

# The Fate of R-Parity

Sogee Spinner  
University of Wisconsin-Madison

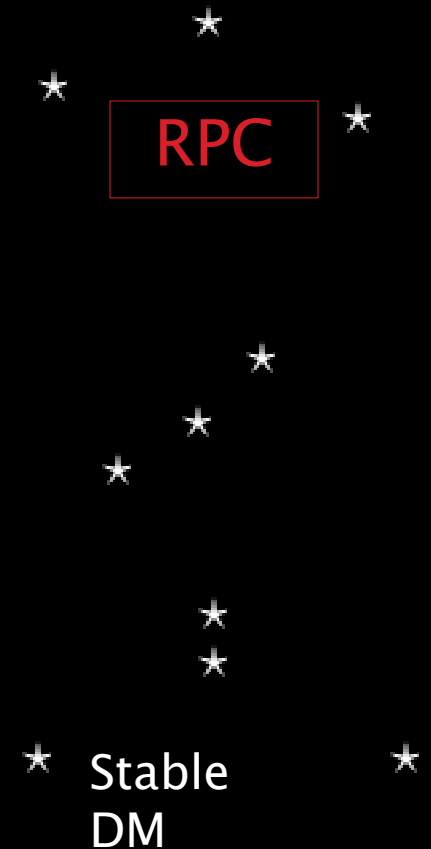
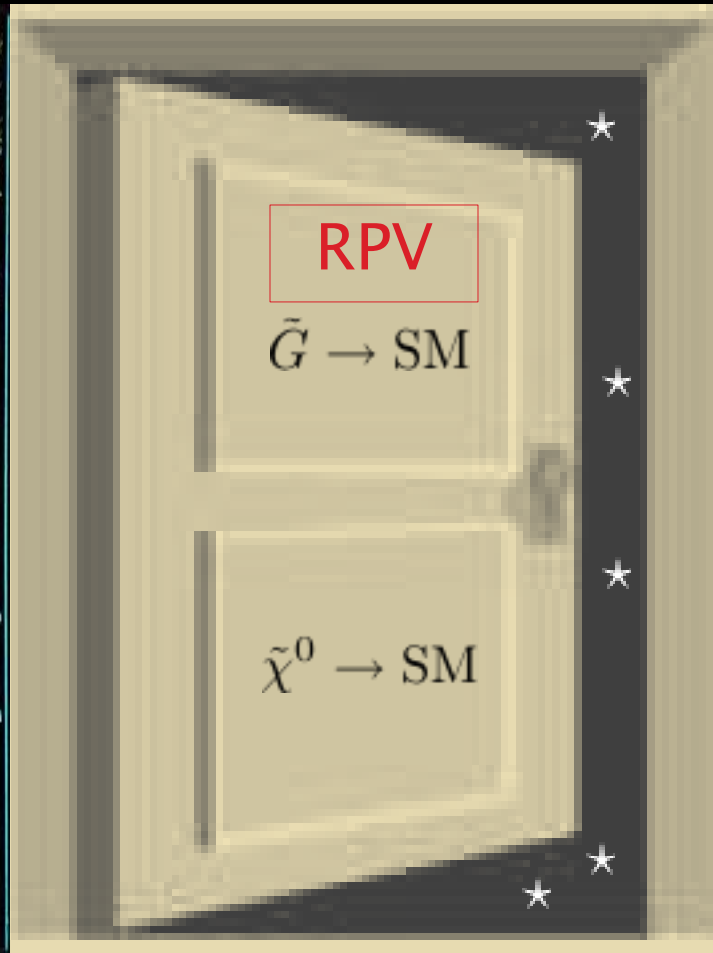
03 June 2010

Planck 2010: CERN

# References

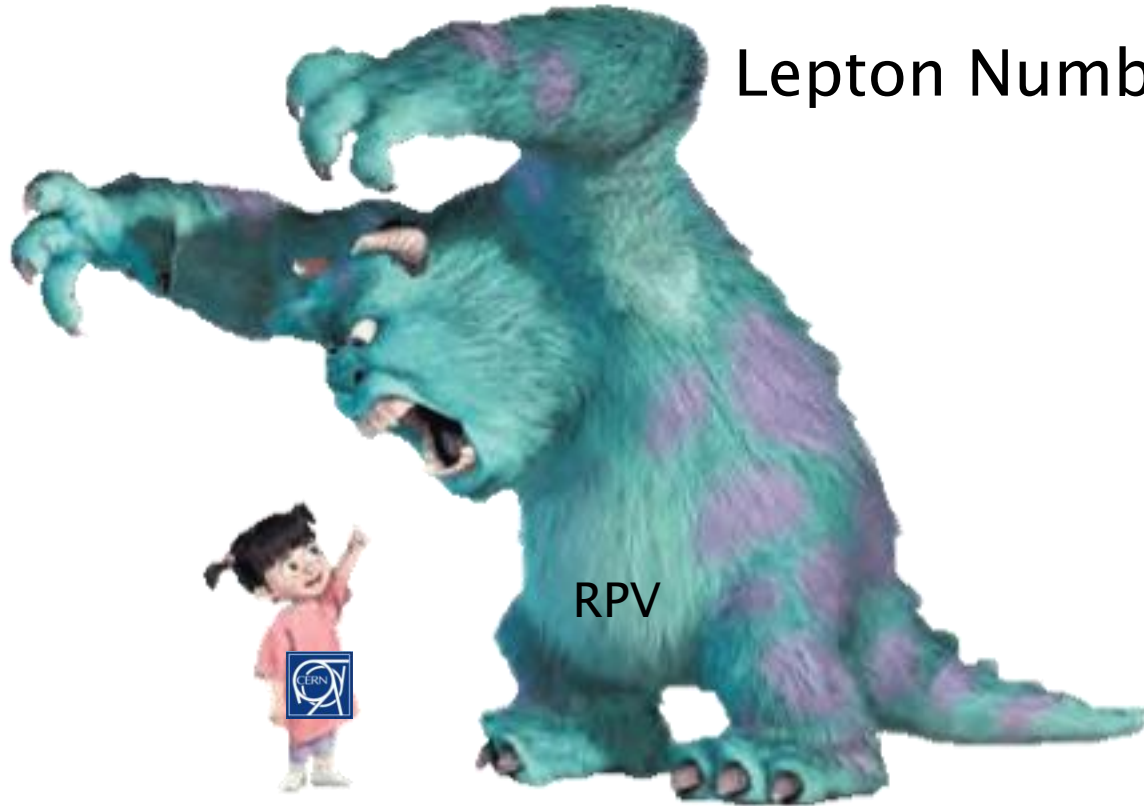
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- P. Fileviez Perez, [S.S.](#), **Phys.Rev.D80:015004**,2009
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- P. Fileviez Perez, [S. S.](#), **arXiv: 1005.4930**

