

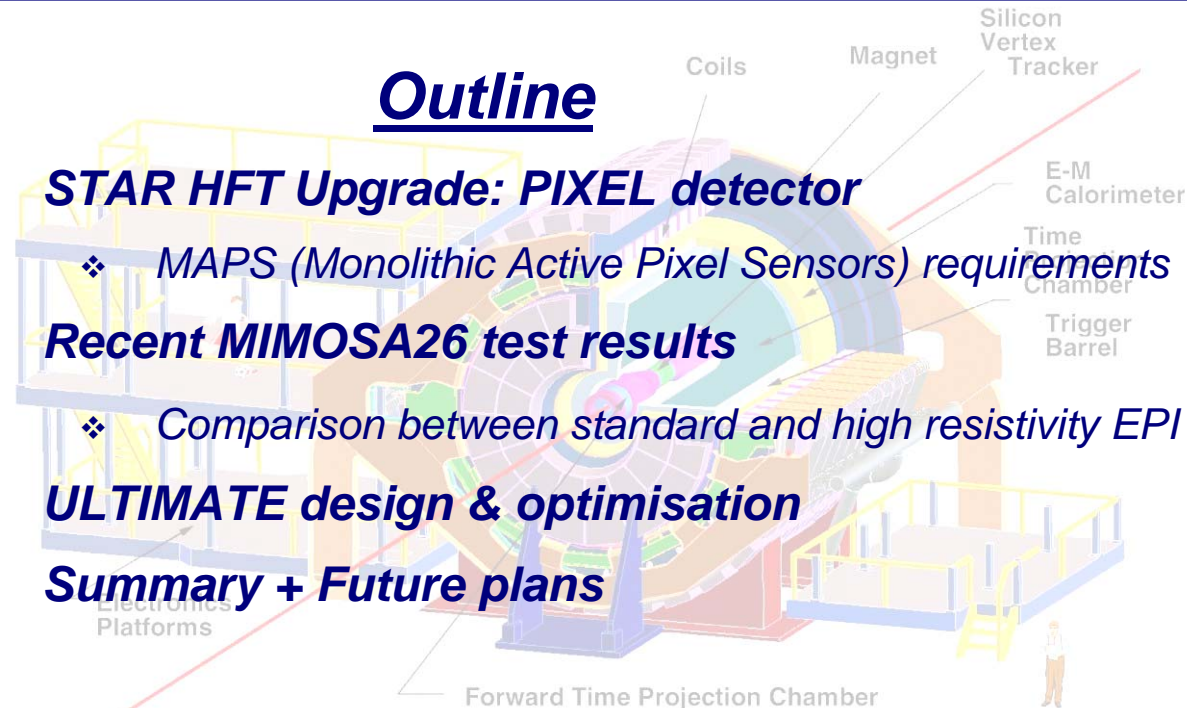
ULTIMATE: a High Resolution CMOS Pixel Sensor for the STAR Vertex Detector Upgrade

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on behalf of the IPHC (Strasbourg) CMOS Sensors group

Outline

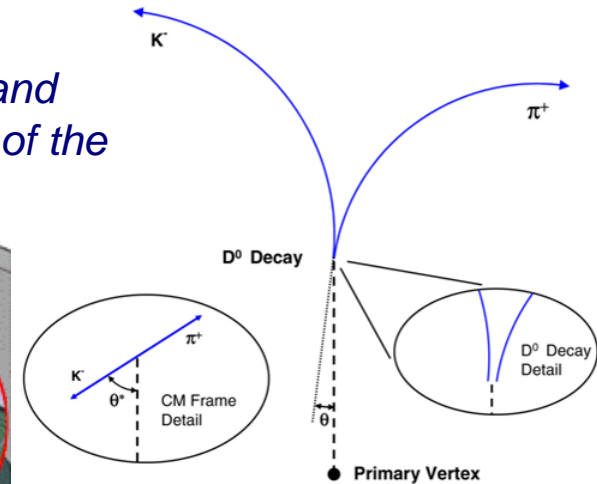
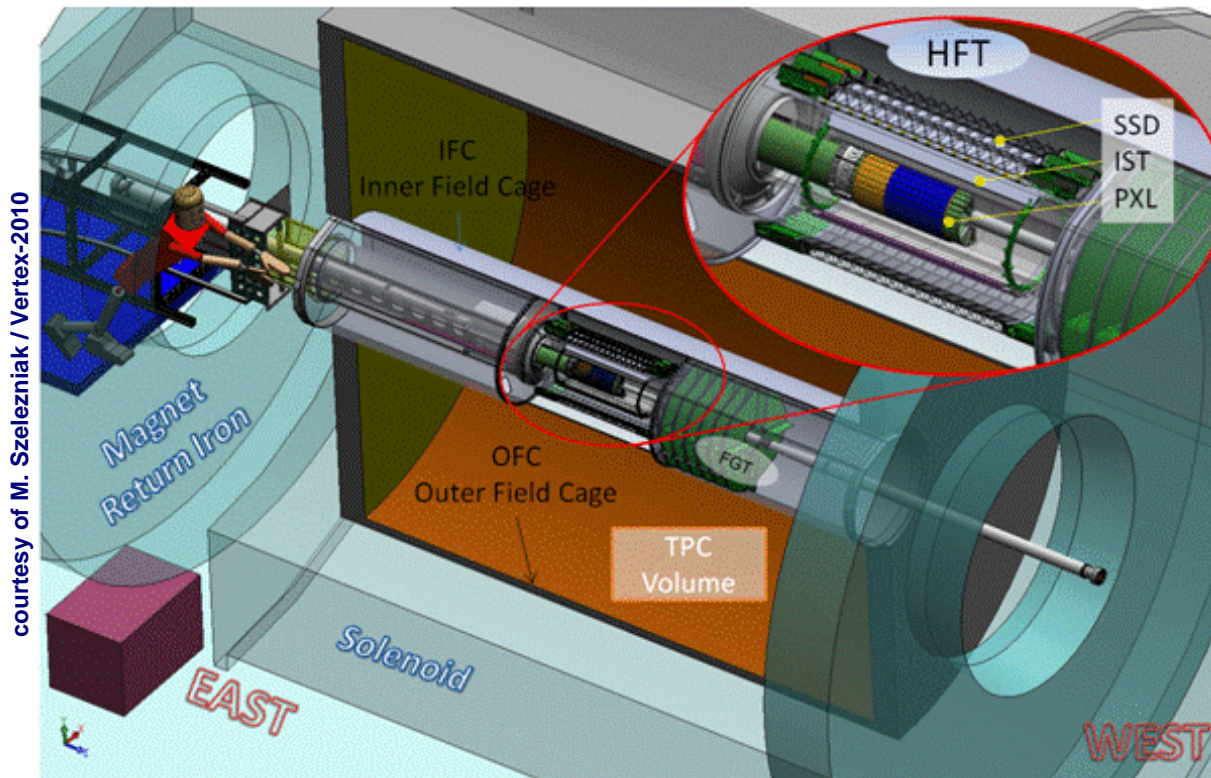
- ❖ **STAR HFT Upgrade: PIXEL detector**
 - ❖ *MAPS (Monolithic Active Pixel Sensors) requirements*
- ❖ **Recent MIMOSA26 test results**
 - ❖ *Comparison between standard and high resistivity EPI layers*
- ❖ **ULTIMATE design & optimisation**
- ❖ **Summary + Future plans**



