

A real time algorithm for track finding in ICARUS experiment

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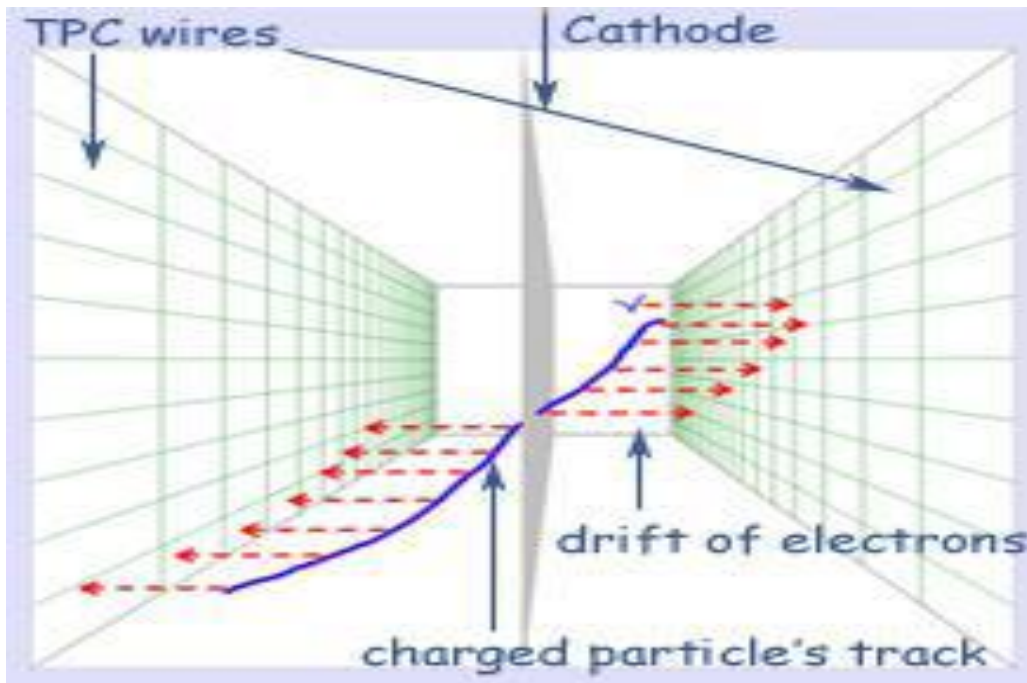
Outline

- Description of the apparatus
- ICARUS physical goals
- The trigger System
- Hit finding algorithm and R.o.I. detection
- Test on Icarino facility

LAr-TPC in a nutshell

Icarus T-600 is the largest LAr TPC operating underground.

It provides calorimetric measurements as well as 3D tracks imaging



LAr acts both as target for neutrinos and as detector itself:

M.i.p. energy loss= 2,2 MeV/cm

electron/ion pairs production
 ≈ 5000 electrons/mm

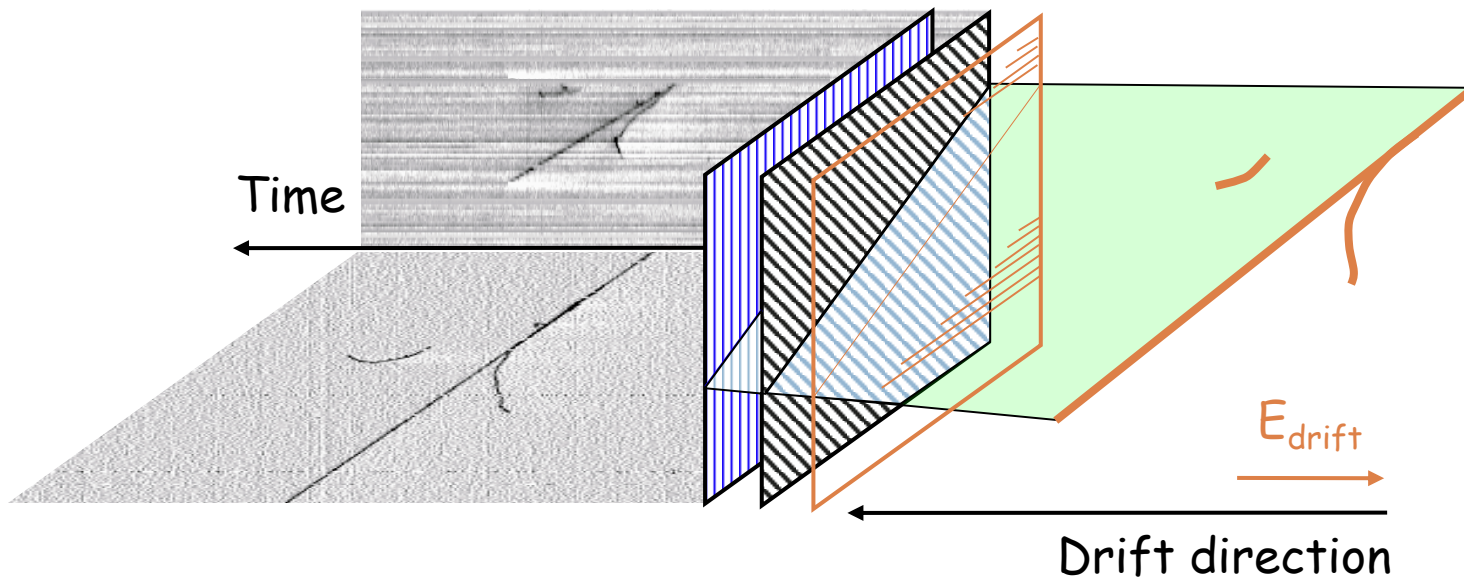
scintillation light
 ≈ 20000 photons/mm

(@500 V/cm external drift electric field)

3D imaging

The electric signal is collected by 3 planes of wires (54000 wires, 3 mm pitch).
2D reconstruction for 20m x 1,5 m

Scintillation light is collected by 74 PMTs placed inside the detector.
Reconstruction of the 3rd dimension 1,5 m



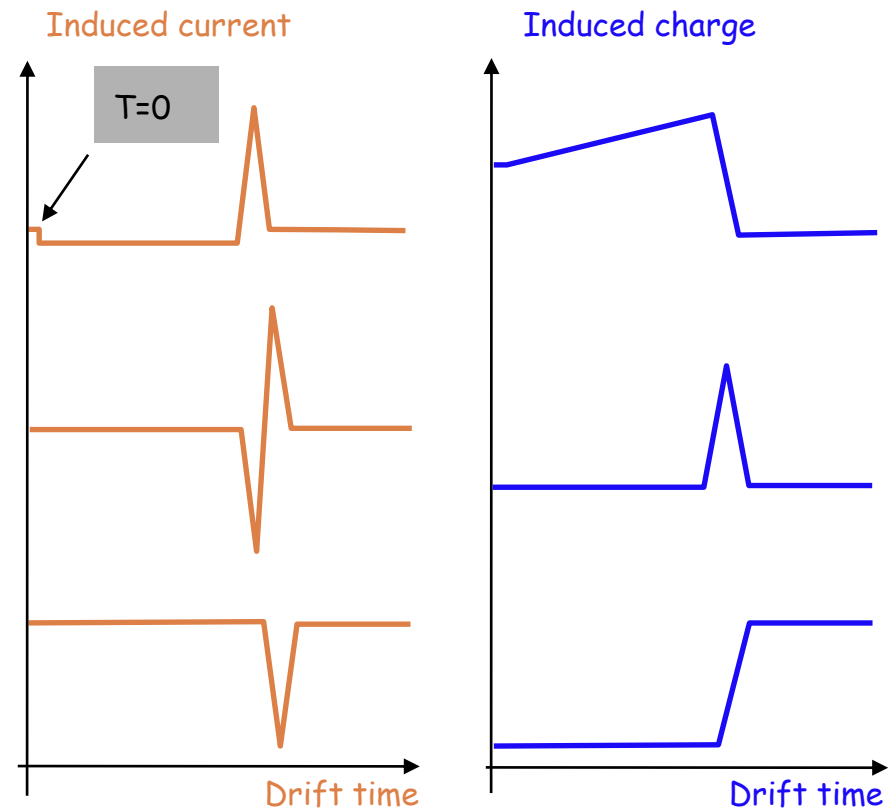
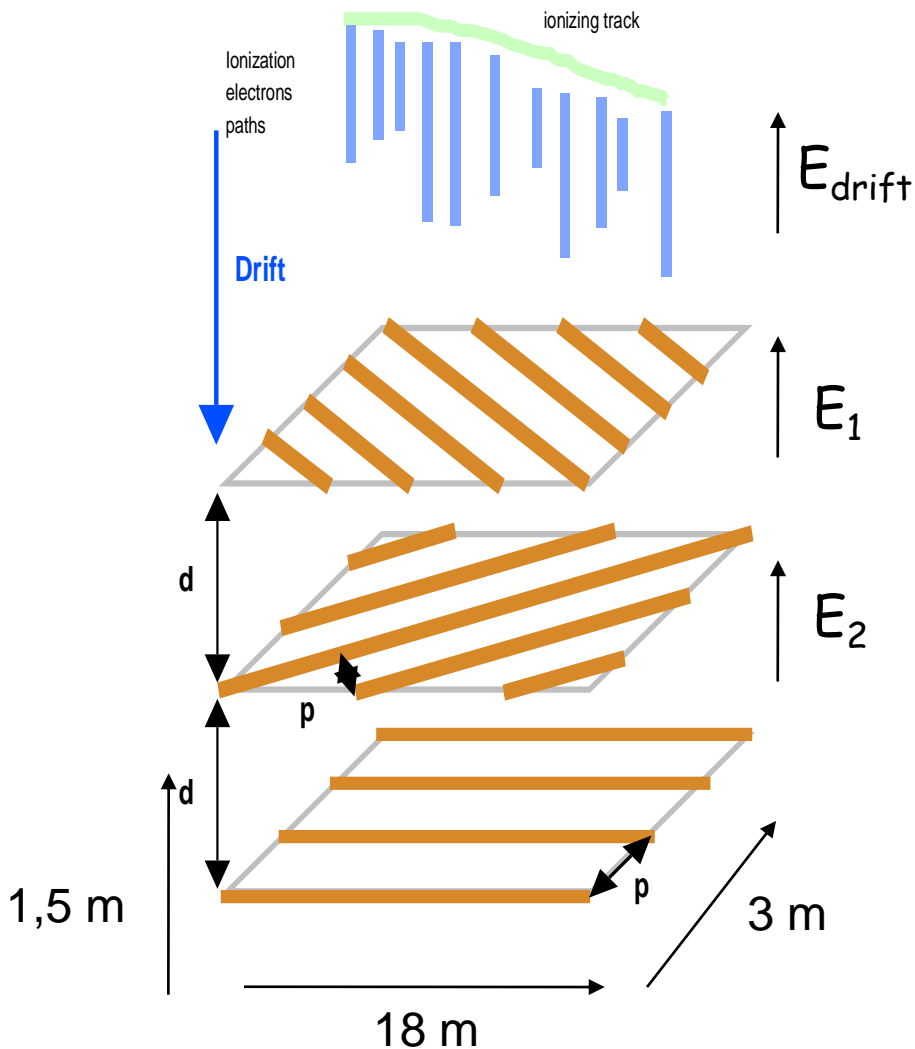
Electrons drift velocity = 1,5 mm/ μ s
(@500V/cm).

