e \neq m_{\tilde{e}}$ \Rightarrow SUSY is broken . . .

. . . via soft SUSY-breaking terms in the Lagrangian (added by hand)

SUSY particles are made heavy: $M_{\text{SUSY}} = \mathcal{O}(1 \text{ TeV})$

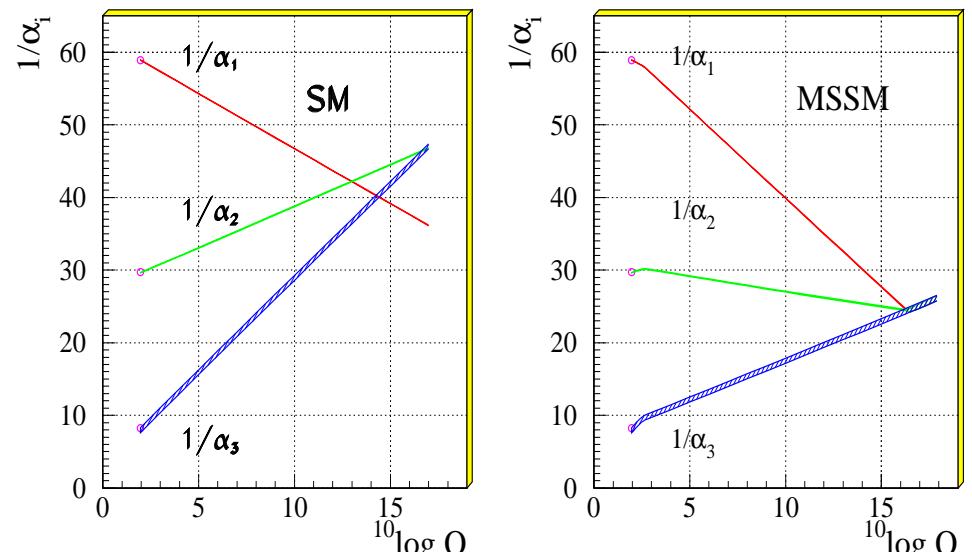
Supersymmetry: Motivation

The SM is in a pretty good shape.

Why MSSM? (Is it worth to double the particle spectrum?)

- 1.) Stability of the Higgs mass against higher-order corr.
- 2.) Unification of gauge couplings:
Not possible in the SM, but in the **MSSM** (although it was **not** designed for it.)
- 3.) Spontaneous symmetry breaking via Higgs mechanism is automatic in **SUSY GUTs**
- 4.) SUSY provides CDM candidate
- 5.) ...

Unification of the Coupling Constants in the SM and the minimal MSSM



[Amaldi, de Boer, Fürstenau '92]

The Minimal Supersymmetric Standard Model (MSSM)

Superpartners for Standard Model particles

$[u, d, c, s, t, b]_{L,R}$	$[e, \mu, \tau]_{L,R}$	$[\nu_{e,\mu,\tau}]_L$	Spin $\frac{1}{2}$
$[\tilde{u}, \tilde{d}, \tilde{c}, \tilde{s}, \tilde{t}, \tilde{b}]_{L,R}$	$[\tilde{e}, \tilde{\mu}, \tilde{\tau}]_{L,R}$	$[\tilde{\nu}_{e,\mu,\tau}]_L$	Spin 0
g	$\underbrace{W^\pm, H^\pm}_{\text{}}$	$\gamma, Z, \underbrace{H_1^0, H_2^0}_{\text{}}$	Spin 1 / Spin 0
\tilde{g}	$\tilde{\chi}_{1,2}^\pm$	$\tilde{\chi}_{1,2,3,4}^0$	Spin $\frac{1}{2}$

Enlarged Higgs sector: Two Higgs doublets

Problem in the MSSM: many scales

→ CPV will mostly be neglected throughout this talk!

\tilde{t}/\tilde{b} sector of the MSSM: (scalar partner of the top/bottom quark)

Stop, sbottom mass matrices ($X_t = A_t - \mu^*/\tan\beta$, $X_b = A_b - \mu^*\tan\beta$):

$$\mathcal{M}_{\tilde{t}}^2 = \begin{pmatrix} M_{\tilde{t}_L}^2 + m_t^2 + DT_{t_1} & m_t X_t^* \\ m_t X_t & M_{\tilde{t}_R}^2 + m_t^2 + DT_{t_2} \end{pmatrix} \xrightarrow{\theta_{\tilde{t}}} \begin{pmatrix} m_{\tilde{t}_1}^2 & 0 \\ 0 & m_{\tilde{t}_2}^2 \end{pmatrix}$$

$$\mathcal{M}_{\tilde{b}}^2 = \begin{pmatrix} M_{\tilde{b}_L}^2 + m_b^2 + DT_{b_1} & m_b X_b^* \\ m_b X_b & M_{\tilde{b}_R}^2 + m_b^2 + DT_{b_2} \end{pmatrix} \xrightarrow{\theta_{\tilde{b}}} \begin{pmatrix} m_{\tilde{b}_1}^2 & 0 \\ 0 & m_{\tilde{b}_2}^2 \end{pmatrix}$$

mixing important in stop sector (also in sbottom sector for large $\tan\beta$)

soft SUSY-breaking parameters A_t, A_b also appear in ϕ - \tilde{t}/\tilde{b} couplings

$$SU(2) \text{ relation} \Rightarrow M_{\tilde{t}_L} = M_{\tilde{b}_L}$$

\Rightarrow relation between $m_{\tilde{t}_1}, m_{\tilde{t}_2}, \theta_{\tilde{t}}, m_{\tilde{b}_1}, m_{\tilde{b}_2}, \theta_{\tilde{b}}$

Enlarged Higgs sector: Two Higgs doublets

$$H_1 = \begin{pmatrix} H_1^1 \\ H_1^2 \end{pmatrix} = \begin{pmatrix} v_1 + (\phi_1 + i\chi_1)/\sqrt{2} \\ \phi_1^- \end{pmatrix}$$

$$H_2 = \begin{pmatrix} H_2^1 \\ H_2^2 \end{pmatrix} = \begin{pmatrix} \phi_2^+ \\ v_2 + (\phi_2 + i\chi_2)/\sqrt{2} \end{pmatrix}$$

$$V = m_1^2 H_1 \bar{H}_1 + m_2^2 H_2 \bar{H}_2 - m_{12}^2 (\epsilon_{ab} H_1^a H_2^b + \text{h.c.})$$

$$+ \underbrace{\frac{g'^2 + g^2}{8}}_{\text{gauge couplings, in contrast to SM}} (H_1 \bar{H}_1 - H_2 \bar{H}_2)^2 + \underbrace{\frac{g^2}{2}}_{\text{gauge couplings, in contrast to SM}} |H_1 \bar{H}_2|^2$$

gauge couplings, in contrast to SM

physical states: h^0, H^0, A^0, H^\pm

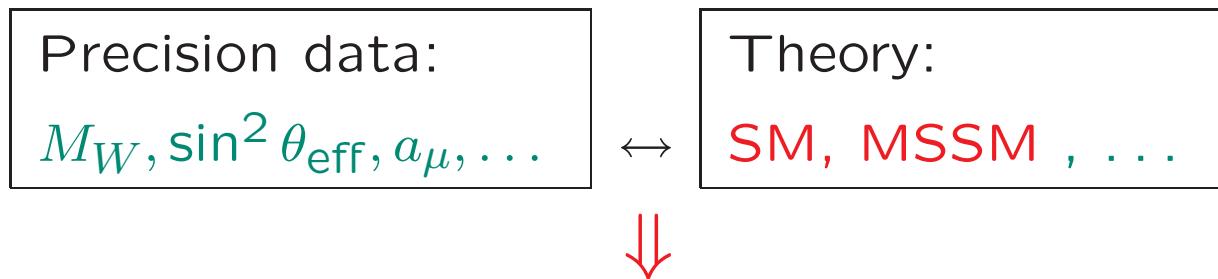
Goldstone bosons: G^0, G^\pm

Input parameters: (to be determined experimentally)

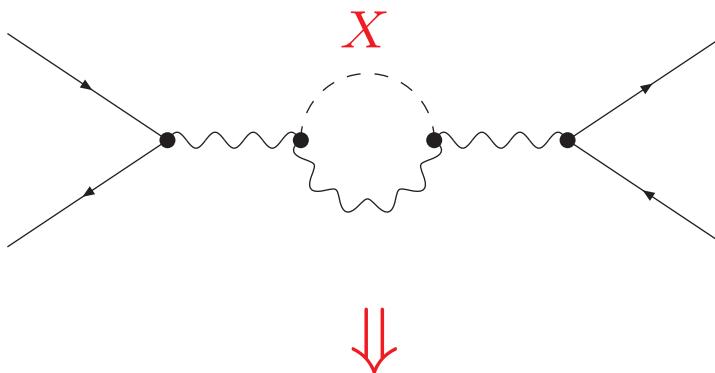
$$\tan \beta = \frac{v_2}{v_1}, \quad M_A^2 = -m_{12}^2(\tan \beta + \cot \beta)$$

How to make a prediction?

Comparison of precision observables with theory:



Test of theory at quantum level: Sensitivity to loop corrections



⇒ Information about unknown parameters

Very high accuracy of measurements and theoretical predictions needed

Example: Prediction for M_W in the **SM** and the **MSSM** :

Theoretical prediction for M_W in terms

of $M_Z, \alpha, G_\mu, \Delta r$:

$$M_W^2 \left(1 - \frac{M_W^2}{M_Z^2} \right) = \frac{\pi \alpha}{\sqrt{2} G_\mu} (1 + \Delta r)$$

\Updownarrow
loop corrections

→ can be approximated with the **ρ -parameter**:

ρ measures the relative strength between
neutral current interaction and charged current interaction

$$\rho = \frac{1}{1 - \Delta \rho}, \quad \Delta \rho = \frac{\Sigma_Z(0)}{M_Z^2} - \frac{\Sigma_W(0)}{M_W^2}, \quad \Delta M_W \approx \frac{M_W}{2} \frac{c_W^2}{c_W^2 - s_W^2} \Delta \rho$$

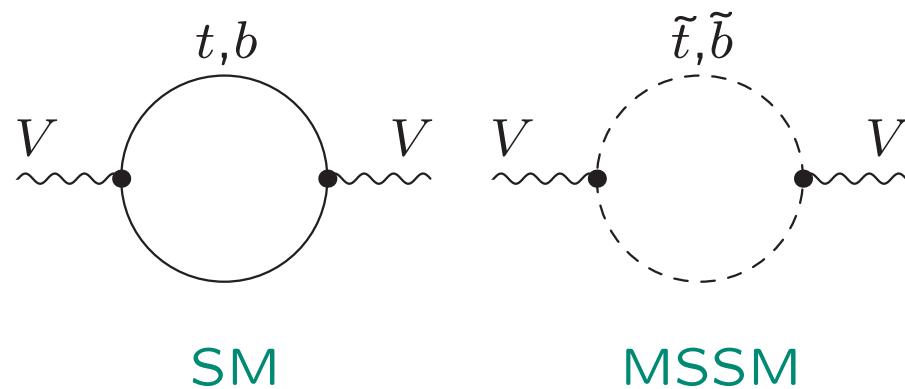
(leading, process independent terms)

→ can be approximated with the **ρ -parameter**:

ρ measures the relative strength between
neutral current interaction and charged current interaction

$$\rho = \frac{1}{1 - \Delta\rho}, \quad \Delta\rho = \frac{\Sigma_Z(0)}{M_Z^2} - \frac{\Sigma_W(0)}{M_W^2}, \quad \Delta M_W \approx \frac{M_W}{2} \frac{c_W^2}{c_W^2 - s_W^2} \Delta\rho$$

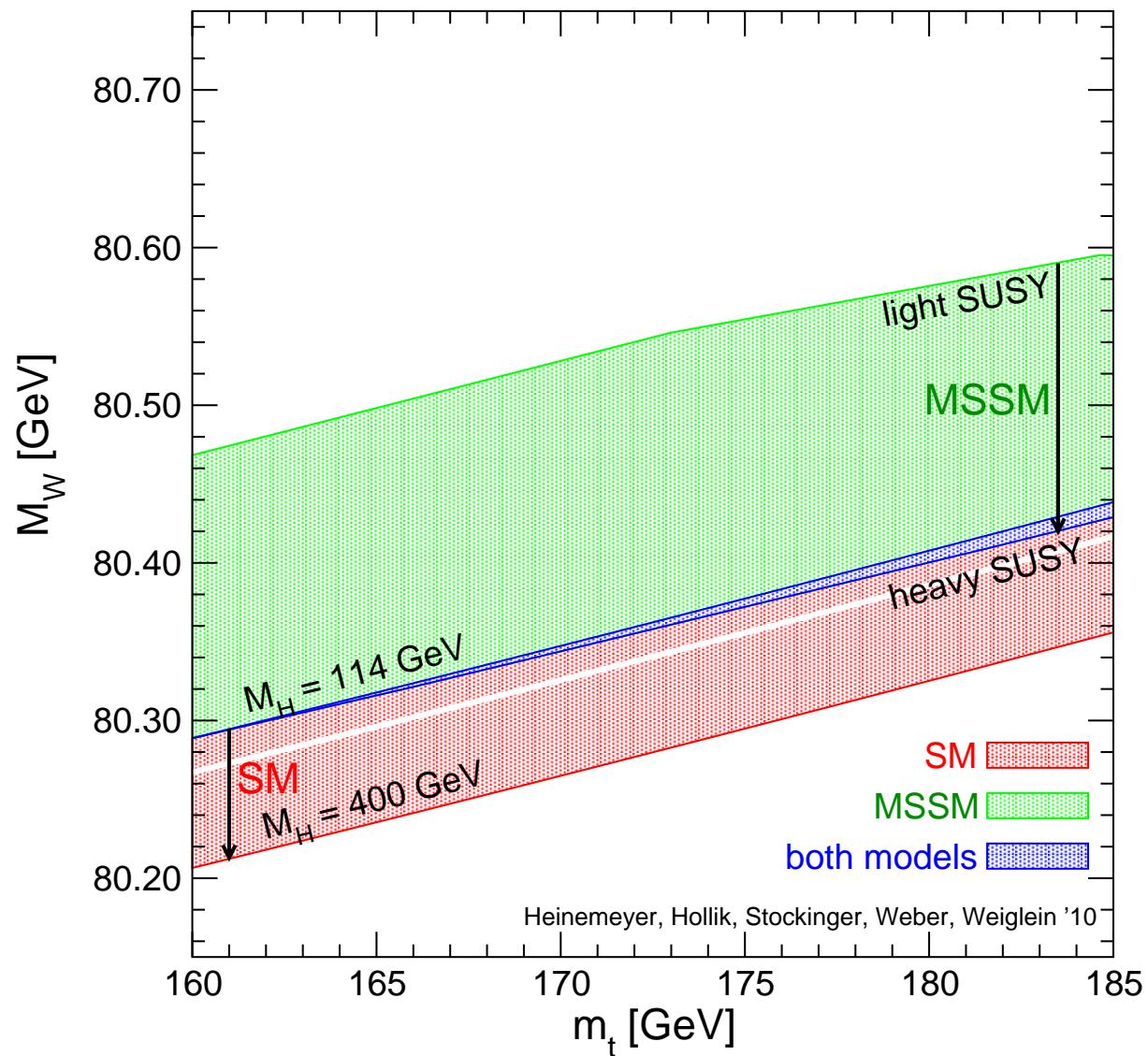
(leading, process independent terms)



$$\Delta\rho^{\text{SUSY}} \text{ from } \tilde{t}/\tilde{b} \text{ loops} > 0 \quad \Rightarrow M_W^{\text{SUSY}} \gtrsim M_W^{\text{SM}}$$

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

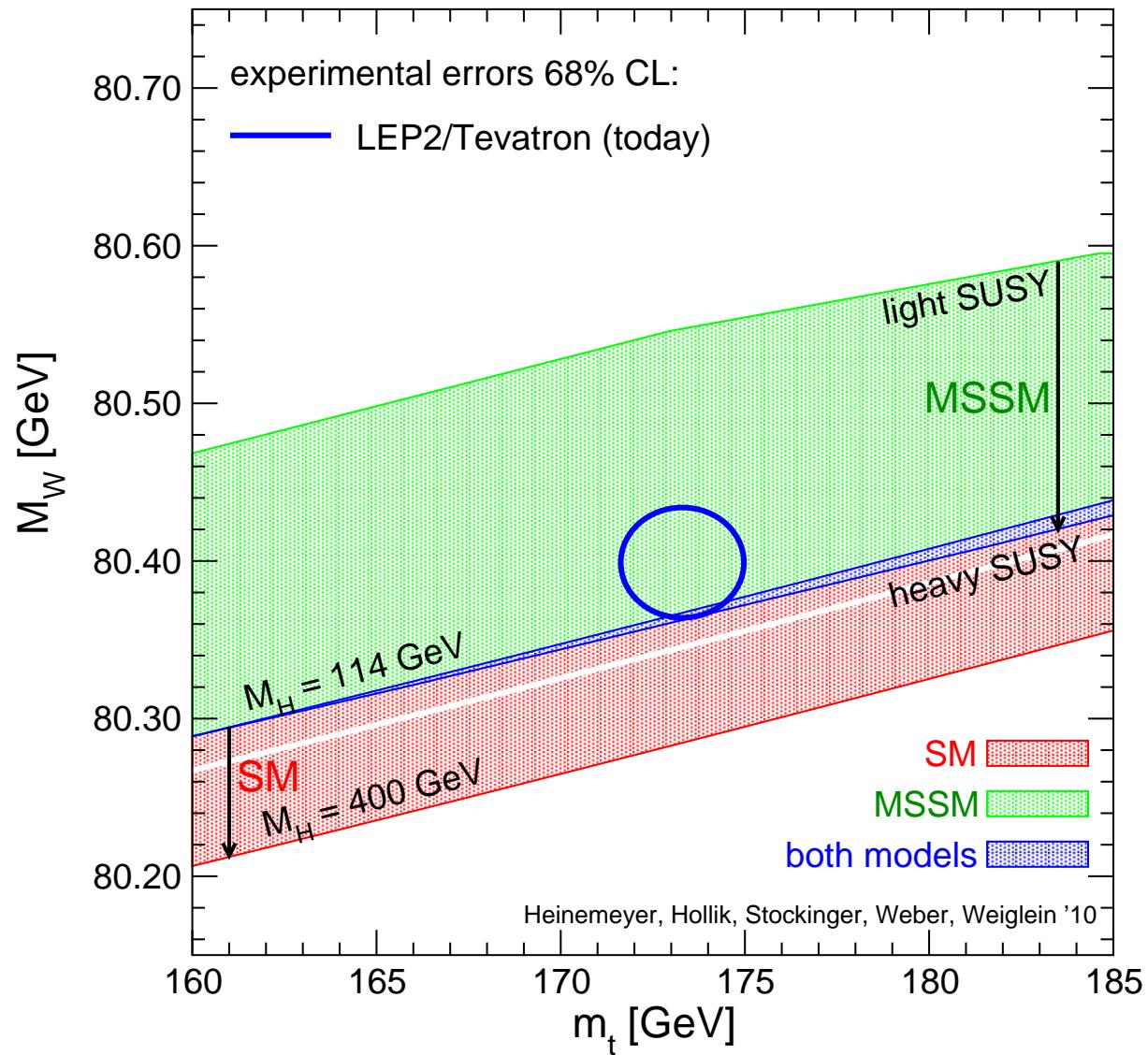
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

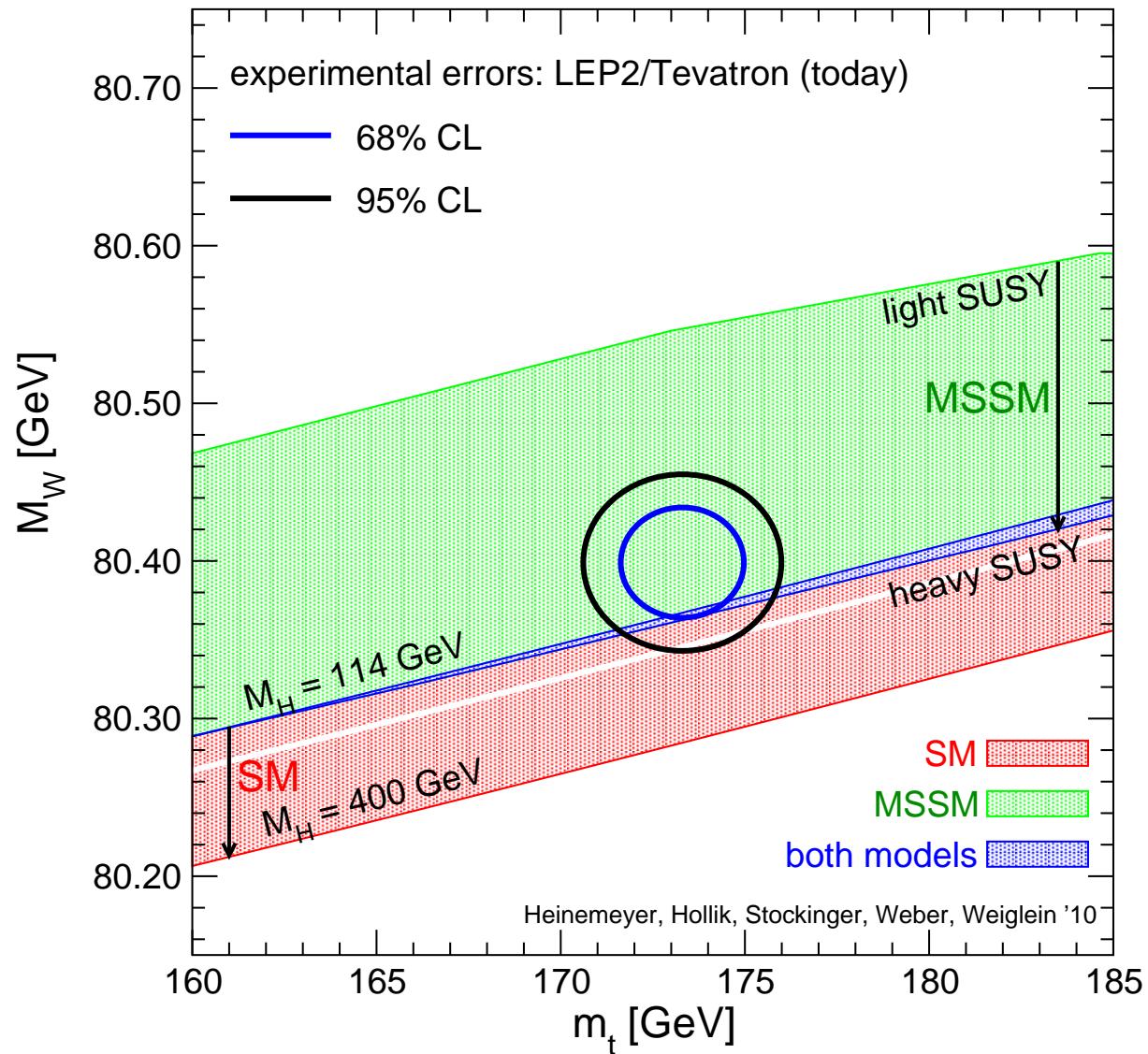
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

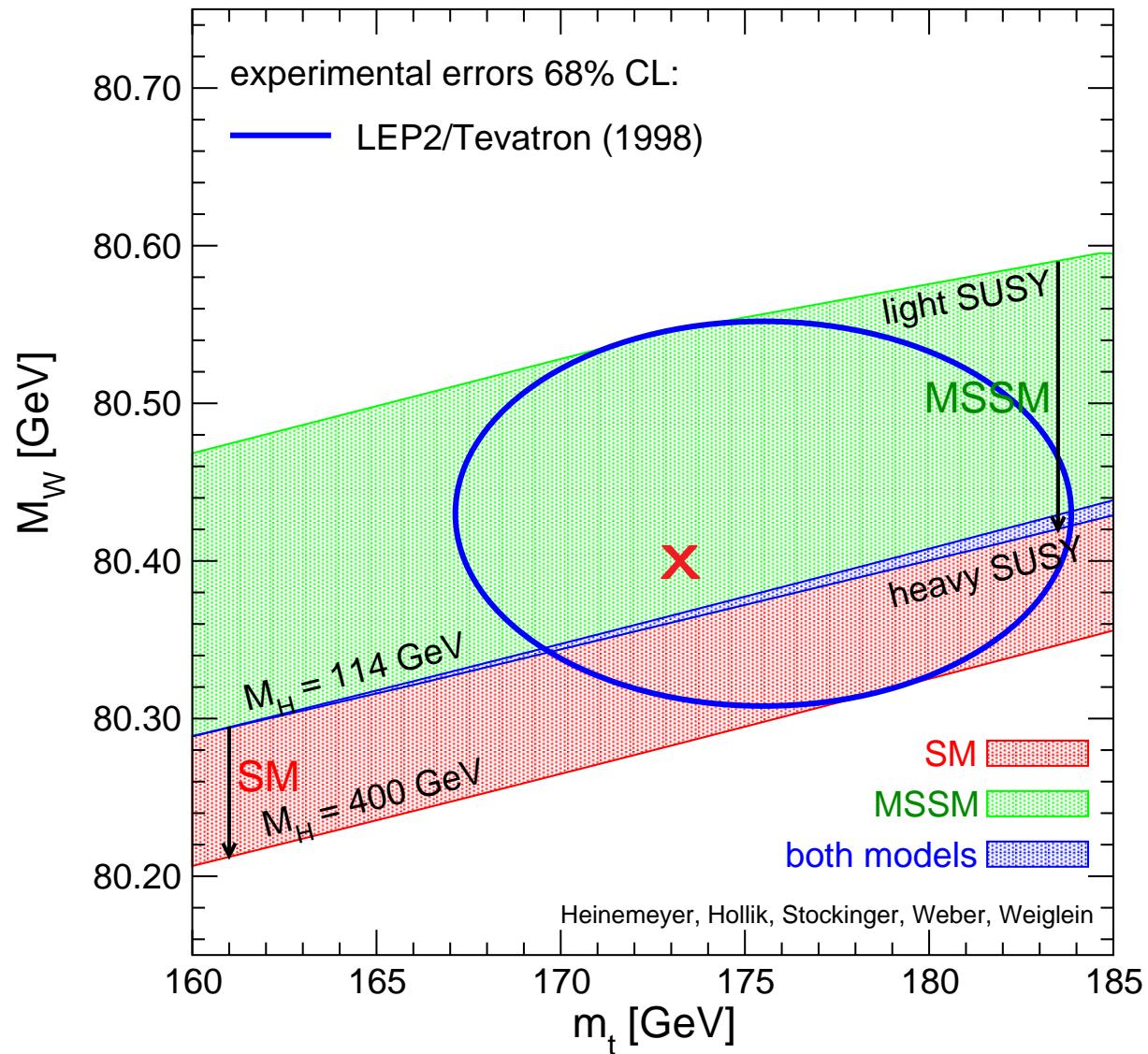
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