STAR Heavy Flavor Tracker (HFT) Upgrade

■ Physics Goals:

- ↪ Identification of mid rapidity Charm and Beauty mesons and baryons through direct reconstruction and measurement of the displaced vertex with excellent pointing resolution



TPC – Time Projection Chamber
(main detector in STAR)

HFT – Heavy Flavor Tracker

- SSD – Silicon Strip Detector
- IST – Inner Silicon Tracker
- PXL – Pixel Detector (PIXEL)

Goal: Increasing pointing resolution from the outside in

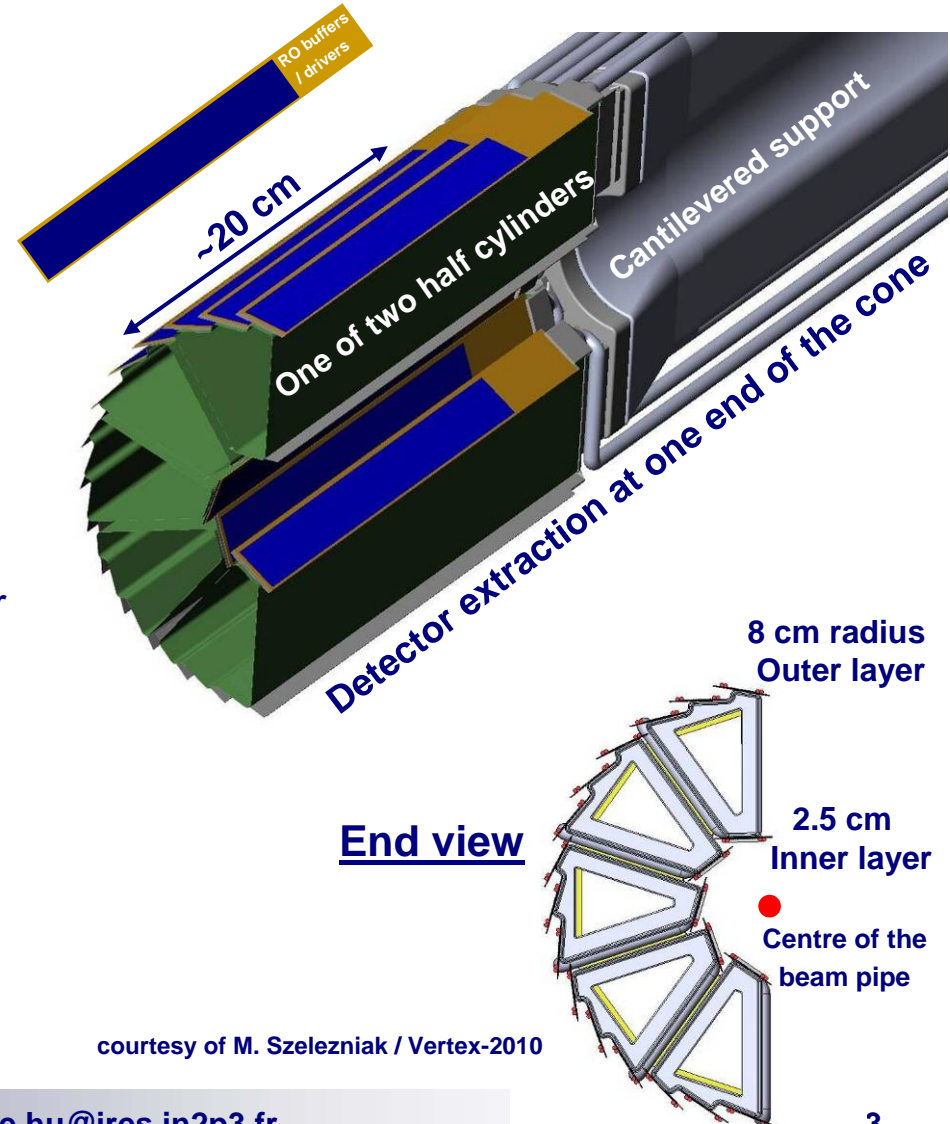


STAR PIXEL Detector

Sensors Requirements

- Multiple scattering minimisation:
 - ↳ Sensors thinned to 50 μm , mounted on a flex kapton/aluminum cable
 - $X/X_0 = 0.37\%$ per layer
- Sufficient resolution to resolve the secondary decay vertices from the primary vertex
 - ↳ $< 10 \mu\text{m}$
- Luminosity = $8 \times 10^{27} / \text{cm}^2 / \text{s}$ at RHIC_II
 - ↳ $\sim 200\text{-}300$ (600) hits / sensor ($\sim 4 \text{cm}^2$) in the integration time window
 - ↳ Shot integration time $\sim < 200 \mu\text{s}$
- Low mass in the sensitive area of the detector
 - airflow based system cooling
 - ↳ Work at ambient ($\sim 35 \text{ }^\circ\text{C}$) temperature
