#### **ISMD 2010**

XL International Symposium on Multiparticle Dynamics 21-25 September 2010, Antwerp, Belgium

# Forward energy and particle flow with the LHCf experiment

Lorenzo Bonechi

University and INFN – Florence (Italy) on behalf of the LHCf collaboration

# The LHCf international collaboration

O.Adriani<sup>1,2</sup>, L.Bonechi<sup>1,2</sup>, M.Bongi<sup>2</sup>, G.Castellini<sup>2,3</sup>, A.Faus<sup>4</sup>, K.Fukatsu<sup>5</sup>, M.Grandi<sup>2</sup>, M.Haguenauer<sup>6</sup>, Y.Itow<sup>5</sup>, K. Kawade<sup>5</sup>, K.Kasahara<sup>7</sup>, D.Macina<sup>8</sup>, T.Mase<sup>5</sup>, K.Masuda<sup>5</sup>, Y. Matsubara<sup>5</sup>, H.Menjo<sup>2</sup>, G.Mitsuka<sup>5</sup>, Y.Muraki<sup>9</sup>, M.Nakai<sup>7</sup>, K.Noda<sup>5</sup>, P.Papini<sup>2</sup>, A.-L.Perrot<sup>8</sup>, S.Ricciarini<sup>2</sup>, T.Sako<sup>6</sup>, Y.Shimitsu<sup>7</sup>, K.Suzuki<sup>5</sup>, Y.Suzuki<sup>7</sup>, K.Taki<sup>5</sup>, T.Tamura<sup>10</sup>, S.Torii<sup>7</sup>, A.Tricomi<sup>11</sup>, W.C.Turner<sup>12</sup>, K.Yoshida<sup>13</sup>, J.Velasco<sup>4</sup>

- 1) University of Florence, Italy
- 2) INFN Section of Florence, Italy
- 3) IFAC CNR, Florence, Italy
- 4) IFIC, Centro Mixto CSIC-UVEG, Spain
- 5) Solar-Terrestrial Environment Laboratory, Nagoya University, Japan
- 6) Ecole Polytechnique, France
- 7) Waseda University, Japan
- 8) CERN, Switzerland
- 9) Konan University, Japan
- 10) Kanagawa University, Japan
- 11) INFN and University of Catania, Italy
- 12) LBNL, Berkeley, USA
- 13) Shibaura Institute of Technology, Japan

# Outline of this presentation

#### • The Physics of LHCf

- Cosmic rays in the atmosphere and hadronic interaction models
- Open issues on the HE Cosmic Ray spectrum

#### Overview of the experiment

- Detection of neutral particles at low angle at LHC
- Description of detectors

#### • Preliminary results (2009/2010 runs)

- 900 GeV
- 7 TeV

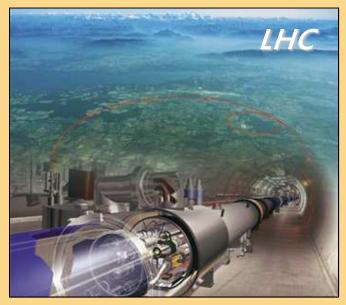
#### Future plan

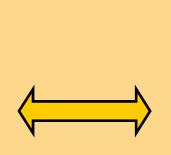
### PART 1

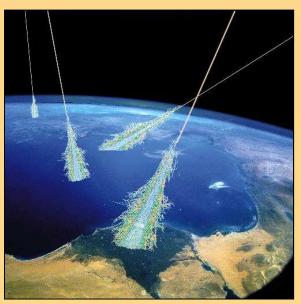
#### **The Physics of LHCf**

# Main topics

- Experimental measurement:
  - Precise measurement of  $\gamma$ ,  $\pi^0$  and n spectra in the very forward region at LHC
- 7 TeV + 7 TeV in the c.m. frame  $\rightarrow 10^{17}$  eV in the laboratory frame:
  - We can better simulate in the biggest's world laboratory what happens in nature when a Very High Energy Cosmic Ray interacts in the atmosphere
- Why in the very forward region?
  - Because the dominant contribution to the energy flux in the atmospheric shower development is carried on by the very forward produced particles

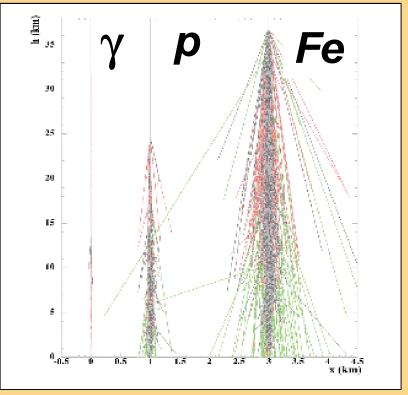






# **Cosmic ray showers**

#### Air shower developments



The hadronic interaction models used in air shower simulations have a big uncertainty due to the lack of experimental data in the energy range over 10<sup>15</sup>eV