# SUSY: Cosmo & Pheno hinges on status of R-parity.



# R-Parity Violation Can Be Scary!

Lepton Number Violation?



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Dark Matter?

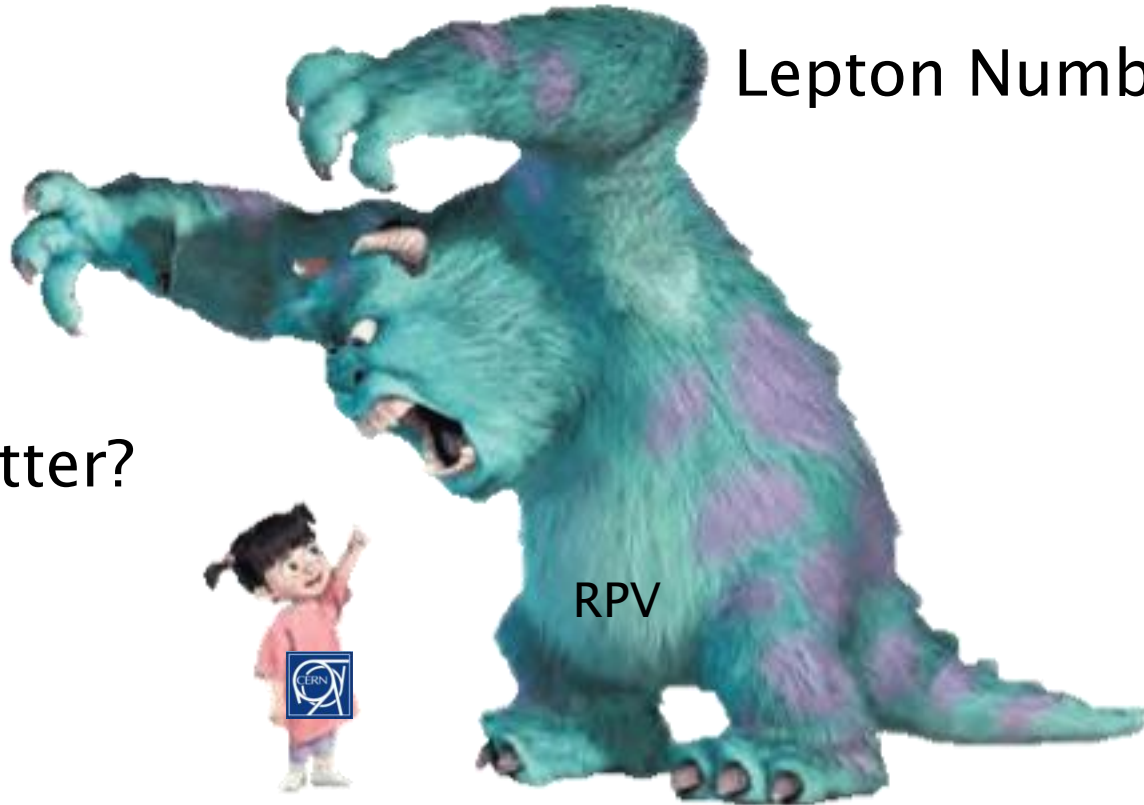




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Lepton Number Violation?

Dark Matter?



RPV

Proton Decay?!?

# But it doesn't have to be!



RPV

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**No Proton decay!**

RPV



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$m_\nu$  and lepton # is Safe!

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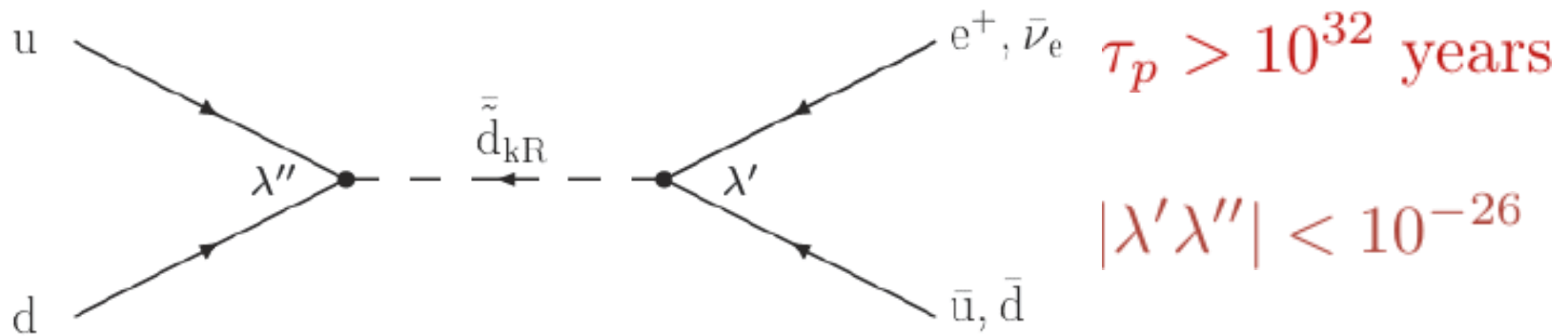
**Dark Matter Candidate: gravitino LSP**