 - ↳ Power consumption $\sim 100 \text{ mW} / \text{cm}^2$
- Sensors positioned close (2.5 - 8 cm radii) to the interaction region
 - ↳ $\sim 150 \text{ kRad} / \text{year}$
 - ↳ few $10^{12} N_{\text{eq}} / \text{cm}^2 / \text{year}$

Total: 40 ladders
Ladder = 10 MAPS sensors ($\sim 2 \times 2 \text{ cm}^2$ each)



courtesy of M. Szelezniak / Vertex-2010

STAR PIXEL Detector

■ 3 steps evolution:

- ↪ 2007: A MimoSTAR-2 sensors based telescope has been constructed and performed measurements of the detector environment at STAR

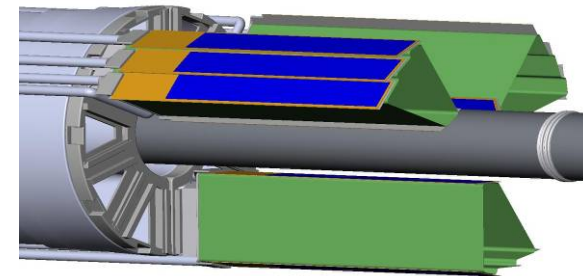
MimoSTAR-2: sensor with analogue output



3 plans telescope with MimoSTAR-2 sensors

- ↪ 2012: The engineering prototype detector with limited coverage (1/3 of the complete detector surface), equipped with PHASE-1 sensors will be installed

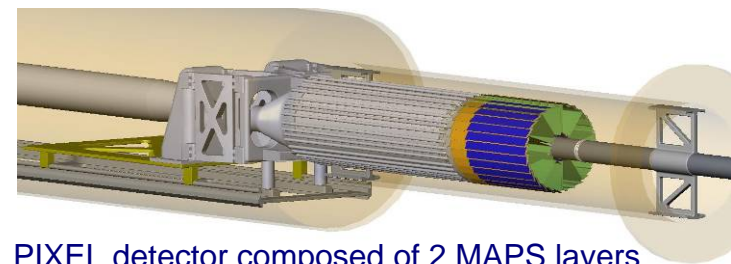
PHASE-1: sensor with binary output without zero suppression



Prototype detector composed of 3 sectors with PHASE-1 sensors

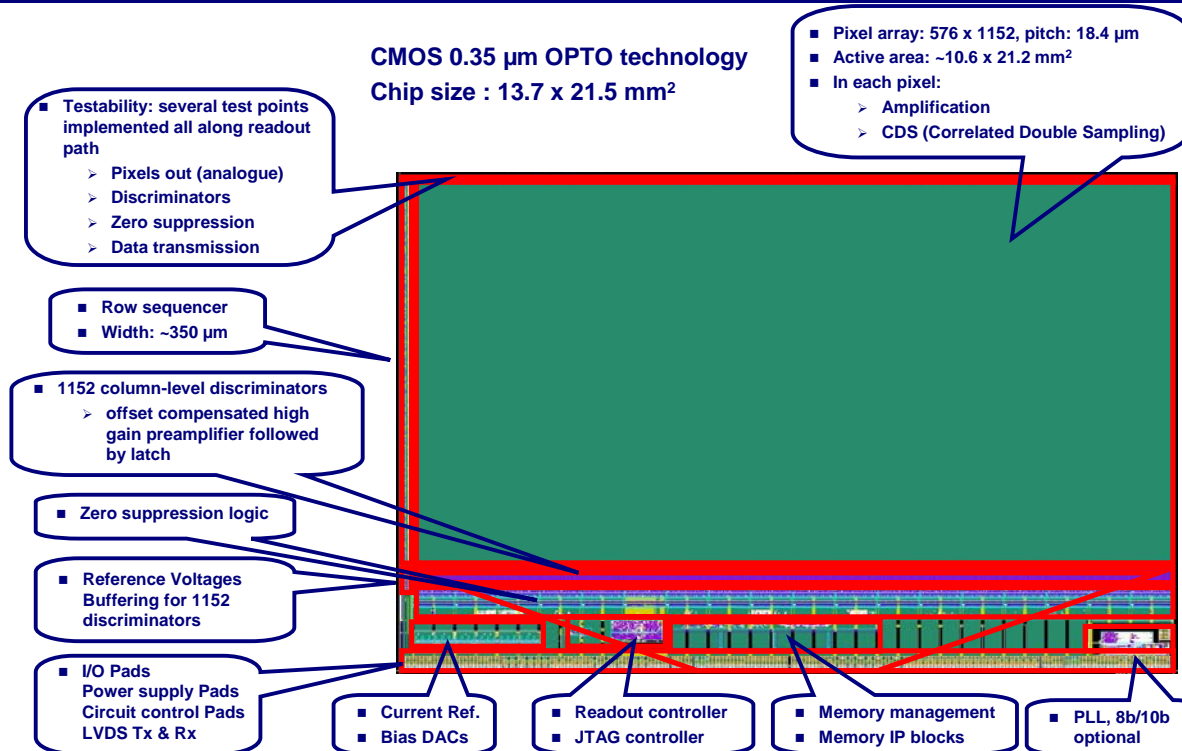
- ↪ 2013: The pixel detector composed with 2 layers of ULTIMATE sensors will be installed