Wires read-out

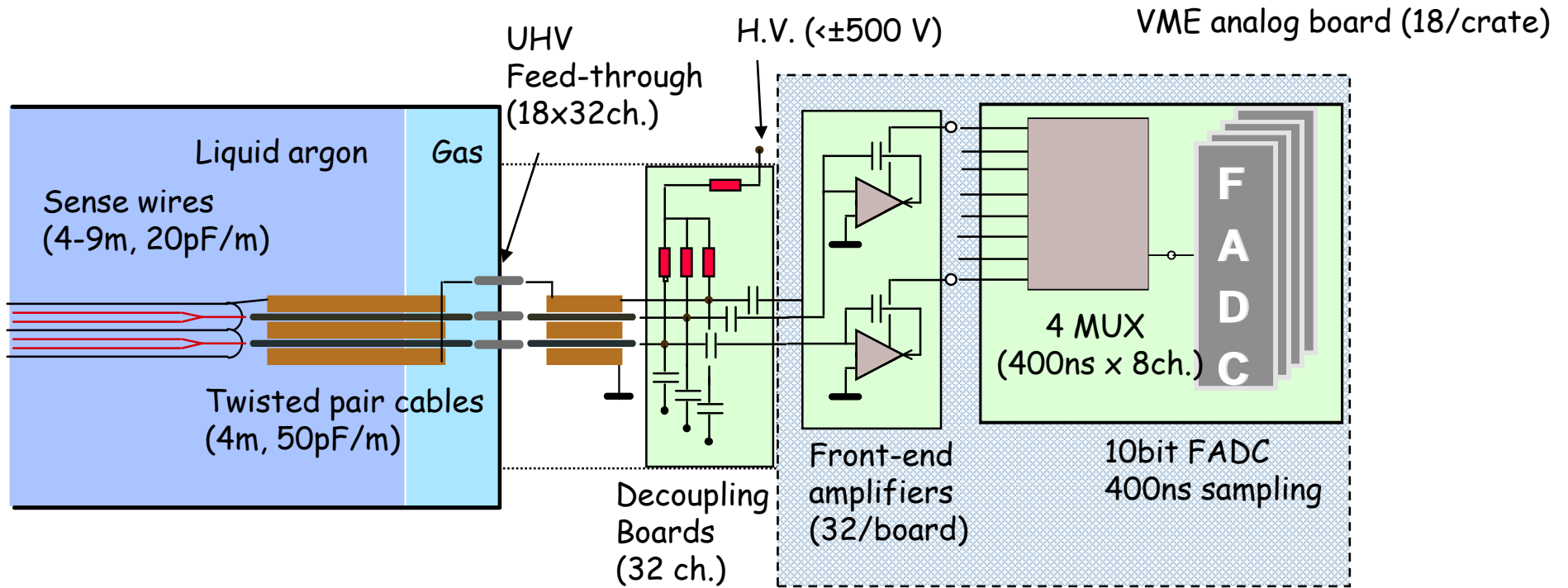
Non-destructive read-out is guaranteed by grid transparency condition:

$$E_1/E_{\text{drift}} = E_2/E_1 > (1+\rho)/(1-\rho)$$

$$\rho = 2\pi r/p \quad (r=\text{wire radius})$$



Front-end electronics



- ≈ 54000 channels
- 1664boards
- 96 crates

High gain, 15 ADC for 3 mm m.i.p. (15000 electrons)
Low noise, r.m.s. = 1ADC (1000 electrons equivalent)

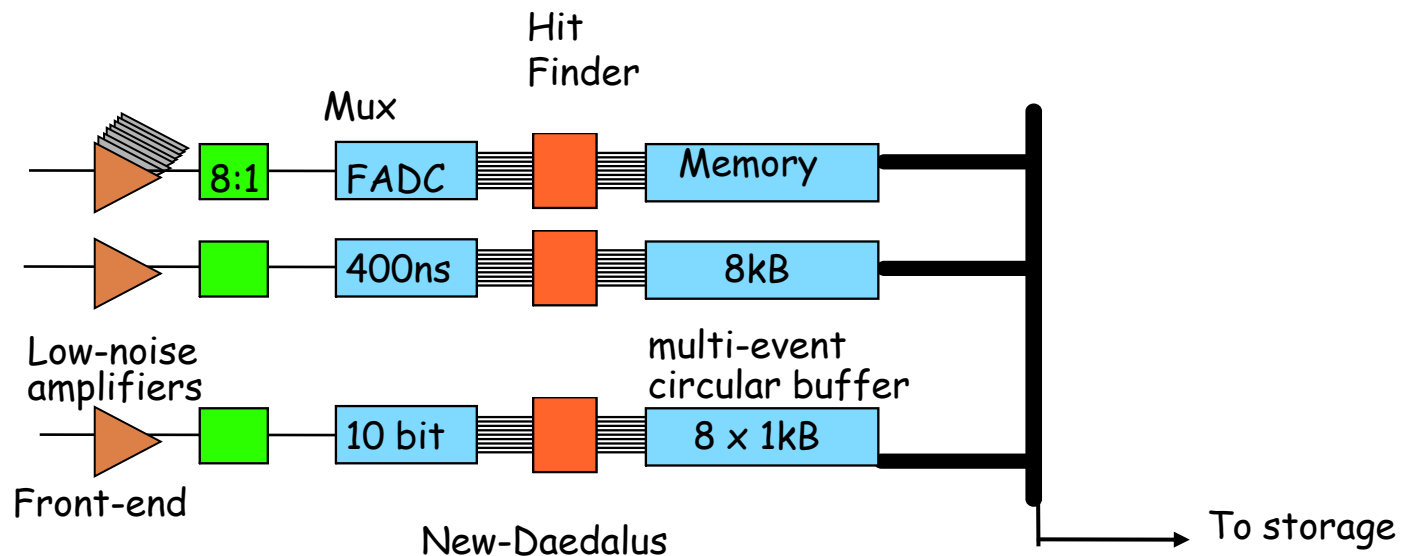
Read-out principle

Requirements:

- Continuously sensitive
- Self-triggering

Solution:

- Multi buffering
- Hit finding
- Boards independency



A trigger for physics

ICARUS physical goals:

- $\nu_{\mu} \rightarrow \nu_{\tau}$ oscillation on CerN to Gran Sasso neutrinos beam
- Proton decay
- Atmospheric and solar neutrinos
- Super Nova explosion

Event type	Events/year	Energy	# wires	# samples
ν_{μ} CNGS	1200	17 GeV	25000	2500
ν_e CNGS	10	17 GeV	25000	2500
ν atm	25/50	10 MeV- 100GeV	25000	2500
ν sol	300	10 MeV	2000	2500
ν SN	10/100	10 MeV	2000	2500
Proton decay	?	800 MeV	2000	2500

Triggering resources

□ PMTs

- Pros: t_0
- Cons: not localized, inefficient for small charge deposition

□ Wires (see below)

- Pros: localized, sensitive to small charge deposition
- Cons: sensitive to neutron capture background

□ CNGS

- Pros: 100% efficiency on CNGS events
- Cons: not localized, high rate of empty events

Triggering modes

□ External trigger:

- Limited in bandwidth (≈ 1 Hz max rate for 1.5 ms drift). Maximum of eight events pile-up before deadtime.

