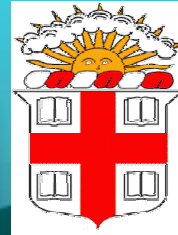


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Vanishing Dimensions and Planar Events at the LHC



Outline

- The Gauge Hierarchy Problem
- Vanishing Dimensions*)
 - Gravity in Reduced Dimensions
 - Astrophysical and Cosmological Consequences
 - Collider Phenomenology
- Conclusions

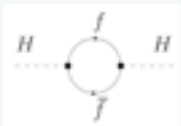
*)This talk is based on the [L. Anchordoqui, D.C. Dai, M. Fairbairn, G.L., and D. Stojkovic, **arXiv:1003.5914**, submitted to PRL]



Gauge Hierarchy Problem

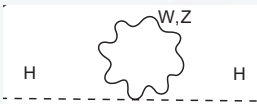
- Loop corrections to the Higgs mass diverge quadratically with the highest scale Λ , making the Higgs mass extremely fine-tuned:

$$L_H = D_\nu \Phi^\dagger D^\nu \Phi - \mu^2 \Phi^\dagger \Phi + \frac{\lambda}{2} (\Phi^\dagger \Phi)^2 - \sum_f Y_f \Phi \bar{\psi}_f \psi_f$$

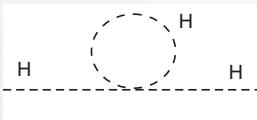


$$\Delta m_H^{(f)} = i \frac{Y_f^2}{2} \int_0^\Lambda \frac{d^4 k}{(2\pi)^4} \text{Tr} \left(\frac{i}{\not{k} - m_f} \frac{i}{\not{k} + \not{p} - m_f} \right) \sim -\Lambda^2 \text{Tr}(I_4) \frac{Y_f^2}{32\pi^2} = -\Lambda^2 \frac{Y_f^2}{8\pi^2}$$

Reduced 4-momentum of the Higgs



$$\Delta m_H^{(V)} = i \frac{g^2(+g'^2)}{4} \int_0^\Lambda \frac{d^4 k}{(2\pi)^4} \frac{1}{k^2 - m_{W(Z)}^2} \sim \Lambda^2 \frac{g^2(+g'^2)}{64\pi^2}$$



$$\Delta m_H^{(H)} = i6\lambda \int_0^\Lambda \frac{d^4 k}{(2\pi)^4} \frac{1}{k^2 - m_H^2} \sim \Lambda^2 \frac{3\lambda}{8\pi^2}$$

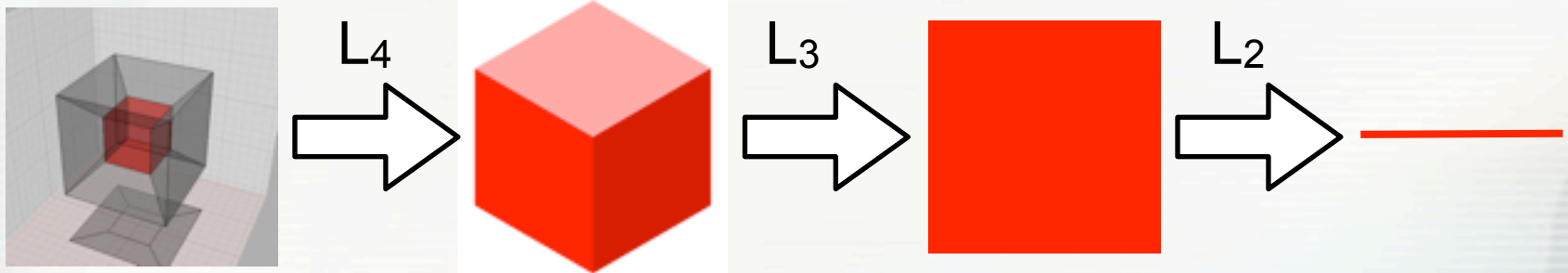
$$\Delta m_H^2 \approx \frac{3\Lambda^2}{16\pi^2 v^2} (2m_W^2 + m_Z^2 + m_H^2 - 4m_t^2)$$