ULTIMATE: sensor with binary output and with zero suppression logic



PIXEL detector composed of 2 MAPS layers

MIMOSA26: Sensor for EUDET Beam Telescope



Collaboration with
IRFU/Saclay

Main characteristics of MIMOSA26 sensor equipping EUDET BT:

- ↪ Column // architecture with in-pixel Amp & CDS and end-of-col. discrimination, followed by \emptyset
- ↪ Active area: 21.2x10.6 mm² ,1152 x 576 pixels, pitch: 18.4 μm $\rightarrow \sigma_{\text{sp.}} < \sim 4 \mu\text{m}$
- ↪ Read out time $< \sim 100 \mu\text{s}$ (10^4 frames/s) \rightarrow suited to $>10^6$ particles/cm²/s
- ↪ Yield $\sim 90\%$ (75% fully functional sensors thinned to 120 μm + 15% (showing one bad row or column)
- ↪ Thinning yield to 50 μm $\sim 90\%$

MIMOSA26 with high resistivity EPI layer (1)

- Charge collection efficiency for the seed pixel, and for 2x2 and 3x3 pixel clusters

EPI layer	Standard (~10 Ω.cm) 14 μm			High resistivity (~400 Ω.cm)			
	Seed	2x2	3x3	EPI thickness	seed	2x2	3x3
CCE (⁵⁵ Fe source)	~21%	~ 54 %	~ 71 %	10 μm	~ 36 %	~ 85 %	~ 95 %
				15 μm	~ 31 %	~ 78 %	~ 91 %
				20 μm	~ 22 %	~ 57 %	~ 76 %

- Signal to noise ratio for the seed pixel before irradiation and after exposure to a fluence of $6 \times 10^{12} n_{eq}/cm^2$

EPI layer	Standard (~10 Ω.cm) 14 μm		High resistivity (~400 Ω.cm)		
	Before irradiation	After $6 \times 10^{12} n_{eq}/cm^2$	EPI thick	Before irradiation	After $6 \times 10^{12} n_{eq}/cm^2$
S/N at seed pixel (¹⁰⁶ Ru source)	~ 20 (230 e ⁻ /11.6 e ⁻)	10.7	10 μm	~ 35	22
			15 μm	~ 41	28
			20 μm	~ 36	-----

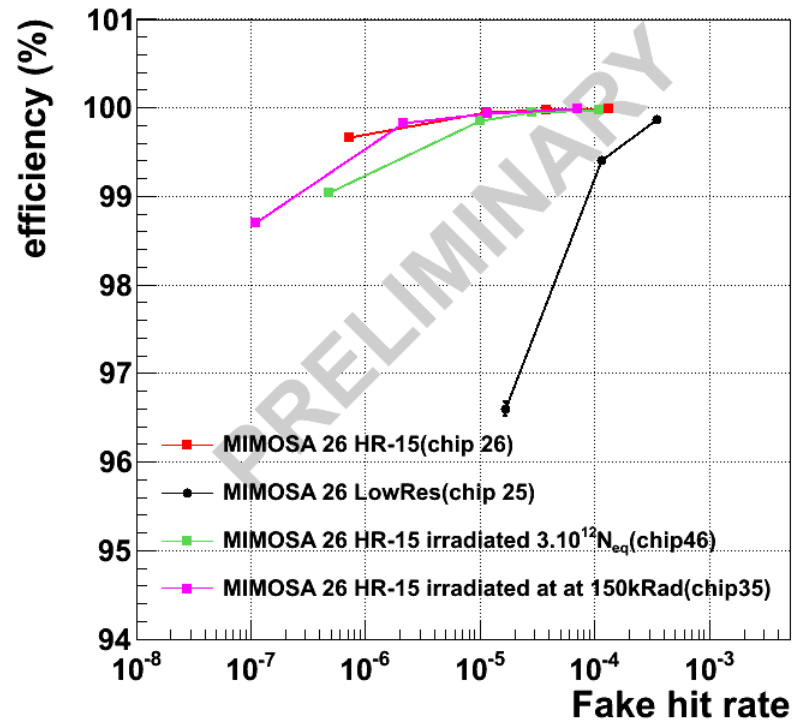
MIMOSA26 with high resistivity EPI layer (2)

■ Beam test at CERN SPS (120 GeV pions)