# B - L violation: proton decay

- MSSM does not have accidental  $B-L$ :

$$W \supset \lambda'' \hat{U}^C \hat{D}^c \hat{D}^C + \lambda' \hat{Q} \hat{L} \hat{D}^C$$

- Rapid proton decay

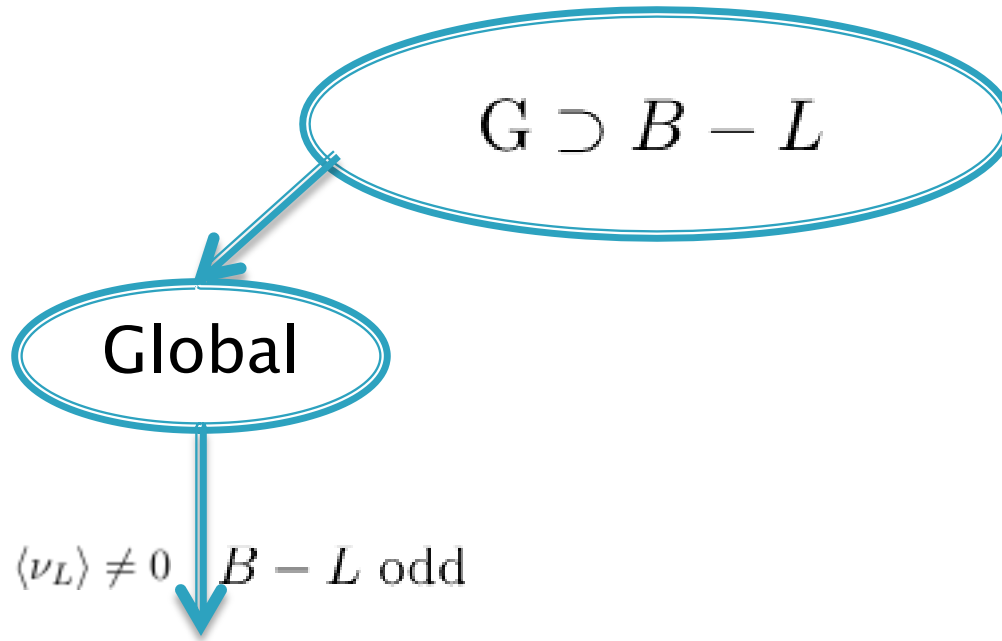


- Impose  $R = (-1)^{3(B-L)+2S}$ 
  - Discrete symmetry is unsatisfying!
  - Remember the connection to  $B - L$

# B-L and R-parity

$$G \supset B - L$$

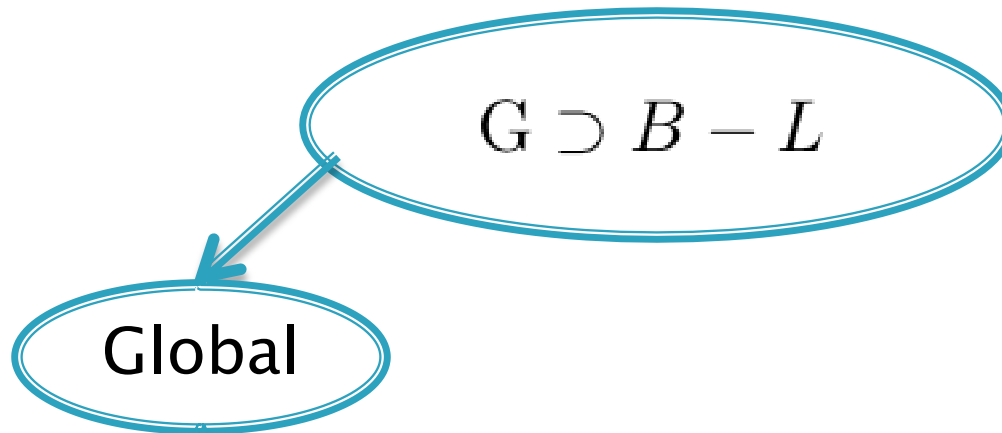
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Breaking:  $\langle \nu_L \rangle \neq 0$   $B - L$  odd