Fine-tuning at the 10^{-32} level is required if Λ is as high as M_{Pl}



A Conceptually New Paradigm

- The effective **dimensionality of space depends on the length scale** we are probing
- **At short distances, the dimensions of space vanish one-by-one:**
 - at the intermediate scale, space is 3-dimensional
 - at scales $\sim 1 \text{ TeV}^{-1}$ space is effectively 2-dimensional
 - at even shorter distances (e.g., in the Big Bang) it is 1-dimensional
- Conversely, **at large distances, the dimensionality increases**; at very large distances space is effectively 4-dimensional



- **Fundamentally new framework**, conceptually opposite to the paradigm of large extra dimensions



Curing the Hierarchy Problem

- Reduce the dimensionality of space!
- In 2+1 space-time the divergence is only linear :

$$\Delta m_H^{(f)} = i \frac{Y_f^2}{2} \int_0^\Lambda \frac{d^3 k}{(2\pi)^3} \text{Tr} \left(\frac{i}{\not{k} - m_f} \frac{i}{\not{k} + \not{p} - m_f} \right) \sim -\Lambda \text{Tr}(I_3) \frac{Y_f^2}{4\pi^2} = -\Lambda \frac{3Y_f^2}{4\pi^2}$$

Reduced (2+1)-momentum of the Higgs

$$\Delta m_H^{(V)} = i \frac{g^2(+g'^2)}{4} \int_0^\Lambda \frac{d^3 k}{(2\pi)^3} \frac{1}{k^2 - m_{W(Z)}^2} \sim \Lambda \frac{g^2(+g'^2)}{8\pi^2}$$

$$\Delta m_H^{(H)} = i6\lambda \int_0^\Lambda \frac{d^3 k}{(2\pi)^3} \frac{1}{k^2 - m_H^2} \sim \Lambda \frac{3\lambda}{\pi^2}$$

- Finally, in 1+1 space-time, it is logarithmic :

$$\Delta m_H^{(f)} = i \frac{Y_f^2}{2} \int_0^\Lambda \frac{d^2 k}{(2\pi)^2} \text{Tr} \left(\frac{i}{\not{k} - m_f} \frac{i}{\not{k} + \not{p} - m_f} \right) \sim -\log \frac{\Lambda}{m_f} \text{Tr}(I_2) \frac{Y_f^2}{4\pi} = -\log \frac{\Lambda}{m_f} \frac{Y_f^2}{2\pi}$$

Reduced (1+1)-momentum of the Higgs

$$\Delta m_H^{(V)} = i \frac{g^2(+g'^2)}{4} \int_0^\Lambda \frac{d^2 k}{(2\pi)^2} \frac{1}{k^2 - m_{W(Z)}^2} \sim \log \frac{\Lambda}{M_{W(Z)}} \frac{g^2(+g'^2)}{8\pi}$$

$$\Delta m_H^{(H)} = i6\lambda \int_0^\Lambda \frac{d^2 k}{(2\pi)^2} \frac{1}{k^2 - m_H^2} \sim \log \frac{\Lambda}{m_H} \frac{3\lambda}{\pi}$$

Fine-tuning is alleviated by reducing dimensionality at high energy!