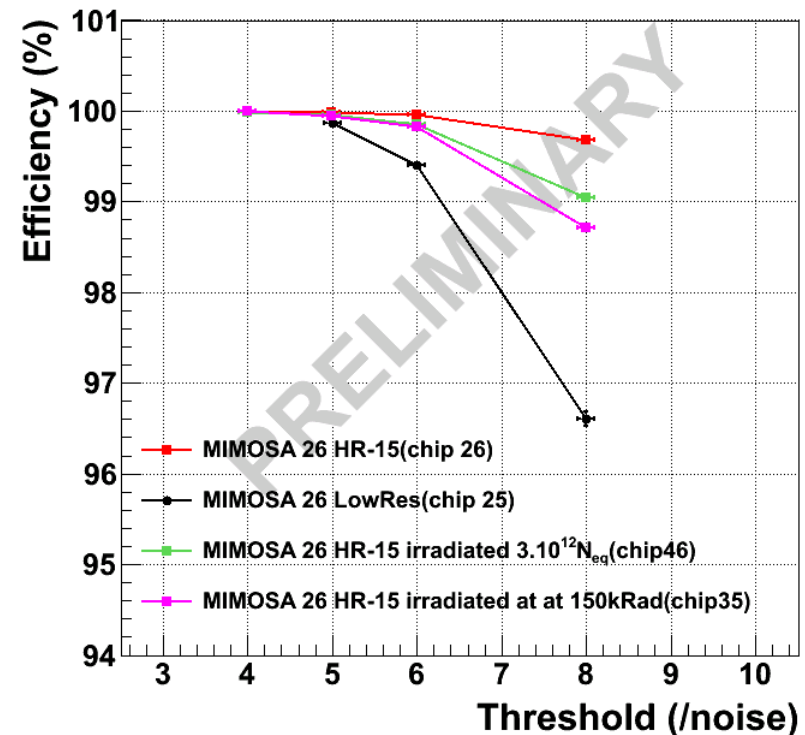
↳ Test conditions:

- 50 MHz to emulate the longer integration time in ULTIMATE
- 35 °C temperature!

Efficiency vs Fake hit rate

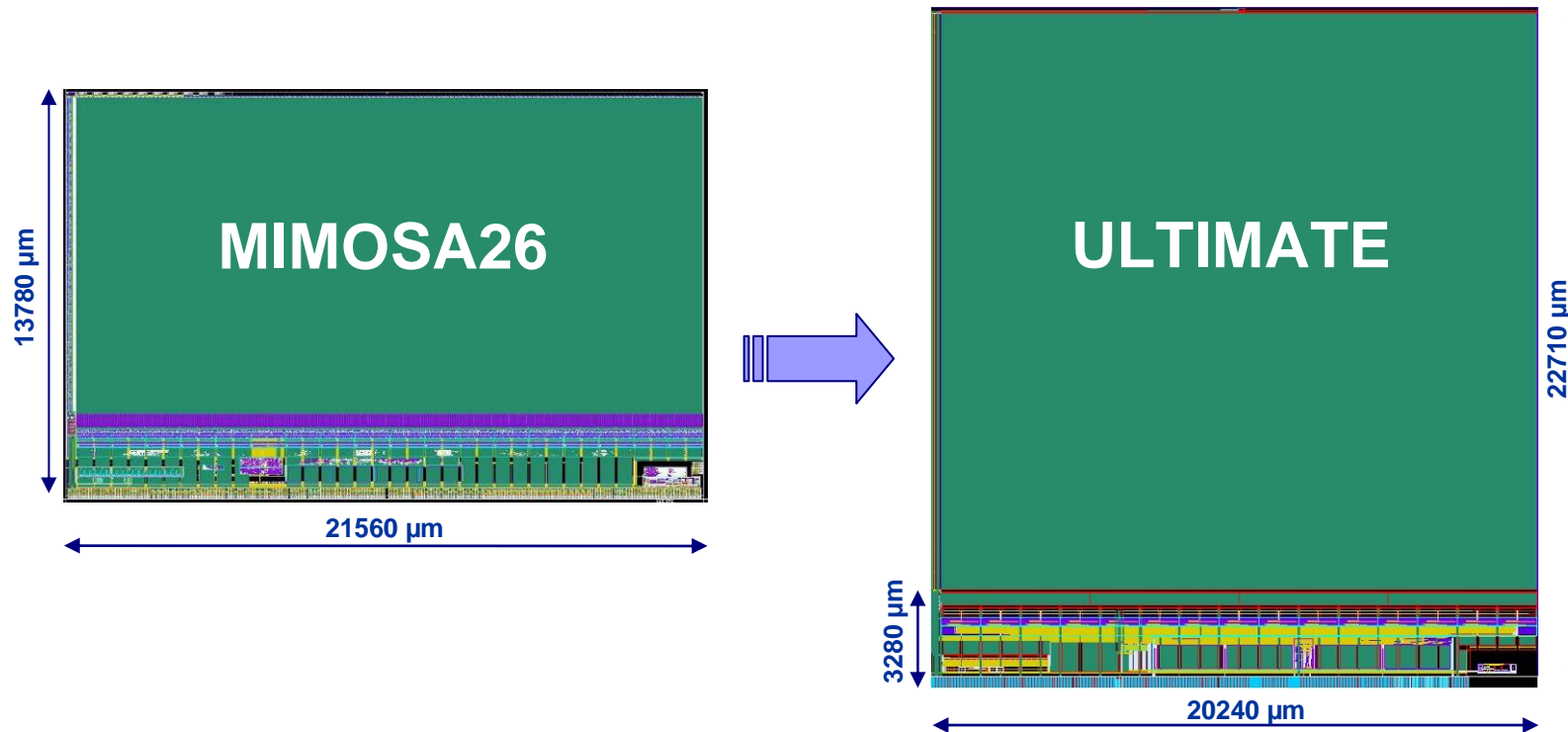


Efficiency vs Threshold



resolution < 5um

ULTIMATE: Extension of MIMOSA26



- Half reticle 1152 x 567 pixel matrix
- Temperature ~20 °C
- Light power consumption constrains
- Space resolution ~4 μm
- No constrains on radiation tolerance

- Full reticle 960 x 928 pixel matrix
 - ↳ Longer integration time ~200 μs
- Temperature 30-35 °C
- Power consumption ~100 mW/cm²
- Space resolution < 10 μm
- 150 kRad / yr & few 10¹² Neq /cm² /yr

