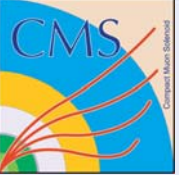


CMS Luminosity

JTREM IV

Valerie Halyo for Lumi Team

Aug 3, 2009



Lumi team Effort for HF based Lumi



**N. Adam, V. Halyo, A. Heister, A. Hunt, Y. Guo,
J. Jones, D. Marlow, J. Werner, D. Stickland, Z. Xie**

- Overview on Luminosity
 - Goals
 - Design strategy
 - Method
- Absolute Lumi normalization
- Lumi dataflow
- User Info
- Conclusions

Design Goals: General Desirables

- Absolute calibration, based on a known cross section with a reliably calculated acceptance.
- Temporal stability against gain changes and other drifts: “countable objects” or self calibrating signals (e.g., MIP peak).
- Linearity over a large range of luminosities.
- Real time operation independent of full DAQ.
- Redundancy
 - There is no perfect method
 - Applies to both real time monitoring and to offline absolute normalization

- Use absolute calibration of machine luminosity or TOTEM measurement as a reference point.
- Use real time techniques (HF, Pixel Telescopes, BRAN) to extrapolate/interpolate to design luminosity
- Normalize the luminosity using processes of \sim known cross section (e.g., W's and Z's)

Signals From HF

Iron fiber calorimeter.

$$3 < \eta < 5$$

HF

T2 CASTOR

T1 & T2 are elements of TOTEM

Minimal additional hardware requirements:

- Mezzanine board to tap into HF data stream and forward bits to a server via Ethernet
- Autonomous DAQ system to provide “always on” operation

Methods:

- Count “zeroes”
- Use also linear E_T sum, which scales directly with luminosity.

Simulations:

Full GEANT4 with realistic representation of photo statistics, electronic noise and quantization, etc. within the framework of CMSSW

The average fraction f , of empty towers per bunch crossing is given by:

$$\langle f \rangle = e^{\mu(p-1)} \Rightarrow -\ln \langle f \rangle = (1-p)\mu$$

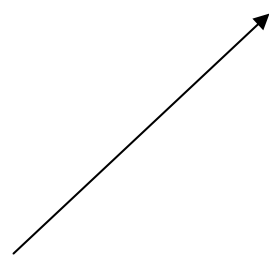
Where:

p = probability that a given tower is empty after *single interaction*,

μ = mean number of interactions per bunch crossing.

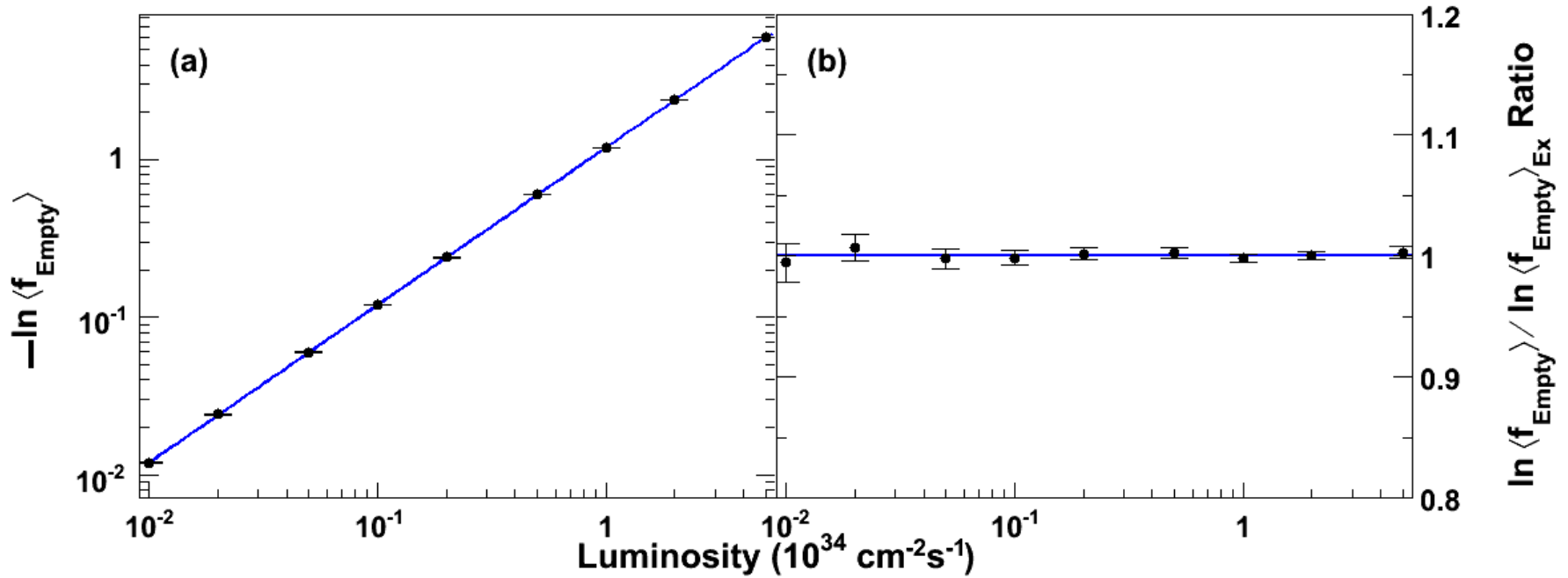
In real life in order to decide whether a tower is empty we have to introduce a threshold cut which would cut somewhat into our signal and therefore introduce a correction to our previous result

$$-\ln\langle f \rangle = (1 - p)(1 - \varepsilon)\mu + N$$



This term is a measure of the overlap between the signal and noise distribution below the threshold

HF Tower Occupancy: ADC > 7 - η Rings 6 - 7



Nadia Adam

we plan to use two sets of two rings.