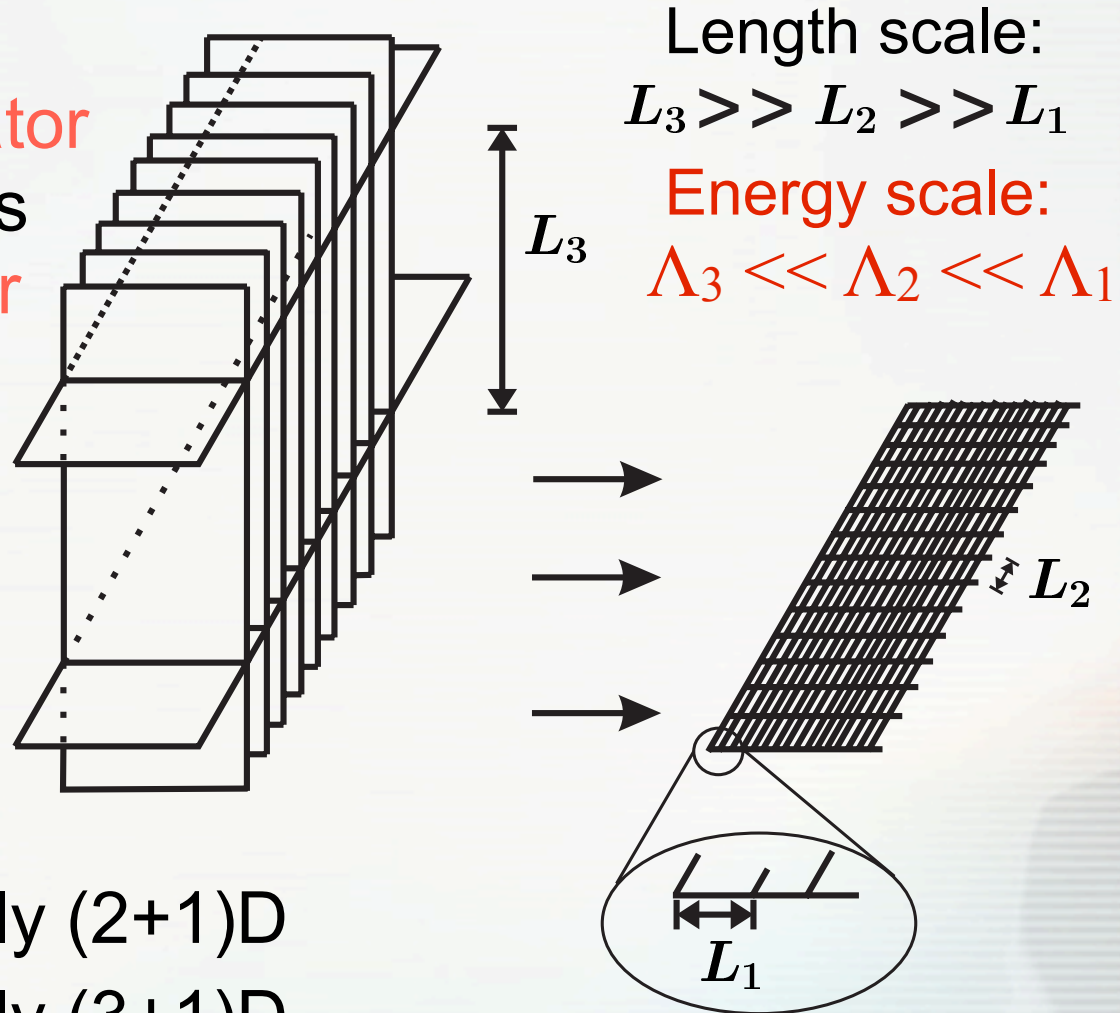
Consequences for Gravity

- Gravity in 3+1-dimensions is:
 - Complicated (hard to quantize!)
 - Non-linear
 - Perturbatively non-renormalizable
- If at short distances space has reduced dimensionality, quantum gravity needs to be described only in 2+1 or 1+1 space-time
 - In (2+1)D and (1+1)D Weyl tensor vanishes: $R_{\mu\nu\rho\sigma} = \varepsilon_{\mu\nu\alpha}\varepsilon_{\rho\sigma\beta} \left(R^{\alpha\beta} + \frac{1}{2}g^{\alpha\beta}R \right)$
- Any solution of the Einstein's equation is locally flat
- Hence in (2+1)D there are no local gravitational degrees of freedom, i.e., no gravitational waves
 - Consequently, there is no graviton in quantum theory
 - QG reduces to QM and is perturbatively renormalizable, see, e.g., [S. Carlip, J. Korean Phys. Soc. 28, **S447** (1995)]
- In (1+1)D situation is even simpler:
 - No dynamics in the metric; theory without additional fields is trivial
 - Connection with recent work on causal dynamical triangulations
 - The “arrow of time” possibly comes from maximum \mathcal{CP} in the BB (given conserved CPT), just as neutrinos are left-handed because of maximum \not{P}