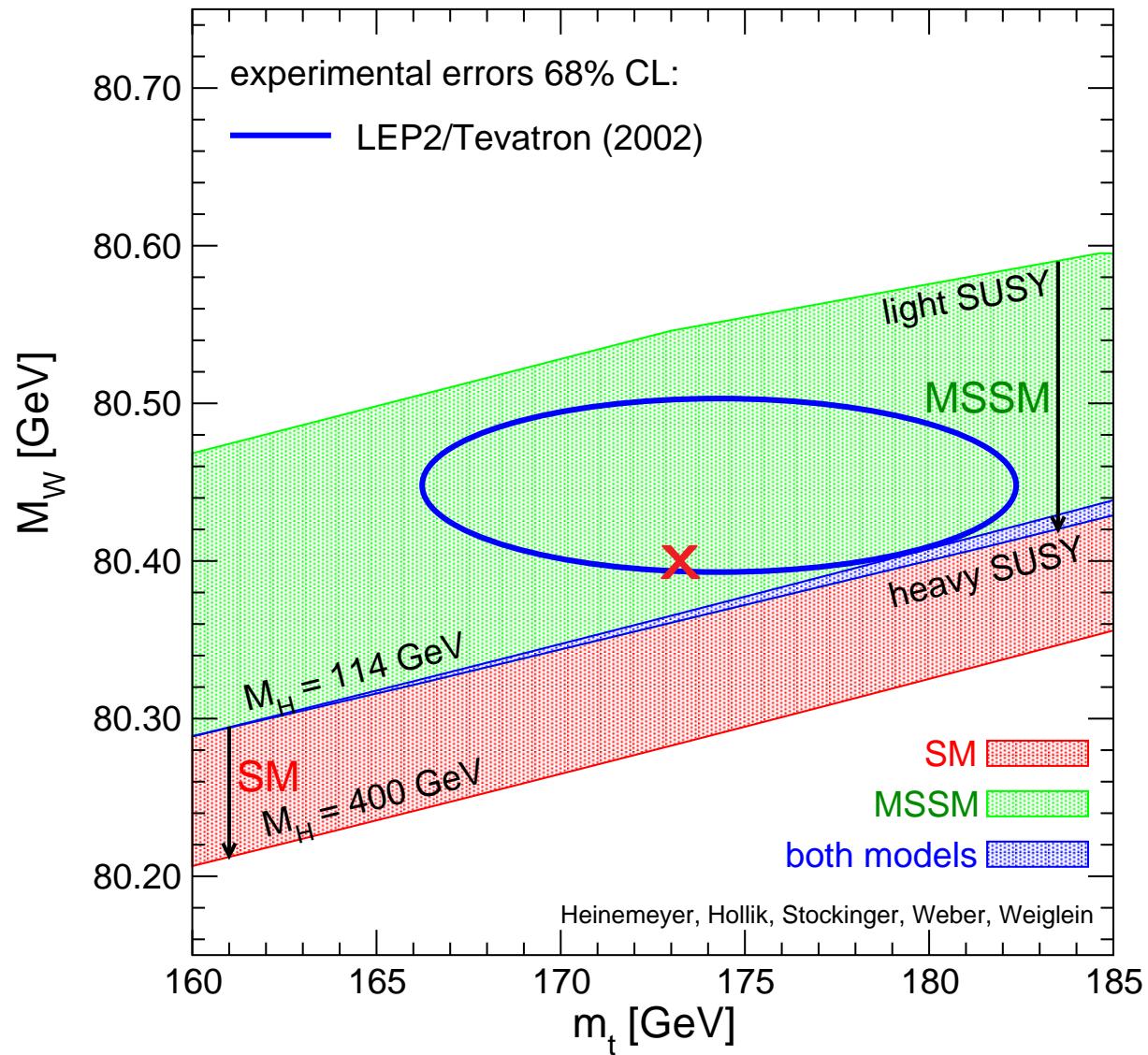
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

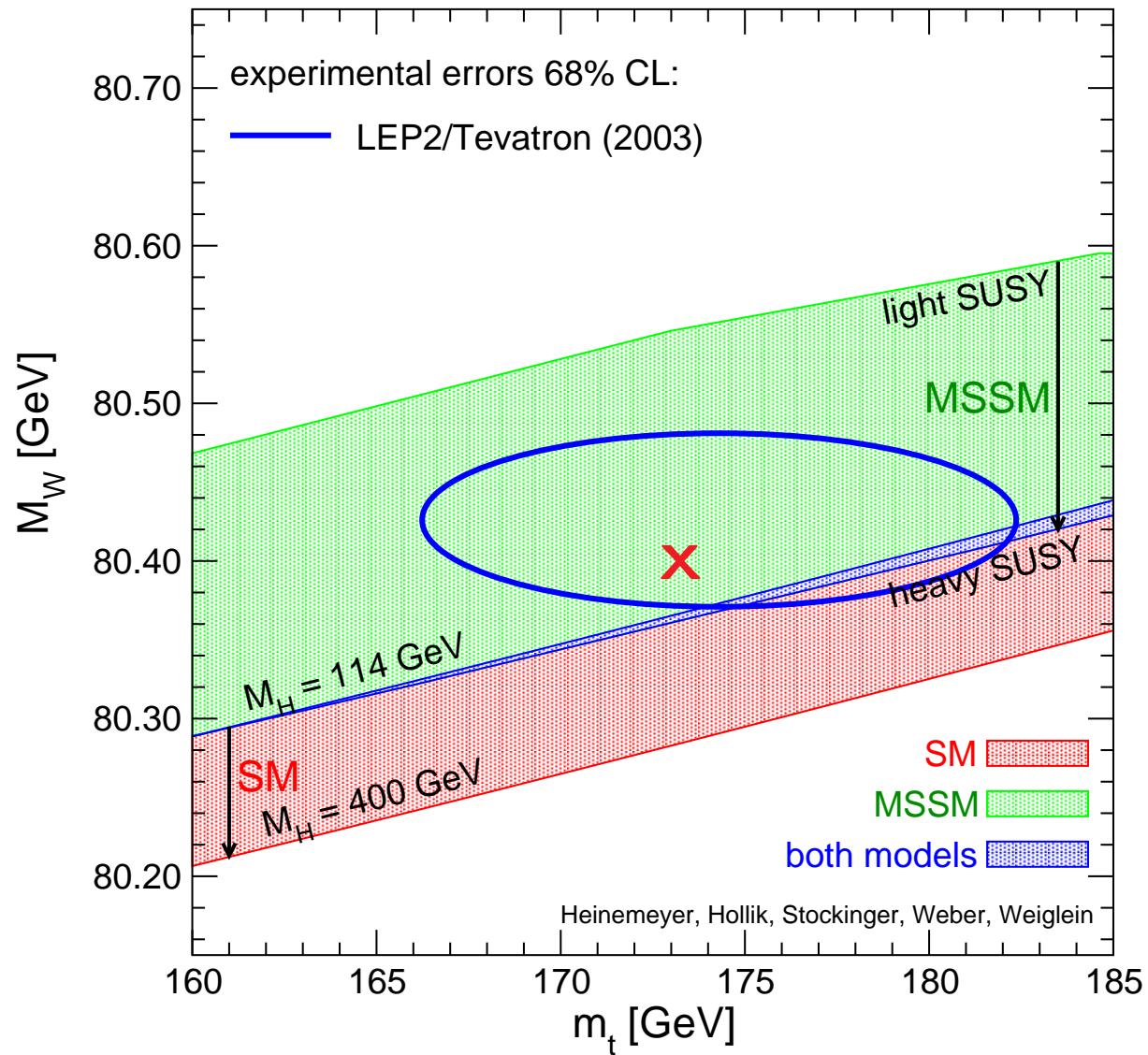
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

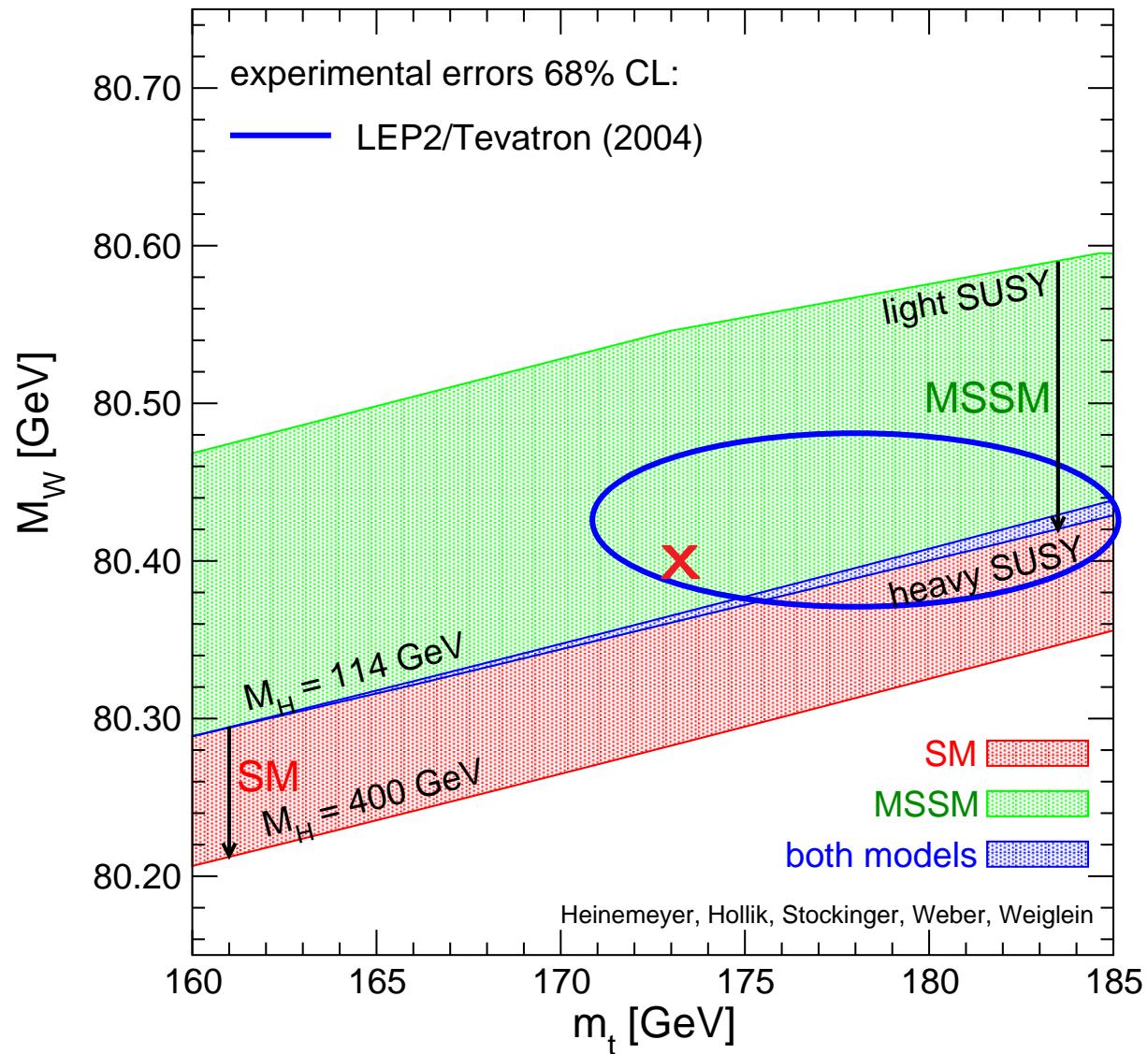
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

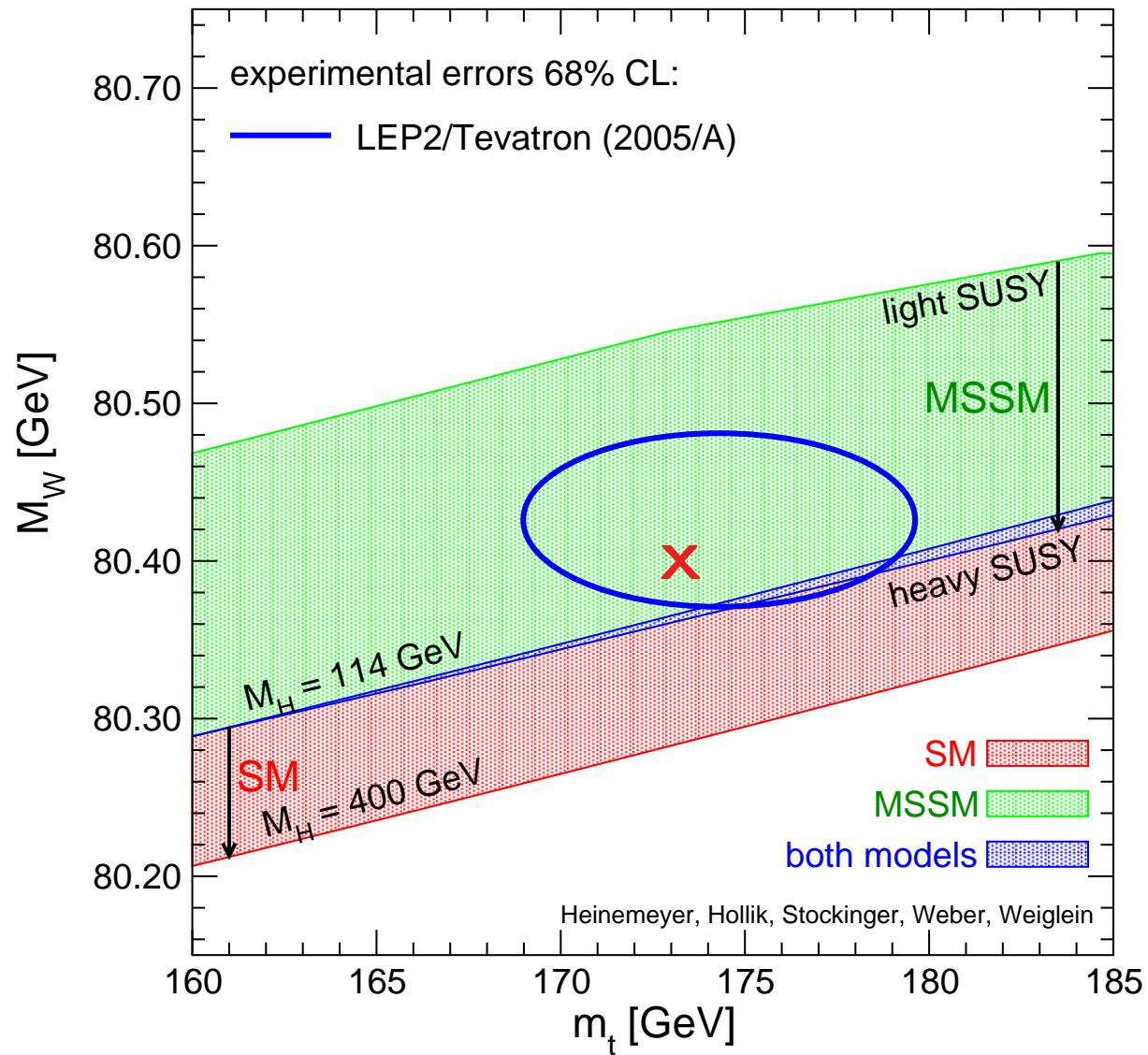
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

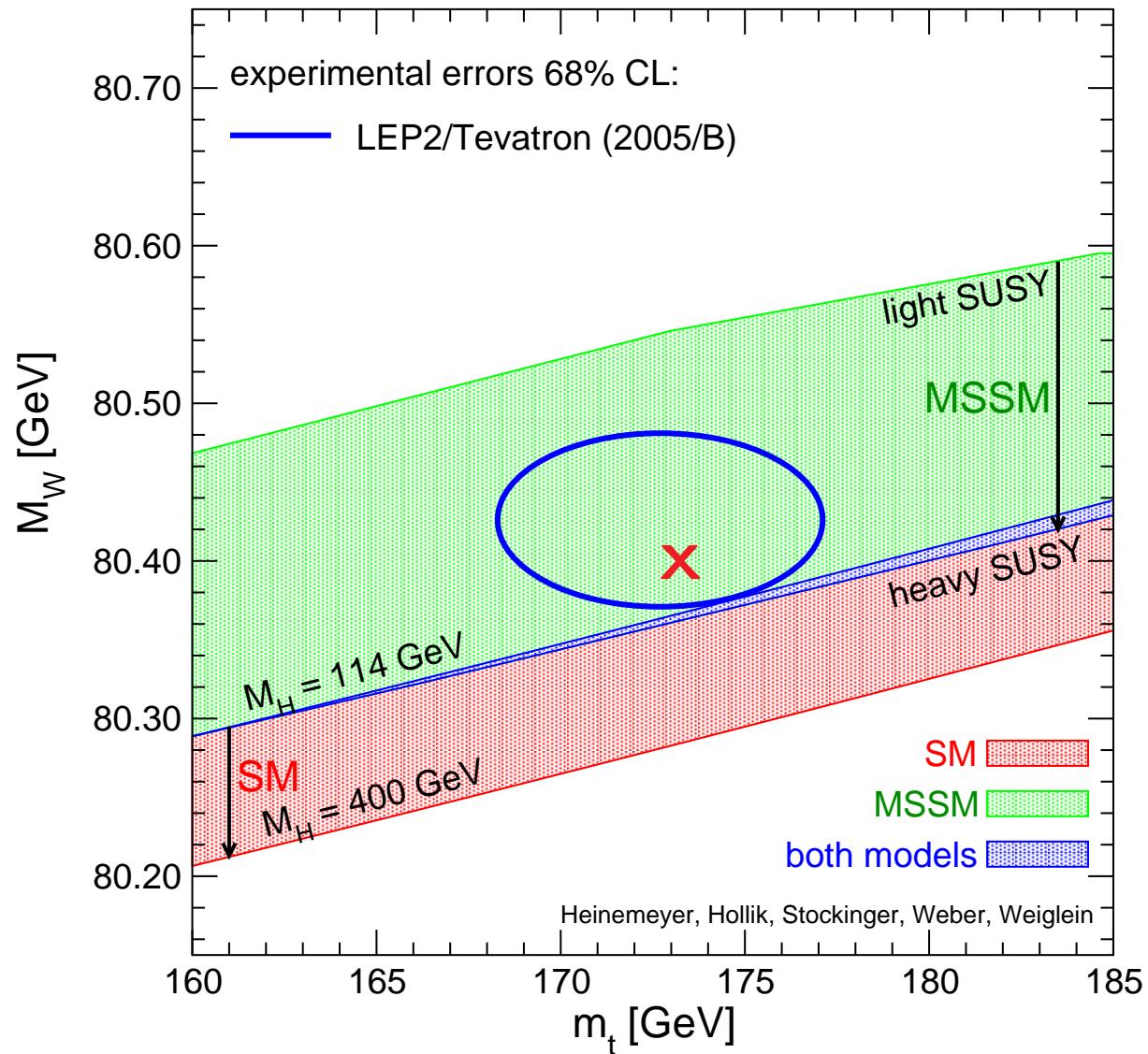
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

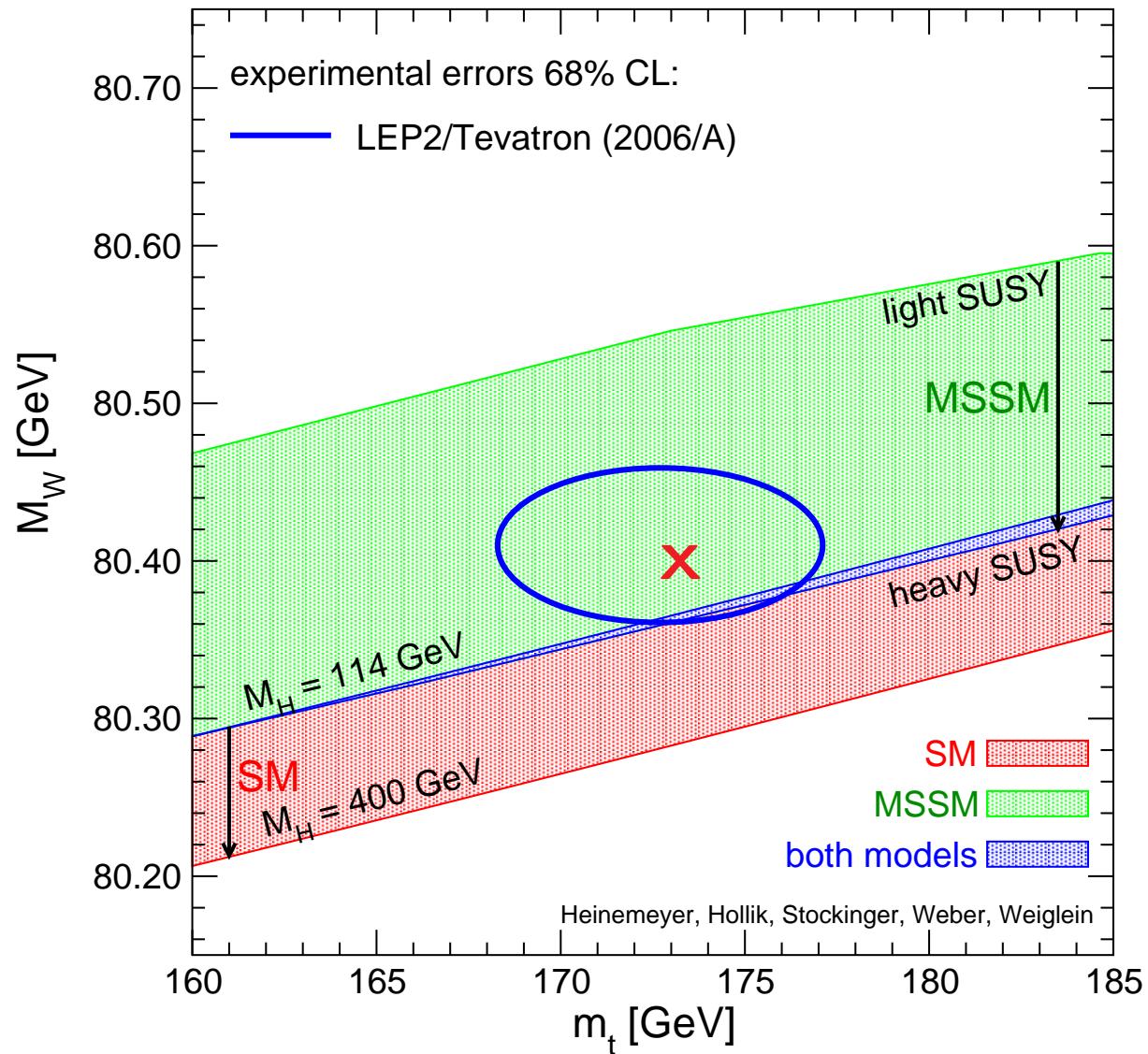
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

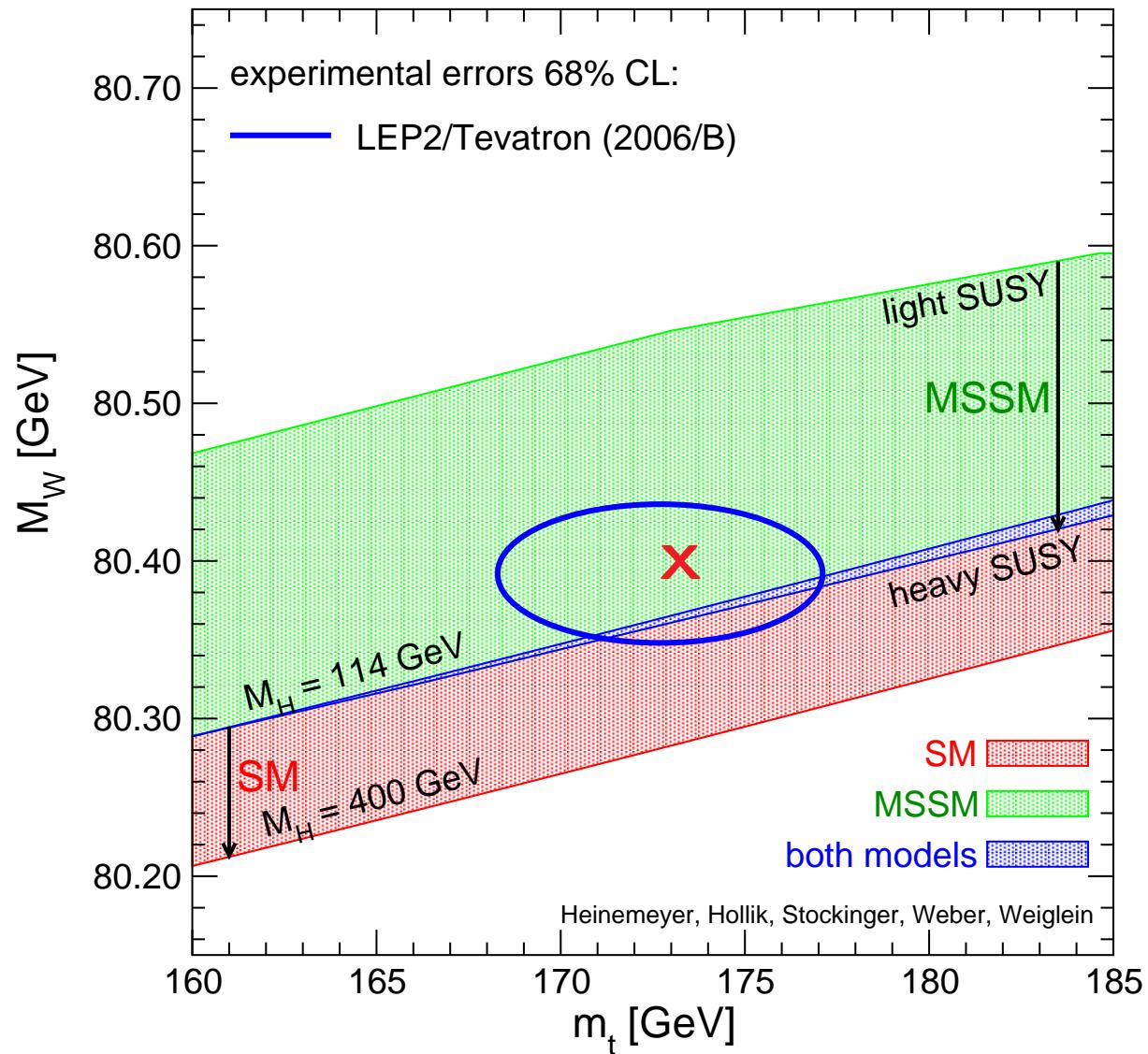
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