#### **Extensive air shower observation**

- longitudinal distribution
- lateral distribution
- Timing

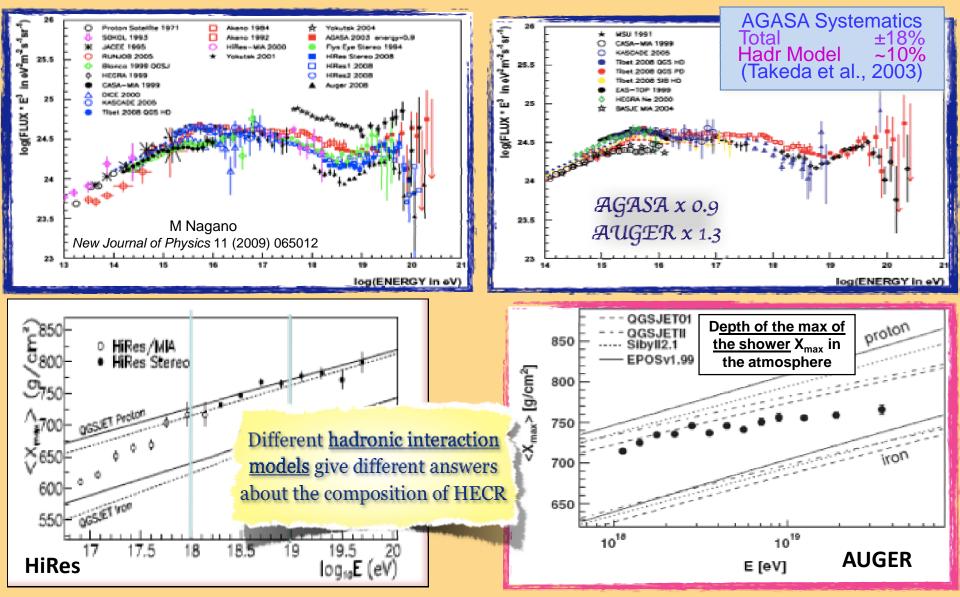
Air shower development

- Type of primary
- Energy
- Arrival direction

#### **Astrophysical parameters**

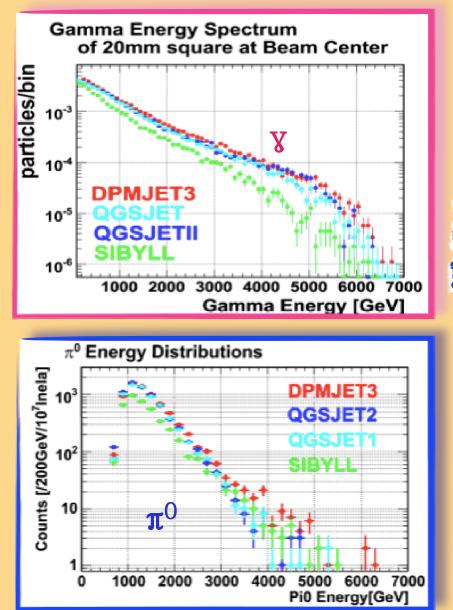
- Spectrum
- Composition
- Source distribution

### **Open Issues on HECR spectrum**



21-25 September 2010

#### What do we expect from LHCf?



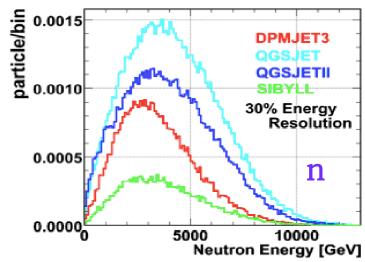
Energy spectra and transverse momentum distribution of

- γ (E>100GeV, ΔΕ/Ε<5%)
- Neutrons (E> few 100 GeV,  $\Delta E/E \sim 30\%$ )
- $\pi^{0}$  (E>500GeV,  $\Delta$ E/E<3%)

in the pseudo-rapidity range  $\eta > 8.4$ 

10<sup>6</sup> collisions ↔ 2min. exposure @ 10<sup>29</sup>cm<sup>-2</sup>s<sup>-1</sup>