Realization of the Framework

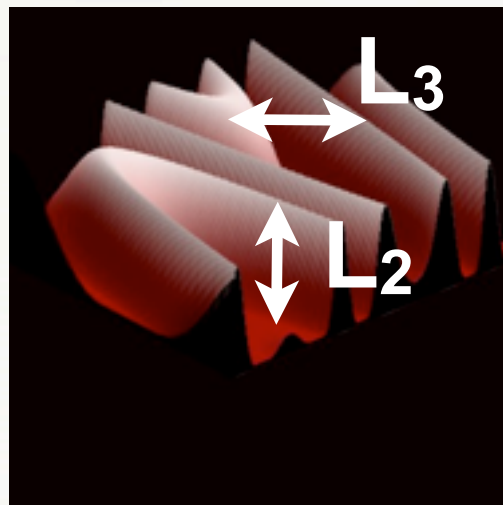
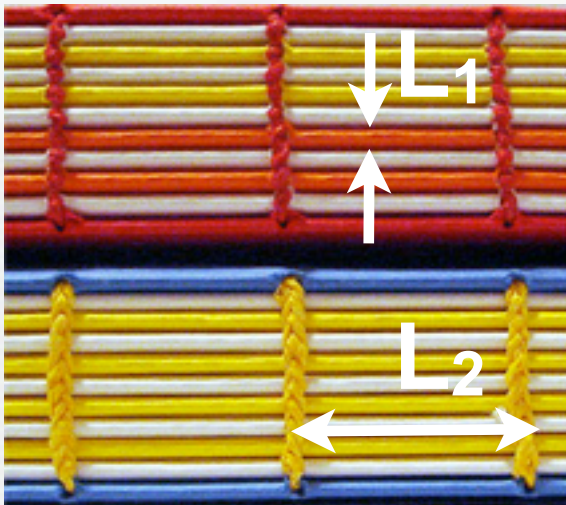
- A **simple realization** of the framework can be done **via an ordered string/brane lattice (no bulk!)**
- This is **inspired by layered metal/insulator structures** ubiquitous in **condensed-matter physics**
- L_1 - fundamental scale of space quantization
- $L_1 \ll x \ll L_2$ - locally (1+1)D
- $L_2 \ll x \ll L_3$ - locally (2+1)D
- $L_3 \ll x \ll L_4$ - locally (3+1)D





Folded Lattice: the Fabric of Space

- At very high energy **universe is a folded string**, with folding given by the fundamental quantization scale L_1
- It then **folds and interweaves forming a 2D structure** with the characteristic scale L_2 , which **in turn folds to make a 3D structure**, etc.
- Just **like a folded tapestry**, which is a 3D object woven with a very long thread, the **fabric of space** can literally be a fabric **made with a single string**



**“Offering of the Heart”
XVth century tapestry in Louvre**



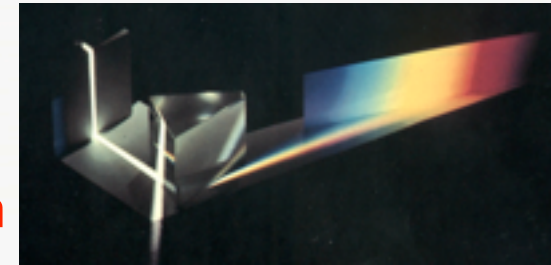
2D Universe

- At length scales $L_2 \ll x \ll L_3$, space is effectively 2D
- It's natural to pick $\Lambda_3 = 1/L_3 \sim 1 \text{ TeV}$, which allows to solve the hierarchy problem
 - $\Lambda_2 = 1/L_2$ can be several orders of magnitude higher, but not too high, as at some point linear ultraviolet divergences in (2+1)D will have to be cured with the logarithmic ones in (1+1)D
- Gravity, as any other force, propagates in 2D:
 - Atomic and nuclear physics constraints on other forces propagating in 2D do not apply given the TeV scale of the dimensional crossover
- The world is truly 2D, in a sense that the only third dimension is the thickness of the brane:
 - Minkowski metric is (1,-1,-1)
 - Gravitational potential $V(r) = 2G_2M \ln(r)$
 - Gravity is still attractive, but stronger, as the force falls off as $1/r$
 - Universe is very hot ($T > 1 \text{ TeV}$), so BBN and CMB not affected



High-Energy Particle Propagation

- Generally, one expect the lattice to fold randomly, thus avoiding creation of a preferred direction in space
- Yet, there is a preferred “cosmic” reference frame in which the lattice is at rest
- When the de Broglie wavelength of a particle in the cosmic frame is less than L_3 , the particle would propagate locally in 2D, and not in 3D!
 - This does not affect the straightness of particle propagation from the source to the observer, as the overall momentum is preserved
 - If the lattice tension is high enough (greater than the particle energy), the particle scatters coherently at the lattice junctions
 - This is similar to the light going straight through a crystal despite scattering off individual atoms via phonon exchange
 - Group velocity is preserved, but the propagation is via a jagged line, which creates an effective refraction index of the media of $1 + \Delta n$, where $\Delta n \sim L_2/L_3$
 - Non-linear dispersion relationship as a Fermi function with the threshold at $1/L_3$ allows to elude astrophysical constraints from γ -rays