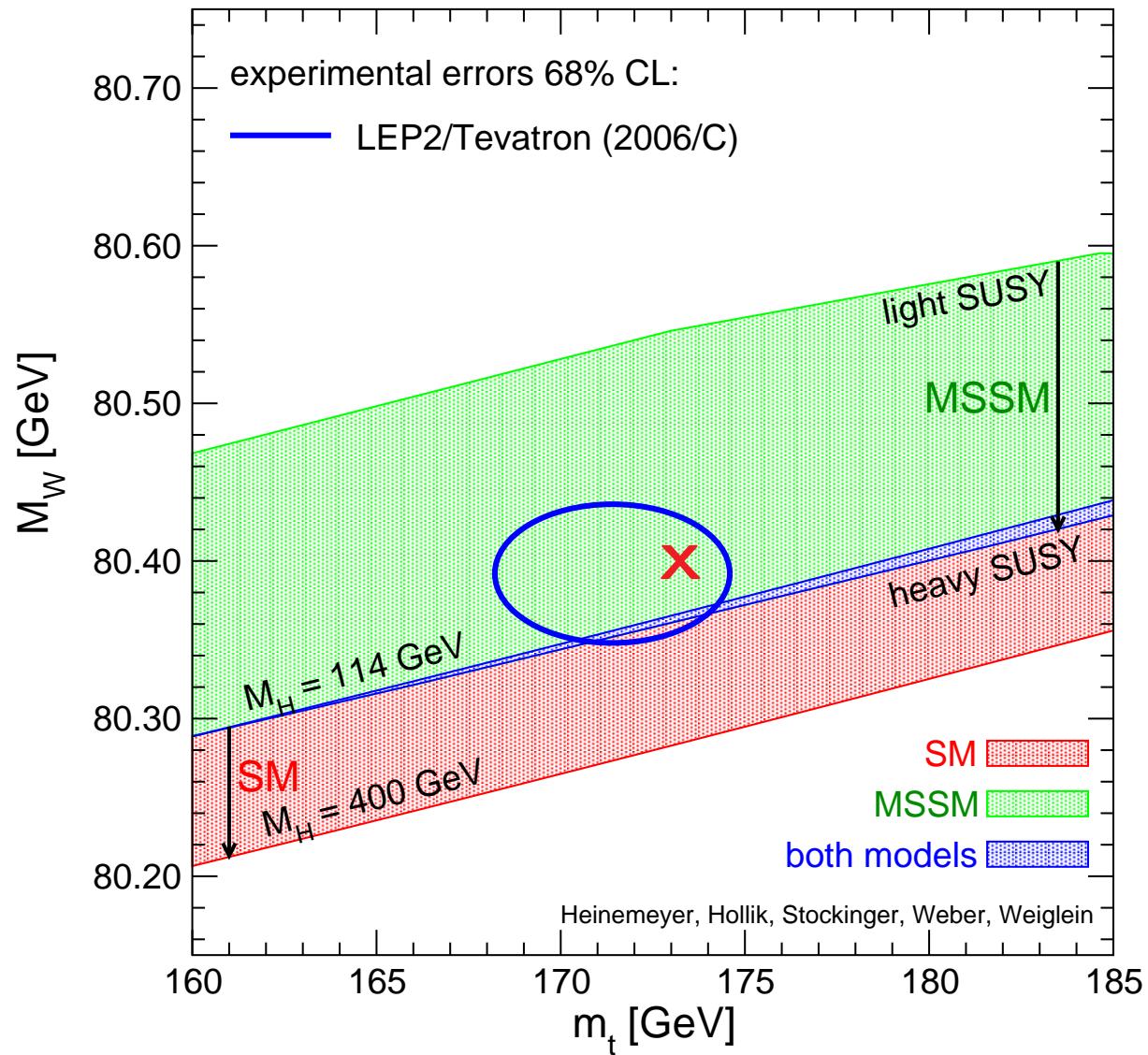
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

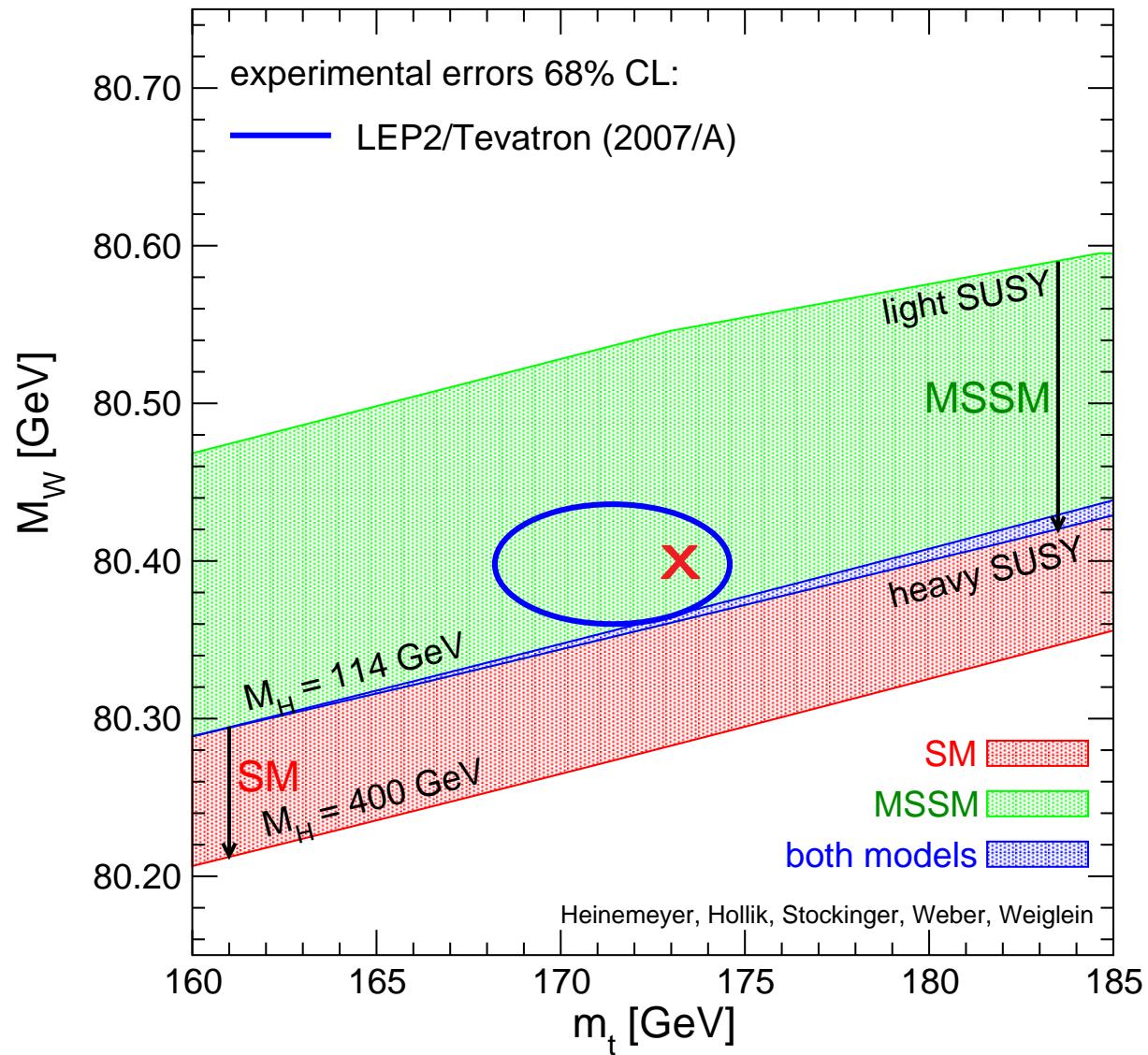
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

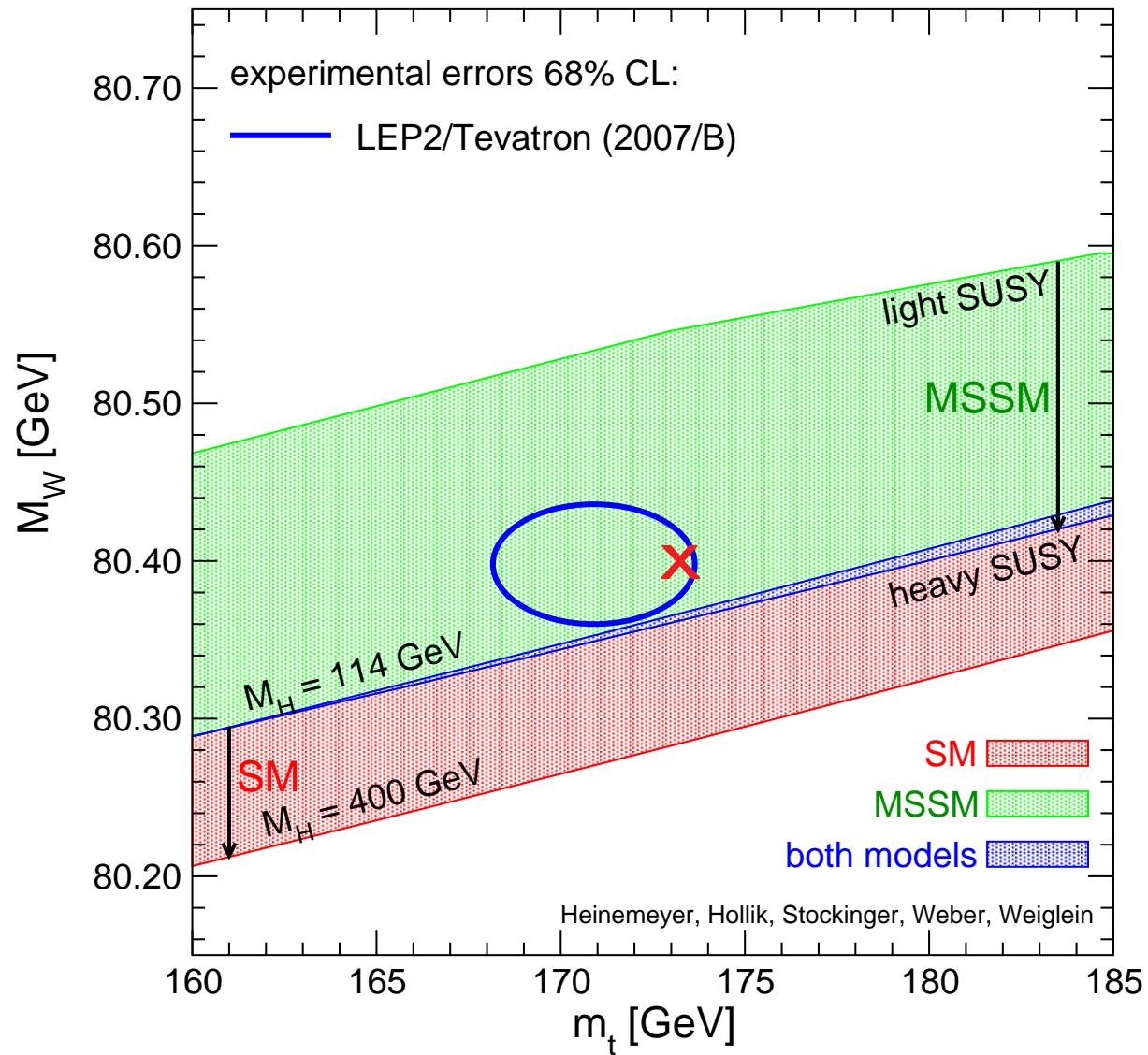
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

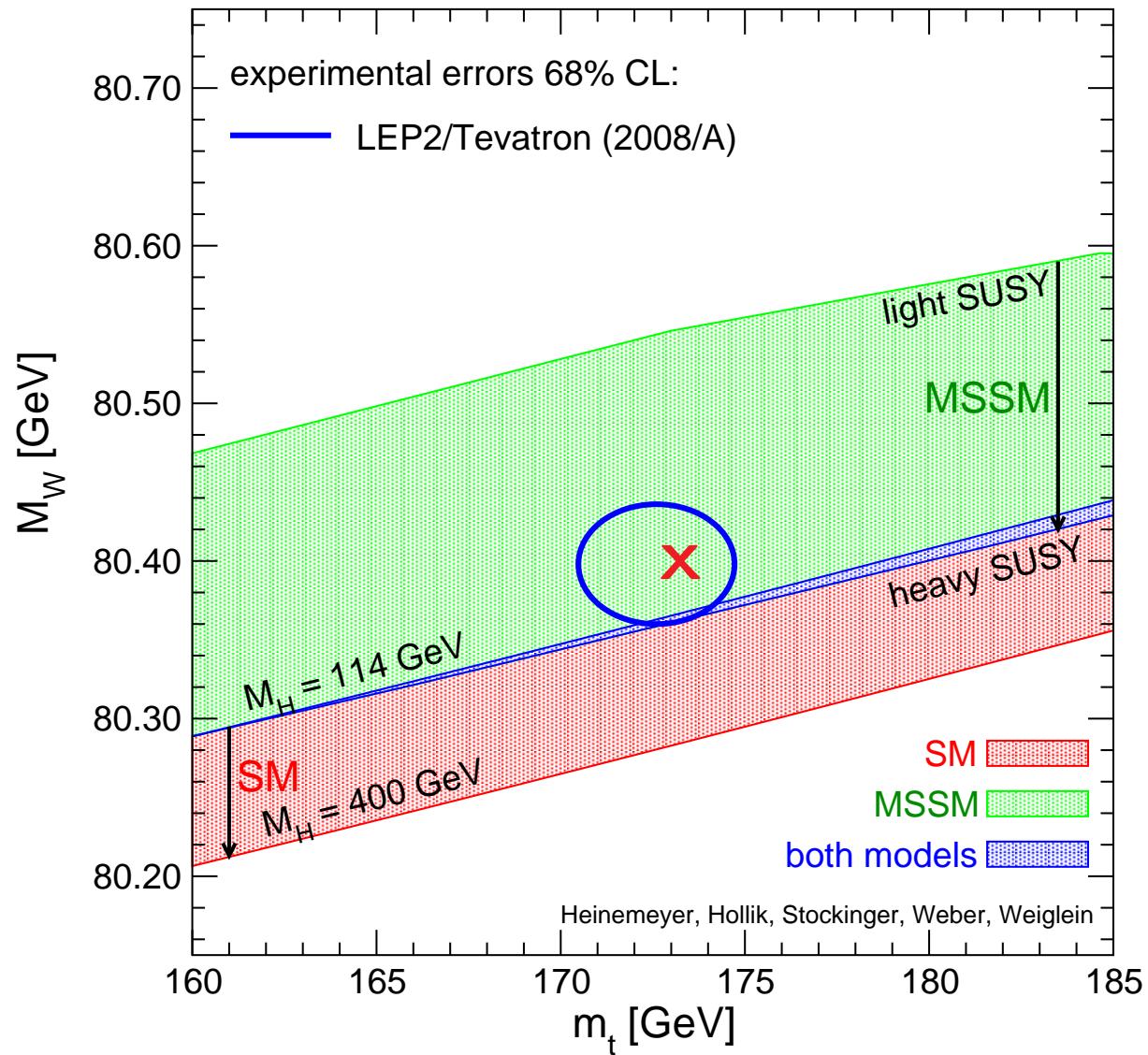
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

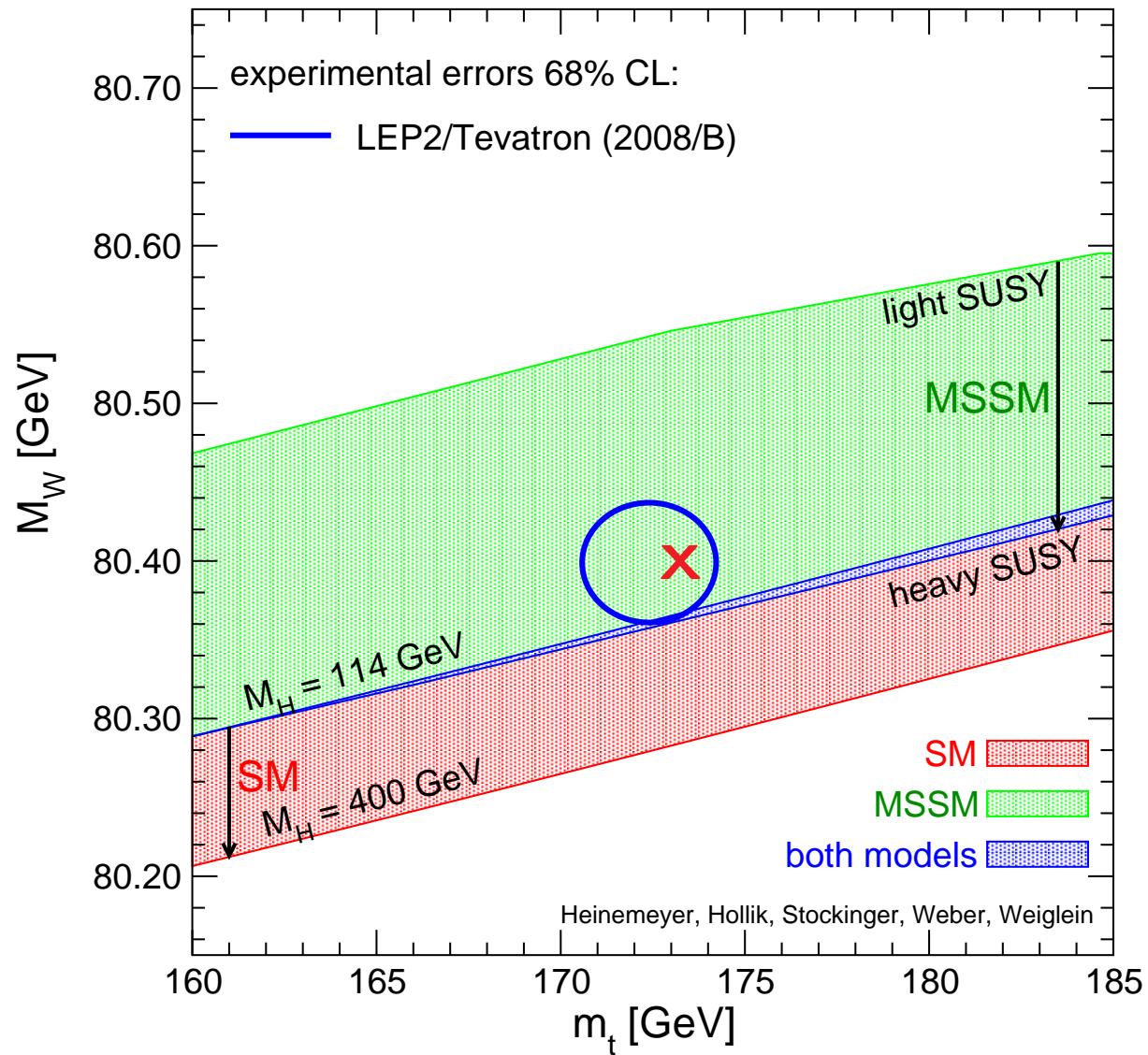
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

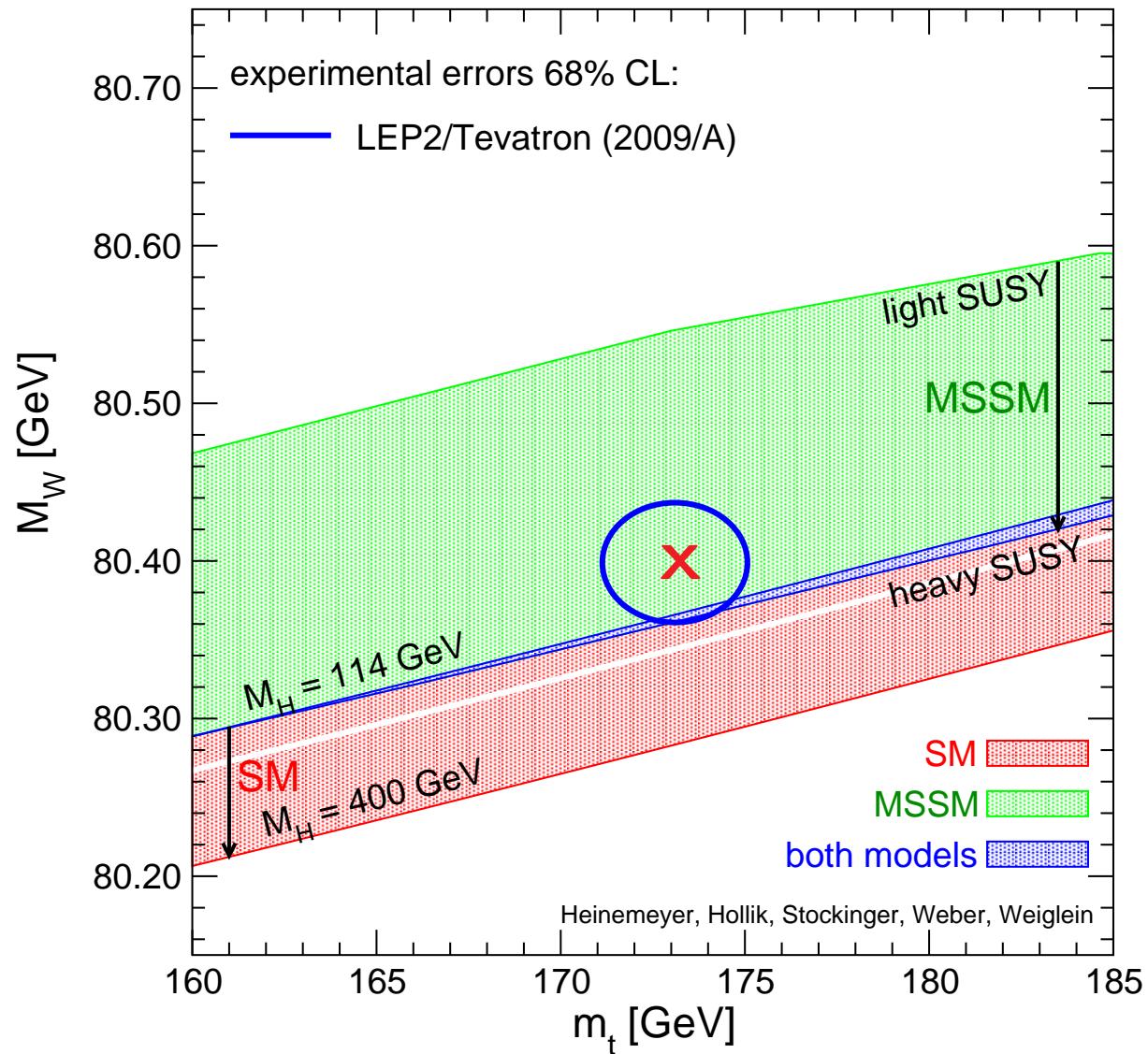
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

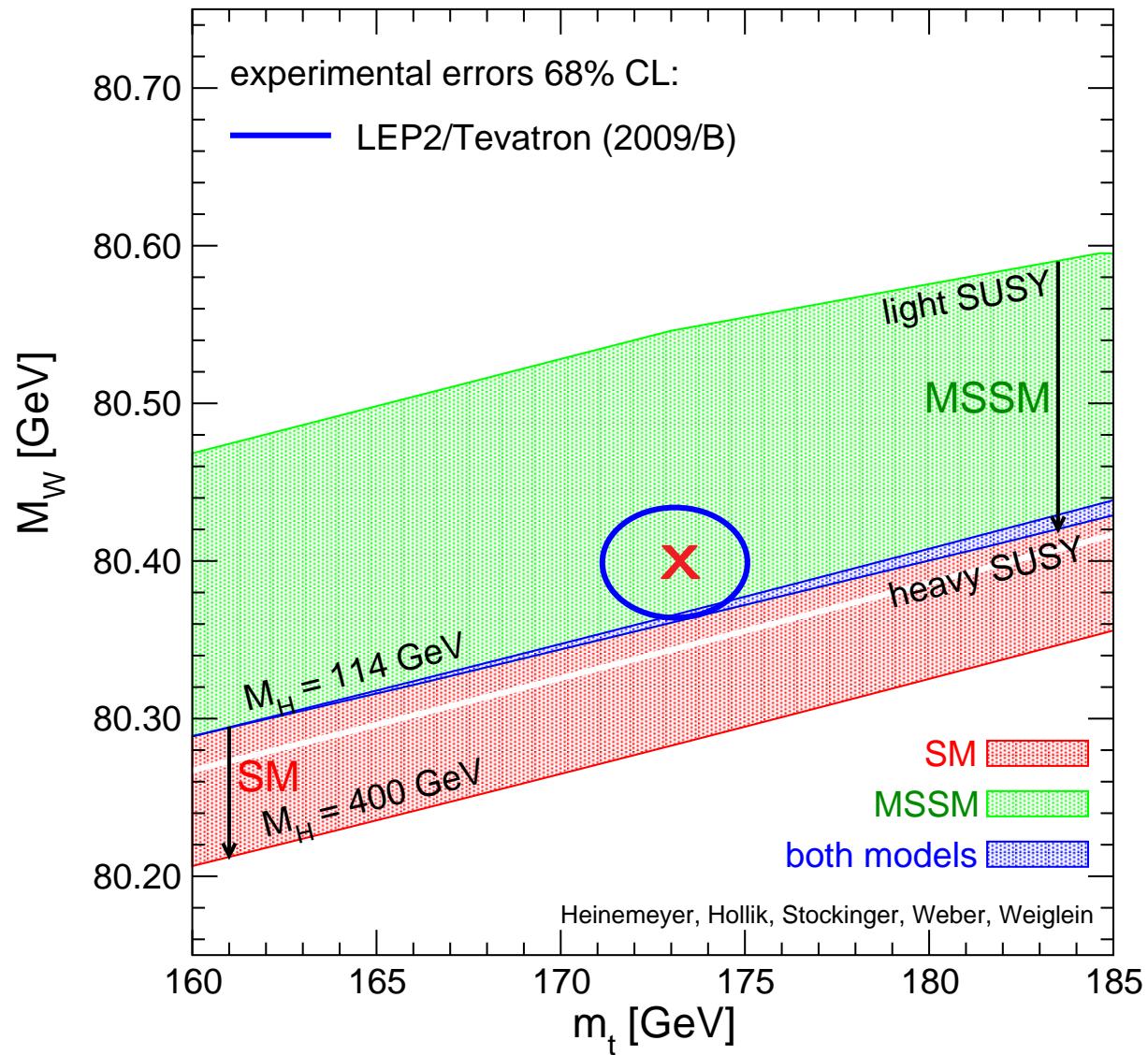
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