> Neutron Energy Spectrum of 20mm Calorimeter at beam center

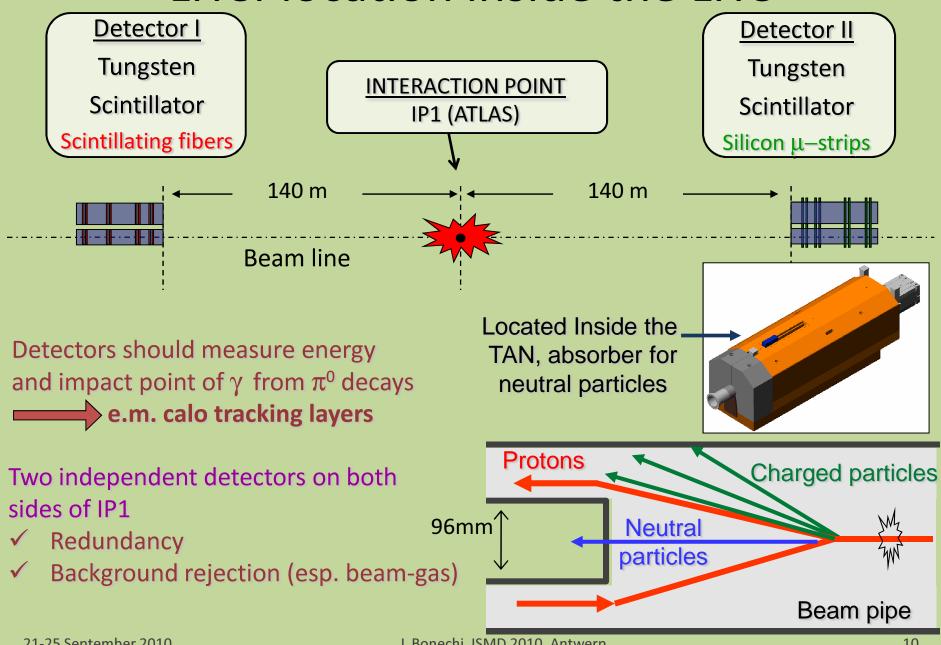


21-25 September 2010

# PART 2

#### **Overview of method and detectors**

# LHCf location inside the LHC



L.Bonechi, ISMD 2010, Antwerp

# The LHCf detectors

### **Sampling E.M. calorimeters**

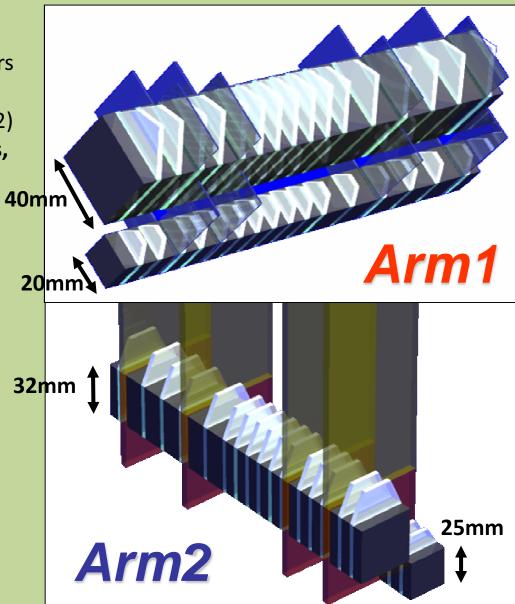
- $\bullet$  W (44 r.l  $\,$  , 1.7  $\!\lambda_{I}$  ) and Scintillator x 16 Layers
- 4 tracking layers: XY-SciFi(Arm#1) and XY-Silicon strip(Arm#2)
- Each detector has two calorimeter towers, which allow to reconstruct  $\pi^0$

#### Expected Performance

Energy resolution (> 100GeV) < 5% for photons 30% for neutrons Position resolution < 200µm (Arm#1) ~ 40µm (Arm#2)

#### **Front Counters**

- thin scintillators with 80x80mm<sup>2</sup>
- To monitor beam condition
- For background rejection of beam-residual gas collisions by coincidence analysis





### PART 3

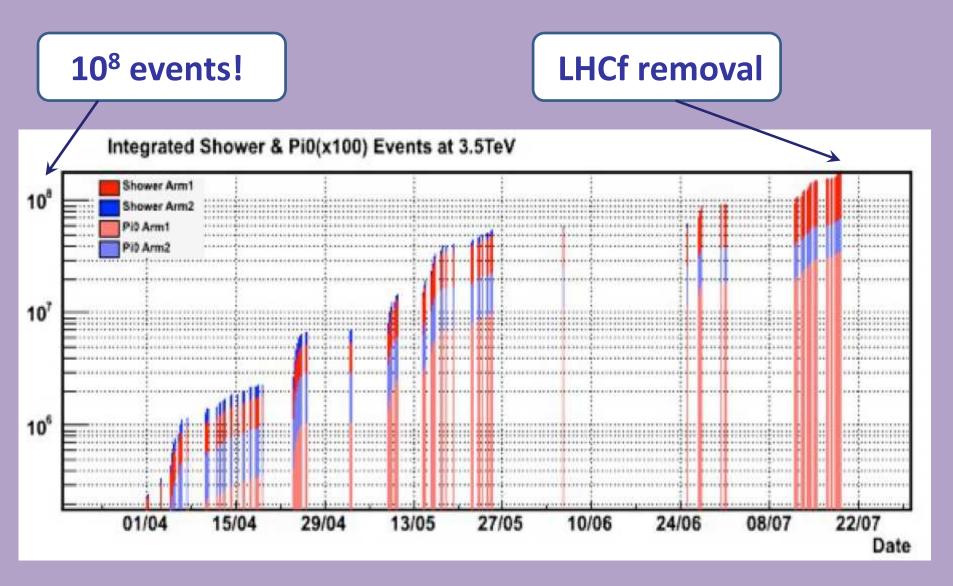
#### Preliminary results (2009 and 2010 runs)