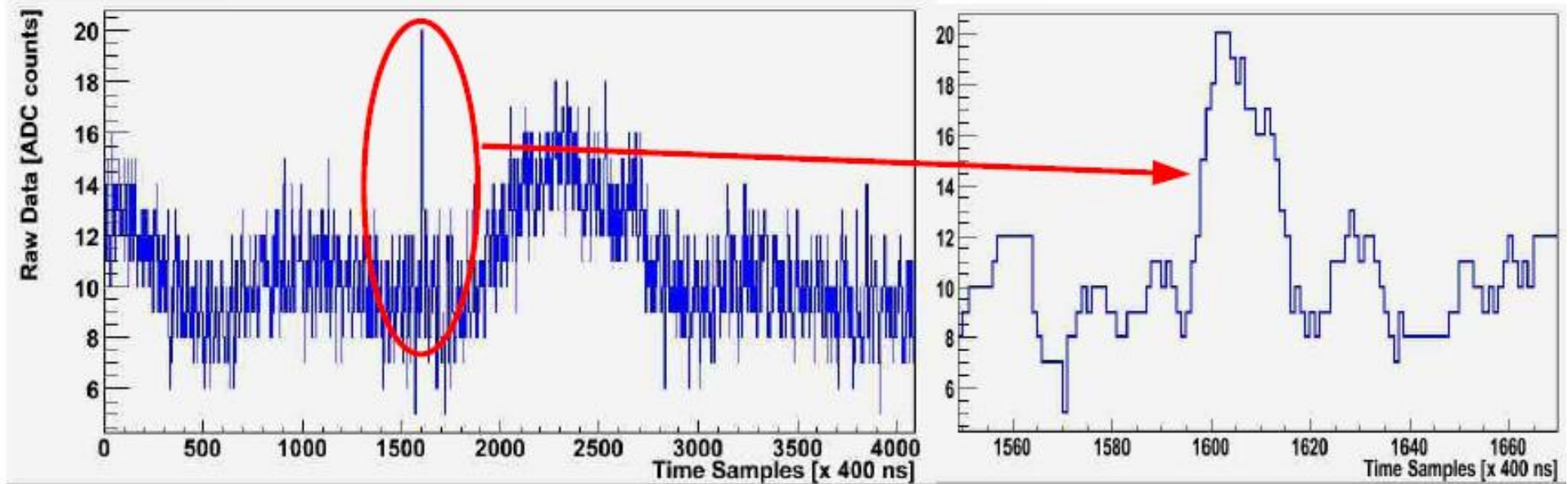
□ External Enable:

- Bandwidth allows up to 1k event “tiles” ($25 \mu\text{s} \cdot 16$ wires) per second per readout crate. NewDaedalus thresholds can be more tolerant without overflowing readout.
- Internal FIFO’s can accept up to 128 fragments.

□ Open Shutter:

- Same bandwidth as above. Useful to collect low energy events.
- Drawback is that correlated noise bursts even at low repetition rate (few per second) would easily saturate the DAQ channel.

M.i.p. signal

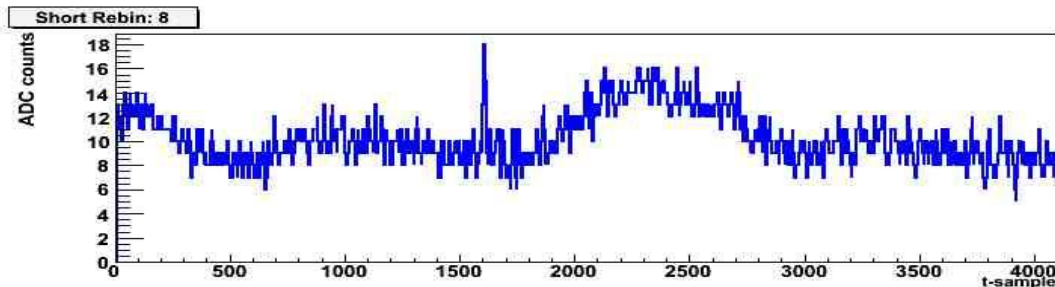


M.i.p. signal: 15 ADC counts, 30/40 t-samples

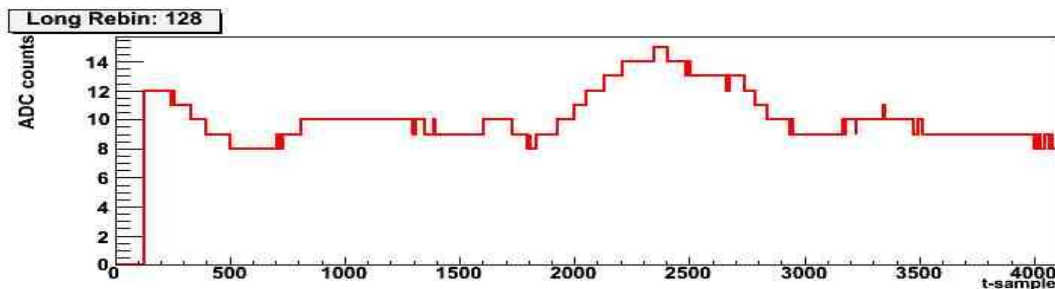
Low frequency noise: ≈ 10 ADC counts, ≈ 2000 t-samples