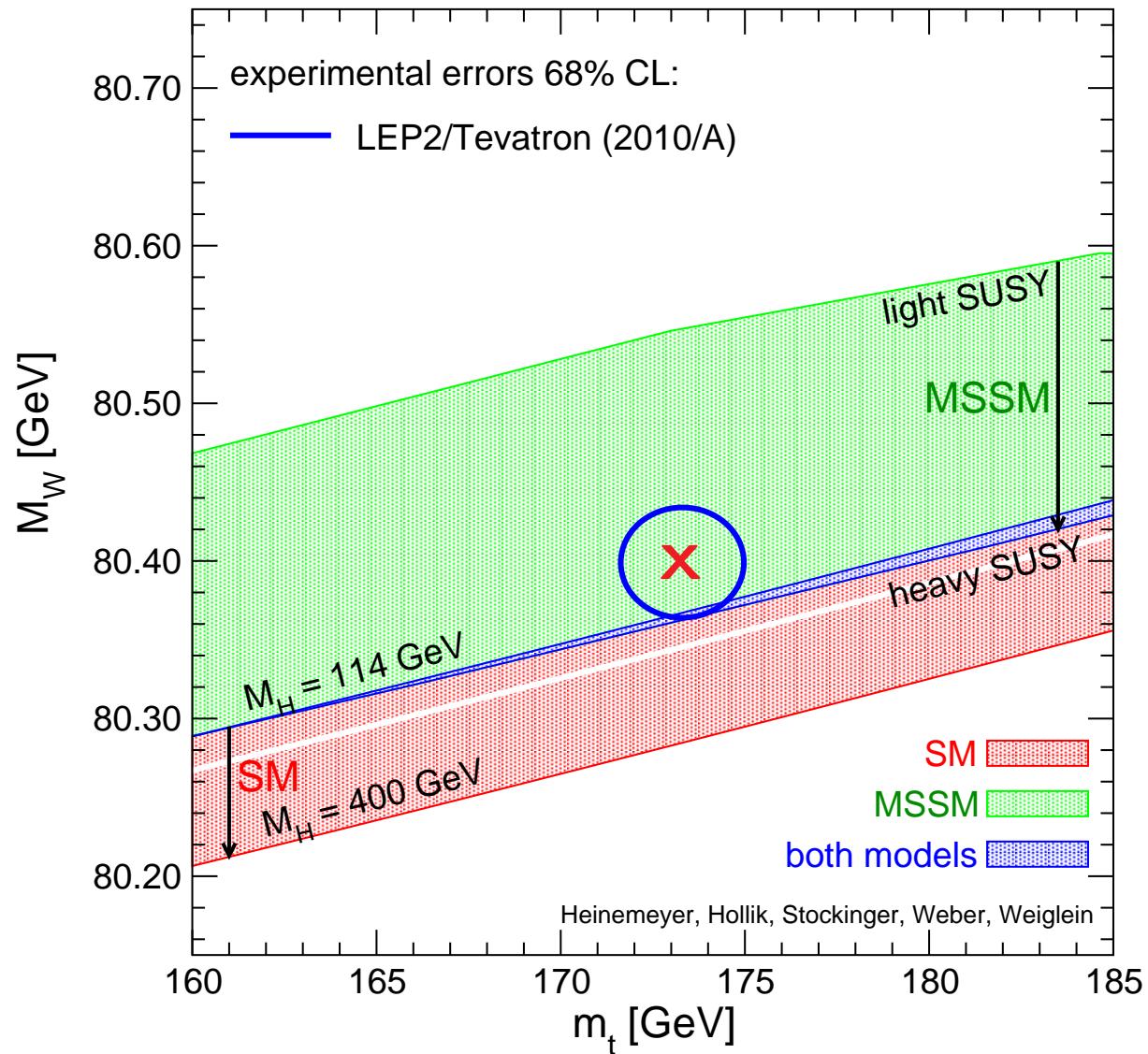
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

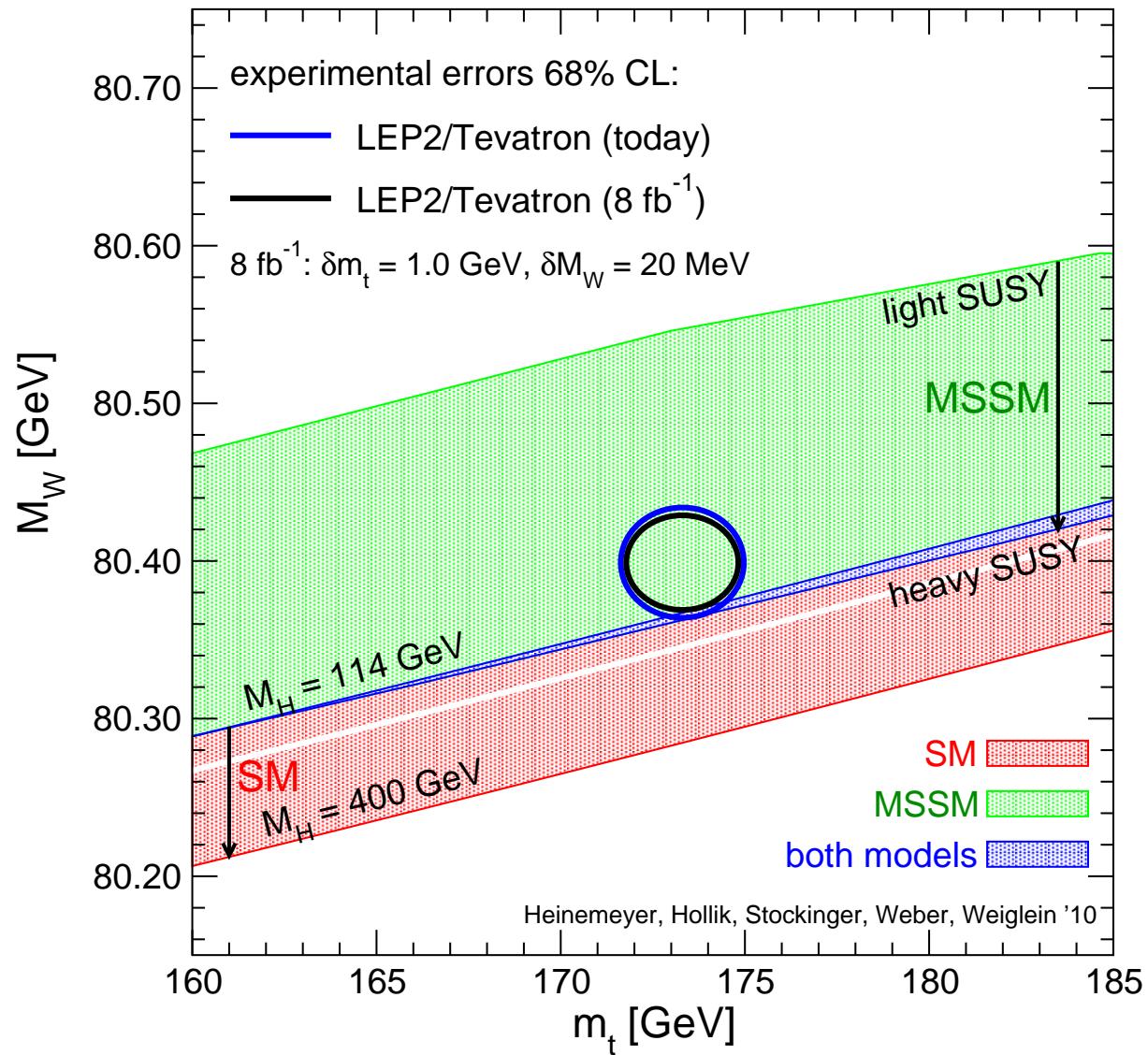
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

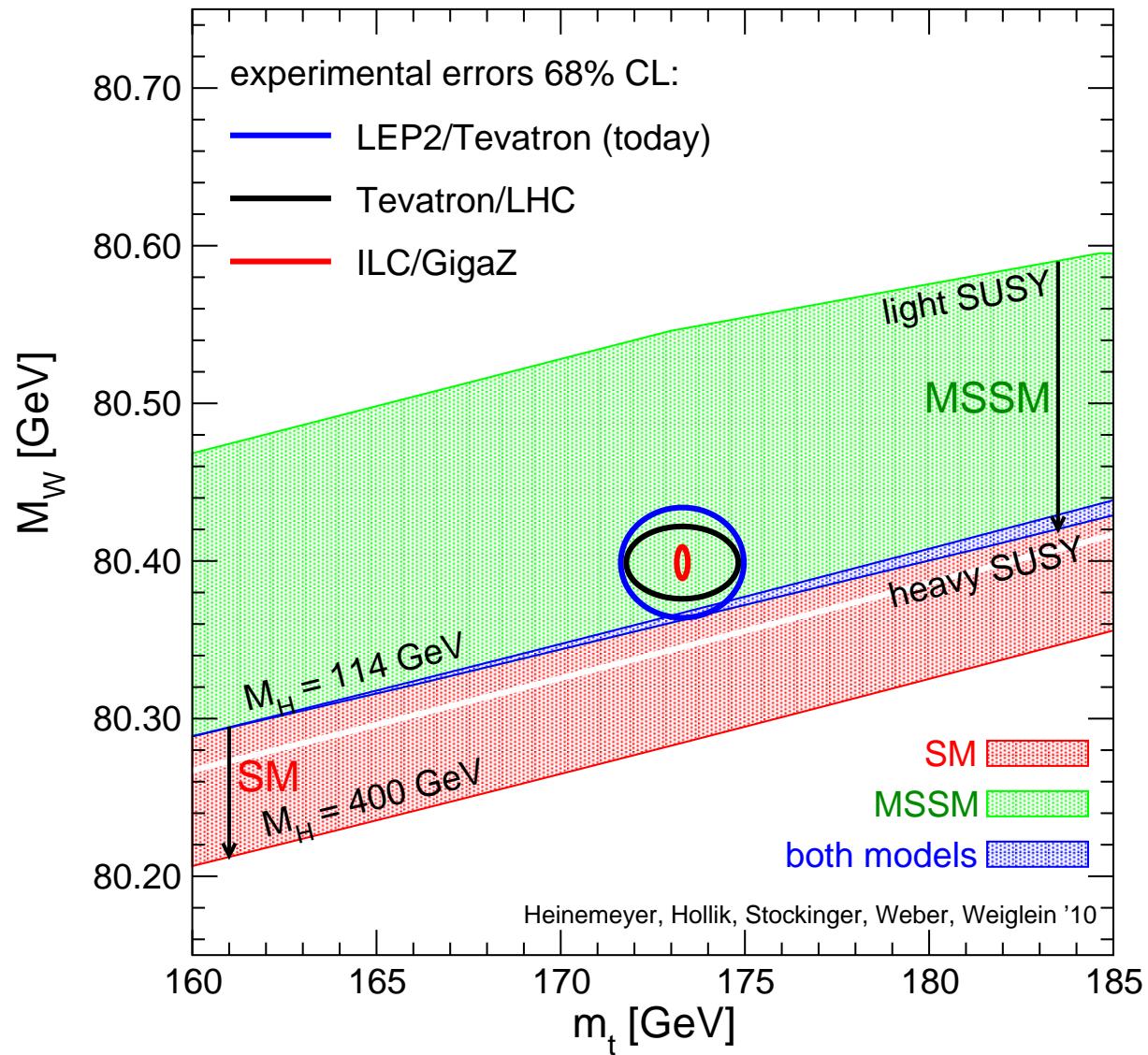
SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Global fit to all SM data:

[LEPEWWG '10]

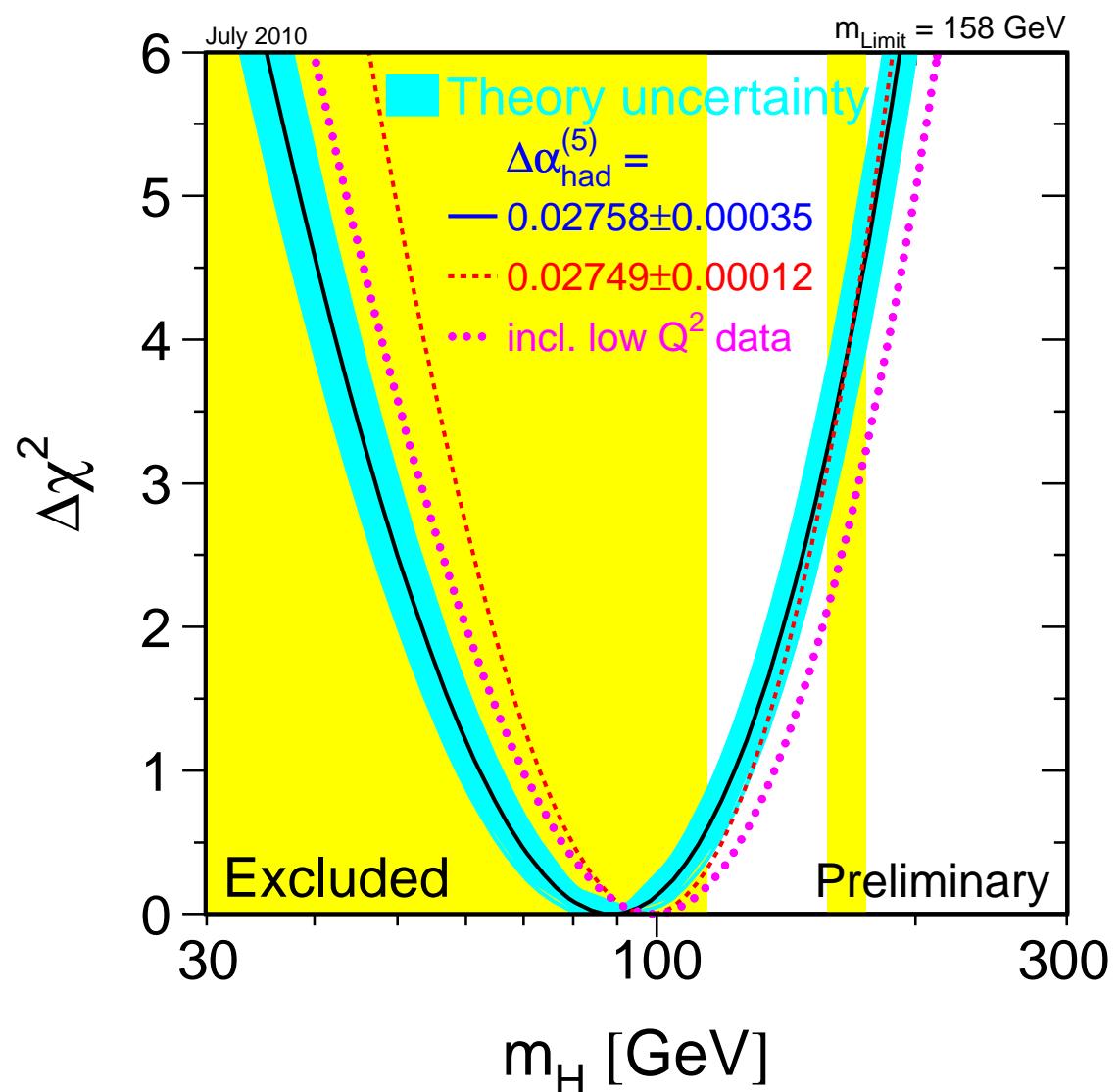
$$\Rightarrow M_H = 89^{+35}_{-26} \text{ GeV}$$

$M_H < 158$ GeV, 95% C.L.

Assumption for the fit:

SM incl. Higgs boson

\Rightarrow no confirmation of
Higgs mechanism



\Rightarrow Higgs boson seems to be light, $M_H \lesssim 160$ GeV

Main idea of analysis:

Combine all existing precision data:

- Electroweak precision observables (**EWPO**)
- B physics observables (**BPO**)
- Cold dark matter (**CDM**)
- ...

Predict:

- best-fit points
- ranges for Higgs masses
- ranges for SM parameters
- ranges for SUSY masses \Rightarrow **LHC reach**

2. The models and the tools

Indirect constraints on M_{SUSY} from existing data?

- Electroweak precision observables (**EWPO**) ?
 - B physics observables (**BPO**) ?
 - Cold dark matter (**CDM**) ?
- ⇒ combination of EWPO, BPO, CDM ?

2. The models and the tools

Indirect constraints on M_{SUSY} from existing data?

- Electroweak precision observables (**EWPO**) ?
- B physics observables (**BPO**) ?
- Cold dark matter (**CDM**) ?
⇒ combination of EWPO, BPO, CDM ?

EWPO M_W : information on $m_{\tilde{t}}$, $m_{\tilde{b}}$ or M_A , $\tan \beta$ or ...

EWPO $(g - 2)_\mu$: information on $\tan \beta$ and/or $m_{\tilde{\chi}^0}$, $m_{\tilde{\chi}^\pm}$ and/or $m_{\tilde{\mu}}$, $m_{\tilde{\nu}_\mu}$

BPO $\text{BR}(b \rightarrow s\gamma)$: information on $\tan \beta$ and/or M_{H^\pm} and/or $m_{\tilde{t}}$, $m_{\tilde{\chi}^\pm}$

CDM (**LSP gives CDM**) : information on $m_{\tilde{\chi}_1^0}$ and $m_{\tilde{\tau}}$ or M_A or ...

2. The models and the tools

Indirect constraints on M_{SUSY} from existing data?

- Electroweak precision observables (**EWPO**) ?
- B physics observables (**BPO**) ?
- Cold dark matter (**CDM**) ?
⇒ combination of EWPO, BPO, CDM ?

EWPO M_W : information on $m_{\tilde{t}}$, $m_{\tilde{b}}$ or M_A , $\tan \beta$ or ...

EWPO $(g - 2)_\mu$: information on $\tan \beta$ and/or $m_{\tilde{\chi}^0}$, $m_{\tilde{\chi}^\pm}$ and/or $m_{\tilde{\mu}}$, $m_{\tilde{\nu}_\mu}$

BPO $\text{BR}(b \rightarrow s\gamma)$: information on $\tan \beta$ and/or M_{H^\pm} and/or $m_{\tilde{t}}$, $m_{\tilde{\chi}^\pm}$

CDM (LSP gives CDM) : information on $m_{\tilde{\chi}_1^0}$ and $m_{\tilde{\tau}}$ or M_A or ...

⇒ combination makes only sense if all parameters are connected!

⇒ GUT based models, ...

Existing analyses for GUT based models: (involving precision observables)

CMSSM/mSUGRA:

- [J. Ellis, S.H., K. Olive, G. Weiglein '04, '06, '07] [J. Ellis, S.H., K. Olive, A. Weber, G. Weiglein '07]
- [E. Baltz, P. Gondolo '04] [R. Ruiz de Austri, R. Trotta and L. Roszkowski '06, '07]
- [B. Allanach, C. Lester and A. Weber '06, '07]
- [F. Feroz, M. Hobson, L. Roszkowski and R. Ruiz de Austri, R. Trotta '08]
- [O. Buchmueller et al. '07] [O. Buchmueller et al. '08] [O. Buchmueller et al. '09]
- [M. Cabrera, A. Casas, R. Ruiz de Austri '09] [Y. Akrami, P. Scott, J. Edsjo, J. Conrad, L. Bergstrom '09]

NUHM (Non-Universal Higgs Mass model):

- [J. Ellis, S.H., K. Olive, G. Weiglein '06] [J. Ellis, S.H., K. Olive, A.M. Weber, G. Weiglein '07]
- [J. Ellis, T. Hahn, S.H., K. Olive, G. Weiglein '07]
- [O. Buchmueller et al. '08] [O. Buchmueller et al. '09]

VCMSSM (Very Constrained MSSM):

- [J. Ellis, S.H., K. Olive, G. Weiglein '06]
- [L. Roszkowski, R. Ruiz de Austri, R. Trotta, Y. Tsai, T. Varley '09]

mSUGRA (GDM) (Gravitino Dark Matter): [J. Ellis, S.H., K. Olive, G. Weiglein '06]

CMSSM, mGMSB, mAMSB: [S.H., X. Miao, S. Su, G. Weiglein '08]

CNMSSM: [D. Lopez-Fogliani, L. Roszkowski, R. Ruiz de Austri, T. Varley '09]

Finite Unified Theories: [S.H., M. Mondragón, G. Zoupanos '07]

Different methods:

1.) Scanning:

- 3-dim scans (possibly with CDM fixing one dimension)
- multi-dim scans
- multi-dim scans (with **Markov Chain Monte Carlo** technique)

⇒ here: results using **last two**

2.) Fitting:

- Frequentist
- Bayesian

⇒ focus on **Frequentist** here

⇒ χ^2 function to include all experimental results

3.) Priors . . . (**none**)

χ^2 calculation:

→ global χ^2 likelihood function

combines all theoretical predictions with experimental constraints:

$$\chi^2 = \sum_i^N \frac{(C_i - P_i)^2}{\sigma(C_i)^2 + \sigma(P_i)^2} + \sum_i^M \frac{(f_{SM_i}^{obs} - f_{SM_i}^{fit})^2}{\sigma(f_{SM_i})^2}$$

N : number of observables studied

M : SM parameters: $\Delta\alpha_{had}, m_t, M_Z$

C_i : experimentally measured value (constraint)

P_i : MSSM parameter-dependent prediction for the corresponding constraint

Assumption: measurements are uncorrelated - fulfilled to a high degree

χ^2 calculation:

→ global χ^2 likelihood function

combines all theoretical predictions with experimental constraints:

$$\chi^2 = \sum_i^N \frac{(C_i - P_i)^2}{\sigma(C_i)^2 + \sigma(P_i)^2} + \sum_i^M \frac{(f_{SM_i}^{obs} - f_{SM_i}^{fit})^2}{\sigma(f_{SM_i})^2}$$

N : number of observables studied

M : SM parameters: $\Delta\alpha_{had}, m_t, M_Z$

C_i : experimentally measured value (constraint)

P_i : MSSM parameter-dependent prediction for the corresponding constraint

Assumption: measurements are uncorrelated - fulfilled to a high degree

What to do if only a lower/upper bound exists?