21-25 September 2010

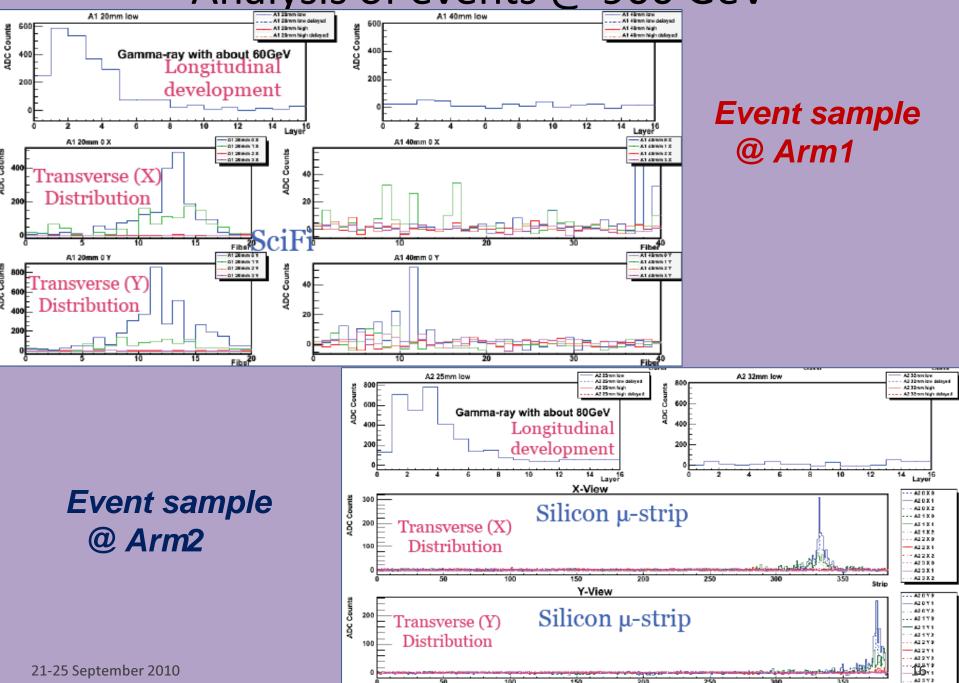
#### Summary of operations in 2009 and 2010 With Stable Beam at 450 + 450 GeV **Total of 42 hours for physics** About 10<sup>5</sup> showers events in Arm1+Arm2 With Stable Beam at 3.5 + 3.5 TeV **Total of 150 hours for physics with different setups** Different vertical position to increase the accessible kinematical range Runs with or without beam crossing angle ~ 4.10<sup>8</sup> shower events in Arm1+Arm2 $\sim 10^6 \pi^0$ events in Arm1+Arm2 Status Completed program for 900 GeV and 7 TeV **Removed detectors from tunnel in July 2010**

- Post-calibration beam test in October 2010
- Upgrade to more rad-hard detectors