High frequency noise: $\approx \pm 2$ ADC counts, ≈ 5 t-samples

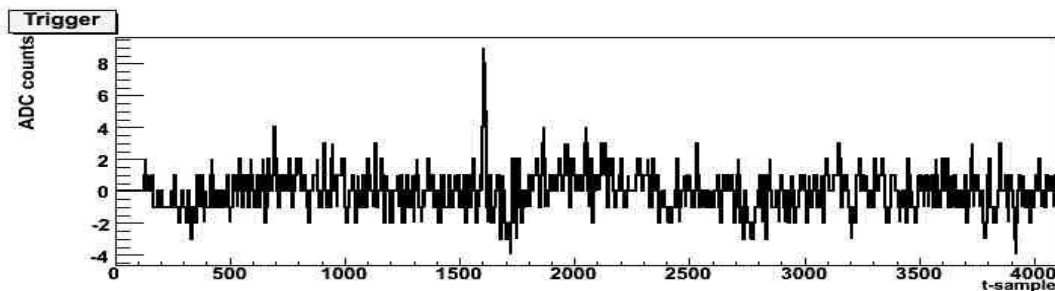
New hit finding algorithm



$$Q_8(t) = \frac{1}{8} \sum_{i=0}^8 Q(t-i)$$



$$Q_{128}(t) = \frac{1}{128} \sum_{i=0}^{128} Q(t-i)$$



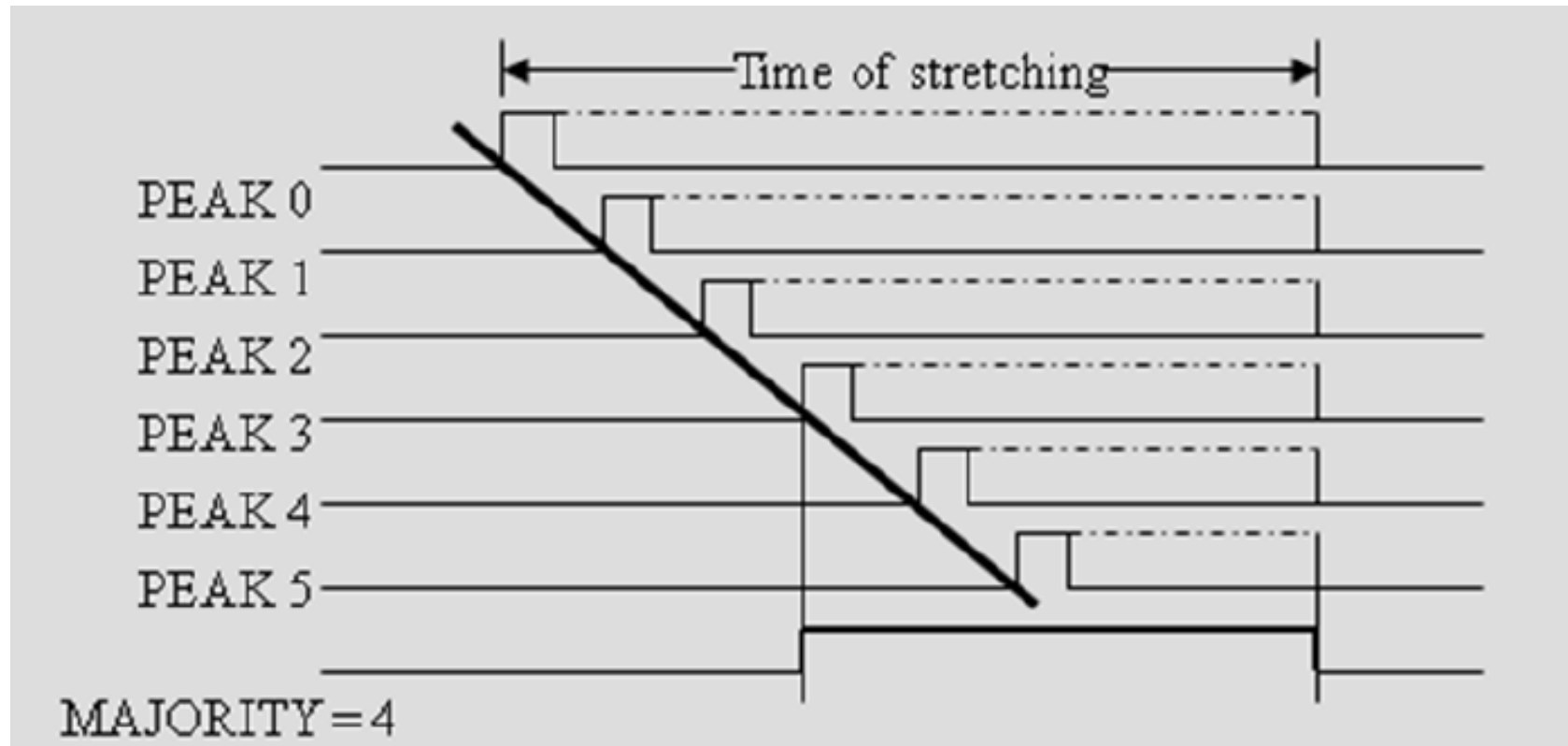
$$S(t) = Q_8(t) - Q_{128}(t)$$

8 samples average to reduce high frequency oscillation

128 samples average to follow baseline modulation

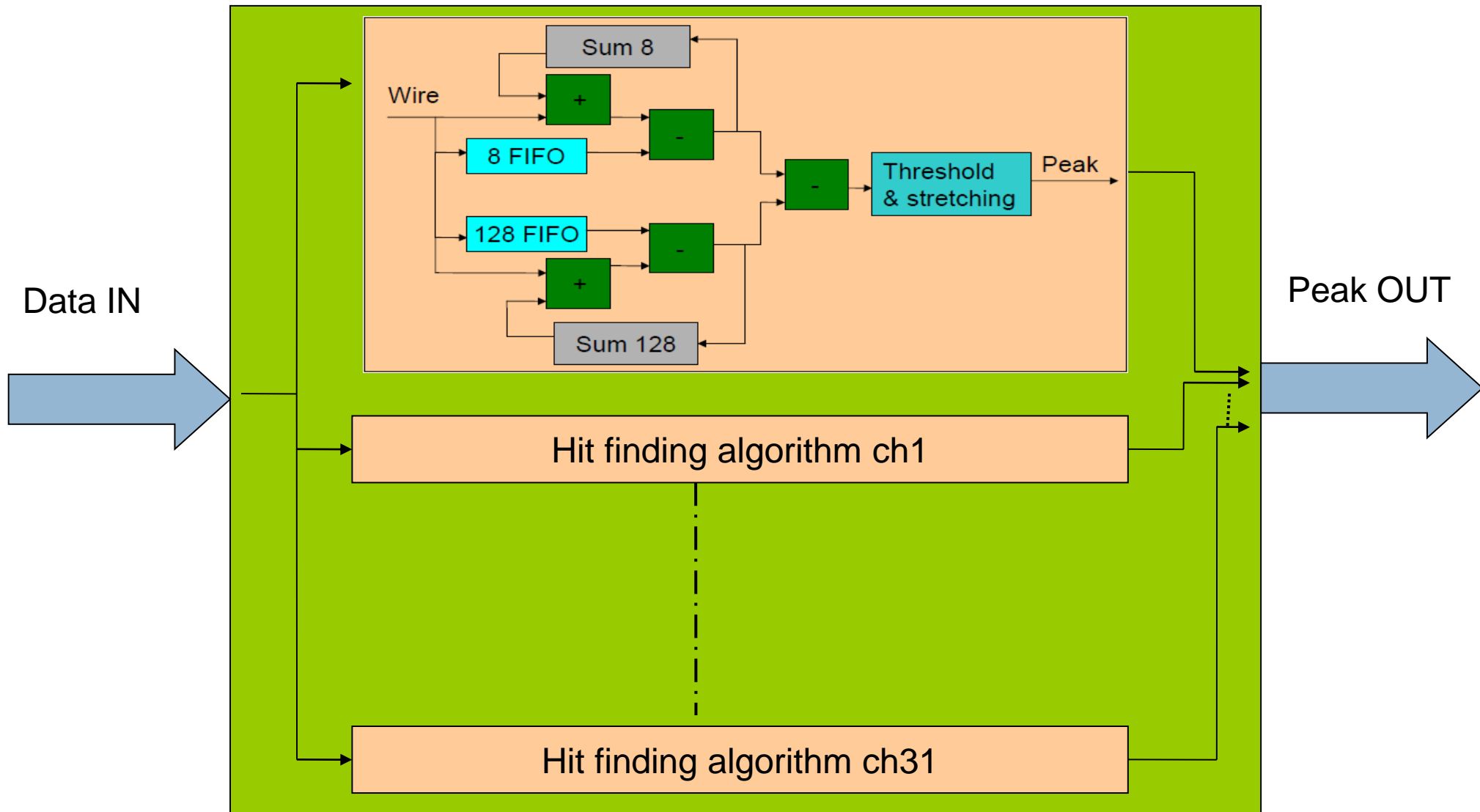
A peak signal is generated when $S(t)$ goes over threshold

2nd step of the algorithm

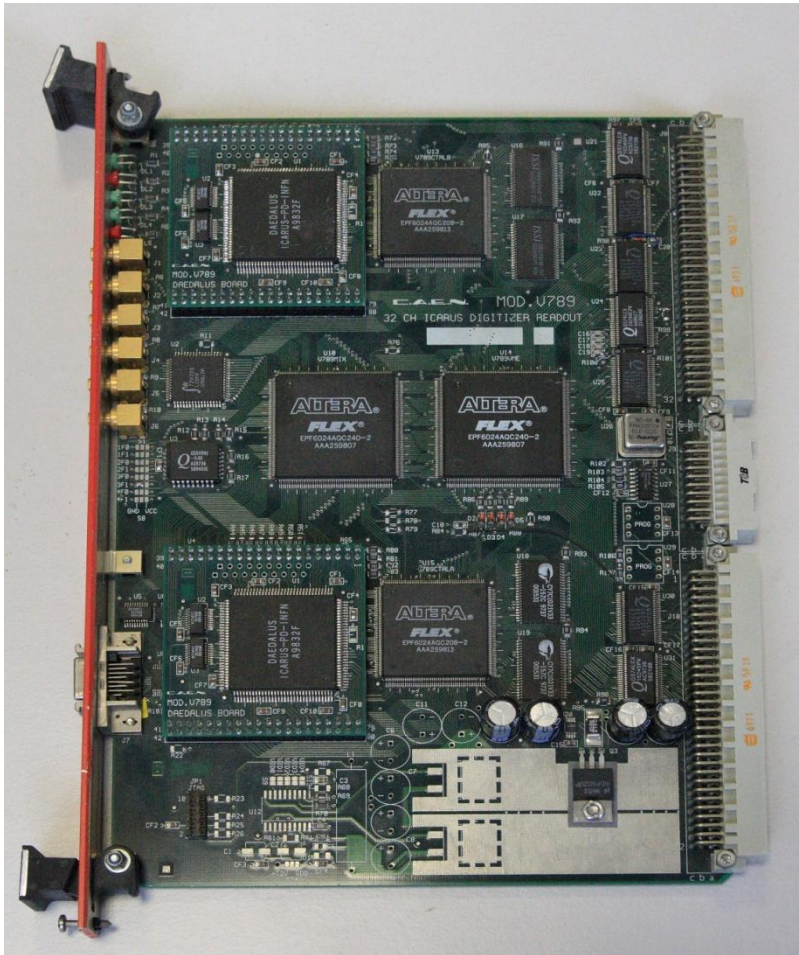


Peak stretching ranging from 25 μ s to 125 μ s to guarantee high efficiency for inclined tracks