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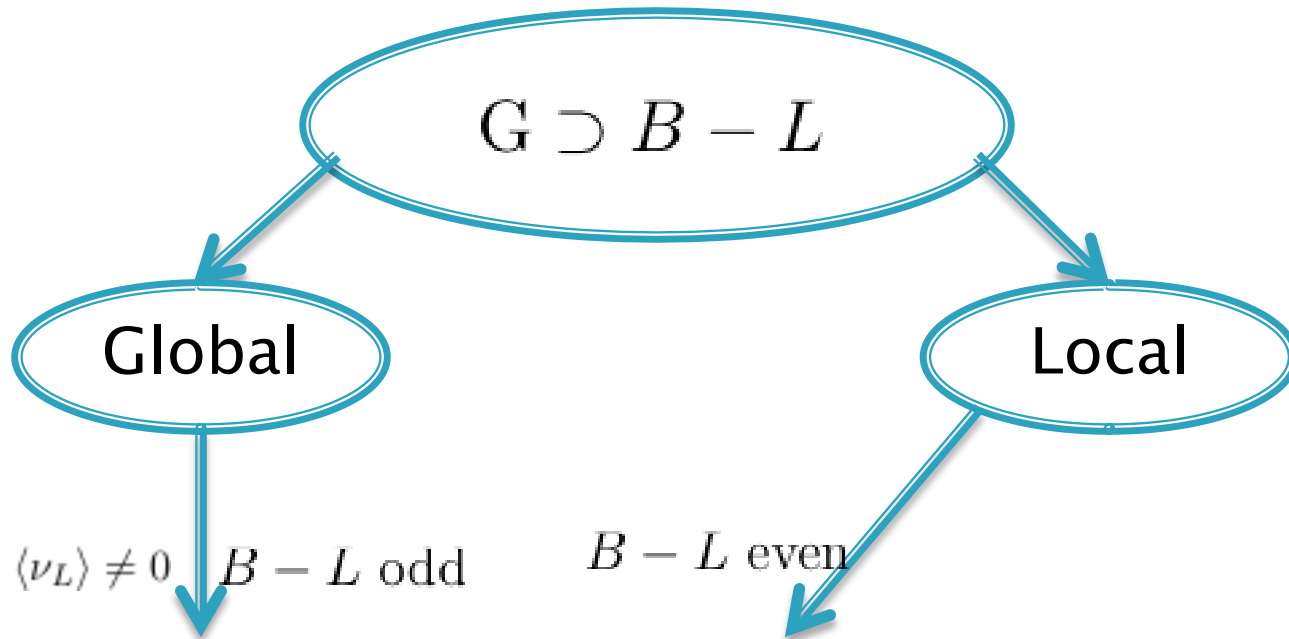


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MSSM + RPV  
Majoron Problem