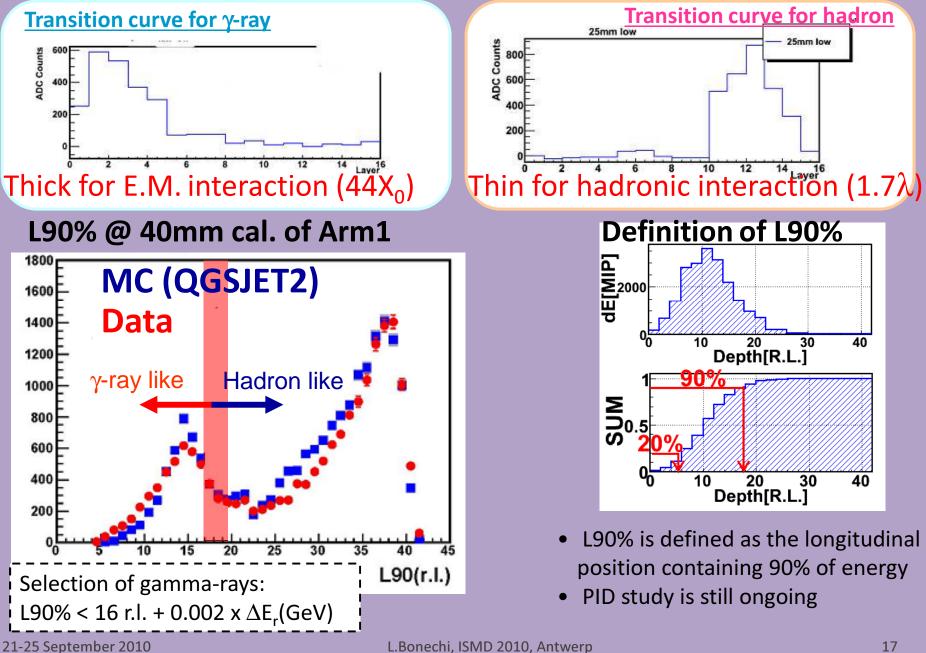
### Accumulated events in 2010



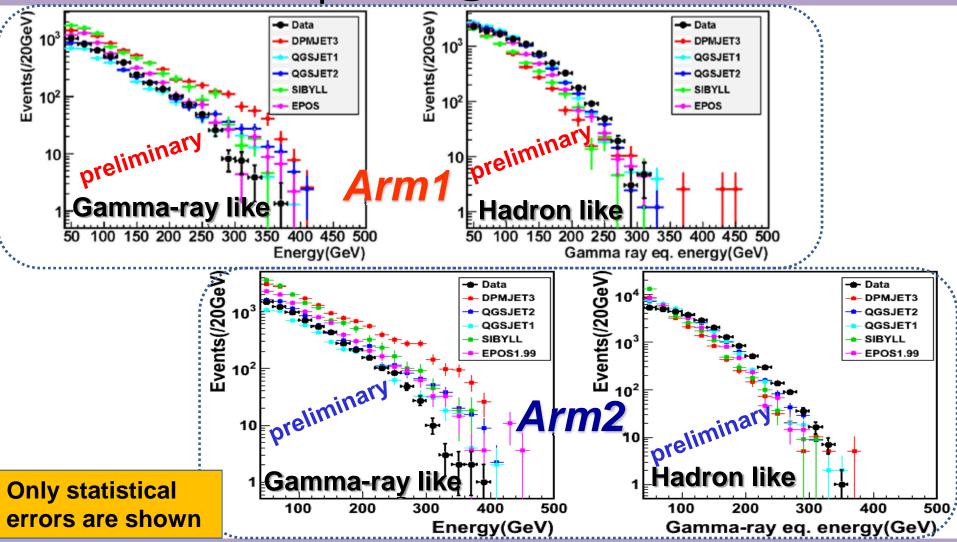
### Analysis of events @ 900 GeV



### Particle identification



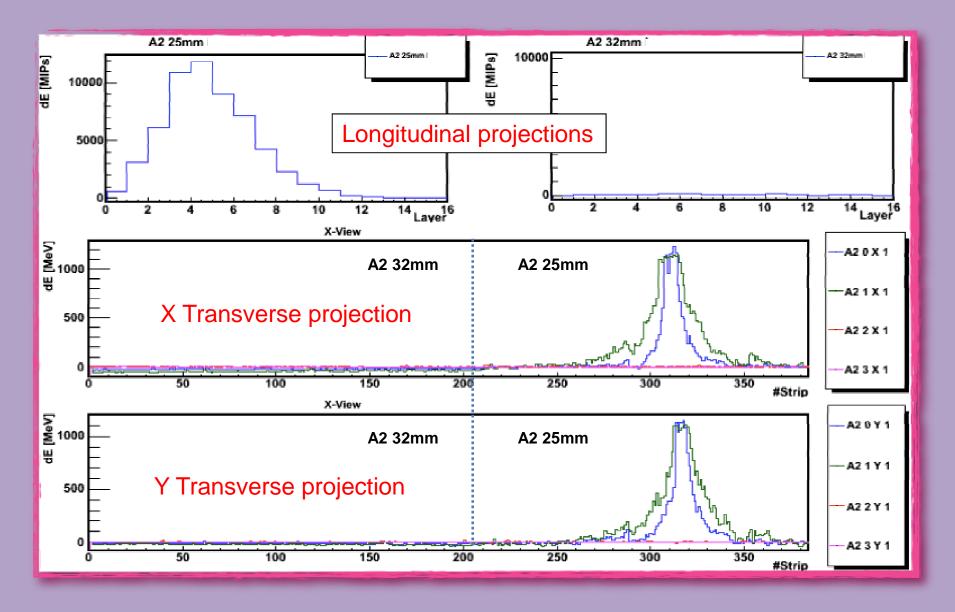




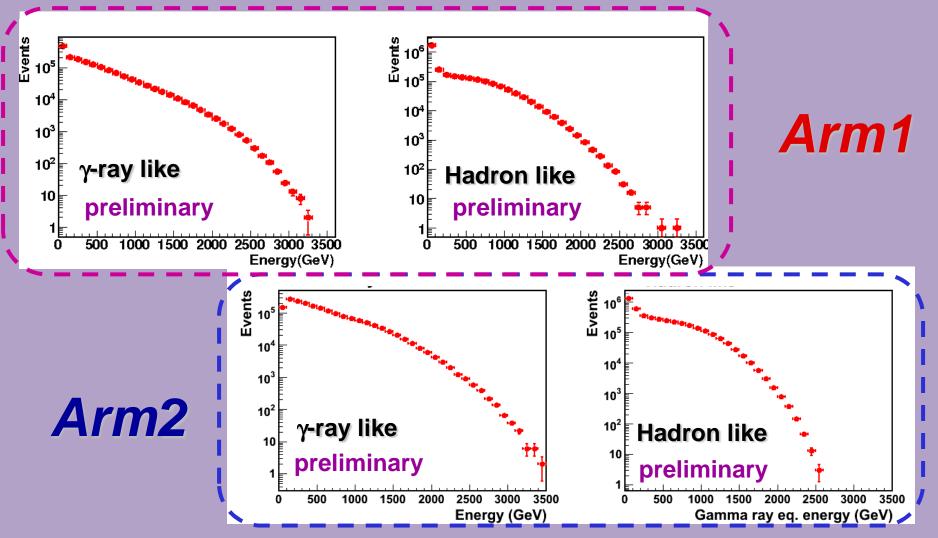
Spectra are normalized by # of gamma-ray and hadron like events. Response for hadrons and systematic errors (mainly absolute energy scale) are under study.

21-25 September 2010

### Beautiful e.m. events @ 7 TeV !!!



### Neutral particle spectra @ 7 TeV



• Very high statistics: only 2% of data is used

• Comparison with M.C. is under development

### Comparison of Arm1 and Arm2 @ 7 TeV

Events/(100GeV) OGeV Events/(10 10<sup>2</sup> 10 preliminary preliminary 10 10 3000 3500 500 2000 2500 2500 1000 1500 500 1500 2000 1000

Gamma-like, Large tower

Red : Arm1 Blue : Arm2 Same runs, same conditions, common rapidity region selected. Spectra corrected for the live time of detectors.