→ especially important: M_h

→ backup

The models: 1.) CMSSM (or mSUGRA):

⇒ Scenario characterized by

$$m_0, m_{1/2}, A_0, \tan\beta, \text{sign } \mu$$

m_0 : universal scalar mass parameter

$m_{1/2}$: universal gaugino mass parameter

A_0 : universal trilinear coupling

$\tan\beta$: ratio of Higgs vacuum expectation values

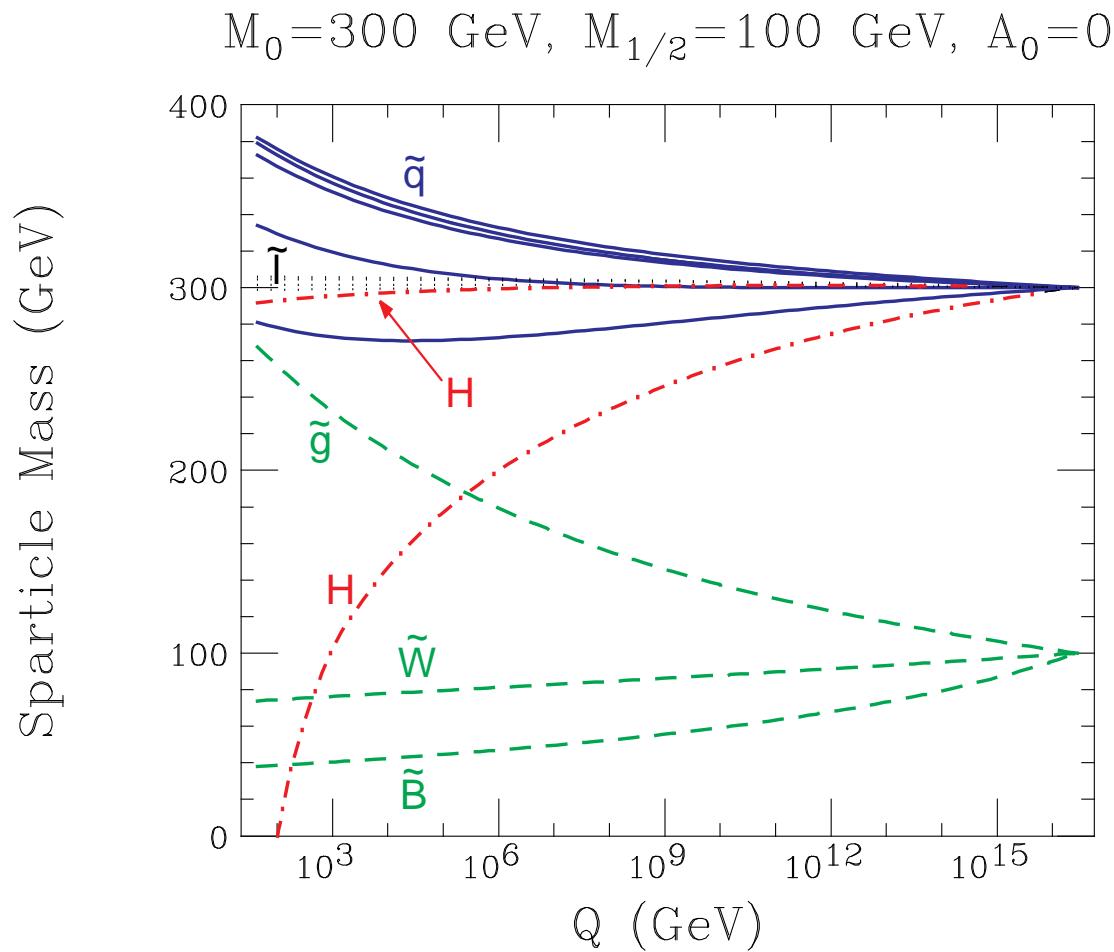
$\text{sign}(\mu)$: sign of supersymmetric Higgs parameter

} at the GUT scale

⇒ particle spectra from renormalization group running to weak scale

⇒ Lightest SUSY particle (LSP) is the lightest neutralino

⇒ particle spectra from renormalization group running to weak scale

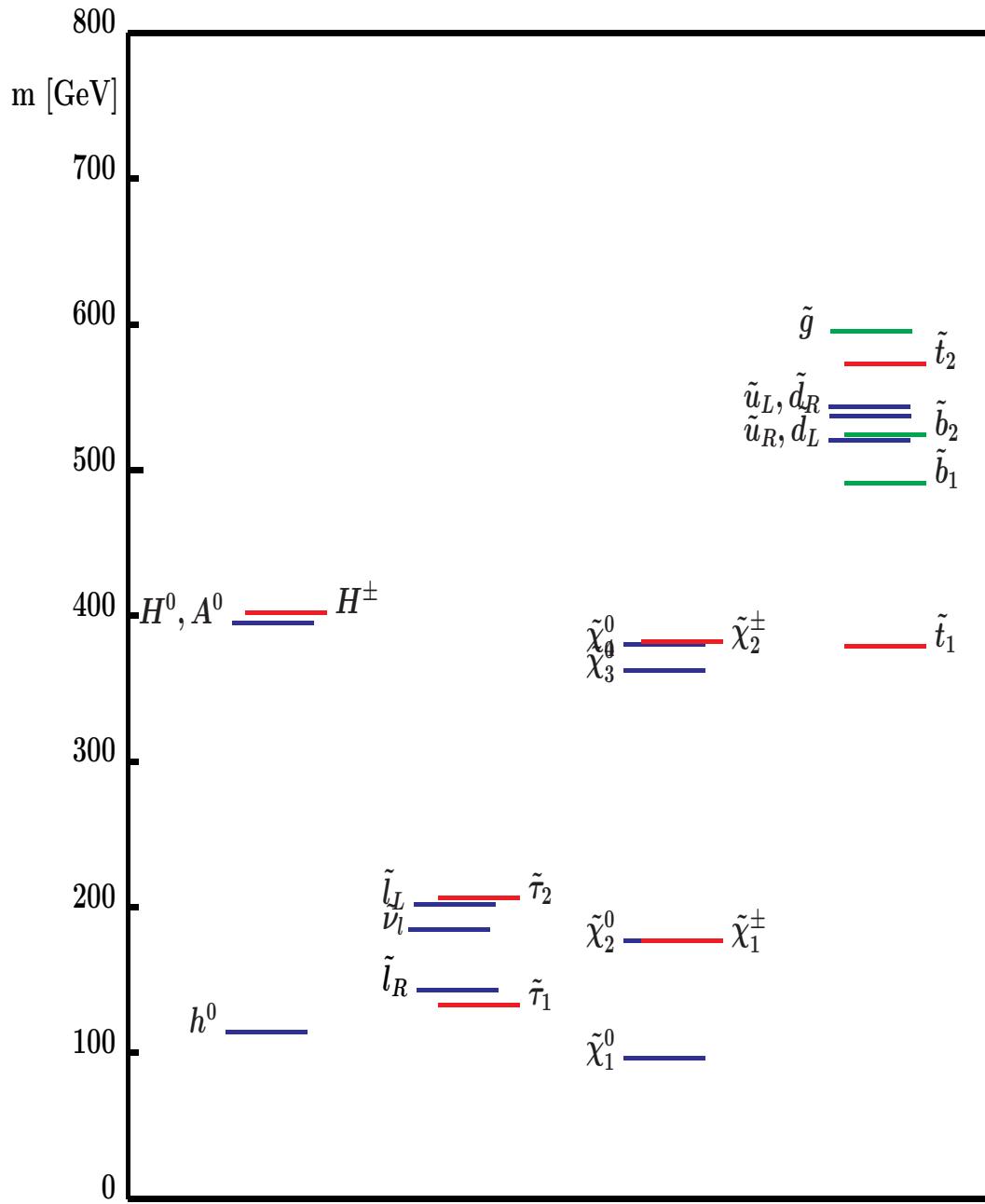


⇒ one parameter turns negative ⇒ Higgs mechanism for free

“Typical” CMSSM scenario
(SPS 1a benchmark scenario):

SPS home page:

www.ippp.dur.ac.uk/~georg/sps



The models: 2.) NUHM1: (Non-universal Higgs mass model)

Assumption: no unification of scalar fermion and scalar Higgs parameter at the GUT scale

⇒ effectively M_A or μ as free parameters at the EW scale

⇒ besides the CMSSM parameters

M_A or μ

Further extension: NUHM2:

Assumption: no unification of the Higgs parameters at the GUT scale

⇒ effectively M_A and μ as free parameters at the EW scale

⇒ besides the CMSSM parameters

M_A and μ

Our tool:

The “MasterCode”



⇒ collaborative effort of theorists and experimentalists

[*Buchmüller, Cavanaugh, De Roeck, Ellis, Flächer, Hahn, SH, Isidori, Olive, Ronga, Weiglein*]

Über-code for the combination of different tools:

- tools are included as **subroutines**
- **compatibility** ensured by collaboration of authors of “MasterCode” and authors of “sub tools” /SLHA(2)
- one “MasterCode” for one model . . .

⇒ evaluate observables of one parameter point consistently with various tools

cern.ch/mastercode

Status of the “MasterCode”:

- one model: (MFV) MSSM (see below)
- tools included:
 - *B-physics* observables [*SuFla*]
 - more *B-physics* observables [*SuperIso*]
 - Higgs related observables, $(g - 2)_\mu$ [*FeynHiggs*]
 - Electroweak precision observables [*FeynWZ*]
 - Dark Matter observables [*MicrOMEGAs*, *DarkSUSY*]
 - for GUT scale models: RGE running [*SoftSusy*]

⇒ all most-up-to-date codes on the market!

- added: χ^2 analysis code [*Minuit*]
- currently being implemented:
 - Higgs constraints (for χ^2 contributions . . .) [*HiggsBounds*]
- planned: inclusion of more tools / more models

Status of the “MasterCode”:

- one model: (MFV) MSSM (see below)
- tools included:
 - *B-physics* observables [*SuFla*]
 - more *B-physics* observables [*SuperIso*]
 - Higgs related observables, $(g - 2)_\mu$ [*FeynHiggs*]
 - Electroweak precision observables [*FeynWZ*]
 - Dark Matter observables [*MicrOMEGAs*, *DarkSUSY*]
 - for GUT scale models: RGE running [*SoftSusy*]

⇒ all most-up-to-date codes on the market!

⇒ crucial for precision!

- added: χ^2 analysis code [*Minuit*]
- currently being implemented:
 - Higgs constraints (for χ^2 contributions . . .) [*HiggsBounds*]
- planned: inclusion of more tools / more models

3. Constraining the lightest MSSM Higgs mass M_h

Contrary to the SM: M_h is not a free parameter

MSSM tree-level bound: $M_h < M_Z$, excluded by LEP Higgs searches

Large radiative corrections:

Dominant one-loop corrections:

→ more in my “discussion session”

$$\Delta M_h^2 \sim G_\mu m_t^4 \log \left(\frac{m_{\tilde{t}_1} m_{\tilde{t}_2}}{m_t^2} \right)$$

The MSSM Higgs sector is connected to all other sector via loop corrections (especially to the scalar top sector)

Measurement of M_h , Higgs couplings ⇒ test of the theory

LHC: $\Delta M_h \approx 0.2$ GeV, ILC: $\Delta M_h \approx 0.05$ GeV

⇒ M_h will be (the best?) electroweak precision observable

Fit of M_h in Supersymmetry?

Fit of M_h in Supersymmetry?

Advantages of fits in the MSSM vs. SM

- $(g - 2)_\mu$ can be used as a constraint
- Cold Dark Matter can be used as a constraint
- $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$ can be used as a constraint
- M_h can be predicted from other parameters
⇒ stronger constraints possible

Fit of M_h in Supersymmetry?

Advantages of fits in the MSSM vs. SM

- $(g - 2)_\mu$ can be used as a constraint
- Cold Dark Matter can be used as a constraint
- $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$ can be used as a constraint
- M_h can be predicted from other parameters
⇒ stronger constraints possible

Disadvantages of fits in the MSSM vs. SM

- many independent mass scales
- M_h can be predicted from other parameters
⇒ more difficult to disentangle effects

Fit of M_h in Supersymmetry?

Advantages of fits in the MSSM vs. SM

- $(g - 2)_\mu$ can be used as a constraint
- Cold Dark Matter can be used as a constraint
- $\text{BR}(B_s \rightarrow \mu^+ \mu^-)$ can be used as a constraint
- M_h can be predicted from other parameters
⇒ stronger constraints possible

Disadvantages of fits in the MSSM vs. SM

- many independent mass scales
- M_h can be predicted from other parameters
⇒ more difficult to disentangle effects

Note: LEP limits on M_h are not included in this (part of the) fit

Prediction of M_h in the CMSSM/NUHM1

[Buchmüller, Cavanaugh, De Roeck, Ellis, Flächer, S.H., Isidori, Olive, Ronga, Weiglein '09]

General idea:

Take the most simple MSSM version: **CMSSM/NUHM1**

→ just three/four GUT scale parameters + $\tan \beta$

Prediction of M_h in the CMSSM/NUHM1

[Buchmüller, Cavanaugh, De Roeck, Ellis, Flächer, S.H., Isidori, Olive, Ronga, Weiglein '09]

General idea:

Take the most simple MSSM version: CMSSM/NUHM1

→ just three/four GUT scale parameters + $\tan\beta$

- combine all electroweak precision data as in the SM (i.e. not M_h)
- combine with B physics observables
- combine with CDM and $(g - 2)_\mu$
- include SM parameters with their errors: m_t, \dots
- scan over the full CMSSM/NUHM1 parameter space

Prediction of M_h in the CMSSM/NUHM1

[Buchmüller, Cavanaugh, De Roeck, Ellis, Flächer, S.H., Isidori, Olive, Ronga, Weiglein '09]

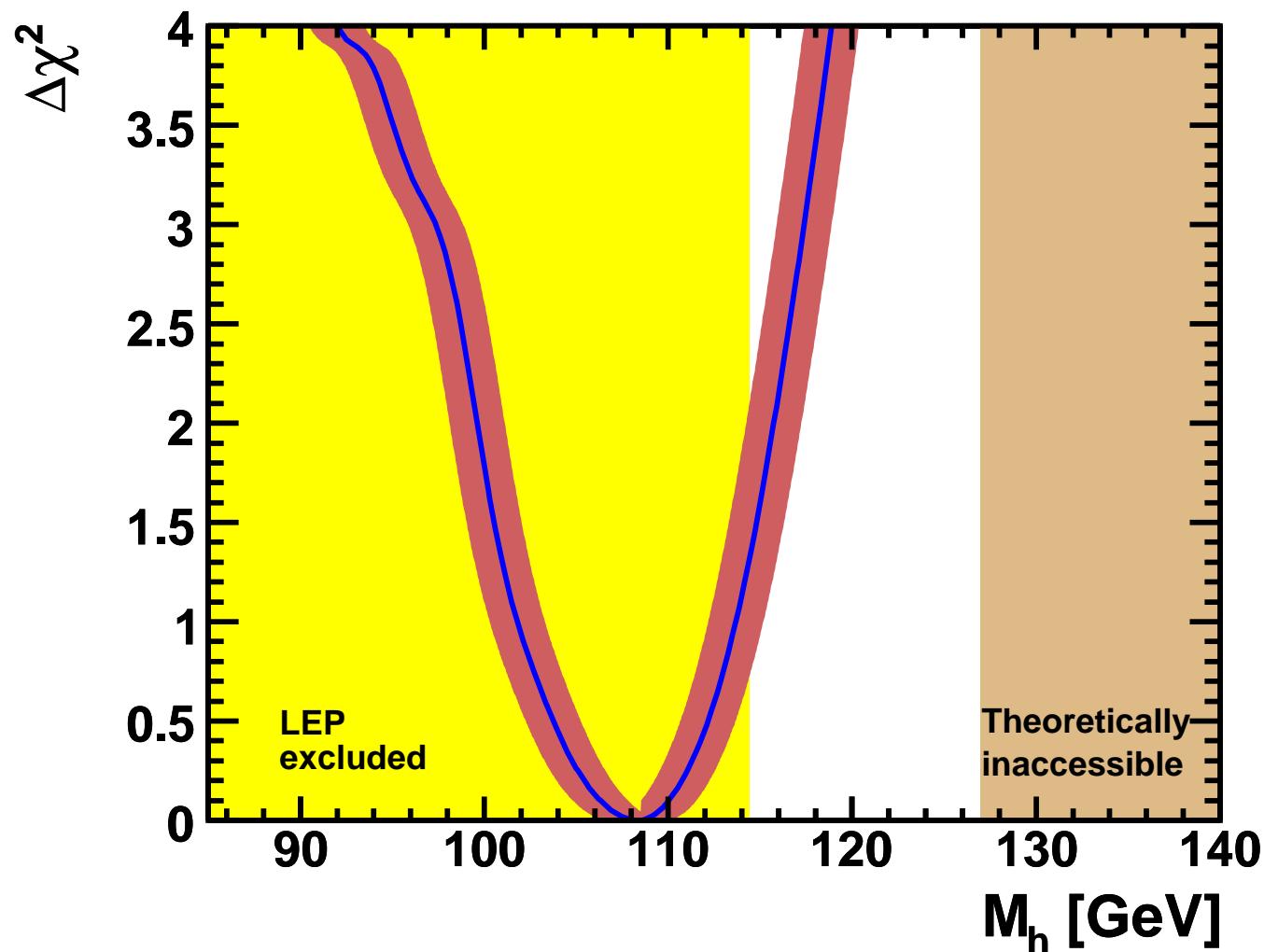
General idea:

Take the most simple MSSM version: CMSSM/NUHM1

→ just three/four GUT scale parameters + $\tan\beta$

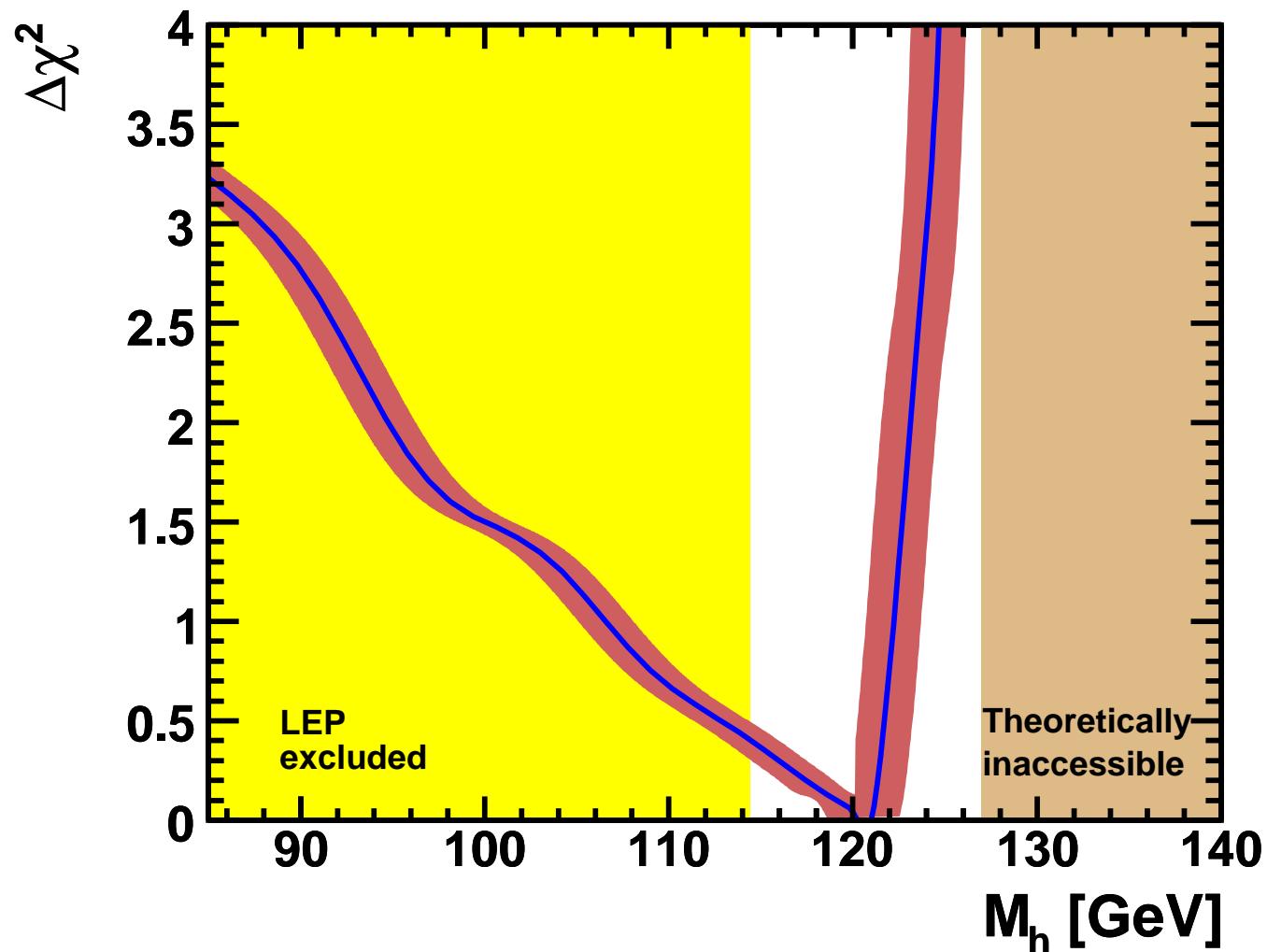
- combine all electroweak precision data as in the SM (i.e. not M_h)
 - combine with B physics observables
 - combine with CDM and $(g - 2)_\mu$
 - include SM parameters with their errors: m_t, \dots
 - scan over the full CMSSM/NUHM1 parameter space
- ⇒ preferred M_h values

CMSSM: red band plot:



$$M_h = 108 \pm 6 \text{ (exp)} \pm 1.5 \text{ (theo)} \text{ GeV}$$

NUHM1: red band plot:

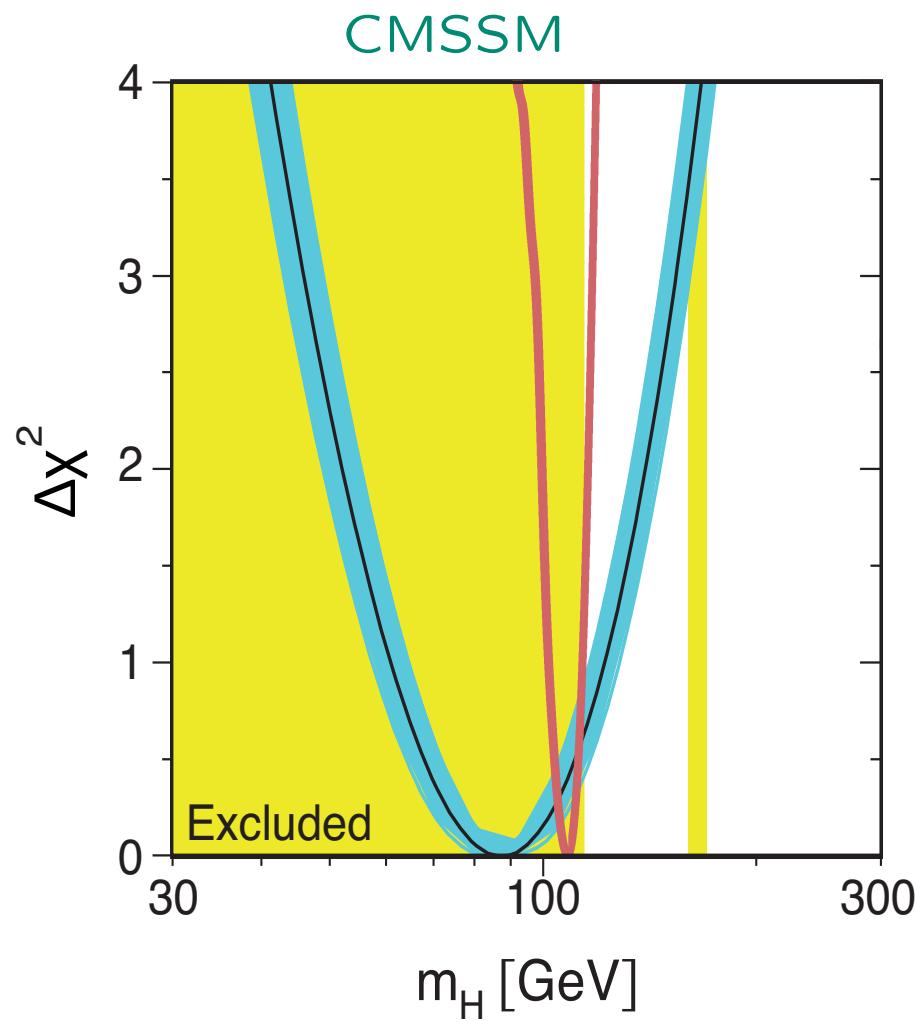


$$M_h = 121^{+1}_{-14} \text{ (exp)} \pm 1.5 \text{ (theo)} \text{ GeV}$$

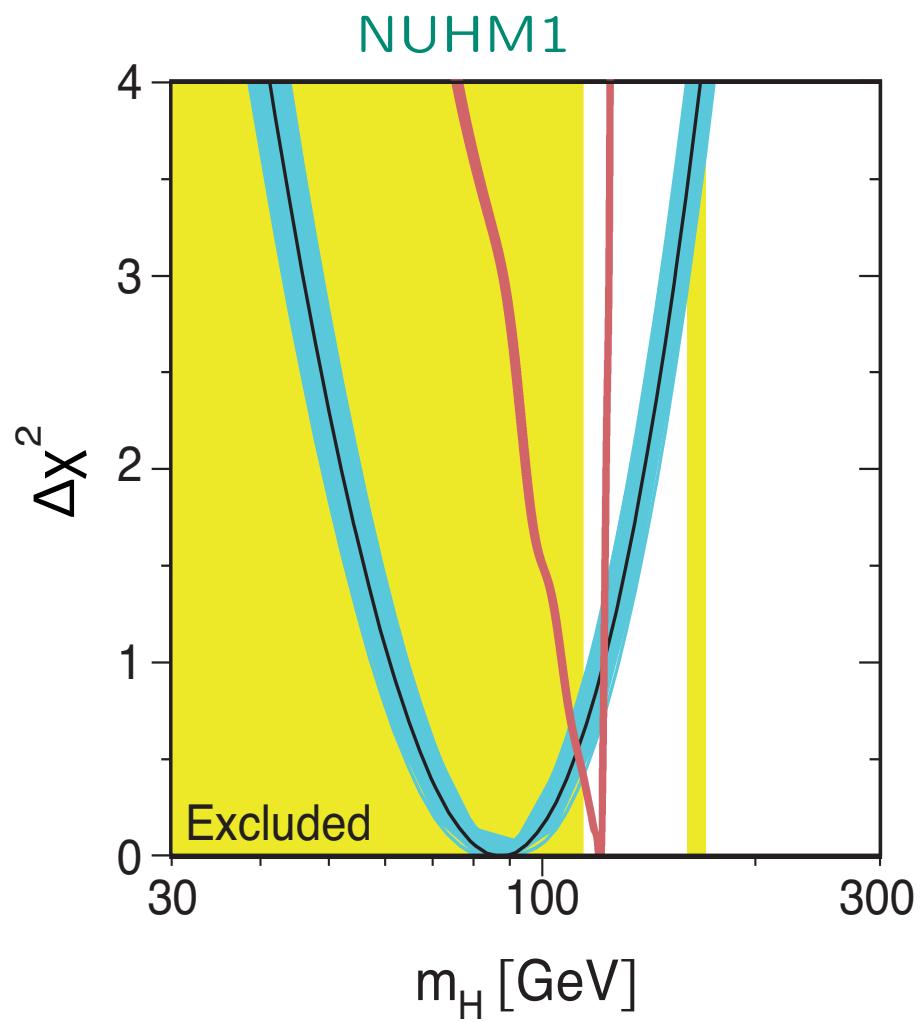
\Rightarrow naturally above LEP limit

Prediction of M_H^{SM} (blue band) and M_h in the MSSM (red band):

[2009]



$M_h^{\text{CMSSM}} = 108 \pm 6 \pm 1.5 \text{ GeV}$
 \Rightarrow as good as the SM



$M_h^{\text{NUHM1}} = 121_{-14}^{+1} \pm 1.5 \text{ GeV}$
 \Rightarrow above the LEP limit

4. LHC reach in the CMSSM/NUHM1

[Buchmüller, Cavanaugh, De Roeck, Ellis, Flächer, S.H., Isidori, Olive, Ronga, Weiglein '09]

- combine all electroweak precision data as in the SM
- combine with B physics observables
- combine with CDM and $(g - 2)_\mu$
- include SM parameters with their errors: m_t , M_Z , $\Delta\alpha_{\text{had}}$

⇒ χ^2 function

→ scan over the full CMSSM/NUHM1 parameter space

~ $2.5 \cdot 10^7$ points samples with MCMC

statistical measure: χ^2 function (Frequentist, no priors)

→ final minimum: Minuit

$\Delta\chi^2$: 68, 95% C.L. contours

⇒ preferred CMSSM/NUHM1 parameters

⇒ $\mathcal{L}_{\text{SUSY}}$

⇒ LHC reach

Best-fit points:

CMSSM:

$m_{1/2} = 310 \text{ GeV}$, $m_0 = 60 \text{ GeV}$, $A_0 = 130 \text{ GeV}$,

$\tan \beta = 11$, $\mu = 400 \text{ GeV}$, $M_A = 450 \text{ GeV}$

$\chi^2/N_{\text{dof}} = 20.6/19$ (36 % probability)

⇒ very similar to SPS 1a :-)

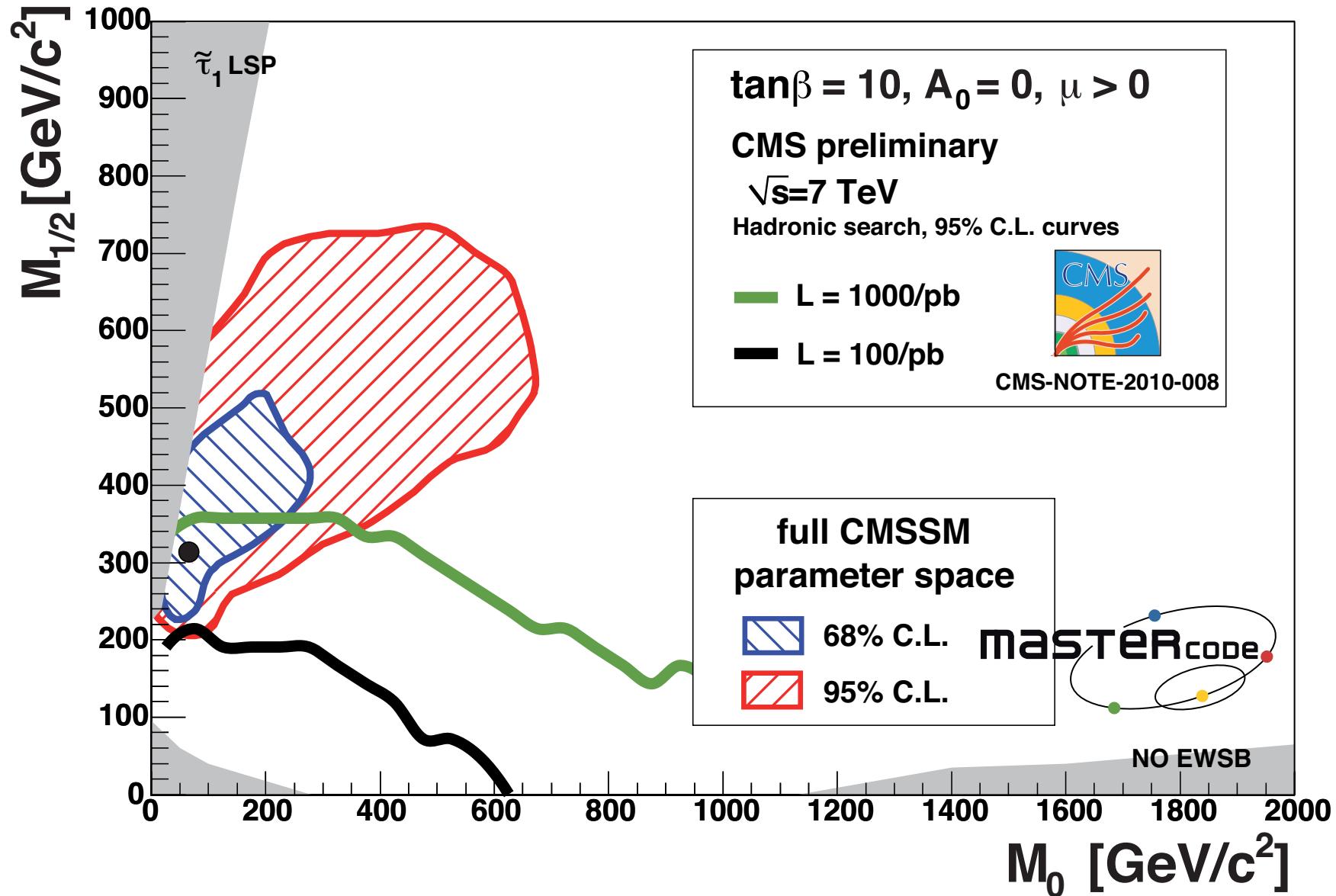
NUHM1:

$m_{1/2} = 270 \text{ GeV}$, $m_0 = 150 \text{ GeV}$, $A_0 = -1300 \text{ GeV}$,

$\tan \beta = 11$, $\mu = 1140 \text{ GeV}$, $M_A = 310 \text{ GeV}$

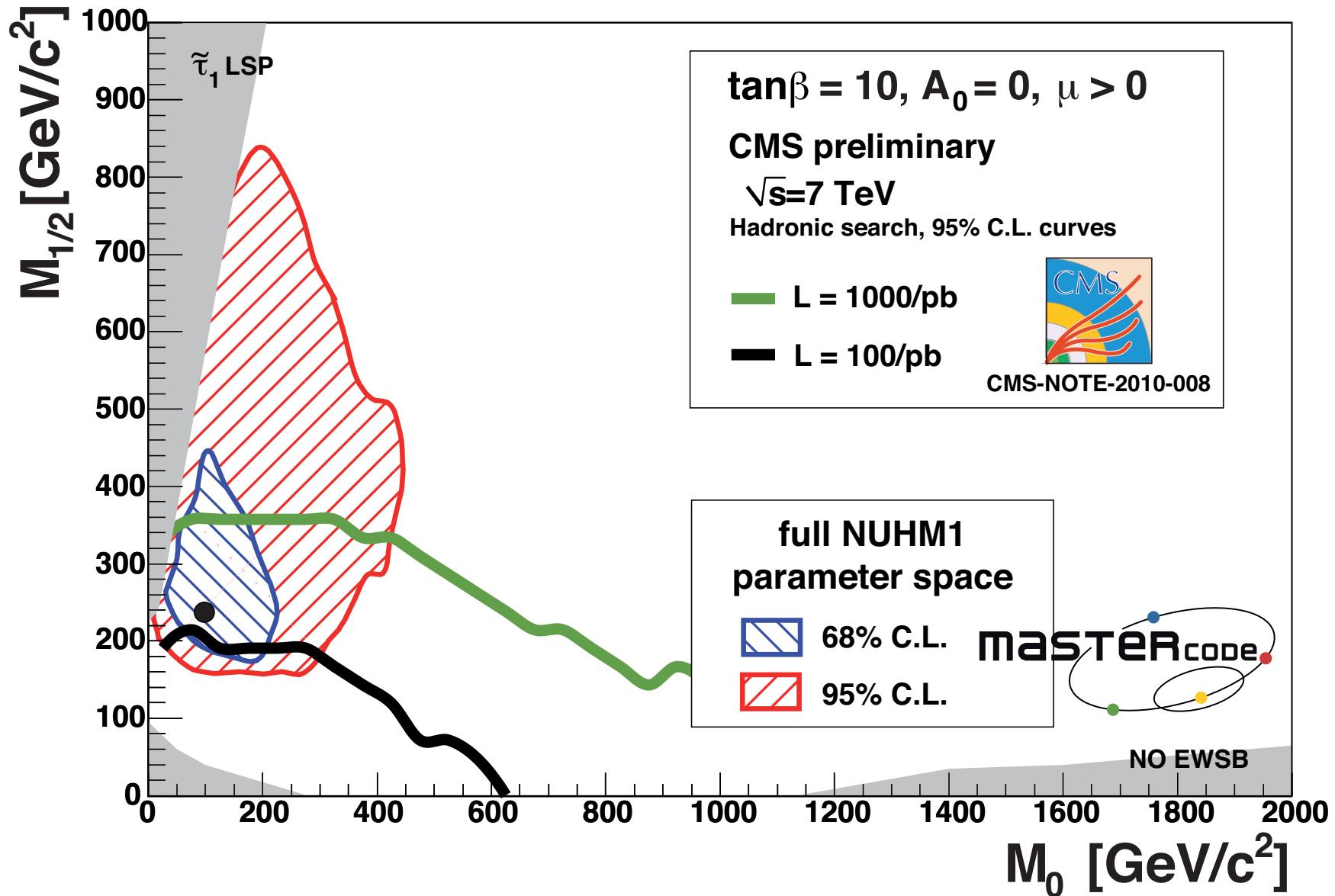
(similar probability)

LHC (CMS) \oplus CMSSM analysis:



⇒ best-fit point and part of 68% C.L. are can be tested in 2011

LHC (CMS) \oplus NUHM1 analysis:

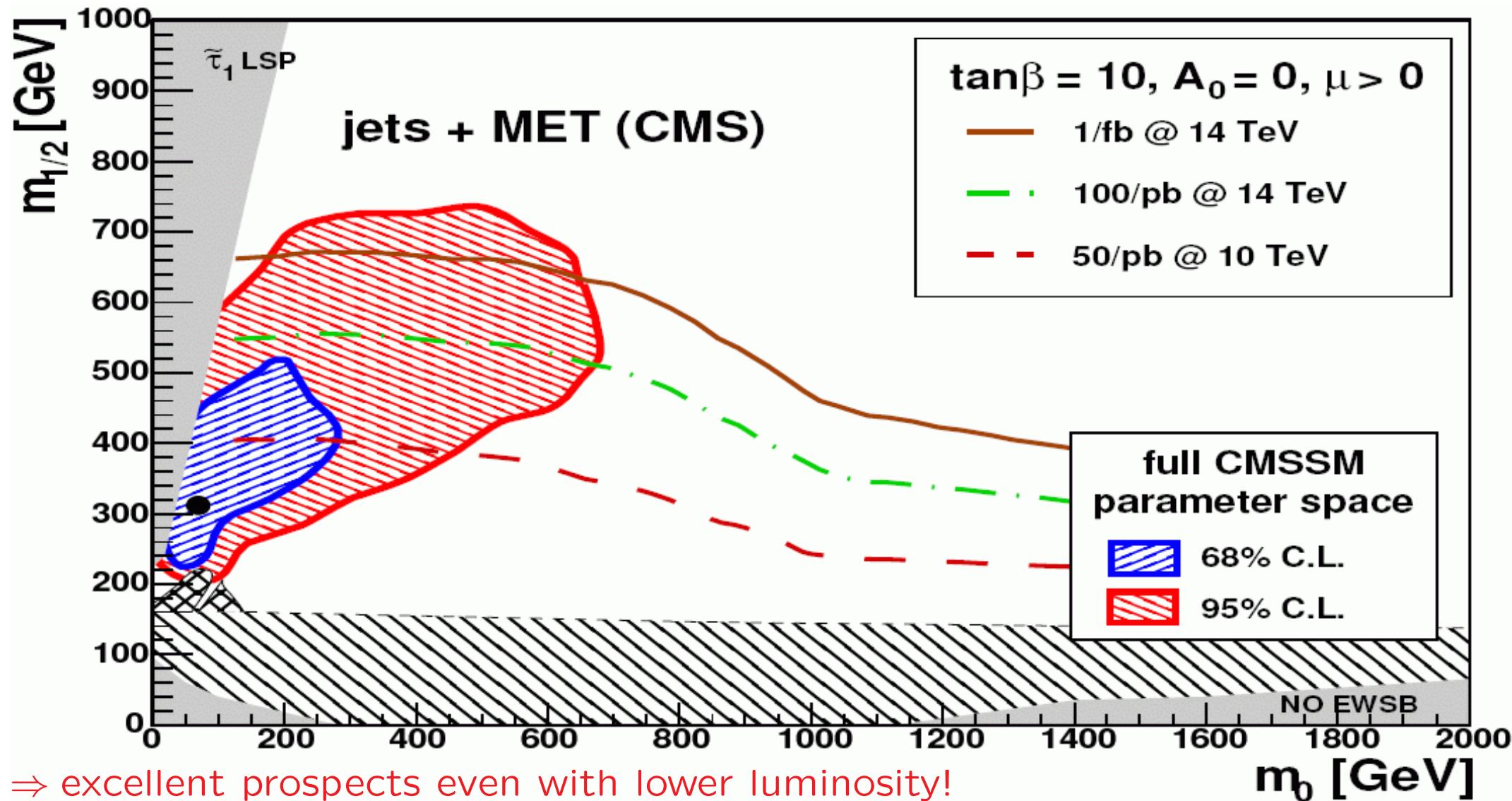


⇒ best-fit point and part of 68% C.L. are can be tested in 2011

LHC (CMS) \oplus CMSSM analysis:

[CMS '07]

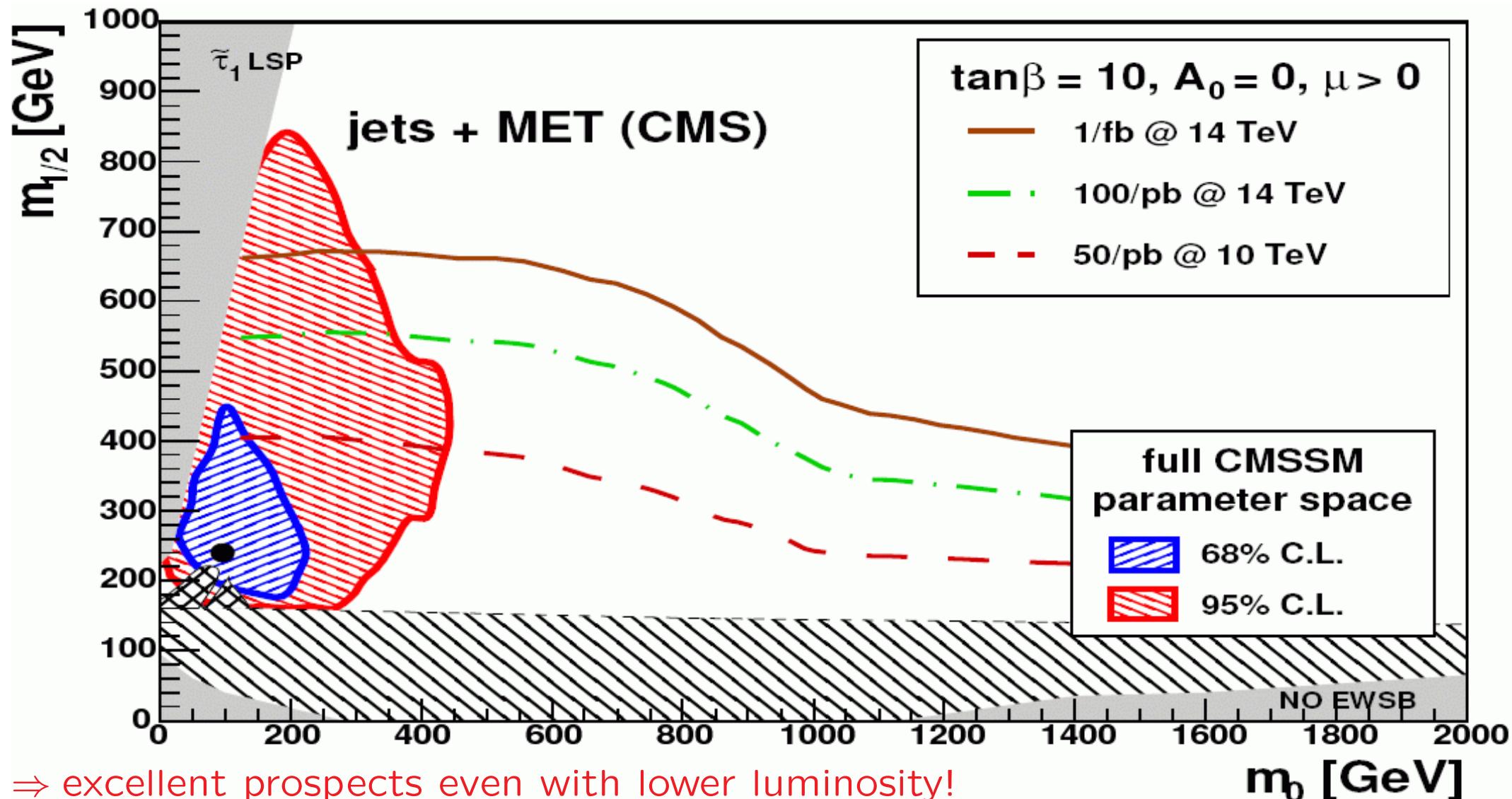
[2008]



LHC (CMS) \oplus NUHM1 analysis:

[CMS '07]

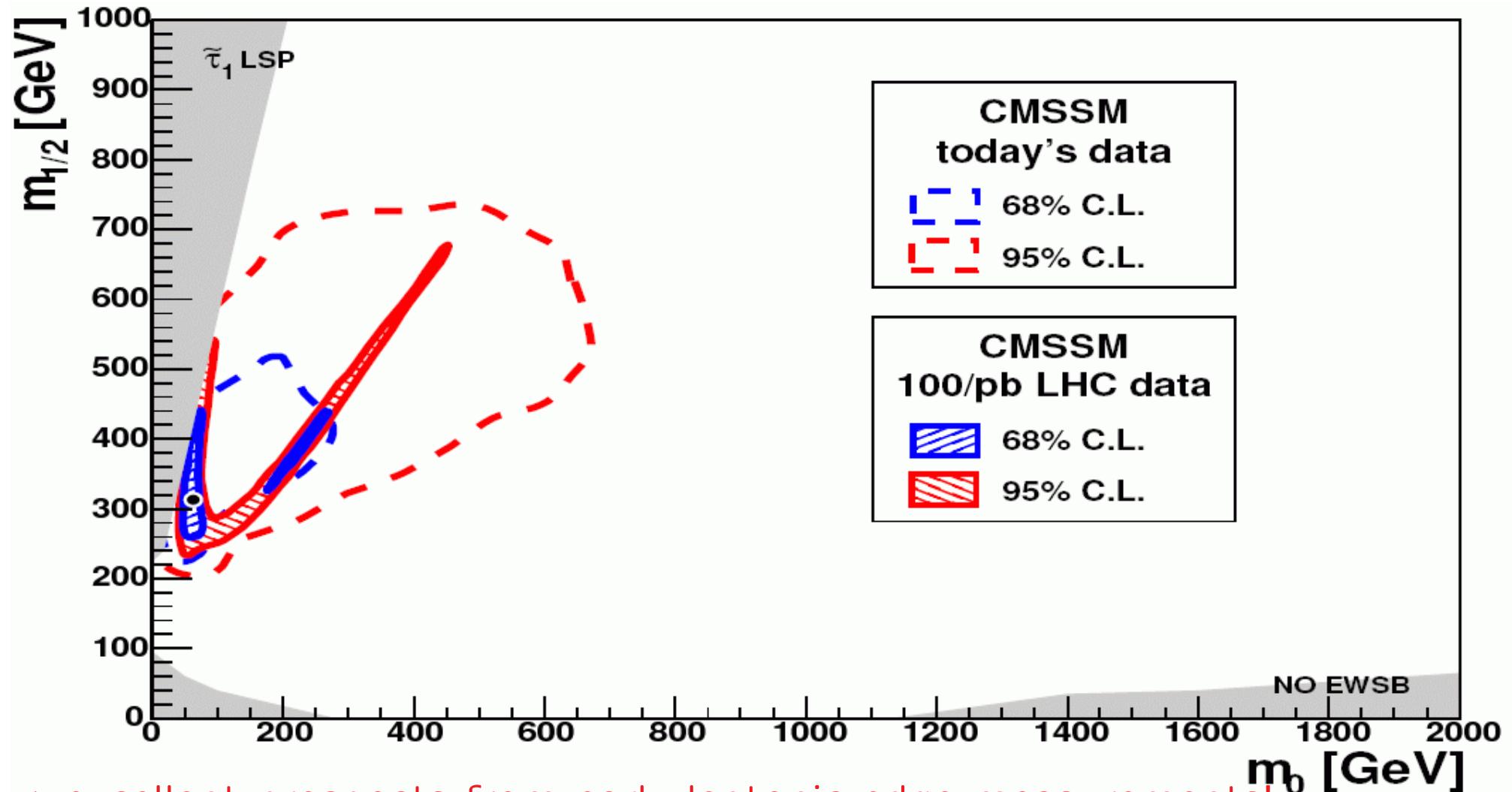
[2008]



[2008]

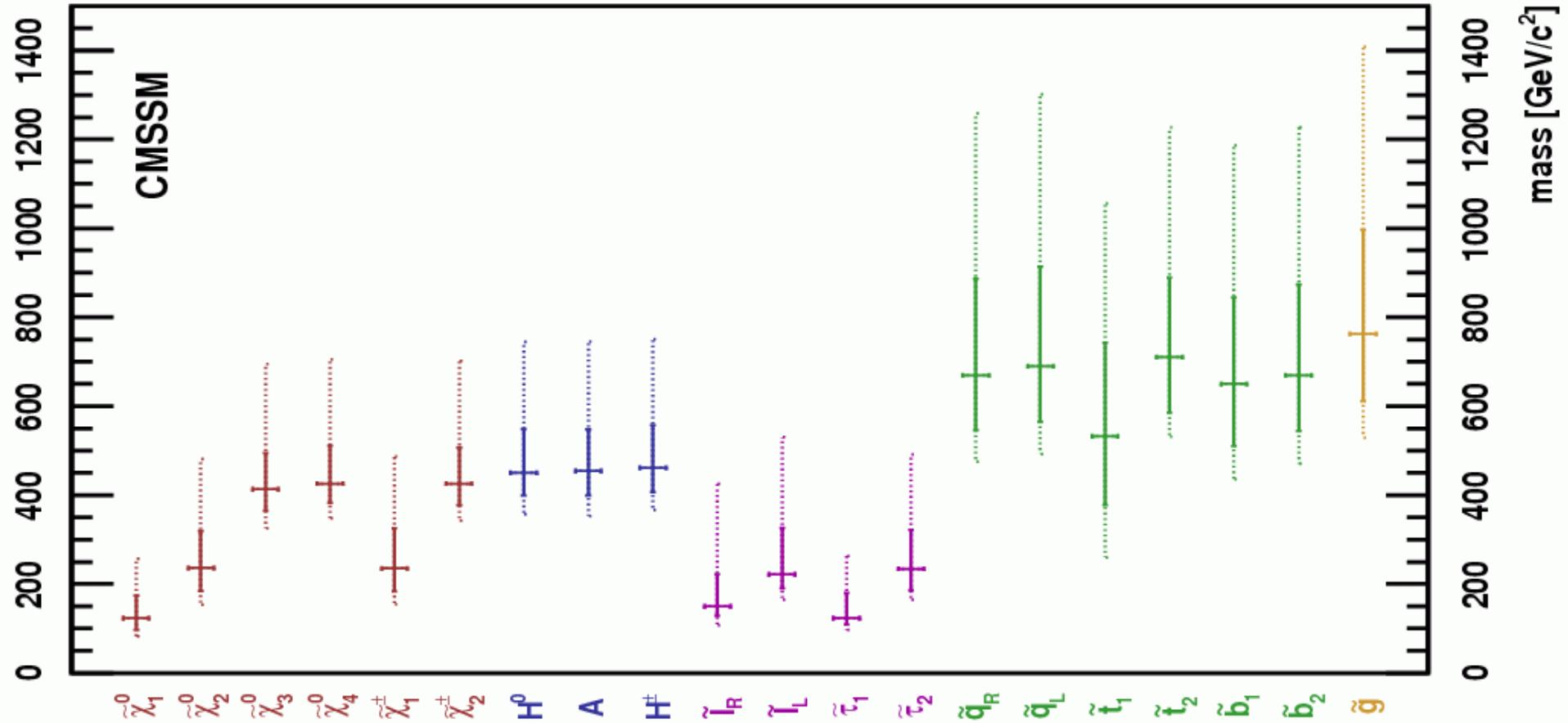
LHC (CMS) \oplus CMSSM analysis:

reach with 1 fb^{-1} @ 14 TeV incl. leptonic edge measurements



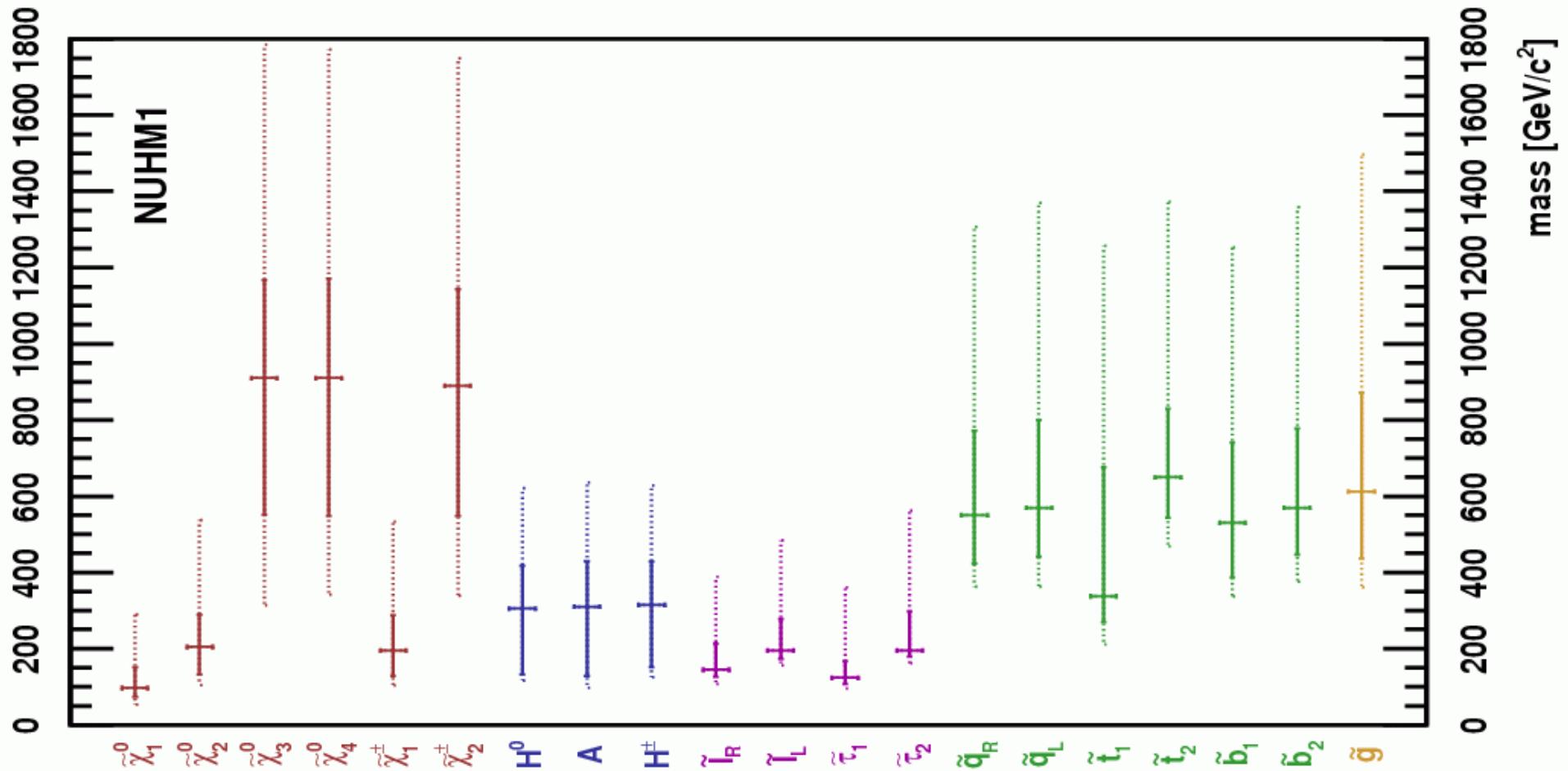
⇒ excellent prospects from early leptonic edge measurements!