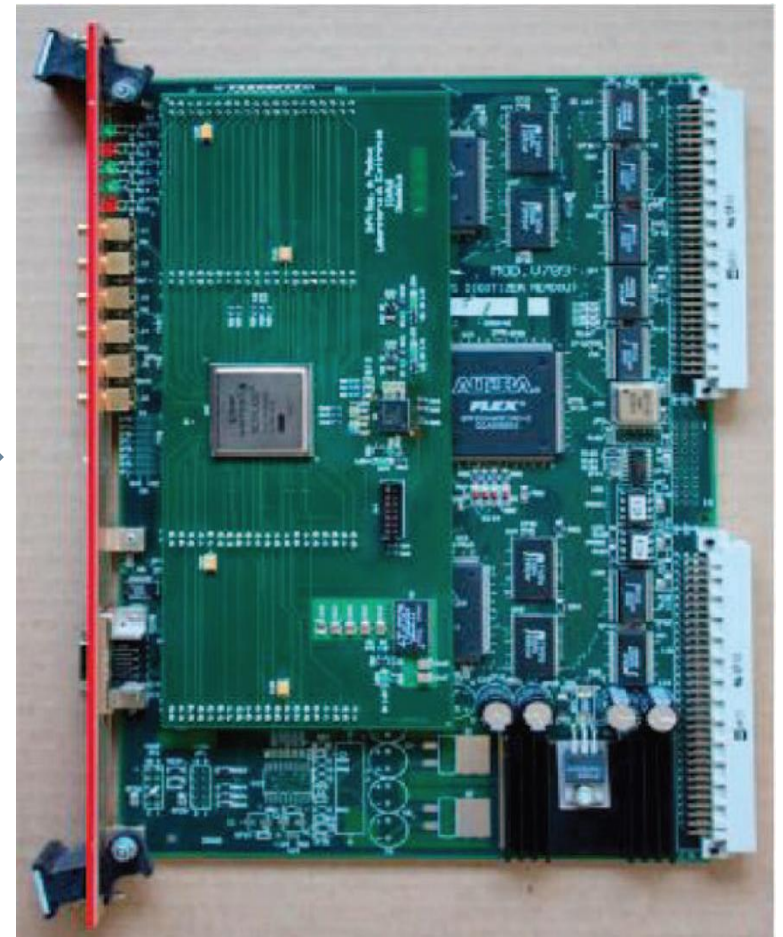
Block diagram



From Daedalus to NewDaedalus

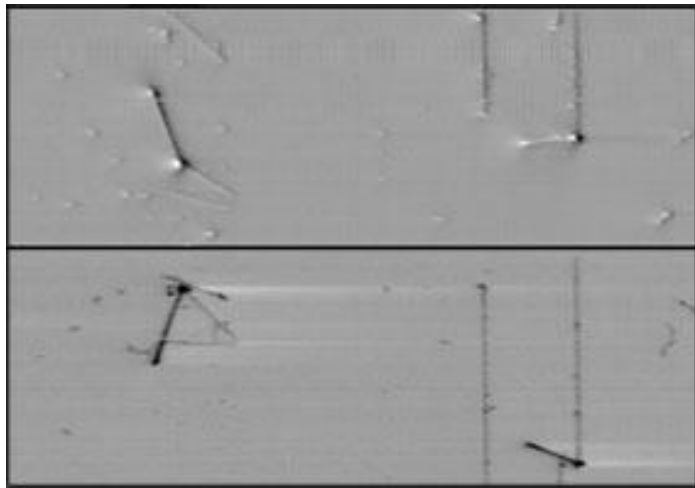


Daedalus:
VLSI 0,5 μm
CMOS gate array



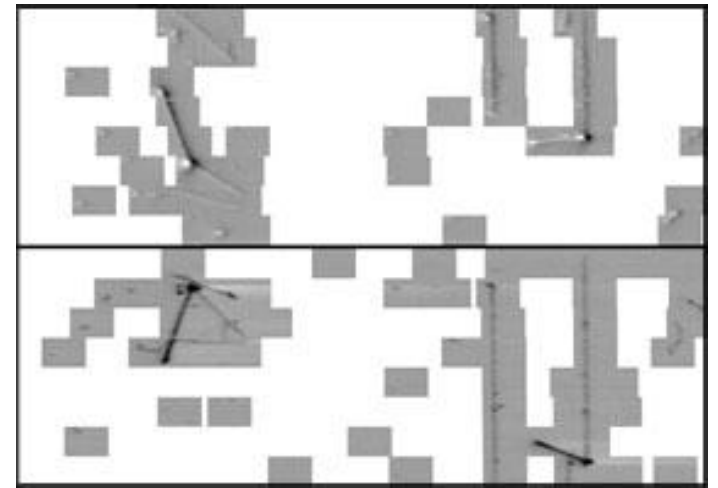
NewDaedalus:
Xilinx Virtex 5 FPGA (prototype)
Xilinx Spartan 6 FPGA (production)

NewDaedalus as R.o.I. selector



Full drift image

Triggered on PMTs



Reduced drift image

Triggered on PMTs

Reduced with NewDaedalus
R.o.I. selection

Icarino test facility



Same electronic chain used for the ICARUS detector

Icarino has been running in 2009 for testing on-line lossless data compression and trigger capabilities of NewDaedalus chip

The Icarino chamber



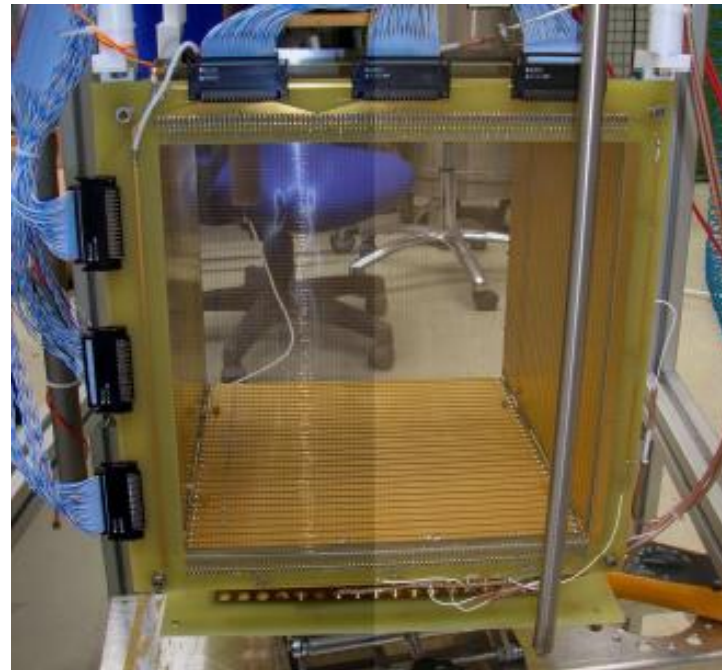
Active mass: 38 Kg

Active volume :32 ,6 x 32,6 x 29,4 cm³

2 planes of wires:

96 horizontal wires (collection)

96 vertical wires (induction)

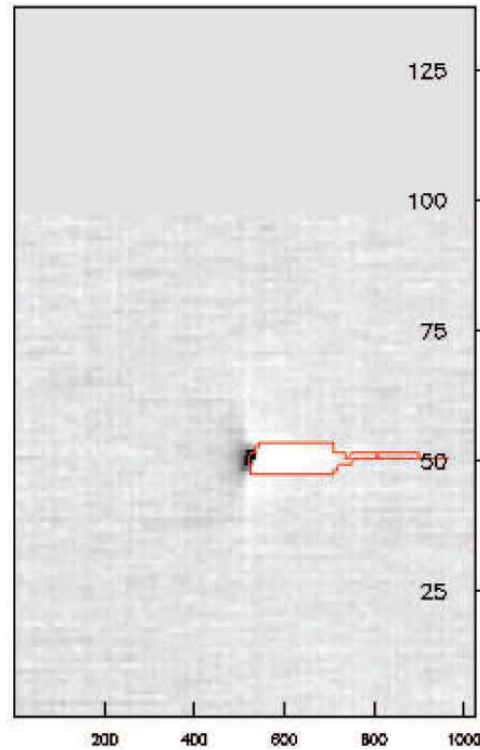
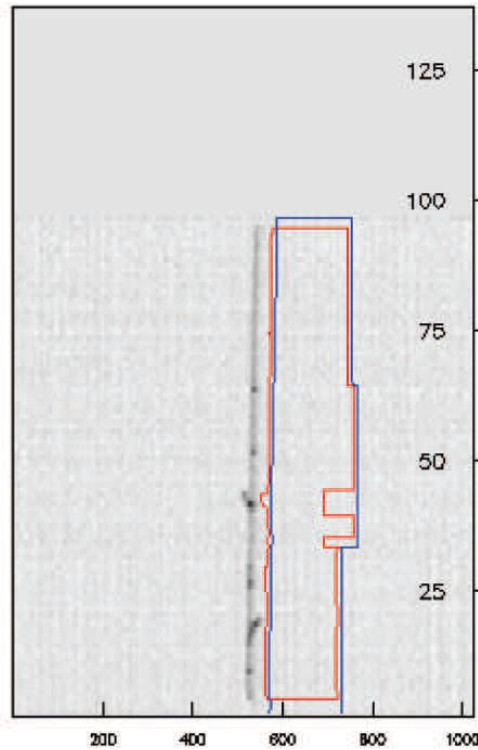