Lorentz-Invariance Violation

- At low energies (below Λ_3) local Lorentz-invariance (LI) is nearly preserved, as particle propagates in 3D
- However, local random orientation of the 2D substructure generates *non-systematic* LI violation in the low-energy effective theory, i.e. even for particles with energy below Λ_3
- Light from distance sources is continually subjected to fluctuations of the layered structure, which introduces uncertainties in determination of its wavelength $\delta\lambda \sim L_3 \left(\frac{\lambda}{L_3}\right)^{1-\alpha}$, where α is a model-dependent parameter
 - See, e.g. [D. Mattingly, Living Rev. Rel. **8**, 5 (2005), gr-qc/0502097]
 - One could think about this phenomena as quantum interference of various paths a particle could take from point A to point B on the lattice, which generates interference fringes similar to a two-slit experiment or Bragg scattering



Constraints from LI Violation

- For a source at a distance L away, an accumulated phase shift between originally coherent photons is going to be: $\Delta\phi = 2\pi a L_3^\alpha L^{1-\alpha} / \lambda$, where a is a parameter of order 1
- Constrained from non-observation of interference fringes (Airy rings) from distance bright objects:
 - e.g., PKS1413+135 ($L = 1.2$ Gpc, $\lambda = 1.6 \mu\text{m}$)
- For $L_3 \sim 1 \text{ TeV}^{-1}$, $\alpha \gtrsim 0.7$ is allowed, leaving reasonable range of model parameter not excluded (including particularly interesting for us case of $\Delta\lambda \approx L_3$, i.e. $\alpha \approx 1$)
- Short pulses of radiation will be **temporally distorted** by $\Delta t \sim L_3^\alpha L^{\alpha-1} / v_\phi$, where v_ϕ is the phase velocity
 - See [Y.J. Ng, H. van Dam, and W.A. Christiansen, *Astrophys. J.* **591**, L87 (2003)]
- Even weaker limits from pulsar signal time-spread:
 - $\alpha \gtrsim 0.49$ from the B1937+21 pulsar timing measurement with $<0.2 \mu\text{s}$ precision



Cosmological Constant Problem

- If L_4 , the distance at which space time becomes $(4+1)D$, is of the order of the present cosmological horizon ($\sim 10^{26}$ m), a very small cosmological constant can be attributed to the lattice size
- Indeed, the Einstein's equations in $(4+1)D$ have the following metric as a vacuum solution:

$$ds^2 = dt^2 - e^{2\sqrt{\Lambda/3}t}(dr^2 + r^2 d\Omega^2) - d\psi^2$$

(here $\Lambda = 3/\psi^2$) [J. Ponce De Leon, Gen. Rel. Grav. **20**, 539 (1988)]

- For a $\psi = \text{const}$ hypersurface, the metric is 3D de Sitter with $\Lambda = \text{const}$
- An observer living on the 3D lattice (i.e., a fixed fold with $\psi = \text{const}$) will measure an effective stress-energy tensor with the equation of state $p = -\rho = -\Lambda \bar{M}_{\text{Pl}}^2$ (\bar{M}_{Pl} - reduced Planck mass)
- The observed vacuum energy density $\rho \approx (2.4 \text{ meV})^4$, corresponds to $\psi \sim 10^{61}/M_{\text{Pl}} \sim 10^{26}$ m (minimum value of ψ , corresponding to maximum value of the cosmological constant)
- One could think about this spatially generated cosmological constant as tiny Casimir force due to the distant parallel folds



Connection with Early Universe

- In terms of the very early Universe, the model seems naively to make the horizon problem more difficult to solve since high-energy particles are restricted to 1D or 2D-surfaces
- However, note that the set up will change the rate at which stress energy is diluted as the Universe expands
- This will change the dynamics of the early Universe and the time it takes for the plasma to cool down, having interesting implications for early cosmological dynamics
- We plan to study this in more details in the upcoming long paper