Energy(GeV)

21-25 September 2010

Gamma-like, Small tower

L.Bonechi, ISMD 2010, Antwerp

3000

Energy(GeV)

3500

### Neutral pions

An example of event (Arm2) Energy spectrum (Arm2) MIPS 25mm 32mm 빀 4000 붱 4000 preliminary 10<sup>4</sup> 2000 Events [/50GeV] 10<sup>3</sup> Silicon strip-X view Lvavd 10<sup>2</sup> 10 500 1000 1500 2000 2500 3000 n R γ<sub>1</sub>(E<sub>1</sub>) E<sub>n°</sub> [GeV] A = 40 m **Reconstructed mass (Arm2)** Events (/1MeV) 600 **∆M/M=2.3%** 500  $\gamma_2(E_2)$ 400 300 preliminary • Pi0's are a main source of electromagnetic 200

- secondaries in high energy collisions.
- The mass peak is very useful to confirm the detector performances and to estimate the systematic error of energy scale.
  21-25 September 2010

100

80

100

120

140

160

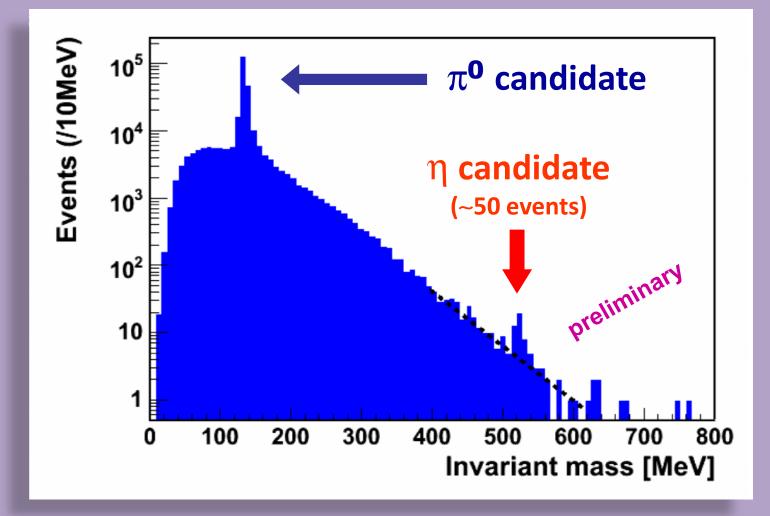
180

Invariant mass (MeV)

200

3500

### $2\gamma$ invariant mass spectrum @ 7 TeV



- The search for  $\eta$  particles is an important tool for discriminating hadronic interaction models, because their spectra differ a lot from model to model
- Important tool also for energy scale calibration

### Schedule and future plan

2010, Oct	Beam test at SPS to confirm the radiation damage and the performance
END 2010	Finalize analysis at 900 GeV (almost completed) and at 7 TeV
2011/2012	Upgrade the detector for radiation hardness: replacement of scintillators and SciFi with GSO
2013	Re-installation of detectors in the tunnel again for the operation at 14TeV

#### Then we are thinking about - Operation at LHC <u>light</u> ion collisions (not Pb-Pb).

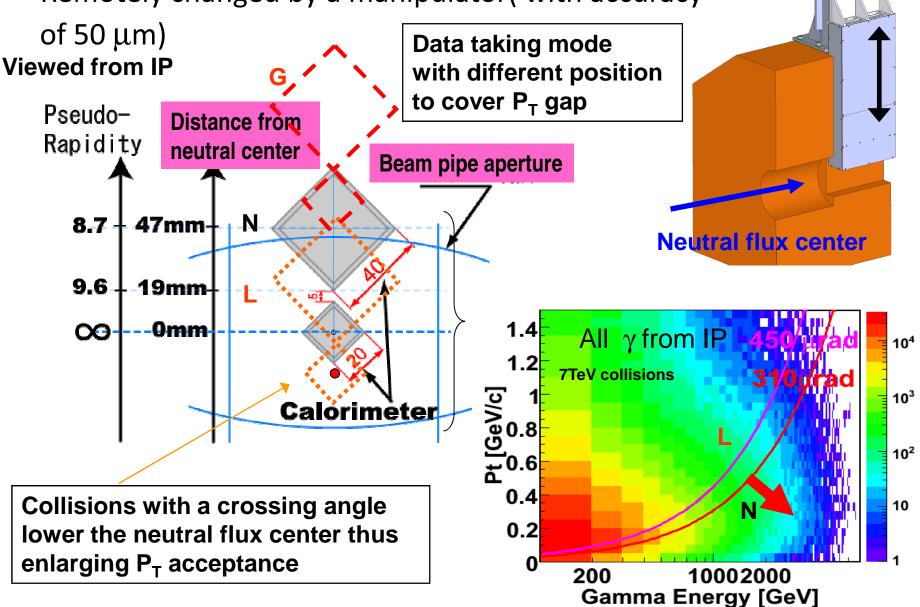
# Conclusions

- LHCf is a forward experiment at LHC; its aim is to measure energy spectra and transverse momentum distributions of very energetic neutral secondaries from pp interactions in the very forward region of IP1 (at pseudo-rapidity greater than 8.4)
- Results will help calibrating the hadronic interaction models; one important field where this measurements are mostly important is the study of atmospheric showers induced by HECR particles
- LHCf successfully completed operations at 900 GeV and 7 TeV. The LHCf detectors has been removed from the LHC tunnel on July 21<sup>st</sup> 2010.
- Analysis of data at 900 GeV is almost completed; we will finalize analysis at 7 TeV before the end of the year.
- Detectors will be **upgrade**d in 2011/2012 for radiation hardness and will be re-installed for data taking at 7 TeV+7 TeV in 2013.