Vertical through-going muon

Run 3362 Event 00004 21 dec 2009 21-55-25

Collection view

Induction view



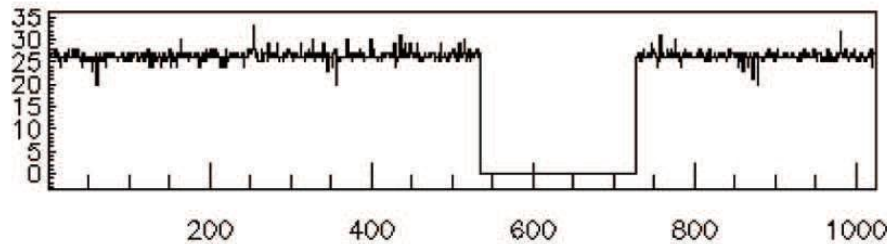
Red line: Peak signal (thre. = 6)

Blue line: Majority signal (maj. = 8)



- Induc.
- Next
- Print
- Zoom
- Wr sel
- Tr.fit

Wire 0102



- Wr fit
- Wr.up
- Wr.dw

Inclined through-going muon

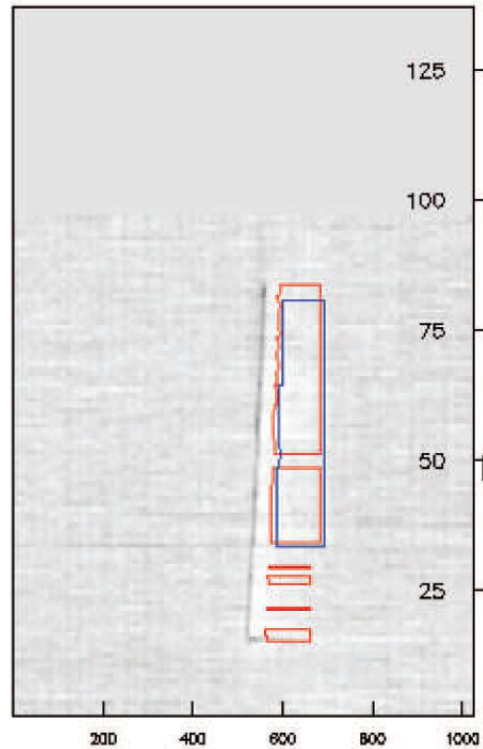
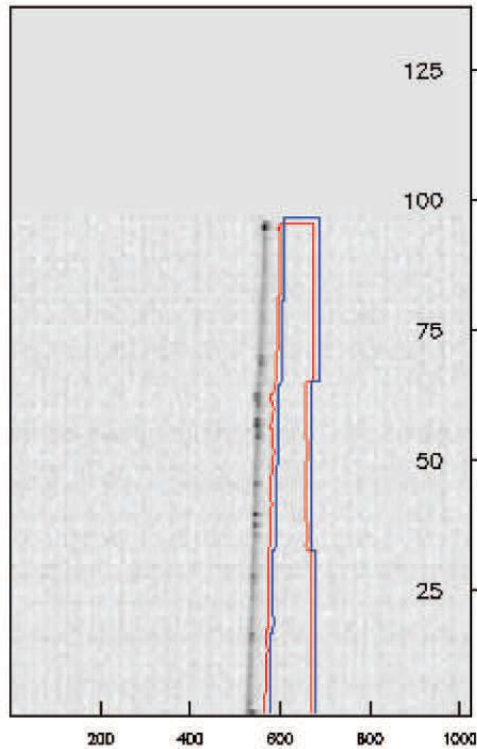
Run 3422 Event 00050 29 dec 2009 15-58-54

Collection view

Induction view



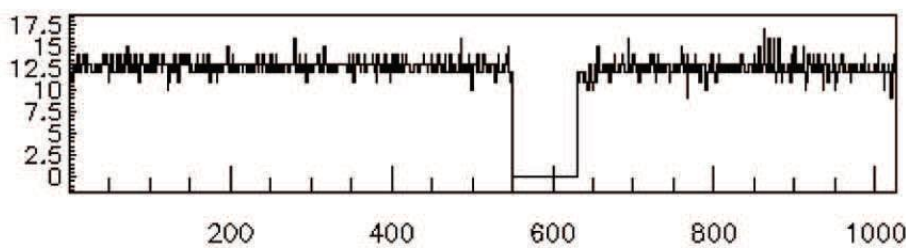
- Induc.
- Next
- Print
- Zoom
- Wr sel
- Find Hit



Red line: Peak signal (thre. = 6)

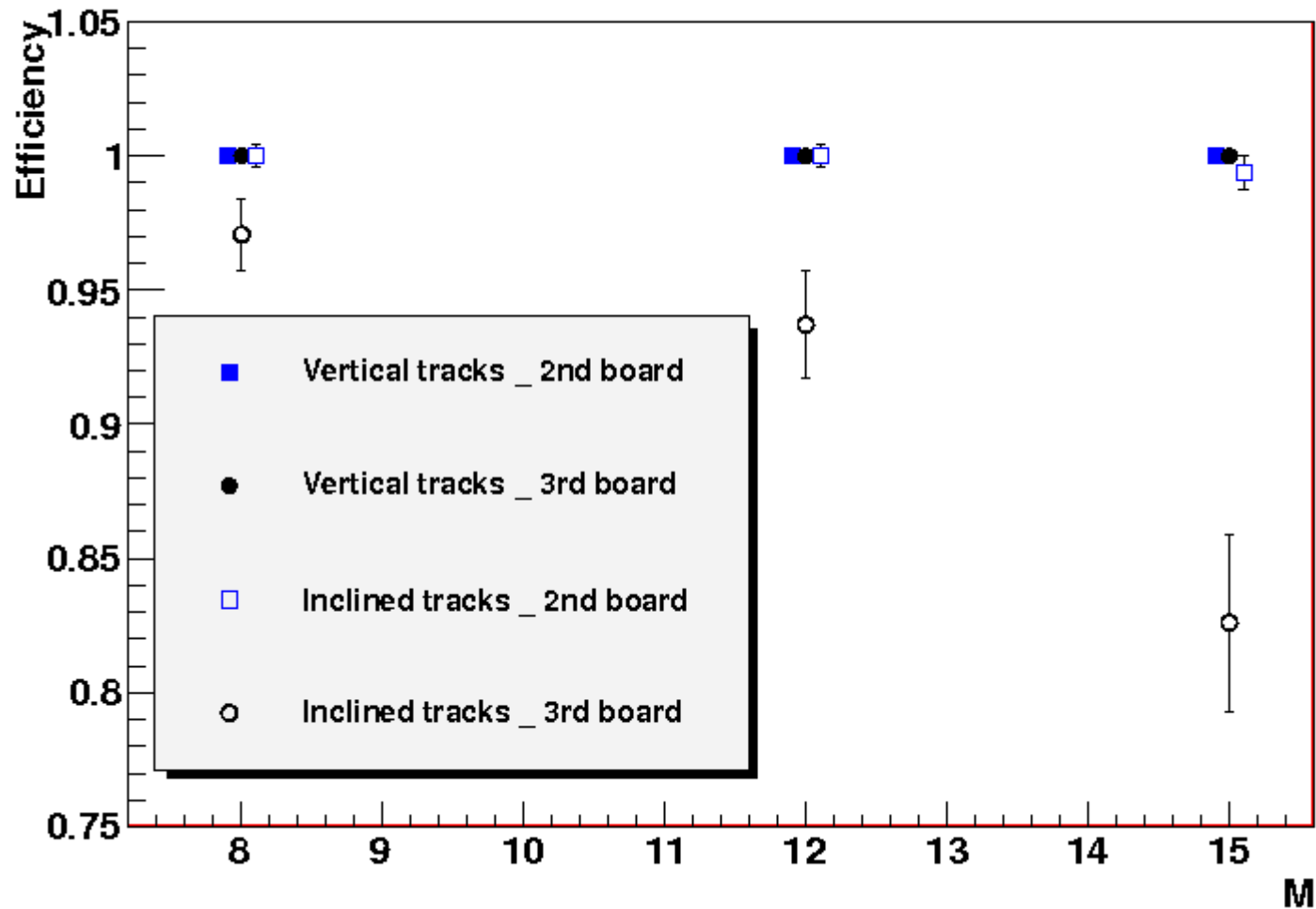
Blue line: Majority signal (maj . = 8)