→ **Optimisation**

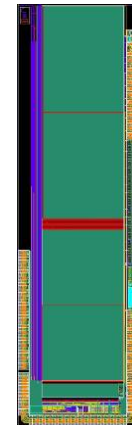
ULTIMATE Design & Optimisation (1)

■ **Reduction of power dissipation:**

- ↪ Power supply voltage reduced from 3.3 to 3 V
- ↪ Optimisation of pixel pitch v.s. non ionising radiation tolerance
 - Larger pitch: 18.4 μm \rightarrow 20.7 μm
 - Shorter integration time: 185.6 μs
 - \rightarrow Validated by a small prototype
- ↪ Optimisation of power consumption of some blocks
- \rightarrow Estimated power consumption ~ 130 mW/cm²

■ **Pixel improvement: charge collection, radiation tolerance**

- ↪ High resistivity EPI substrate & radiation tolerance design
 - Lab test with ⁵⁵Fe at 35 °C and integration time imposed by the STAR requirements shows:
 - Conversion gain is improved by a factor of two
 - ENC $>\sim 10$ e⁻ before irradiation
 - ENC $>\sim 13$ e⁻ after irradiation with 150 kRad
 - ENC $>\sim 16$ e⁻ after irradiation with 3×10^{12} Neq/cm²
 - SNR up to 30 after irradiation
 - SNR ~ 30 for standard resistivity EPI before irradiation
 - Beam test measurements are in progress
 - Preliminary results show further performance improvements

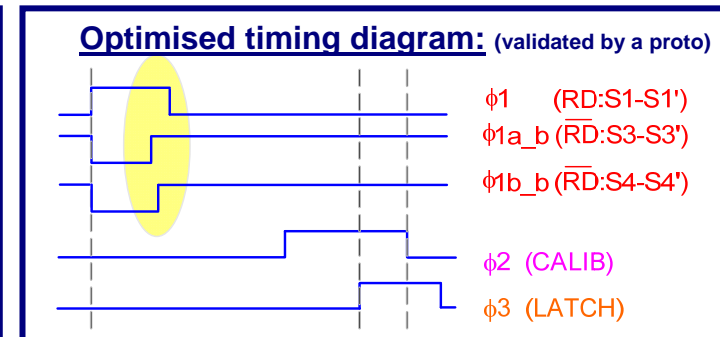
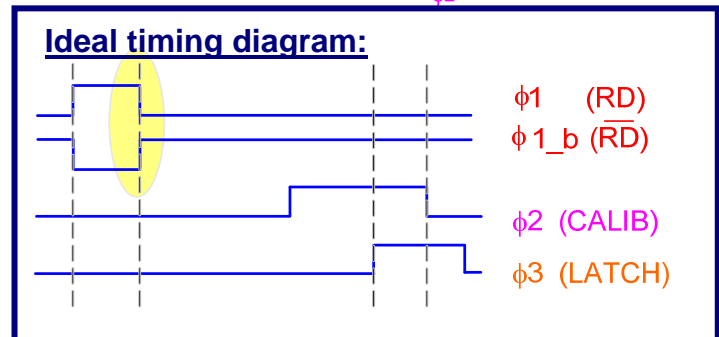
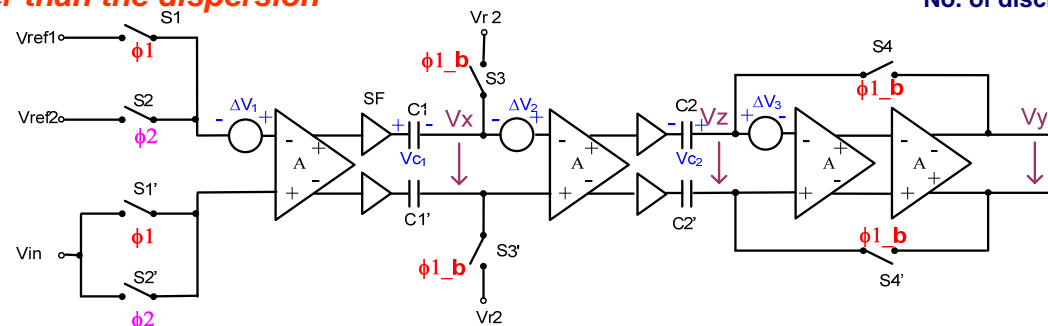
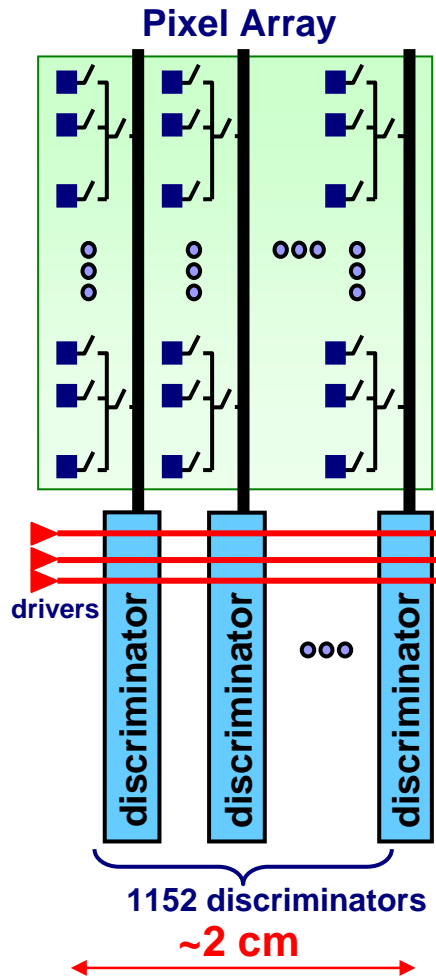
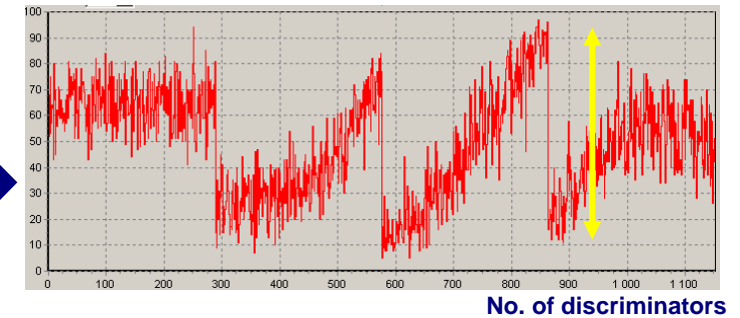