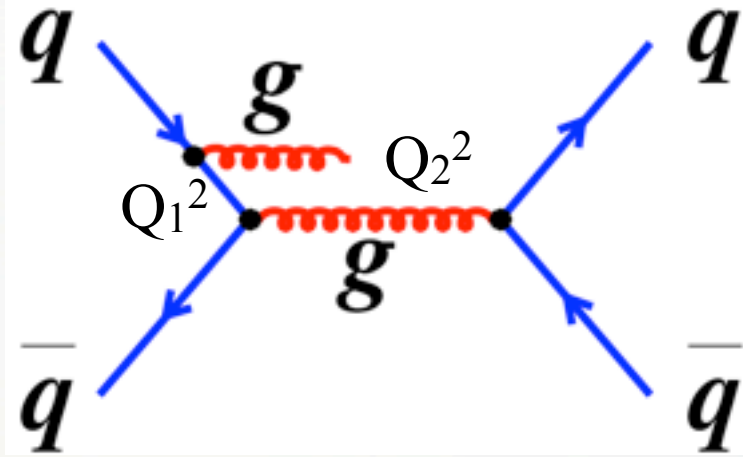
Collider Signatures

- Once the collision energy in the c.o.m. frame exceeds $\Lambda_3 \sim 1/L_3$, the properties of the collision change:
 - Propagator is confined to $(2+1)D$
 - Phase space is modified
 - The parton-level amplitudes change their form (drop faster)
 - Longitudinal polarization is suppressed (and so is $V_L V_L$ scattering - could help unitarization without a light Higgs)
 - N.B.: spin-statistics theorem does not hold in $(2+1)D$ or $(1+1)D$
- Example: consider 4-point effective operator, e.g. high-mass Drell-Yan process
 - Coulomb potential:
 - $(3+1)D$: $V(r) \sim \alpha_3/r$; $[V] = E = 1/[L] \Rightarrow \alpha_3$ is dimensionless
 - $(2+1)D$: $V(r) \sim \alpha_2 \log(r/r_0) \Rightarrow [\alpha_2] = [E] = 1/[L]$
 - Cross-section:
 - $(3+1)D$: $[\sigma_3] = [L^2] \Rightarrow \sigma_3 \sim \alpha_3^2/E^2$
 - $(2+1)D$: $[\sigma_2] = [L] \Rightarrow \sigma_2 \sim \alpha_2^2/E^3$