Wire 0101



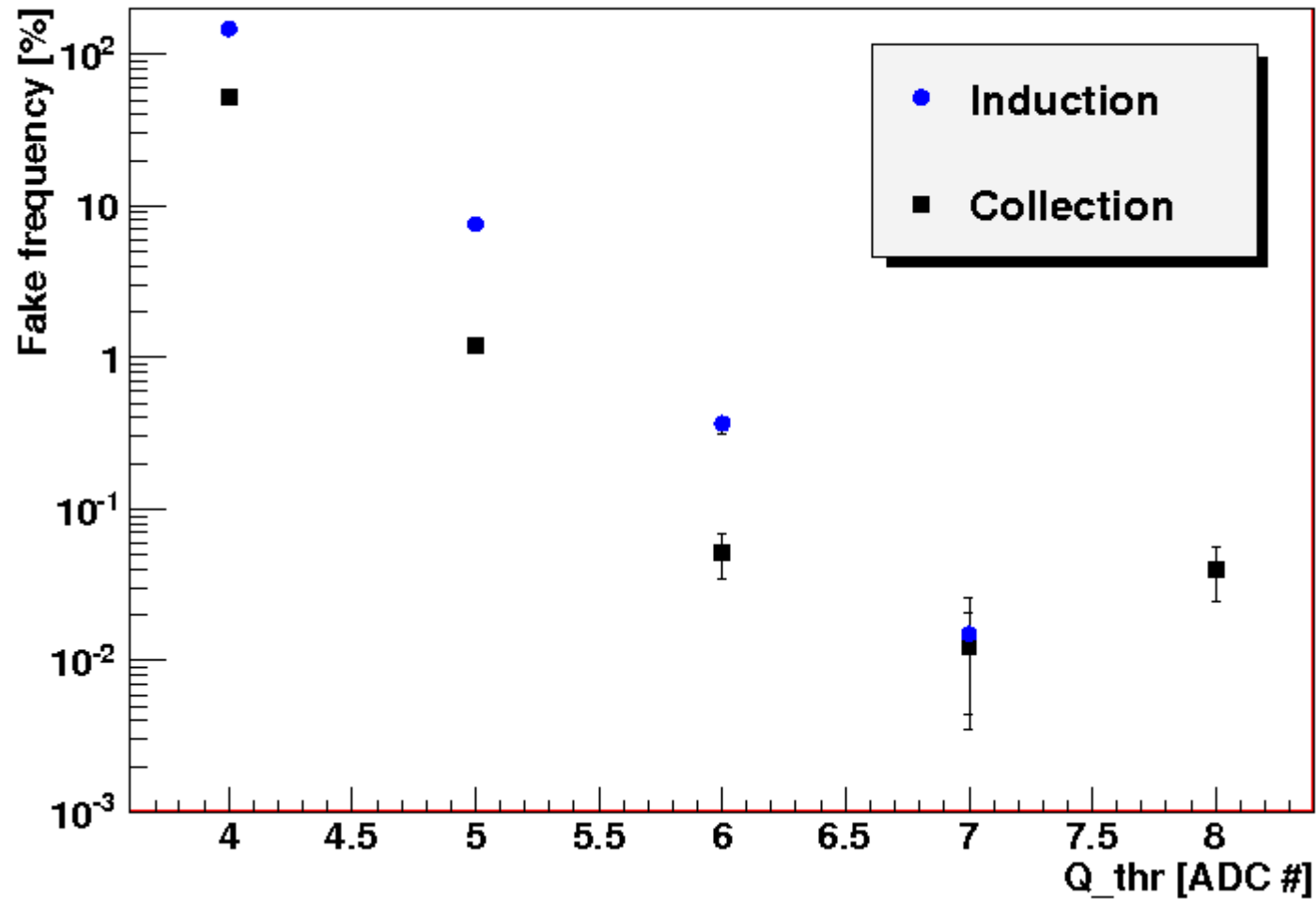
- Wr fit
- Wr.up
- Wr.dw

NewDaedalus efficiency



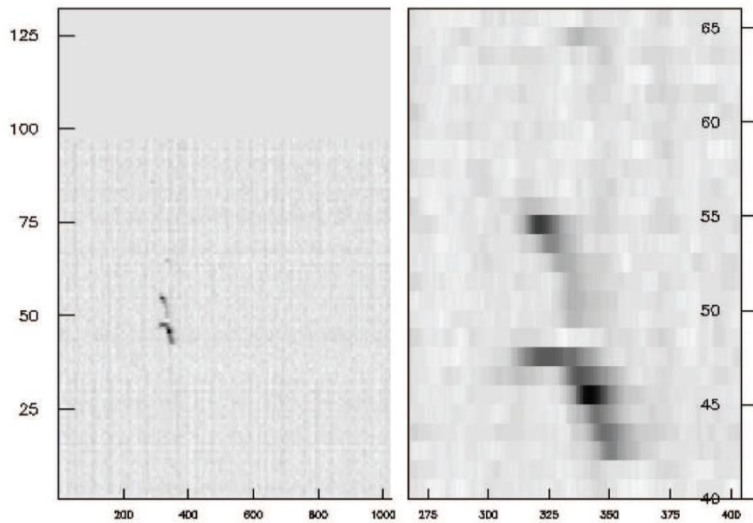
Threshold = 6 ADC

NewDaedalus rate of fake



Small charge deposition setup

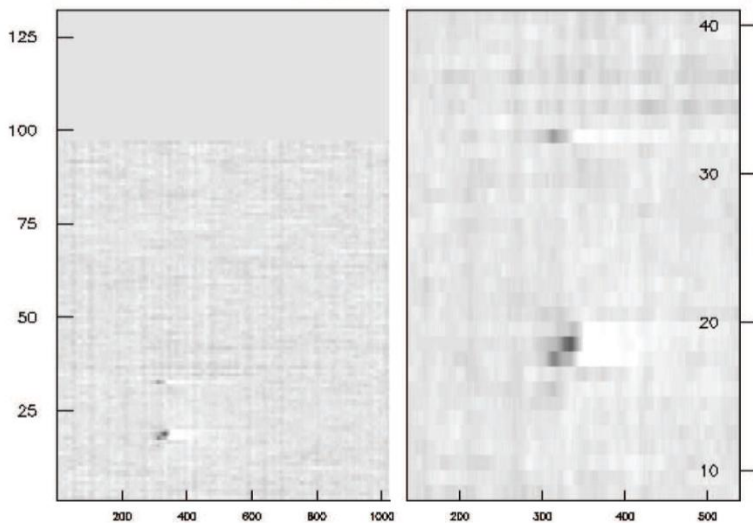
Run 409 Event 00007 28 dec 2009 15-20-19
Collection view



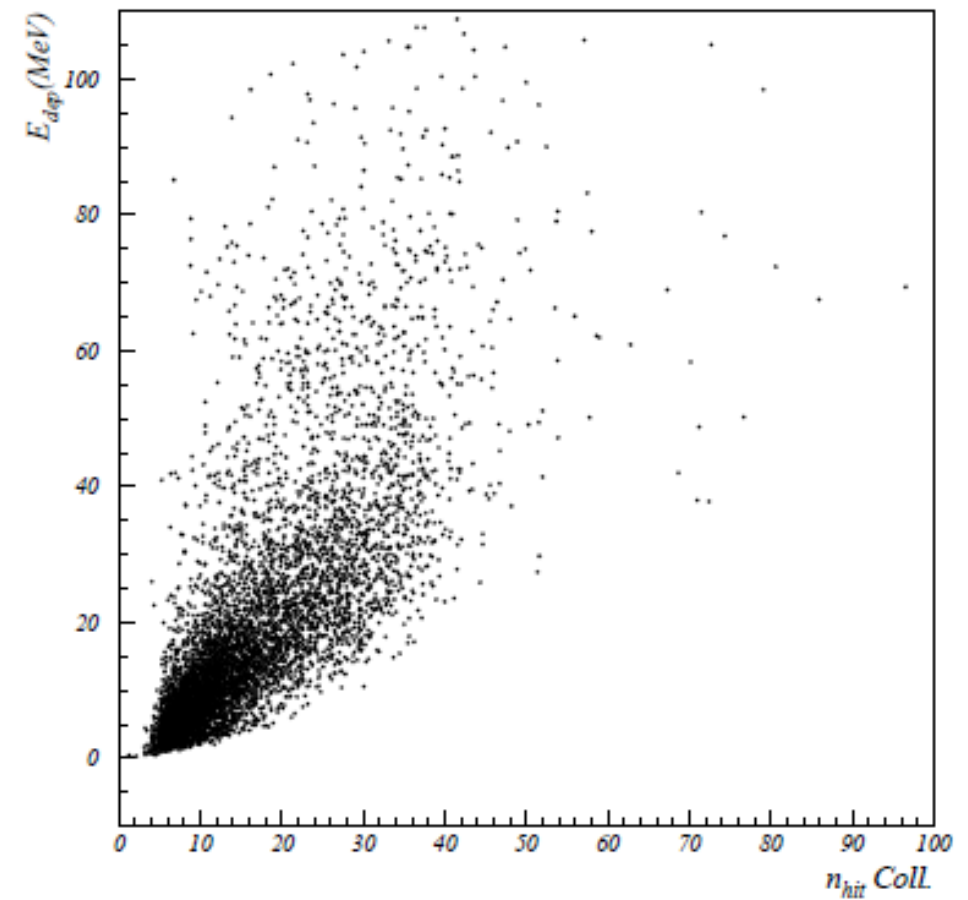
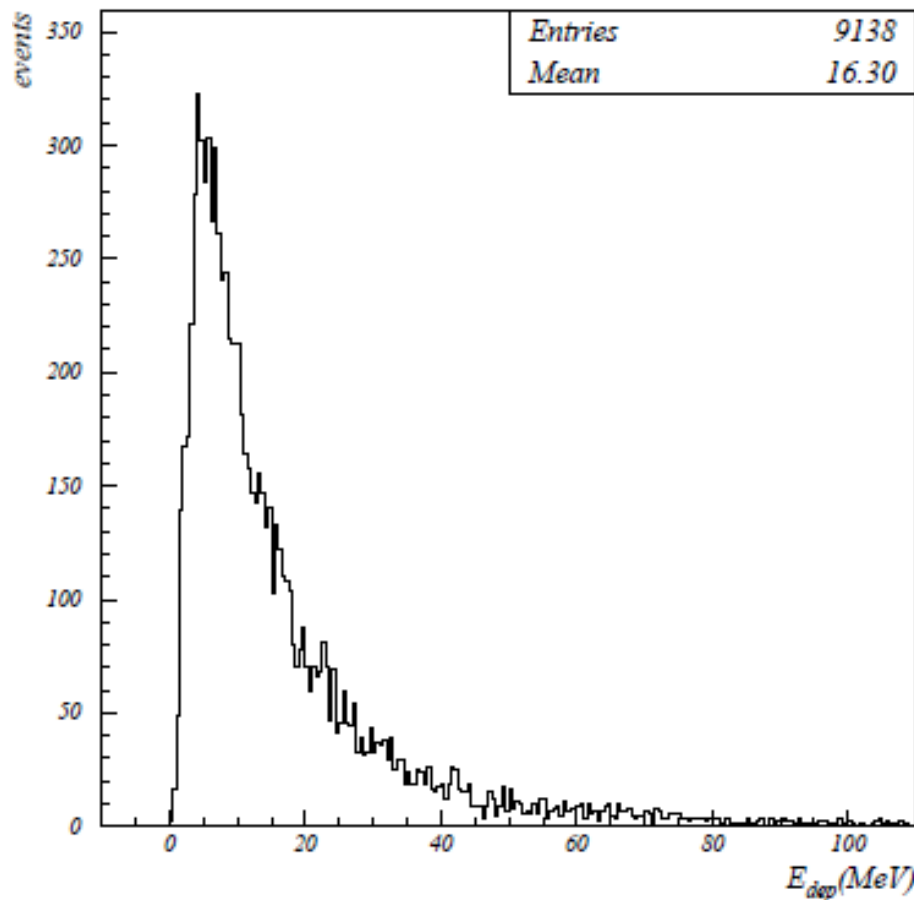
Central collection board NewDaedalus signal used as trigger