Aulakh, Mohapatra '82  
Masiero, Valle '90

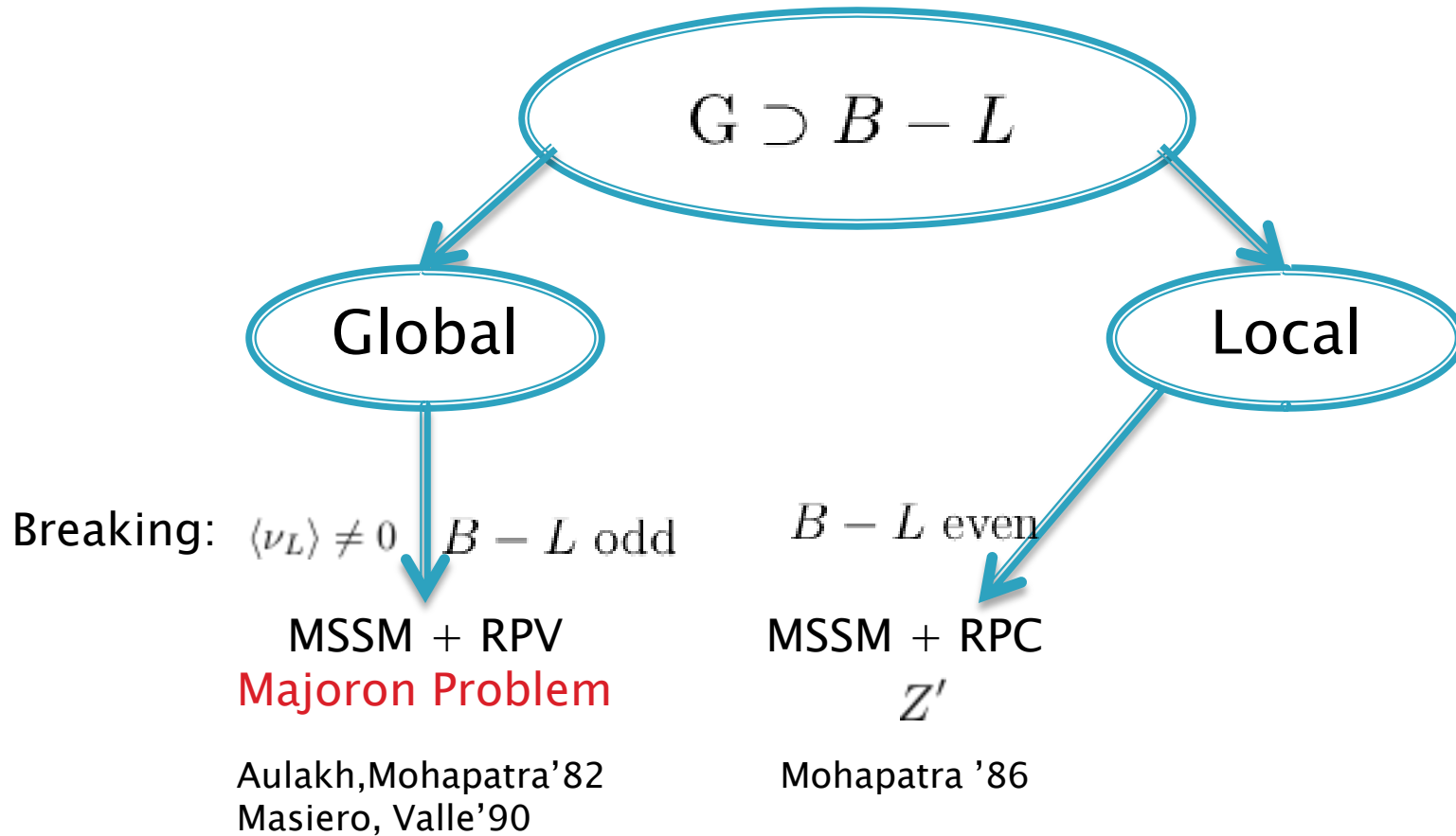
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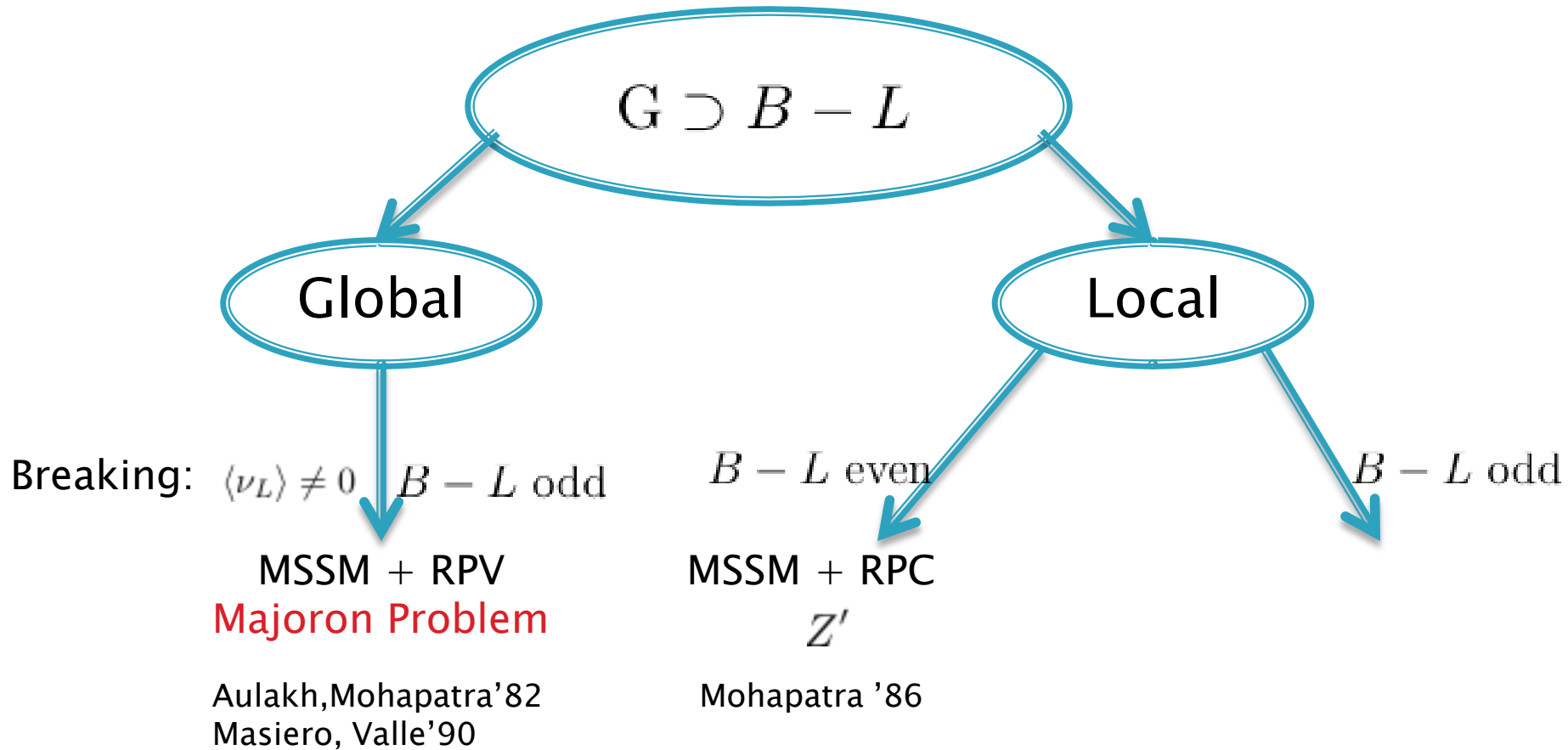