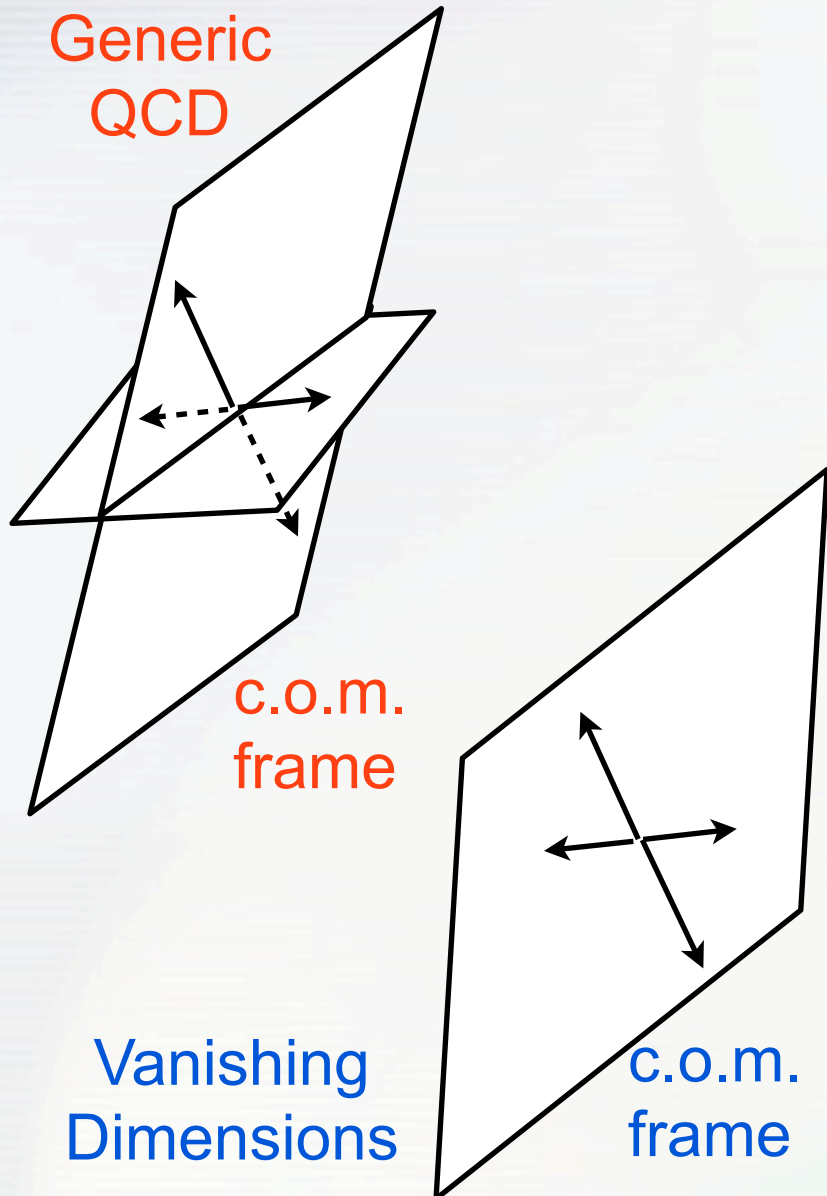
Planar Events at the LHC

- Important parameter is the **momentum transfer**, which **determines the distance scale**, which the interaction probes
- Consider **$2 \rightarrow 3$ scattering**: if **both Q_1^2 and Q_2^2 are $> \Lambda_3^2$** , the **three jets become planar in the c.o.m. frame** with the orientation given by the local fold
- Unfortunately, since the overall 3-momentum of the system is preserved due to elastic interactions with the lattice, **the events are not planar in the lab frame** (otherwise the model would violate T & CPT)
- Since **every $2 \rightarrow 3$ process is planar in the c.o.m. frame** (due to momentum conservation), **three-jet events can not be used** to easily probe for vanishing dimensions





Planar Four-Jet Events at the LHC

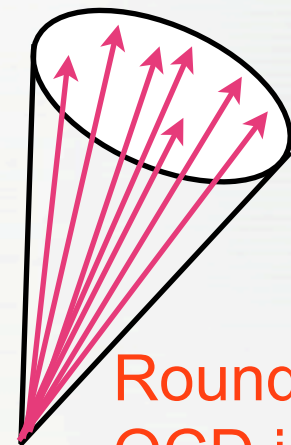


- But **four-jet events could be used** to distinguish the two cases!
- If all three propagators have $Q^2 \gtrsim \Lambda_3^2$, **four-jet events become planar**, unlike the QCD ones (bi-planar)
- Preliminary estimates (work in progress in collaboration with Gabe Shaughnessy) show that it **may be observable at 14 TeV LHC with $O(10 \text{ fb}^{-1})$** if $\Lambda_3 \sim 1 \text{ TeV}$

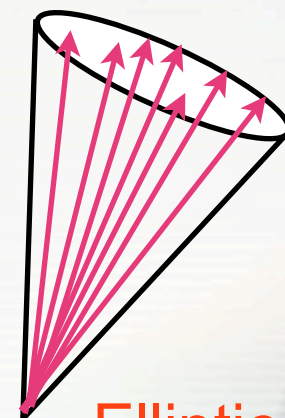


Elliptic Jets

- By the same token, if several consequent splittings of a parton shower creating a jet are above the dimensional crossover scale, a **high-energy jet will have an elongated shape** after full development of the parton shower and hadronization
- **Ellipticity of the energy flow** in a jet therefore may be another **important variable**
- **Detailed studies are needed** to see how often regular QCD jets fluctuate to have an elliptic shape, but in general one should look for an increase of ellipticity with jet energy as a signature for vanishing dimensions
- Possible connection with **elongated UHECR showers** observed by Pamir collaboration in the 80-ies [see, e.g. A. De Roeck et al, in Proc. 13th Blois Workshop, 2009, arXiv:1002.3527]



Round
QCD jet



Elliptic
jet



Conclusions



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- This paradigm may allow to address fundamental problems of astro-particle physics and cosmology - the hierarchy problem, unitarization of scattering cross sections, cosmological constant problem, arrow of time, etc. from a completely new angle
- The existing constraints on this framework appear to be weak from the current set of experimental data
- Detectable consequences for astrophysical observations and the LHC are discussed

Thank you!



Technical Details...





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- Whether it holds up or not will be clear after detailed studies
- Yet, we think it's worth bringing up to the attention of the astro-particle community with the goal to spawn these studies