# Backup slides

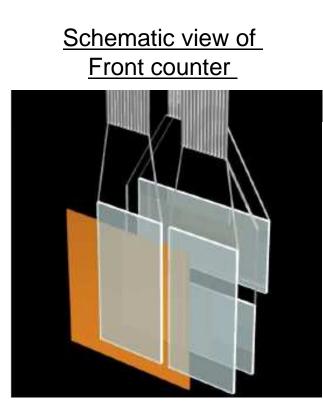
### Detector vertical position and acceptance

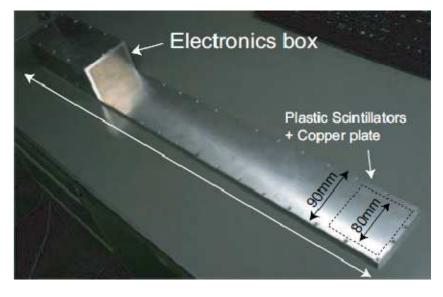
• Remotely changed by a manipulator( with accuracy



### Front counters

 Thin scintillators with 8x8cm<sup>2</sup> acceptance, which have been installed in front of each main detector.



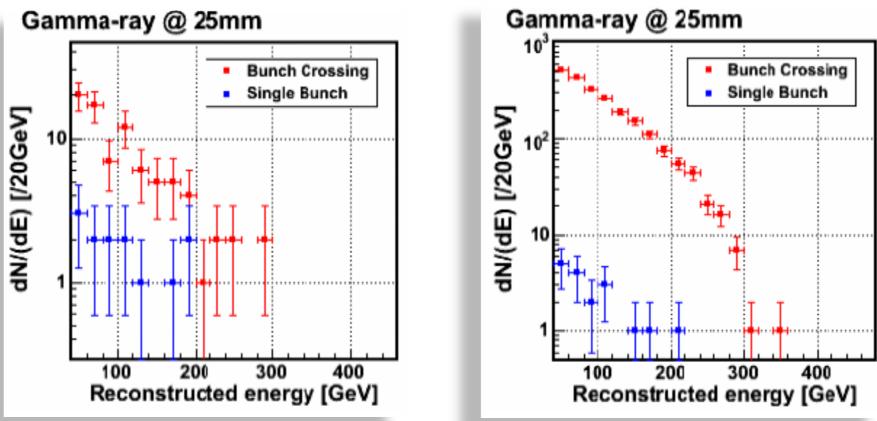


- To monitor beam condition.
- For background rejection of beam-residual gas collisions by coincidence analysis

### Beam-gas backgroud @ 900 GeV

2009

2010

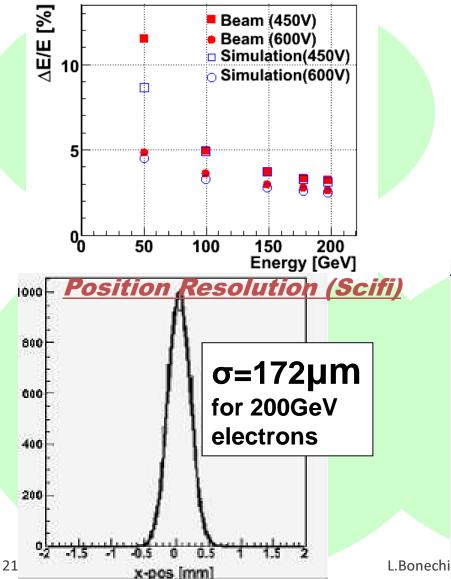


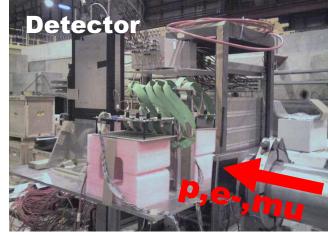
#### Very big reduction in the Beam Gas contribution!!!! Beam gas ~ I, while interactions ~ I<sup>2</sup>

21-25 September 2010

### Beam test @ SPS

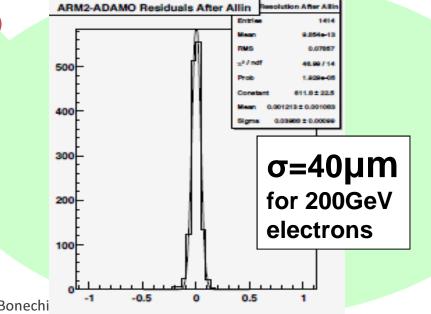
#### <u>Energy Resolution</u> for electrons with 20mm cal.