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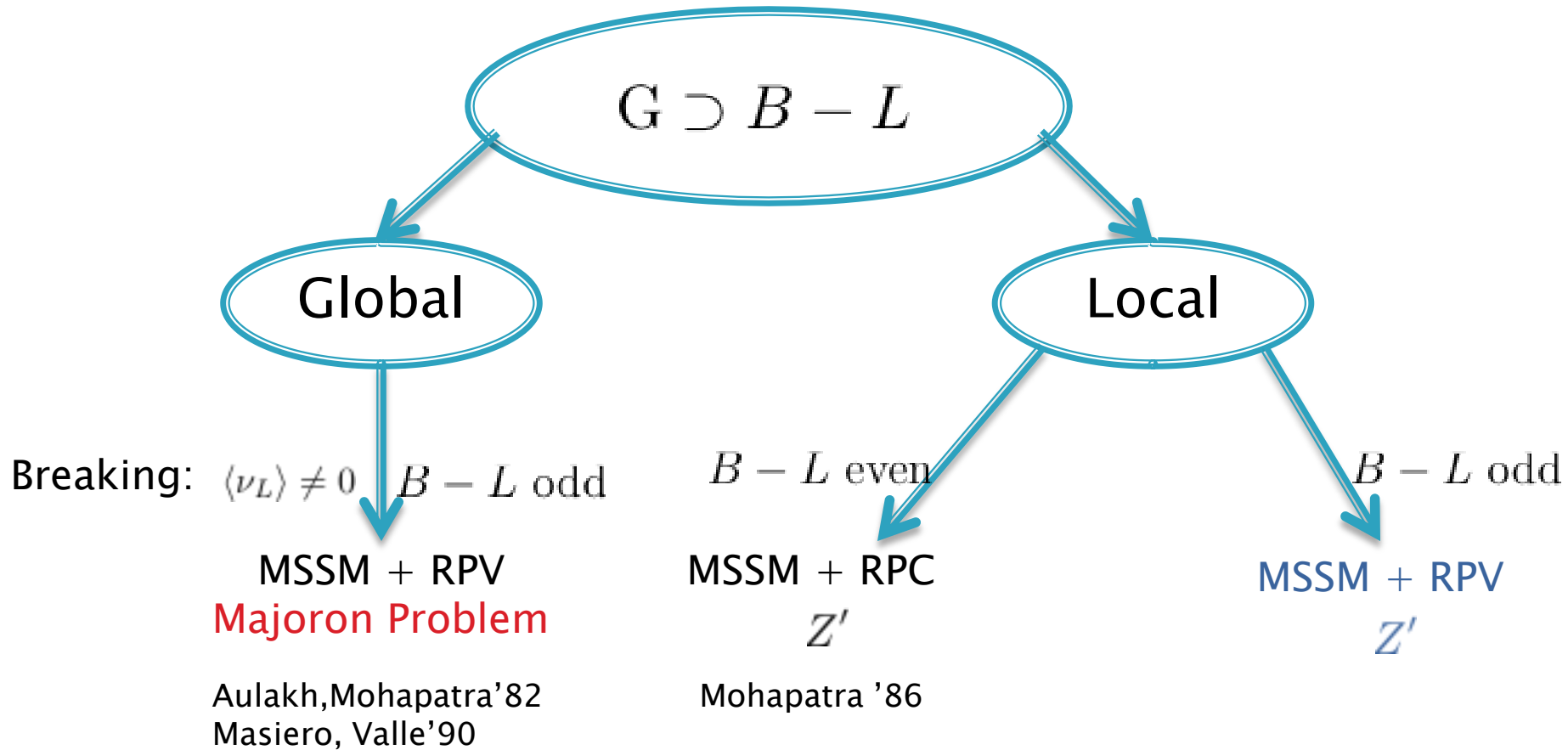
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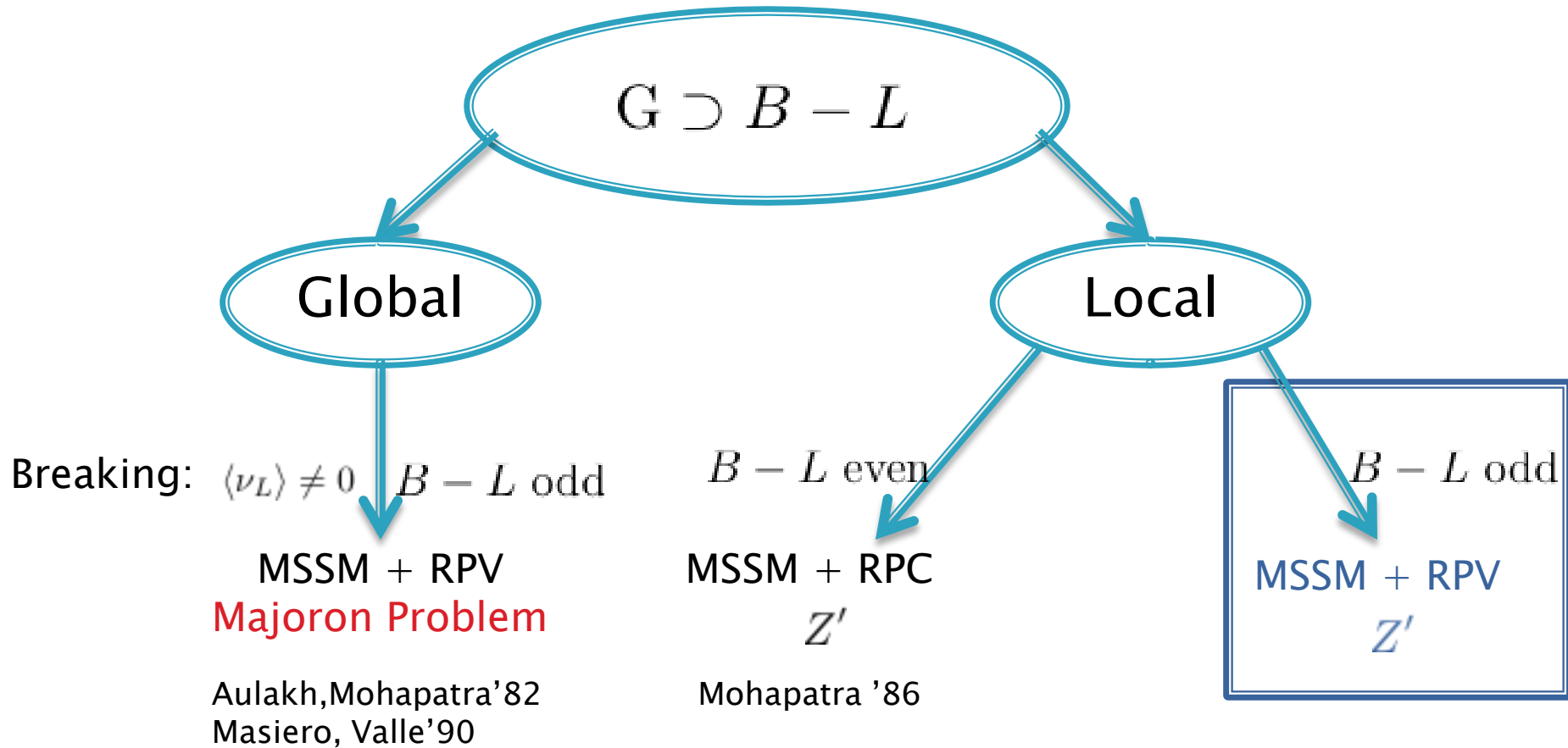