Masses for best-fit points: CMSSM



⇒ largely accessible spectrum for LHC (and ILC)

Masses for best-fit points: NUHM1

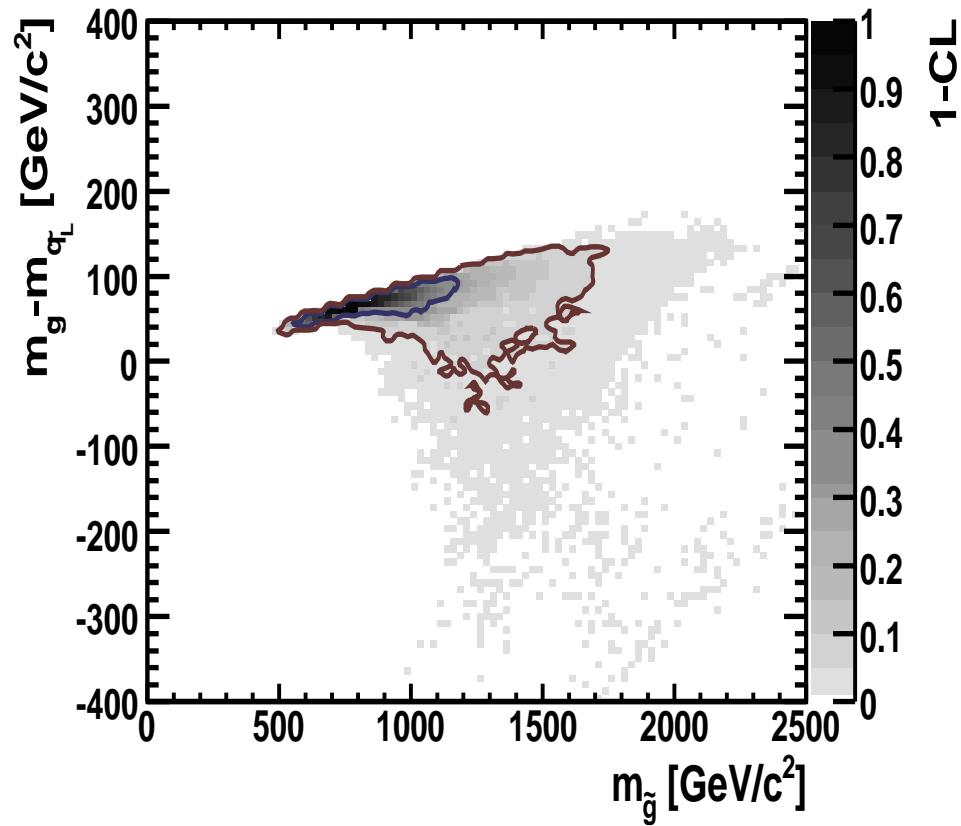


⇒ largely accessible spectrum for LHC (and ILC)

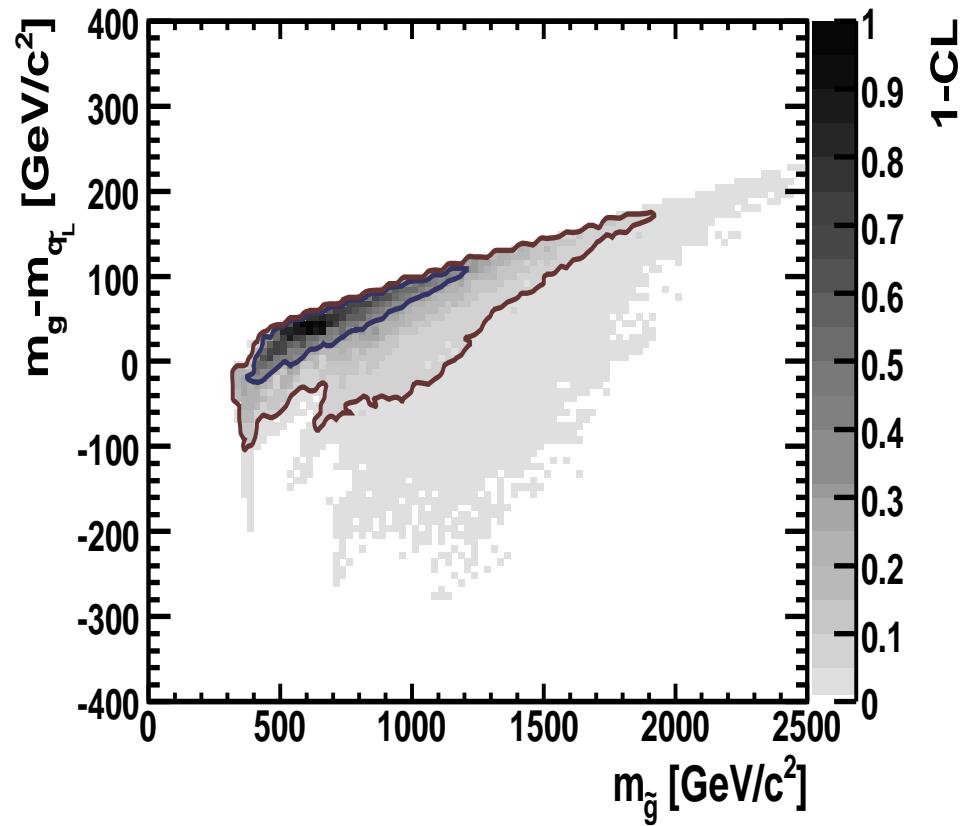
[2009]

Some more predictions: $m_{\tilde{g}} - m_{\tilde{q}_L}$

CMSSM

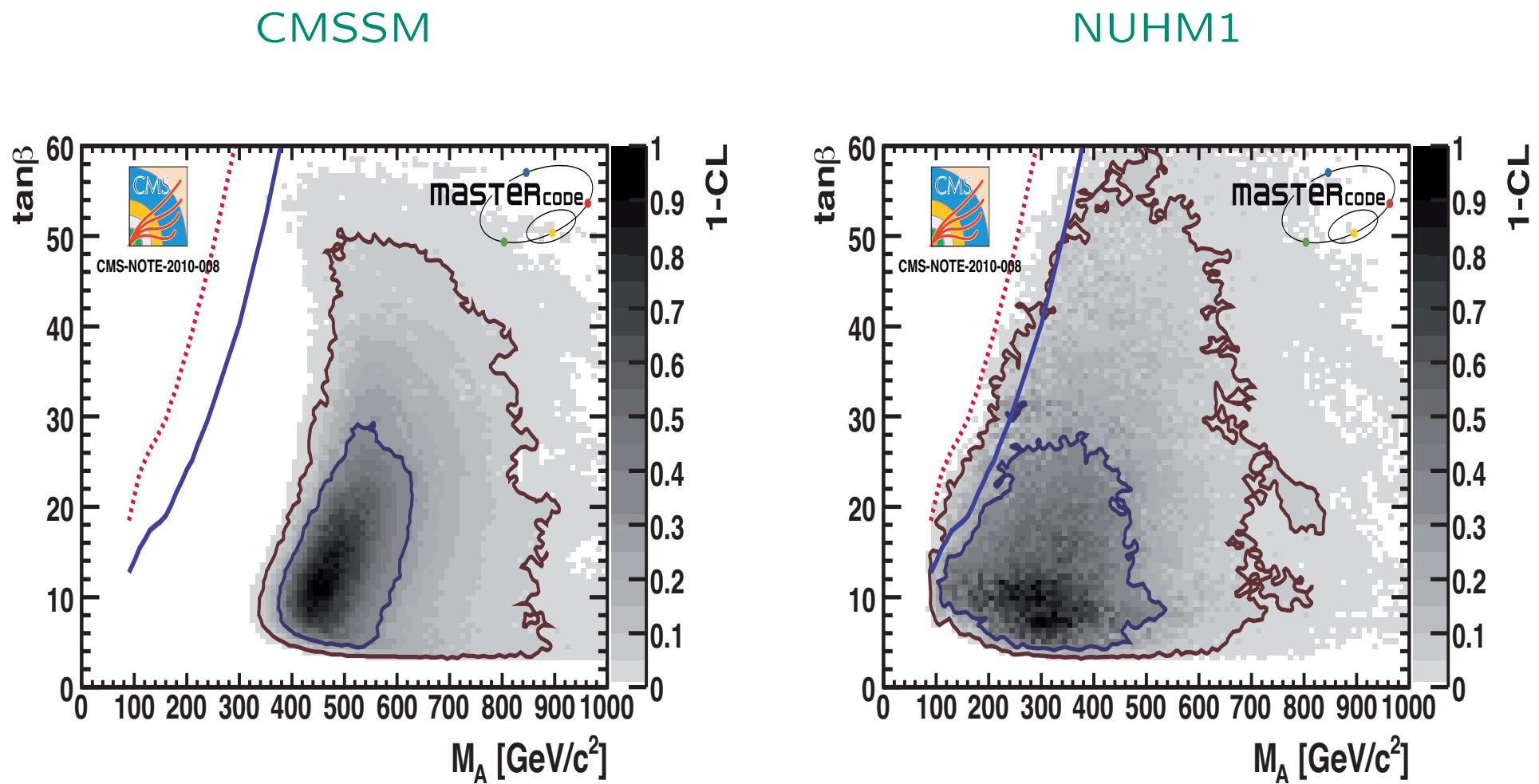


NUHM1



⇒ $m_{\tilde{g}}$ often largest mass, but exceptions are possible

Some more predictions: preferred M_A – $\tan\beta$ parameter space



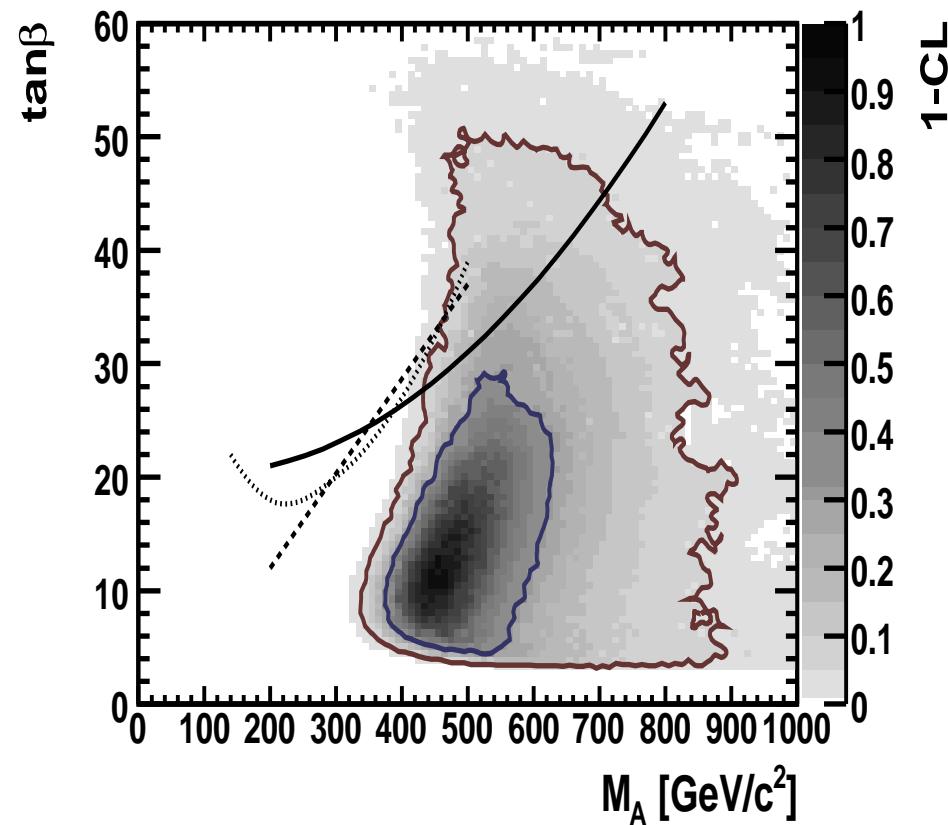
red dotted: discovery with 1 fb^{-1} @ 7 TeV

blue solid: 95% C.L. exclusion with 1 fb^{-1} @ 7 TeV

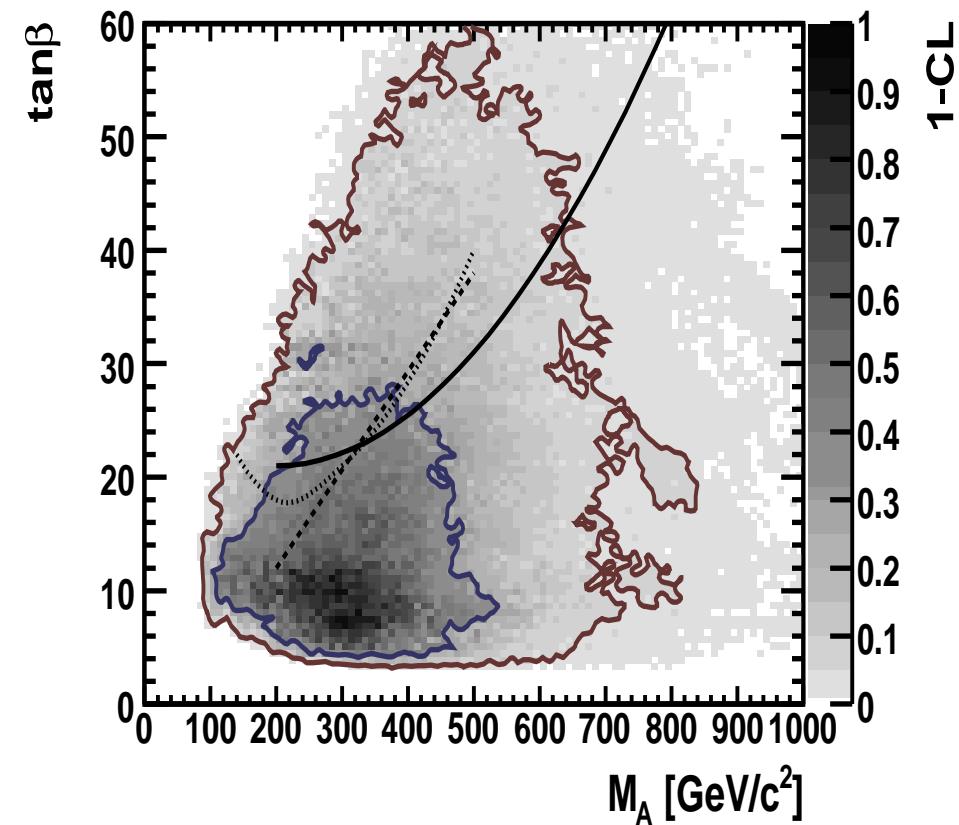
⇒ preferred regions missed in 2010-2011 run

Some more predictions: preferred M_A – $\tan\beta$ parameter space

CMSSM



NUHM1



CMS analysis for 30 fb^{-1} @ 14 TeV

⇒ still best-fit regions missed by LHC (better for ILC(1000))

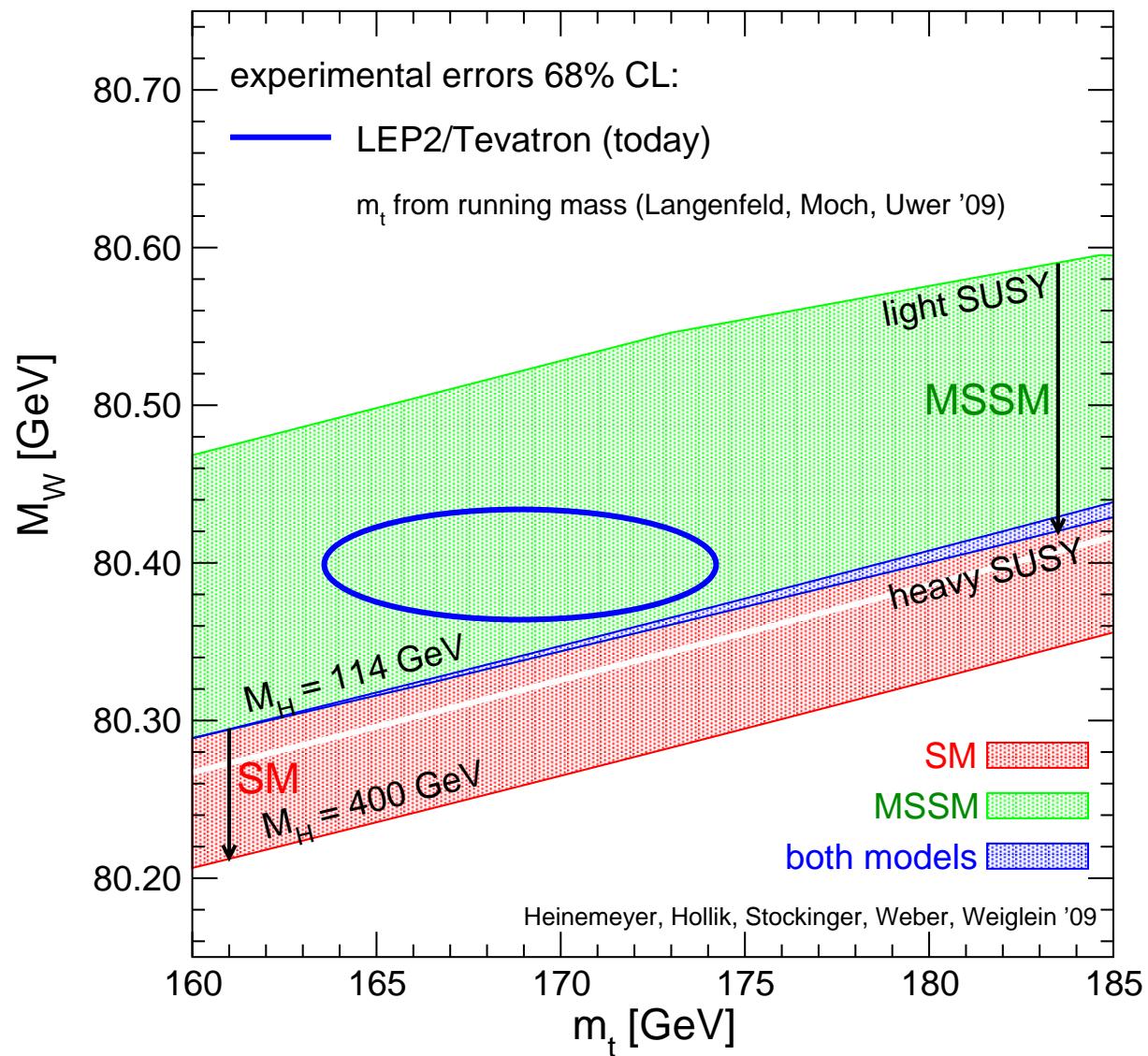
5. Conclusions

- Idea: Predict most probable MSSM parameter regions using existing data: EWPO, BPO, CDM, ...
- Models: CMSSM, NUHM1
- statistical measure: χ^2 function (Frequentist, no priors)
 $\sim 2.5 \cdot 10^7$ points samples with MCMC
 $\Delta\chi^2$: 68, 95% C.L. contours
- Best-fit points:
CMSSM: $m_{1/2} = 310$ GeV, $m_0 = 60$ GeV, $A_0 = 240$ GeV,
 $\tan\beta = 11$, $\mu = 380$ GeV, $M_A = 410$ GeV
 \Rightarrow very similar to SPS 1a :-)
Prediction of M_h (no LEP bound): $M_h = 108 \pm 6 \pm 1.5$ GeV
- NUHM1: $m_{1/2} = 270$ GeV, $m_0 = 150$ GeV, $A_0 = -1300$ GeV,
 $\tan\beta = 11$, $\mu = 1140$ GeV, $M_A = 310$ GeV
Prediction of M_h (no LEP bound): best fit: $M_h \approx 121$ GeV
- 68% C.L. areas: partially covered with $\sim 1 \text{ fb}^{-1}$ @ 7 TeV (u.d.!)
 \Rightarrow early LHC data could be very conclusive!

Back-up

Example: Prediction for M_W in the SM and the MSSM :

[S.H., W. Hollik, D. Stockinger, A. Weber, G. Weiglein '07]



MSSM band:

scan over
SUSY masses

overlap:

SM is MSSM-like
MSSM is SM-like

SM band:

variation of M_H^{SM}

Global fit to all SM data incl. direct searches:

[*GFitter* '09]

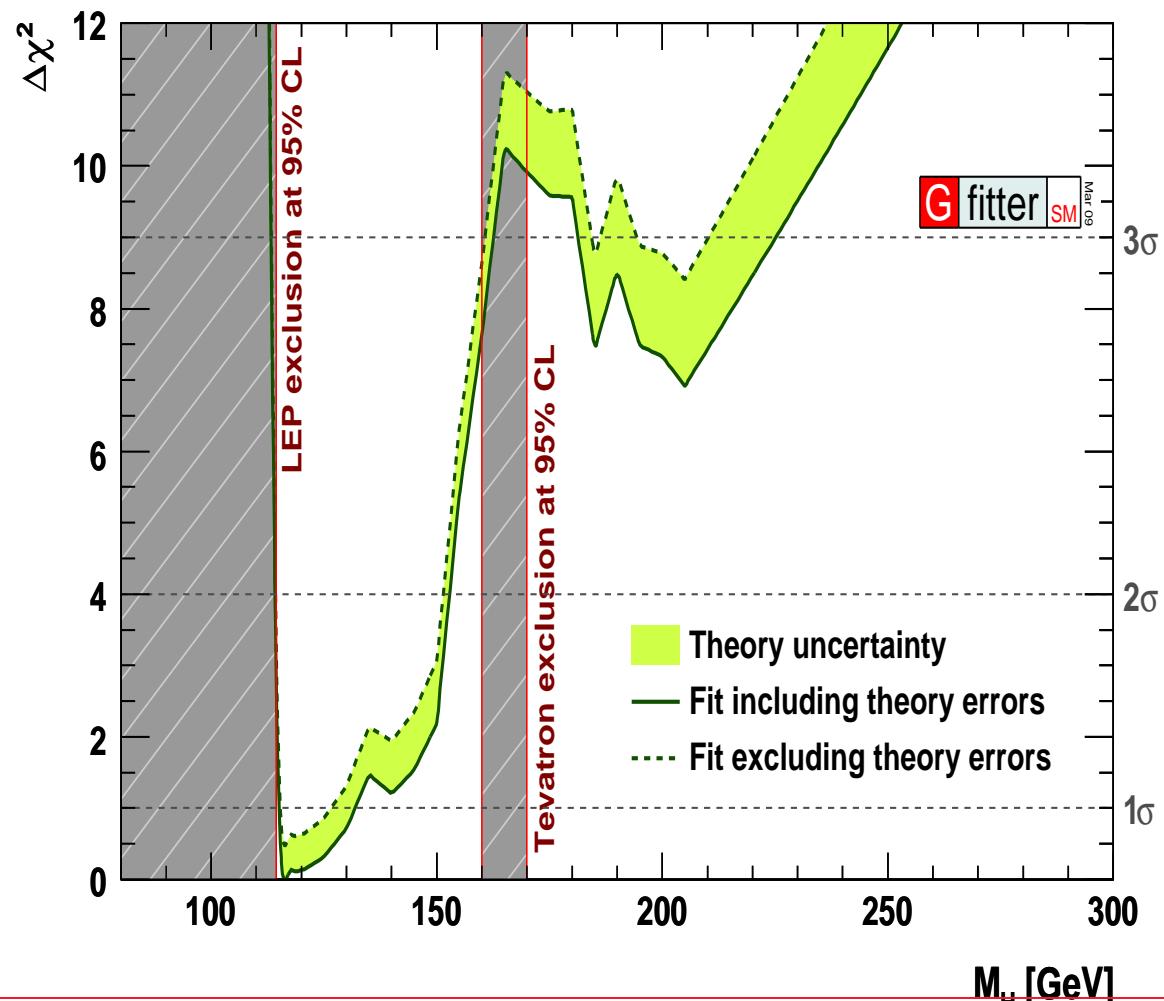
$$\Rightarrow M_H = 116.4^{+18.3}_{-1.4} \text{ GeV}$$

$$M_H < 152 \text{ GeV, 95% C.L.}$$

Assumption for the fit:

SM incl. Higgs boson

\Rightarrow no confirmation of
Higgs mechanism

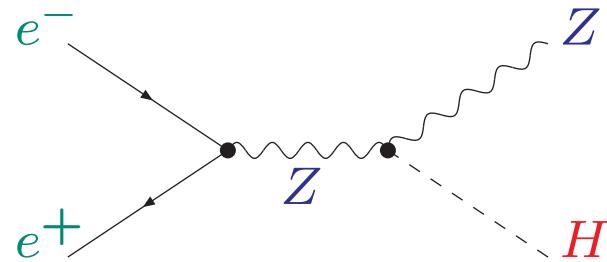


\Rightarrow Higgs boson seems to be light, $M_H \lesssim 150 \text{ GeV}$

SM Higgs search at LEP:

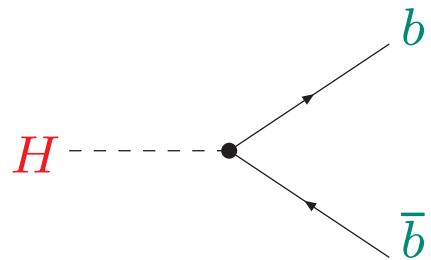
Dominant SM production process:

$$e^+ e^- \rightarrow ZH$$



Dominant decay process:

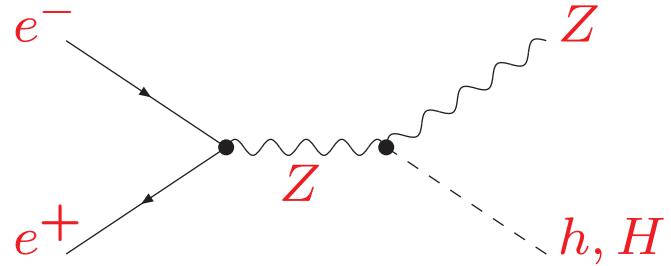
$$H \rightarrow b\bar{b}$$



Bounds valid in the CMSSM? NUHM1? MSSM?

