# FROM SOFT WALLSTO INFRARED BRANES 

Gero von Gersdorff (École Polytechnique) Planck 2010

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Soft Wall models are generalizations of RS that do not possess an IR brane

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Singularities are "good" if
$e^{-A(y)} \rightarrow 0$ faster than $\left(y_{s}-y\right)^{\frac{1}{4}}$
Gubser'02, Forste et al '02 Cabrer, GG \& Quirós '09

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- Problem of IR brane: Planck suppressed operators become only TeV suppressed, predictions for $\mathrm{p}>\mathrm{TeV}$ hard (KK masses)
- Need to specify infinitely many operators?
- Soft Walls give "resolution" of IR brane


## CLASSES OF SOFT WALLS

Soft Walls are rather restricted by demanding:

- A Gap in the spectrum should exist
- The singularity should be a "good"

Then, asymptotics near singularity fall into 2 classes

CLASSES OF SOFT WALLS

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Proper Length coordinates
$d s^{2}=e^{-2 A(y)} d x^{\mu} d x^{\nu} \eta_{\mu \nu}+d y^{2} \quad d s^{2}=e^{-2 A(z)}\left(d x^{\mu} d x^{\nu} \eta_{\mu \nu}+d z^{2}\right)$

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\begin{gathered}
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Integrate over the region near the singularity
$S=\int_{0}^{y_{s}} d^{5} x \mathcal{L}_{\text {bulk }} \longrightarrow S=\int_{0}^{y_{1}} d^{5} x\left[\mathcal{L}_{\text {bulk }}+\mathcal{L}_{\mathrm{SW}} \delta\left(y-y_{1}\right)\right]$
Equivalent description of SW in terms of IR brane


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Equivalent description of SW in terms of IR brane


- For $y_{1}$ close to $y_{s}$ close $\mathbb{R}$ Lagrangian "universal"
- Facilitates comparison with standard 2 brane compactif.
- IR Lagrangian makes sense even for $p>T e V$
- Useful approximation scheme: Approximate "new" bulk by RS metric

SCALAR FIELDS

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\mathcal{L}_{\text {bulk }}=\frac{1}{2} \sqrt{-g}\left(g^{M N} \partial_{M} \psi \partial_{N} \psi+M^{2} \psi^{2}\right)
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\begin{aligned}
& \left(e^{2 A(y)} \partial_{\mu} \partial^{\mu}+\partial_{y}^{2}-4 A^{\prime}(y) \partial_{y}-M^{2}\right) K\left(x, x^{\prime} ; y\right)=0 \\
& K\left(x, x^{\prime} ; y_{1}\right)=\delta\left(x-x^{\prime}\right) \quad \psi(x, y)=\int d^{4} x^{\prime} K\left(x, x^{\prime} ; y\right) \psi\left(x^{\prime}, y_{1}\right)
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\begin{gathered}
\mathcal{F}(p)=K^{\prime}\left(p, y_{1}\right) \\
\mathcal{S}_{\mathrm{SW}}=\frac{1}{2} \int \frac{d^{4} p}{(2 \pi)^{4}} \sqrt{-g(y)} \mathcal{F}(p) \psi(-p, y) \psi(p, y) \delta\left(y-y_{1}\right)
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SWI $\mathcal{F}(p)=e^{A\left(z_{1}\right)} \frac{\sqrt{-p^{2}} J_{\alpha+1}\left(\sqrt{-p^{2}} \Delta_{z}\right)}{J_{\alpha}\left(\sqrt{-p^{2}} \Delta_{z}\right)} \quad \alpha=\frac{4-\nu^{2}}{2\left(\nu^{2}-1\right)}$ :

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SW2 $\mathcal{F}\left(p^{2}\right)=e^{A\left(z_{1}\right)} \sqrt{-p^{2}} \tan \left(c_{\beta}\left[-\frac{p^{2}}{\rho^{2}}\right]^{\frac{1}{4 \beta}}\right) \quad \beta=\frac{\sigma-1}{2 \sigma}$

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Asymptotic Spectrum (n large)
Approximation $\quad m_{n} z_{s} / \pi=n+0.34$
Numerical $\quad m_{n} z_{s} / \pi=(10.40,50.37,100.36,500.35)$

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Compare with exact solution (analytic only for $\beta=\frac{1}{4}$ )
Exact spectrum is $m_{n}=\sqrt{12 n} \rho$
Conicides with previous method since $c_{\frac{1}{4}}=\frac{\pi}{12}$

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- Future work: higher spin, symmetry breaking...