# B-L and R-parity





# B-L and R-parity



## Local B-L: $SM \times U(1)_{B-L}$

Quarks:  $B - L = \pm \frac{1}{3}$ ;    Leptons:  $B - L = \pm 1$

$\hat{N}^c \sim (1, 1, 0, 1)$  - For anomaly cancellation

$$W_{B-L} = W_{\text{MSSM}} + Y_\nu \hat{L} \hat{H}_u \hat{N}^c$$

But, how is  $B-L$  broken?

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But, how is  $B-L$  broken?

$$\langle \tilde{N}^c \rangle \equiv \frac{v_R}{\sqrt{2}} \neq 0 : SM \otimes U(1)_{B-L} \longrightarrow SM$$

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- VEV:  $v_R = \sqrt{\frac{-8M_{\tilde{N}^c}^2}{g_{BL}^2}}$  SM like VEV

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- Generate neutrinos masses.
- Only bilinear RPV; B conserved, no proton decay!

$$\lambda'' = 0 \text{ in } \lambda'' \hat{U}^C \hat{D}^C \hat{D}^C,$$

# Higgs Sector For Left-Right: RPC

$$SU(2)_L \times SU(2)_R \times U(1)_{B-L}$$

$$\begin{aligned} \Delta &= \begin{pmatrix} \frac{1}{\sqrt{2}}\Delta^+ & \Delta^{++} \\ \Delta^0 & -\frac{1}{\sqrt{2}}\Delta^+ \end{pmatrix} \sim (3, 1, 2) & \Delta^c &= \begin{pmatrix} \frac{1}{\sqrt{2}}\Delta^{c+} & \Delta^{c++} \\ \Delta^{c0} & -\frac{1}{\sqrt{2}}\Delta^{c+} \end{pmatrix} \sim (3, 1, -2) \\ \bar{\Delta} &= \begin{pmatrix} \frac{1}{\sqrt{2}}\bar{\Delta}^- & \bar{\Delta}^0 \\ \Delta^{--} & -\frac{1}{\sqrt{2}}\Delta^- \end{pmatrix} \sim (3, 1, -2) & \bar{\Delta}^c &= \begin{pmatrix} \frac{1}{\sqrt{2}}\bar{\Delta}^{c-} & \bar{\Delta}^{c0} \\ \Delta^{c--} & -\frac{1}{\sqrt{2}}\Delta^{c-} \end{pmatrix} \sim (3, 1, 2) \end{aligned}$$

Plus  $S \sim (1, 1, 0)$  or nonrenormalizable term

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$$\text{Again, } \langle \tilde{L}^C \rangle = \langle \tilde{N}^c, \tilde{E}^C \rangle = (v_R, 0)$$

The Simplest Supersymmetric Left-Right Model

# Remaining Questions

- Why is right-handed sneutrino tachyonic?
  - Would like to mimic the success of REWSB in MSSM.
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- Why is right-handed sneutrino tachyonic?
  - Would like to mimic the success of REWSB in MSSM.
  - But RSB not possible in the minimal model (no large Yukawas).
- What about even  $B-L$  Higgs (canonical Model)
  - Maybe RPC more important than minimalism.
- Answers are related!

# Canonical B-L Model

- Simplest extension with Yukawa couplings

$$\hat{X}, \hat{\bar{X}} \sim -2, 2$$
$$W \supset f \hat{N}^C \hat{N}^C \hat{X} + \mu_X \hat{X} \hat{\bar{X}}$$

- $f$  can drive radiative symmetry breaking but

$$m_X^2 < 0 \text{ or } m_{\tilde{N}^C}^2 < 0$$

- Or what is [The Fate of R-parity?](#)

# The Fate of $R$ -parity

- Depends solely on  $m_{\tilde{N}c}^2$

$$m_{\tilde{N}c}^2 > 0 \rightarrow \text{RPC}$$

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- RGEs (one family approximation):