ULTIMATE Design & Optimisation (2)

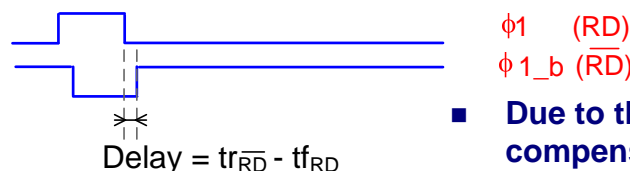
Discriminator timing diagram optimisation

↪ Threshold non-uniformity reduction

- *Mimosa26: Threshold dispersion of 1152 discriminators (divided in 4 groups)*
- *It doesn't disturb chip operation if threshold is set to be higher than the dispersion*



Obtained timing diagram due to long track:

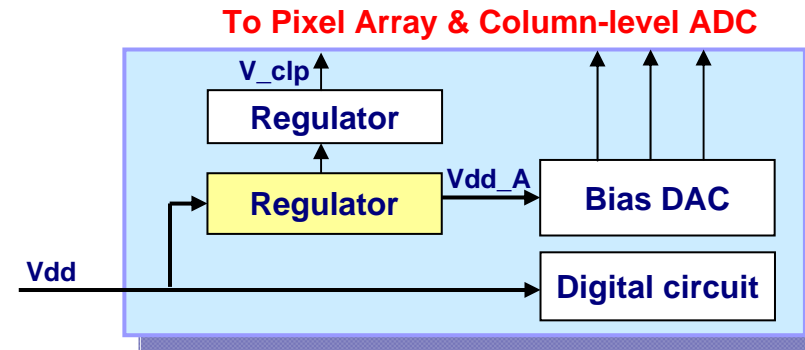
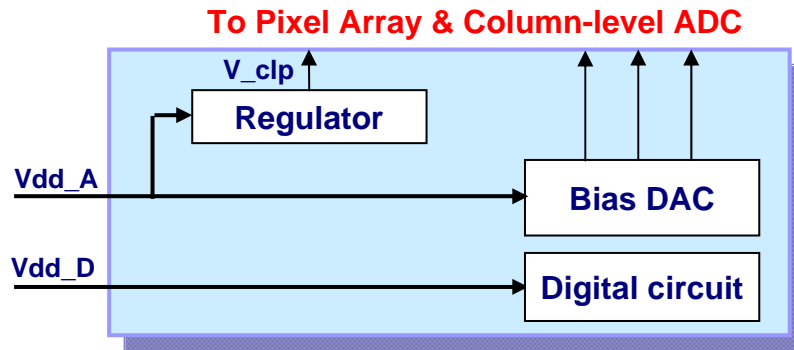


- Due to the delay, ex. charge injections by S3, S4 cannot be compensated by the auto 0 phase

ULTIMATE Design & Optimisation (3)

■ On-chip voltage regulators

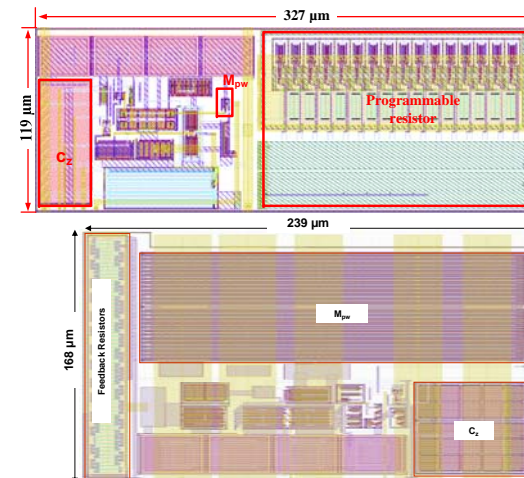
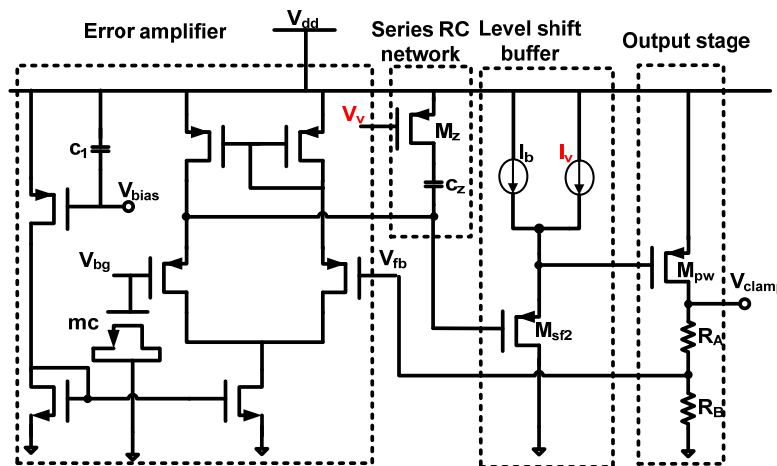
- ↪ Interferences minimisation on critical nodes
- ↪ Common power supply for A & D