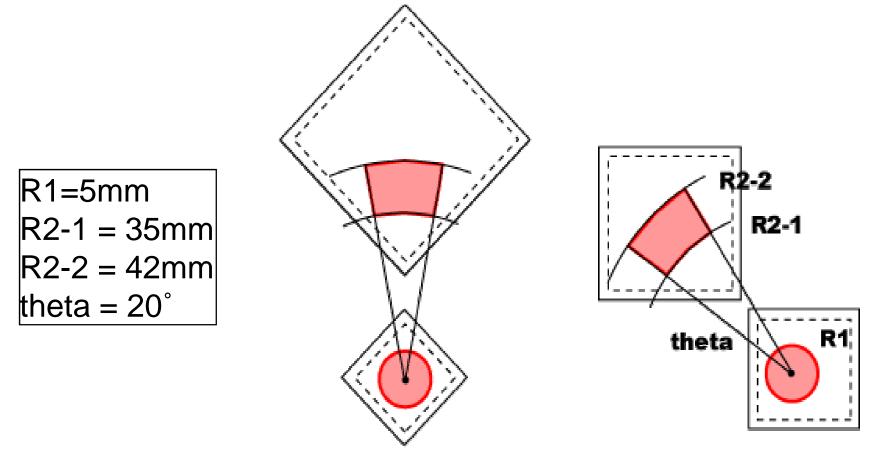


- Electrons 50GeV/c 200GeV/c
- Muons 150GeV/c
- Protons 150GeV/c, 350GeV/c

#### Position Resolution (Silicon)



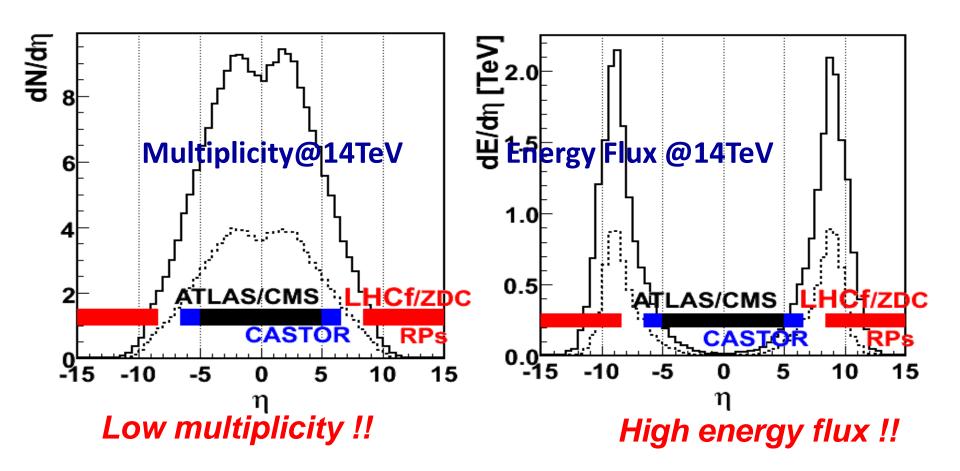
### Selection of rapidity region (comparison Arm1/2)



Both Arm1 and Arm2 cover the same rapidity area in small and large tower. Here the beam center is determined by our measurements.

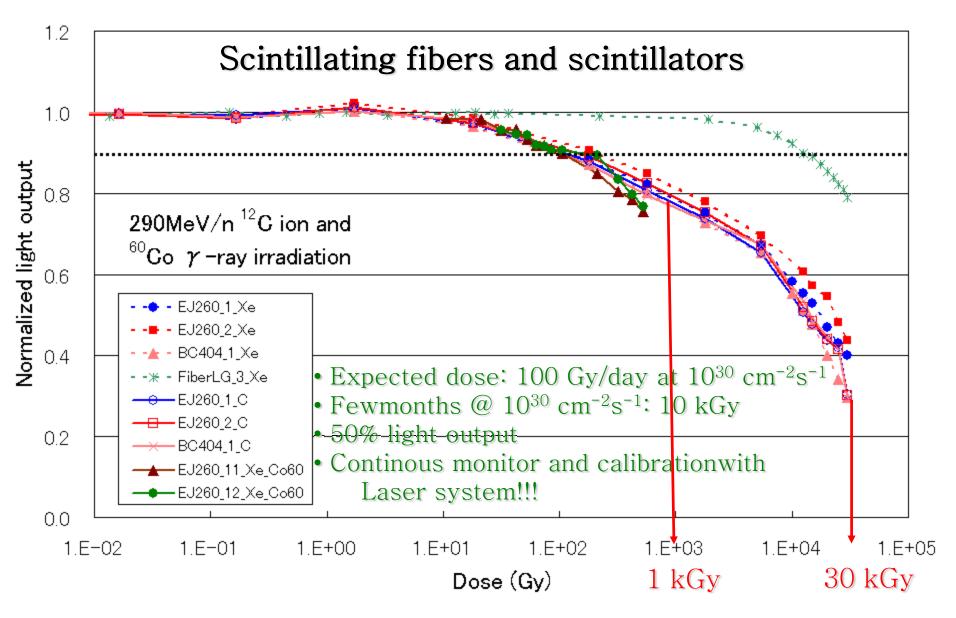
21-25 September 2010

### Particle and energy flow vs rapidity



simulated by DPMJET3

### **Radiation damage**



### Status of detectors and upgrade

LHCf detectors removed from the TAN on 20<sup>th</sup> July 2010

-Plastic scinitillator degradation of a few % due to > 5 Gy dose

- "Post"-calibration by a SPS test beam are planned on 15<sup>th</sup> to 28<sup>th</sup> October
- Re-installation at LHC at the next energy upgrade. R&D and fabrication of radiation-hard GSO scinitillators and fibres are on-going for the "phase-2" of the LHCf detector.