$$X_X \equiv m_X^2 + 2m_{\tilde{N}^c}^2 + 4a_f^2$$

$$16\pi^2 \frac{dm_{\tilde{N}^c}^2}{dt} = 8 f^2 X_X - 3 g_{BL}^2 M_{BL}^2$$

$$16\pi^2 \frac{dm_X^2}{dt} = 4 f^2 X_X - 12 g_{BL}^2 M_{BL}^2$$

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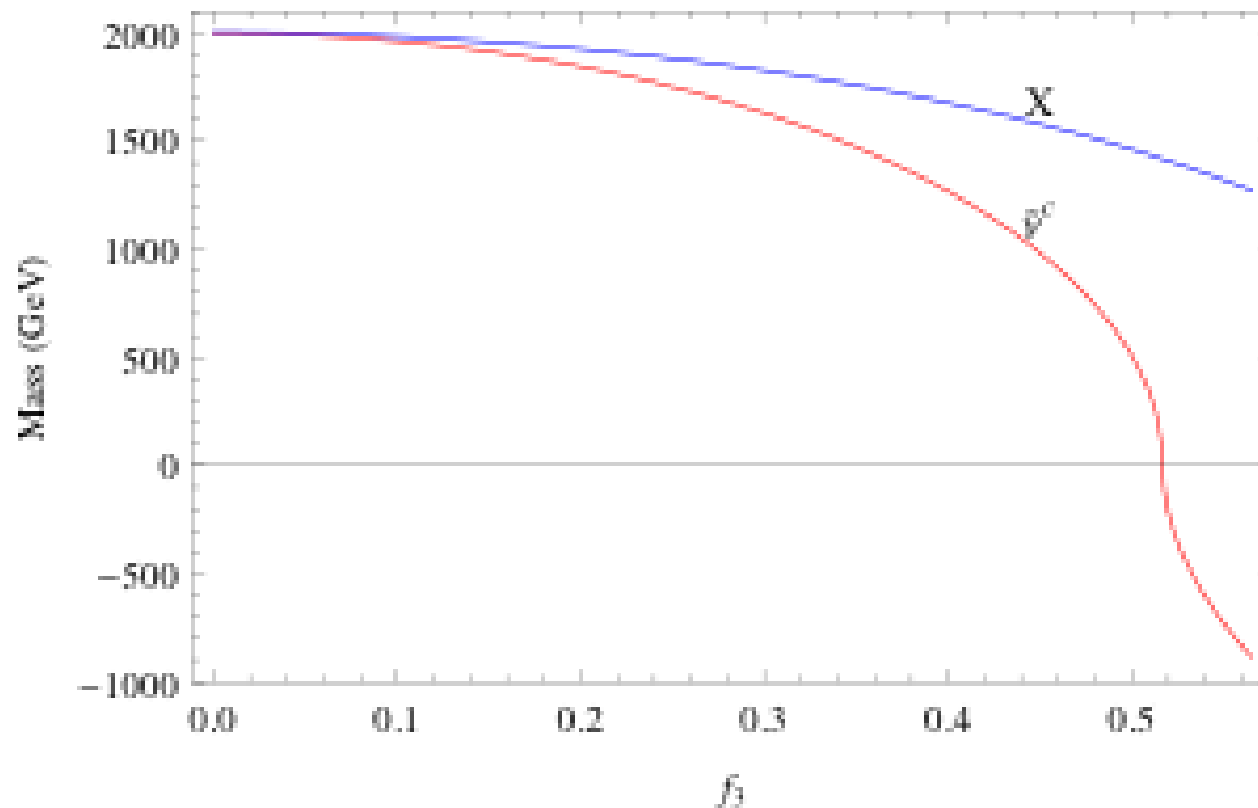
$$16\pi^2 \frac{dm_{\tilde{N}^c}^2}{dt} = \boxed{8} f^2 X_X - \boxed{3} g_{BL}^2 M_{BL}^2$$

$$16\pi^2 \frac{dm_X^2}{dt} = \boxed{4} f^2 X_X - \boxed{12} g_{BL}^2 M_{BL}^2$$


***Favors  $R$ -parity Violation!!!***

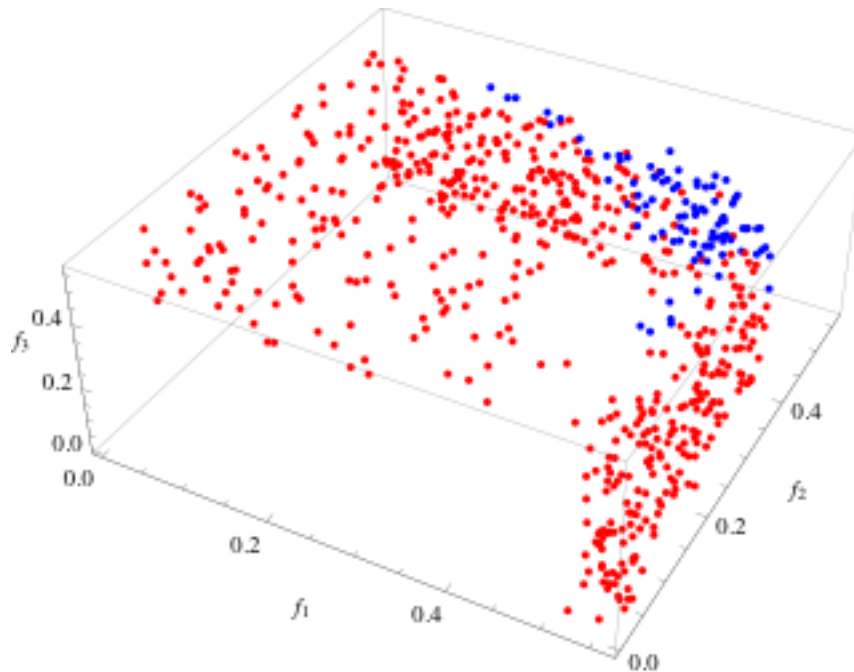
- Check if RSB is possible.
- MSUGRA boundary conditions
- Soft masses verses  $f_3$ :

$$m_0 = 2000 \text{ GeV}; M_{1/2} = 200 \text{ GeV}; A_0 = 0$$



# Three Families

- $m_X^2$  enhanced by  $tr f$ ; RSB and RPC possible



$$\begin{aligned} m_0 &= 2000 \text{ GeV}, \\ M_{1/2} &= 200 \text{ GeV}, \\ A_0 &= 0 \end{aligned}$$

RPV 5 times more likely:  
even in the canonical model, RPV is probable.

# Conclusion

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- Minimal  $B-L$  models:  $R$ -parity violation.
- Add  $B-L$  even Higgs: radiative symmetry breaking but  $R$ -parity violation still very likely.
- In both cases  $R$ -parity violation is viable:
  - no proton decay.
  - TeV scale theories great for LHC.
  - dark matter still possible.