Lateral collection boards NewDaedalus signals used as veto

Induction view



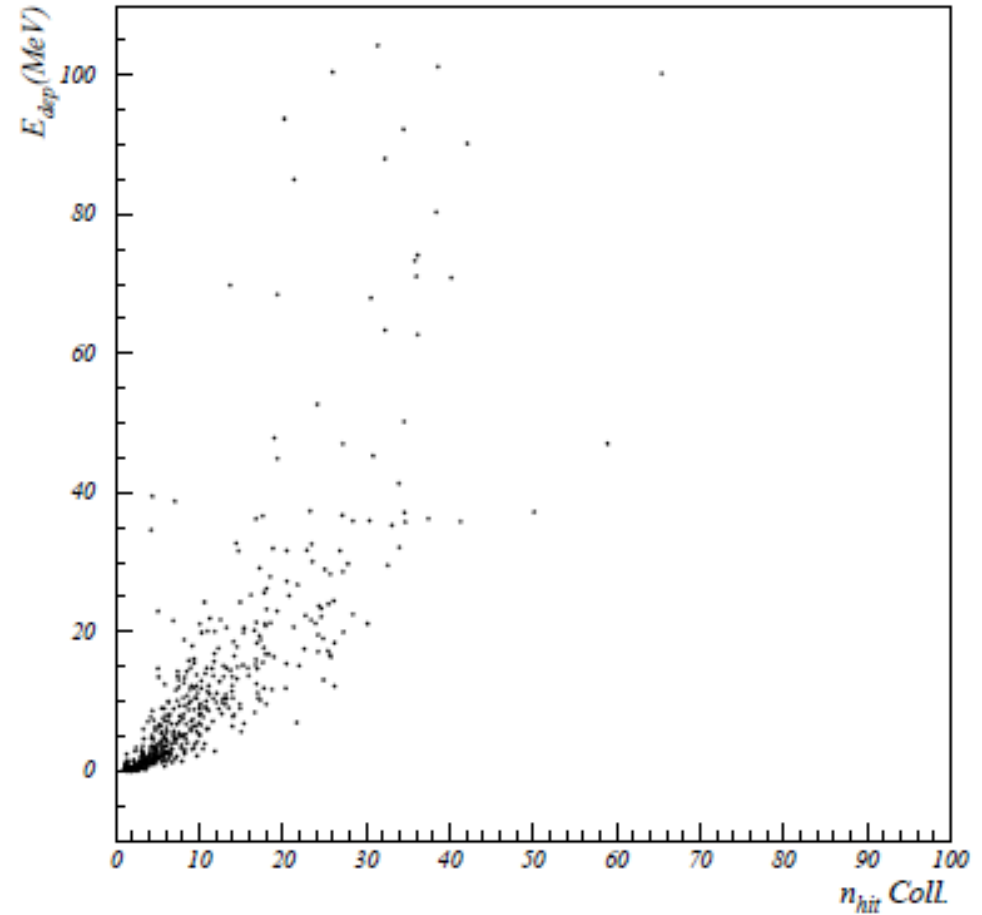
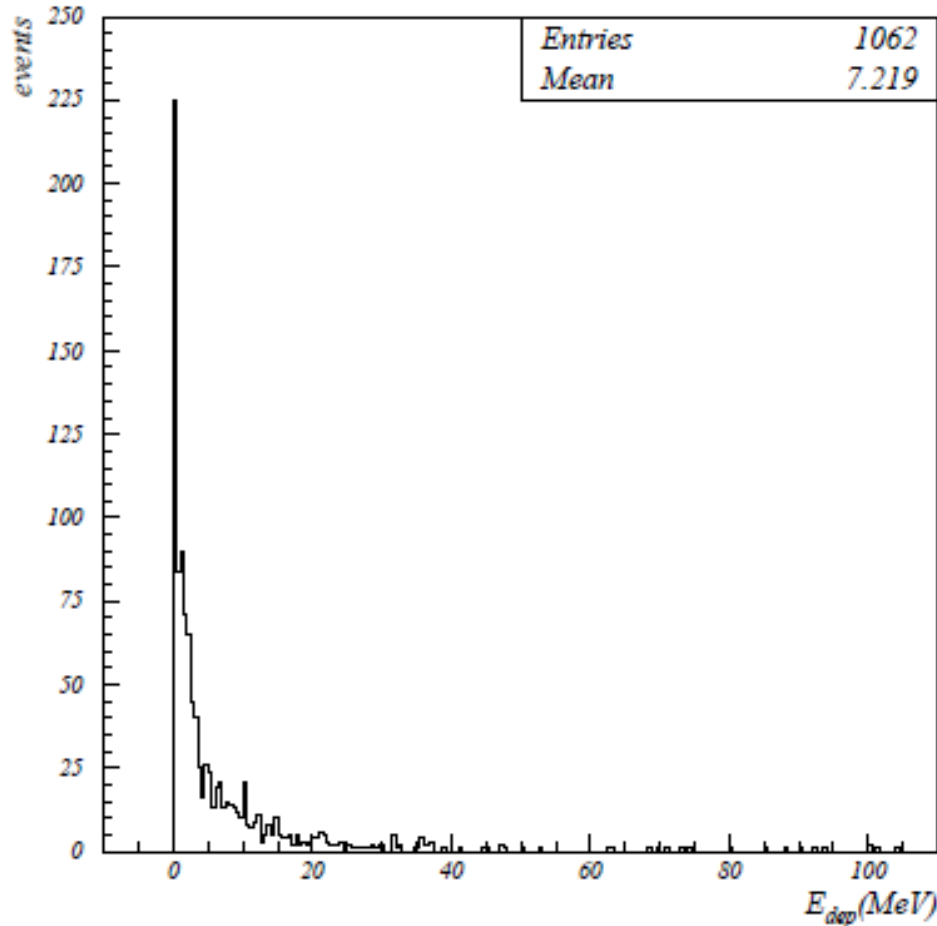
Small charge deposition spectrum



Histogram of the energy deposition of the events taken with the small charge deposition setup

Threshold = 6 ADC Majority = 4

Small charge deposition spectrum



Histogram of the energy deposition of the events
taken with the small charge deposition setup

Threshold = 5 ADC Majority = 3

Conclusions

- R.o.I. algorithm has been implemented and successfully tested
- Sensitivity of the detector can be pushed to few MeV
- First prototype will be implemented in ICARUS after the commissioning phase
- Application to install NewDaedalus chip on the whole detector before March 2011, before CNGS beam start