See Jia Wang's poster

↪ Regulator: Low Dropout ($V_{in} = 3 - 3.3 V$)

- Characteristics: Dropout voltage: $\sim 0.3 V$, PRSS: $\sim 40 dB$, Noise: $< 1 \mu V/\sqrt{Hz}$ (freq at 10 KHz), Power consumption: $< \sim 1mW$,



For Pixels Clamping

For Analogue Power Supply

ULTIMATE Design & Optimisation (4)

■ Zero Suppression circuit (SuZe) adapted to STAR condition

- ↪ Higher hit density → larger memories
- ↪ 2 memories of 2048x32 bits
- ↪ 2 LVDS data out at 160 MHz

■ Latch up free memory may be integrated in future chip

↪ Prototype: SRAM memory module

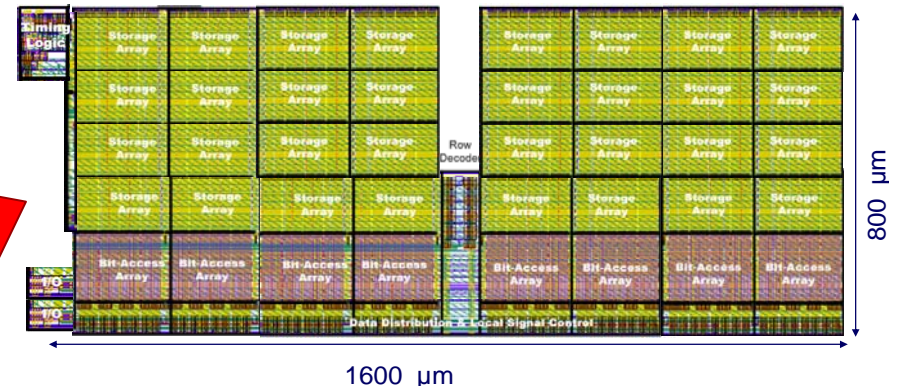
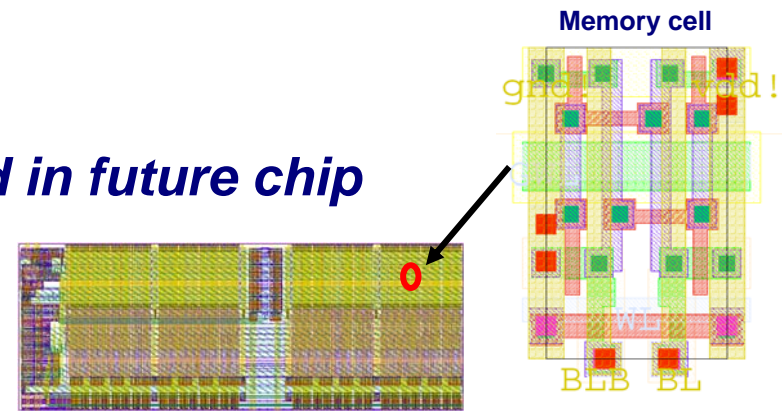
- Size: 256 x 8 bits
- Area: 1000 x 375 μm^2
- Self-timing for radiation tolerance

- Write Access Cycle: 4.3 ns
- Read Access Cycle: 4.4 ns
- Power consumption: 0.15 mA/MHz

↪ A larger memory can be built by reorganizing the designed module

[1]

Compatible with
IP block



[1] K. Kloukinas, G. Magazzu, A. Marchioro, "A Configurable Radiation Tolerant Dual-Ported Static RAM macro, designed in a 0.25 μm CMOS technology for applications in the LHC Environment", 8th Workshop on Electronics for LHC Experiments, 2002

Summary + Future plans

- **MAPS with high-resistivity epitaxial layer show excellent detection performances:**
 - ↪ Detection efficiency >99.8% (at 10^{-5} fake hit rate)
 - ↪ Improvement of radiation tolerance, non ionising radiation tolerance increased by $O(10^2)$
- **PIXEL detector for the STAR experiment based on ULTIMATE sensors will fulfil the STAR HFT requirement**
 - ↪ ULTIMATE sensor is planned to be delivered to LBL in Q1 2011
 - ↪ 400 sensors per detector
- **Engineering prototype detector equipped with PHASE-1 sensors covering 1/3 of the complete detector surface, will be installed in 2012 (physics in 2013)**
- **Nominal detector installation is scheduled 2013 (physics in 2014)**



Milestone: CMOS Sensors for the STAR PXL detector

- *LBNL-IPHC collaboration on vertex detector upgrade of STAR experiment (RHIC)*
 - ↳ *PXL: 2 layers equipped with 10/ in + 30/ out ladders, each made of 10 sensors designed by IPHC*
 - ↳ *CMOS pixel sensors → breaking edge technology: granular, thin, swift, low power*
- *November 2009: technological choice of PXL discussed in CD-1 review*
 - ↳ *formally approved in Spring 2010 → 1st vertex detector ever operated with CMOS sensors*
- *2009/2010 activities on 2 generations of sensors :*
 - ↳ *PHASE-1 (physics in 2013): validated & integrated on ladders at LBNL*
 - ↳ *ULTIMATE (physics in 2014): 4 times faster, more radiation tolerant, etc.*
 - *design under way in a new fabrication technology*
 - *main innovative features w.r.t. PHASE-1 validated on prototype tested at CERN-SPS*
- *Next steps :*
 - ↳ *submission of ULTIMATE sensor in December 2010 tests in Spring 2011*
 - ↳ *contribution to ladder testing*
 - ↳ *contribution to STAR central software development for tracking in PXL*