Search for neutral SUSY Higgs bosons:

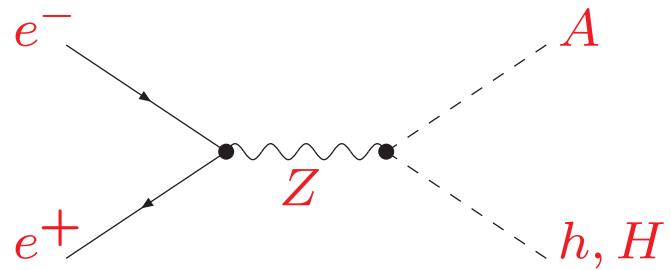
$e^+e^- \rightarrow Zh, ZH$



$$\sigma_{hZ} \approx \sin^2(\beta - \alpha_{\text{eff}}) \sigma_{hZ}^{\text{SM}}$$

$$\sigma_{HZ} \approx \cos^2(\beta - \alpha_{\text{eff}}) \sigma_{hZ}^{\text{SM}}$$

$e^+e^- \rightarrow Ah, AH$



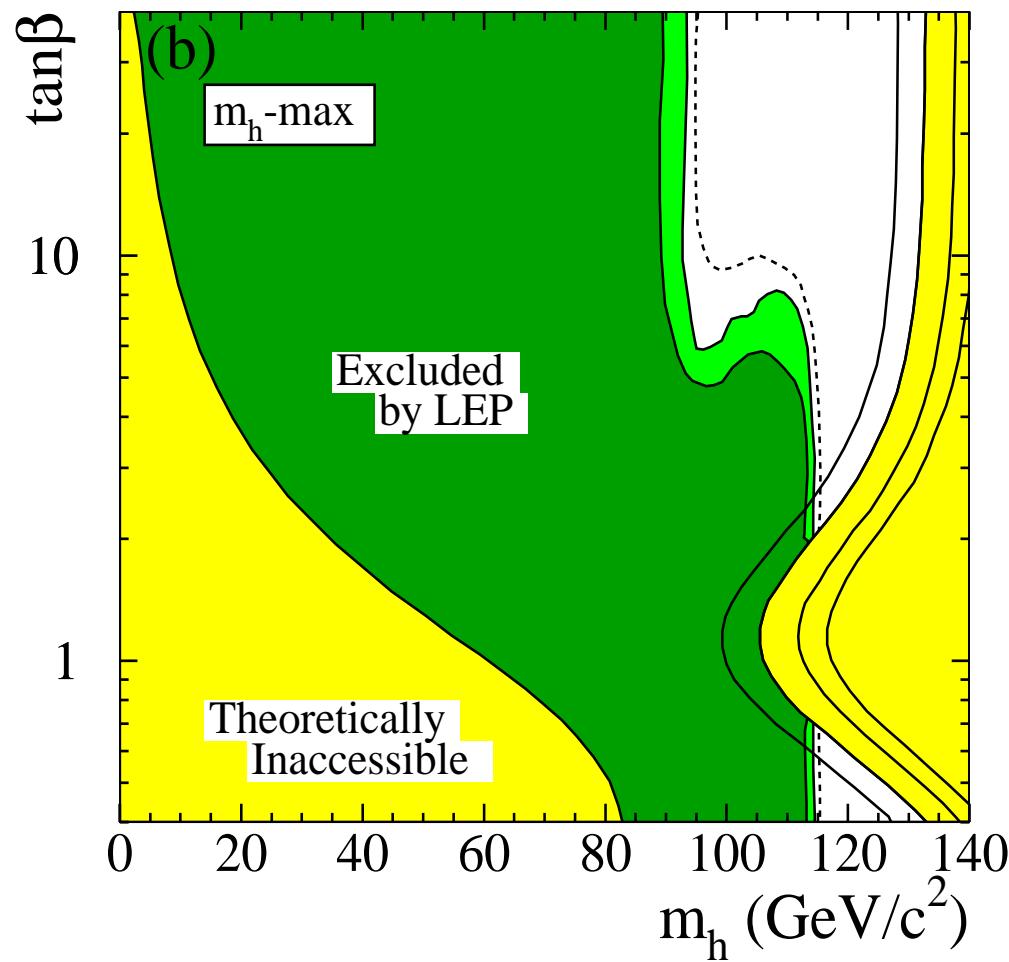
$$\sigma_{hA} \propto \cos^2(\beta - \alpha_{\text{eff}}) \sigma_{hZ}^{\text{SM}}$$

$$\sigma_{HA} \propto \sin^2(\beta - \alpha_{\text{eff}}) \sigma_{hZ}^{\text{SM}}$$

Constraints from the Higgs search at LEP [LEP Higgs Working Group '06]

Experimental search vs. upper M_h -bound (FeynHiggs 2.0)

m_h^{\max} -scenario ($m_t = 174.3$ GeV, $M_{\text{SUSY}} = 1$ TeV):

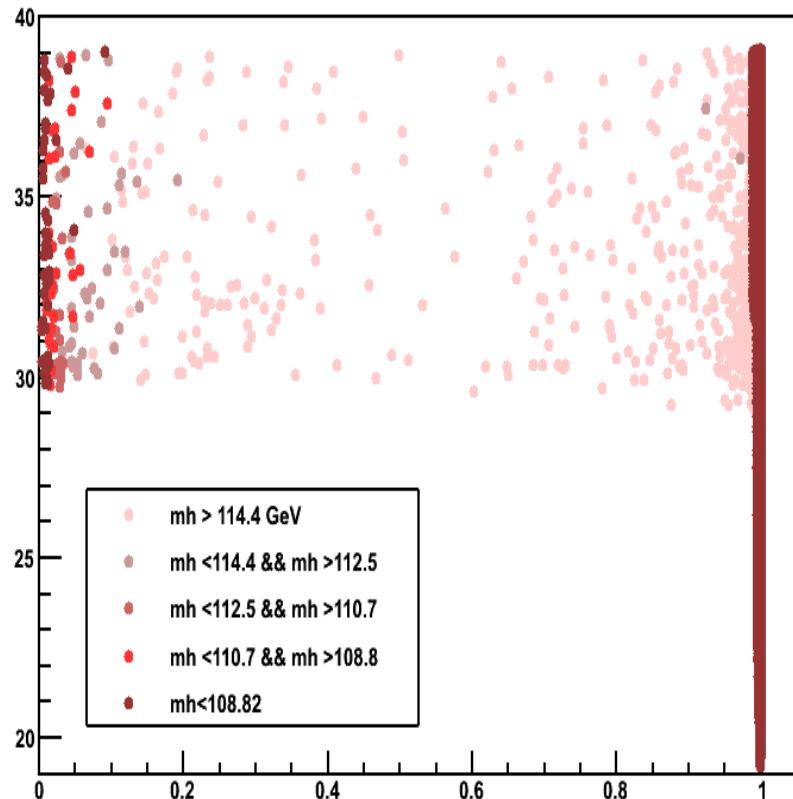


$m_h > 92.8$ GeV
(expected: 94.9 GeV), 95% C.L.

$M_A > 93.4$ GeV
(expected: 95.2 GeV)

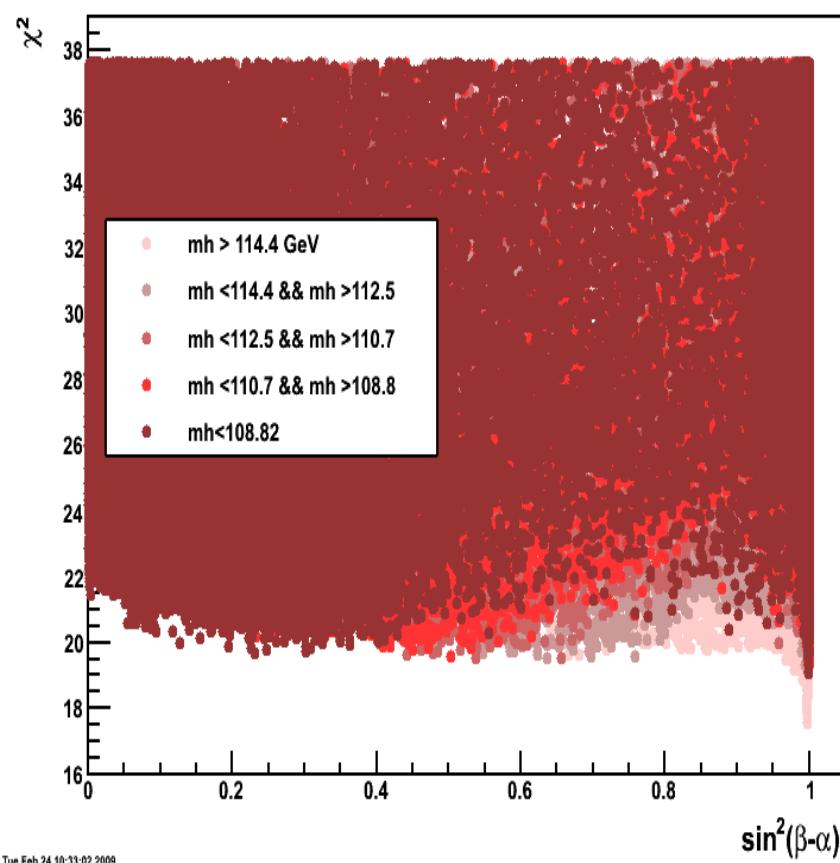
$\sin^2(\beta - \alpha_{\text{eff}})$ in the CMSSM, NUHM1:

CMSSM



Tue Feb 24 14:53:20 2009

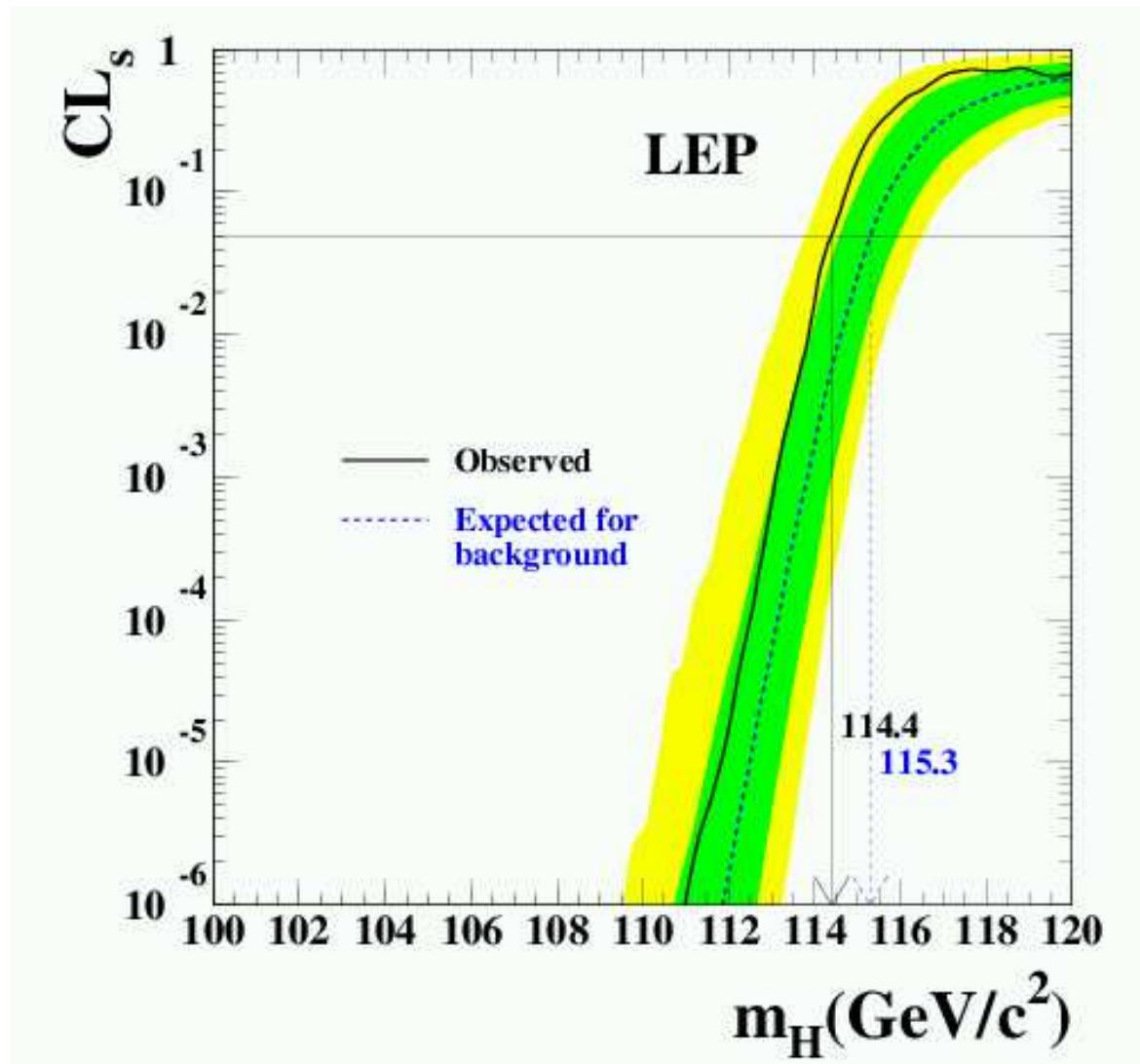
NUHM1



Tue Feb 24 10:33:02 2009

In CMSSM:

SM bound of M_H search can be used [LEP Higgs Working Group '03]



CL_s can be
used/transformed
into χ^2 values

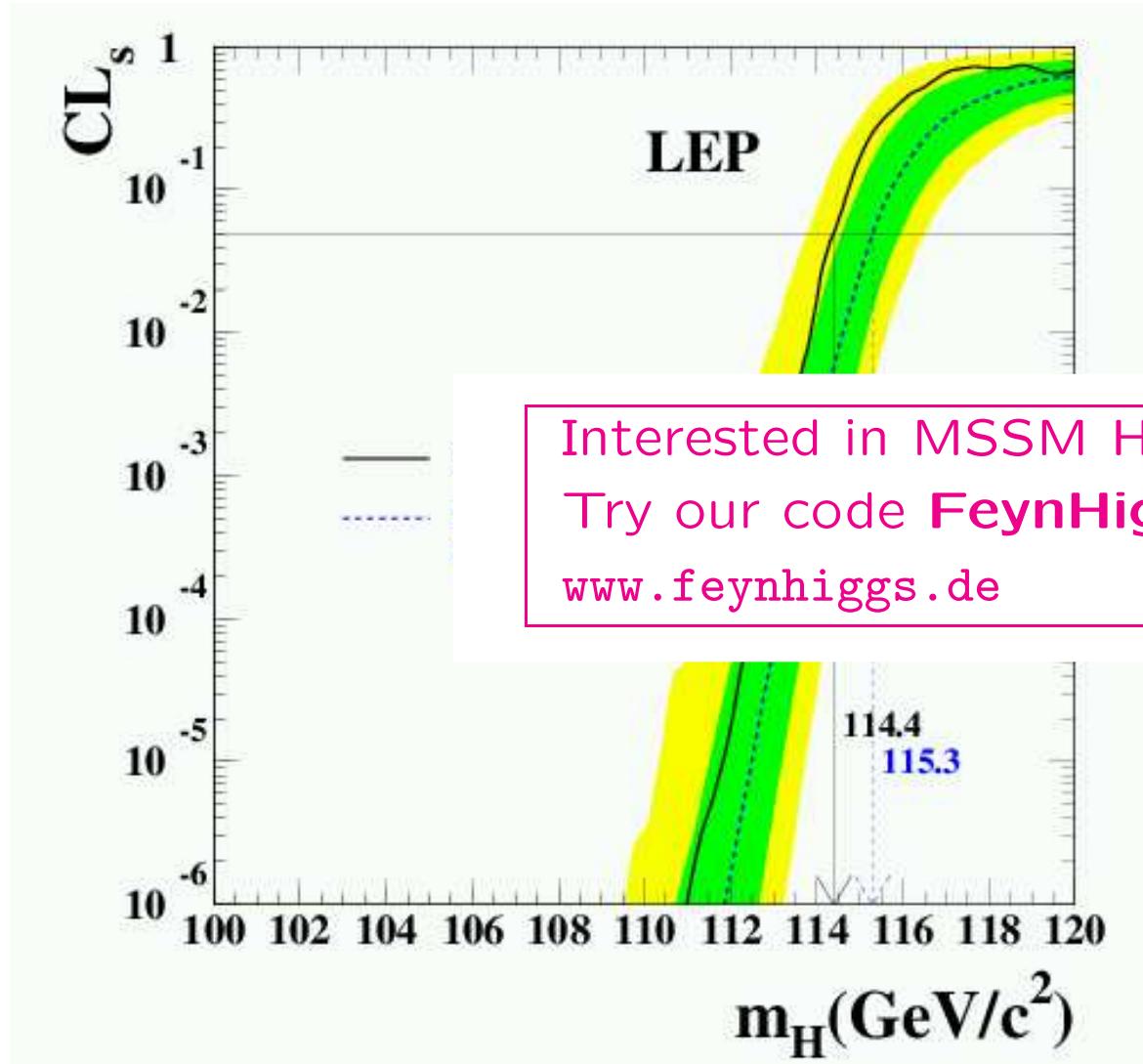
⇒ can be included into
 χ^2 evaluation

$$\delta M_h^{\text{intr.}} \approx 3 \text{ GeV}$$

We use *FeynHiggs*

In CMSSM:

SM bound of M_H search can be used [LEP Higgs Working Group '03]



CL_s can be
used/transformed
into χ^2 values

Interested in MSSM Higgs physics?
Try our code **FeynHiggs**
www.feynhiggs.de

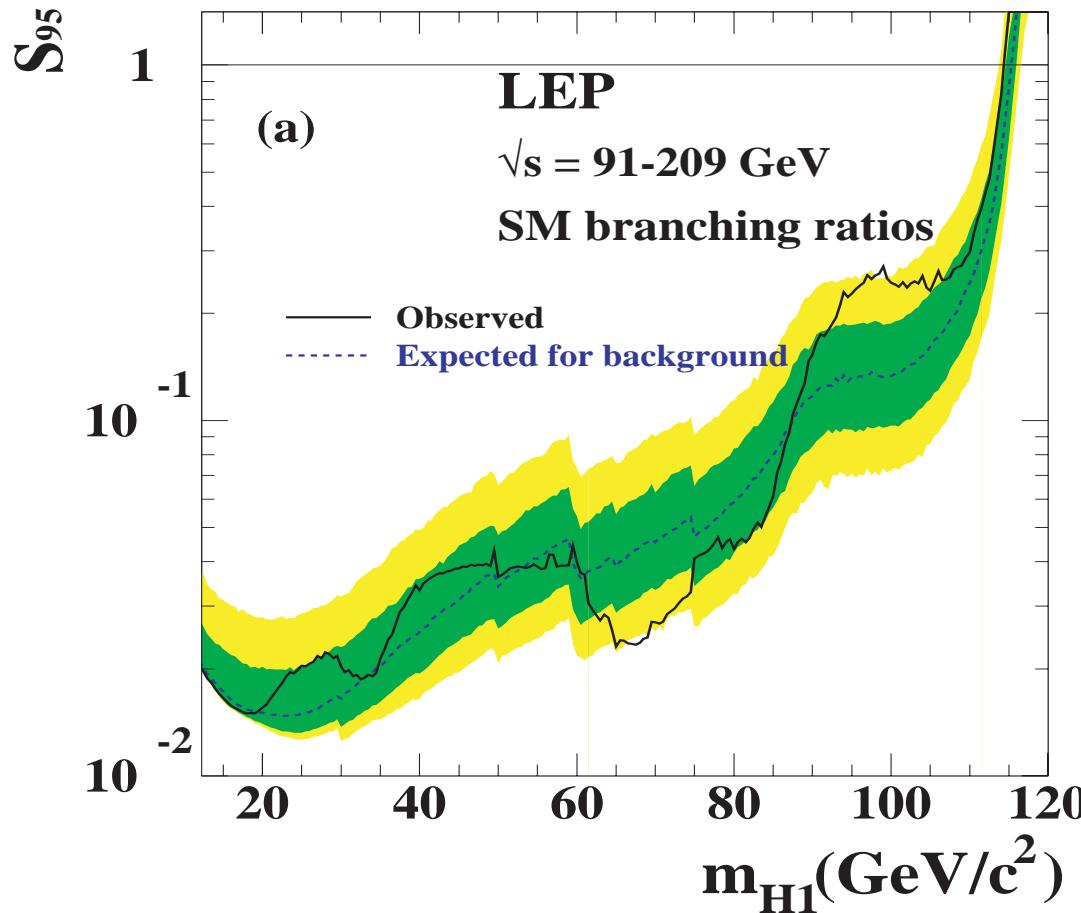
ed into

$$\delta M_h^{WW} \approx 3 \text{ GeV}$$

We use **FeynHiggs**

In the NUHM1:

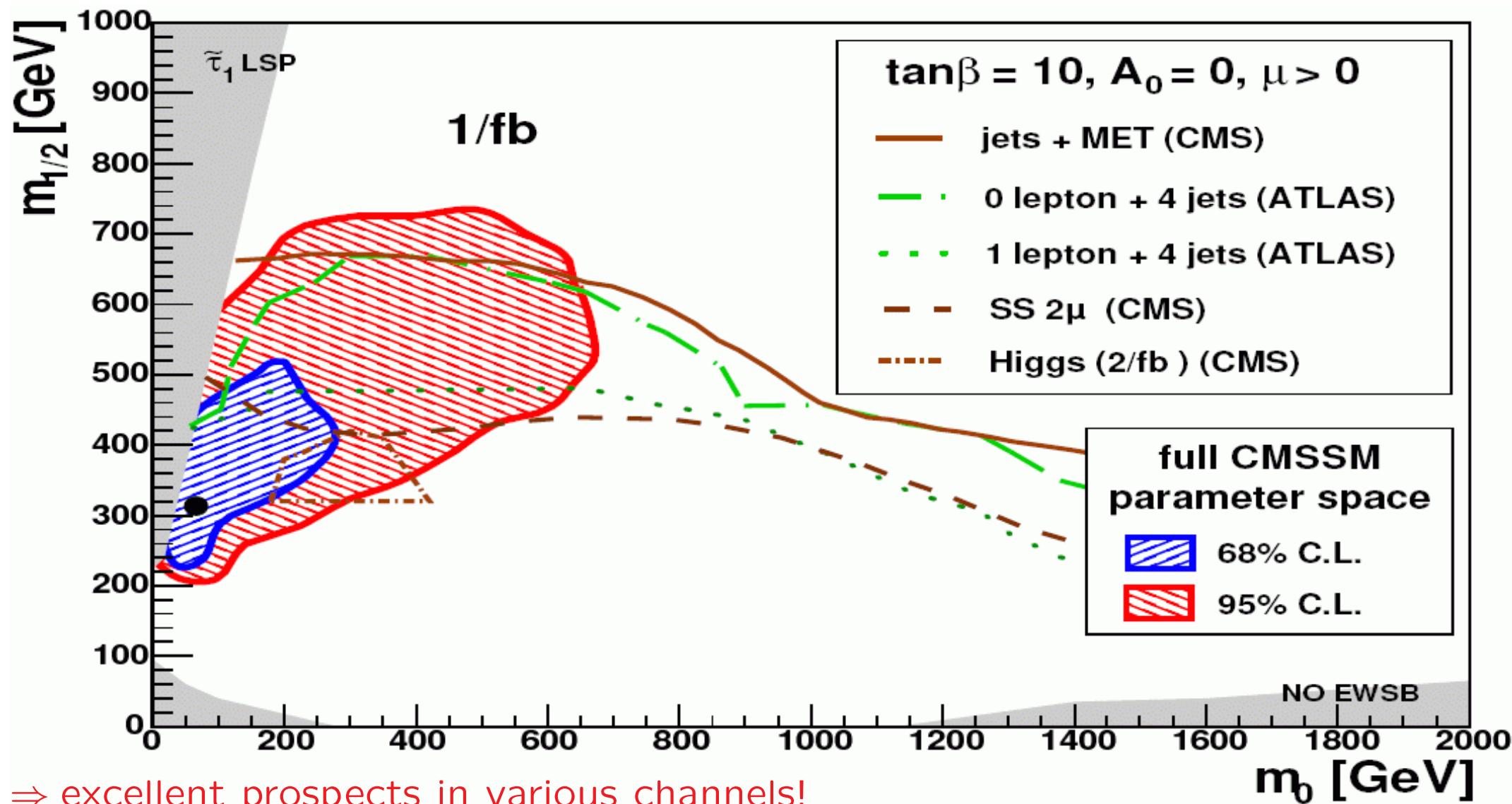
SM bound on M_H is reduced: $S_{95} \sim \sin^2(\beta - \alpha_{\text{eff}})$



⇒ take into account the LEP SM Higgs bound . . .
... but shifted according to the reduced coupling

LHC (CMS) reach with 1 fb^{-1} :

[MasterCode '08] [CMS '07]



LHC (CMS) reach with 1 fb^{-1} : NUHM1 analysis

[MasterCode '08] [CMS '07]