Average transverse energy per tower per BX

$$\langle E_T \rangle = \nu(1 - p)\mu + N$$

Where:

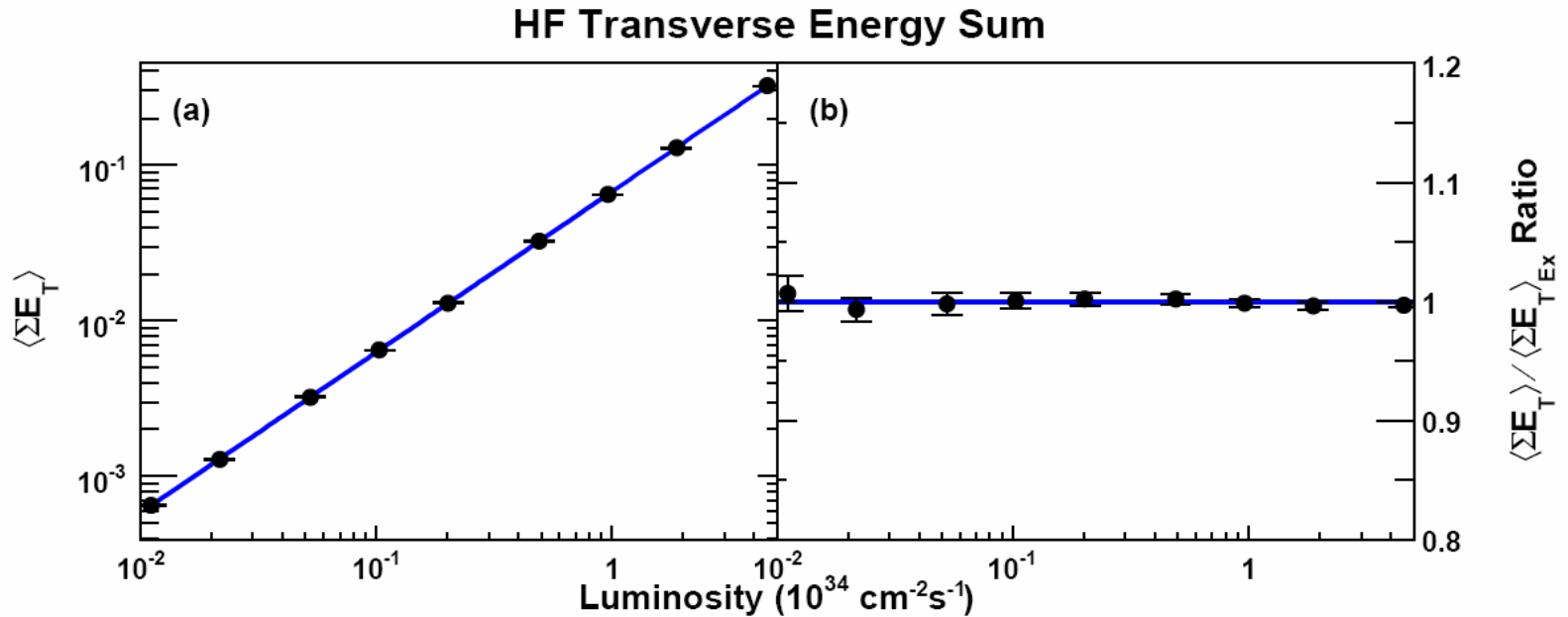
p = probability that a given tower
is empty after *single*
interaction,

μ = mean number of
interactions per bunch crossing

N = Noise contribution.

ν = $\langle E_T \rangle$ for a single occupied
tower in a single interaction

The average EtSum is linear over all the expected luminosity dynamic range

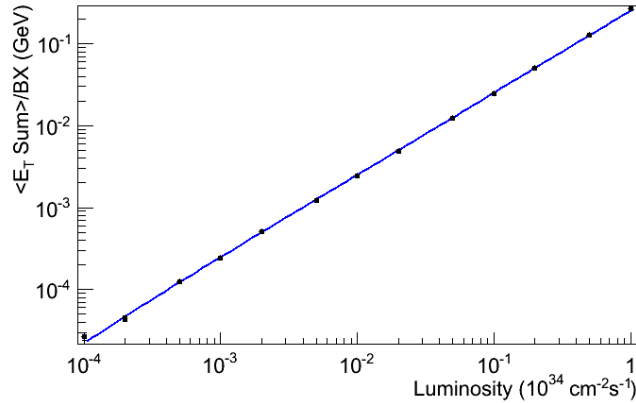


Nadia Adam

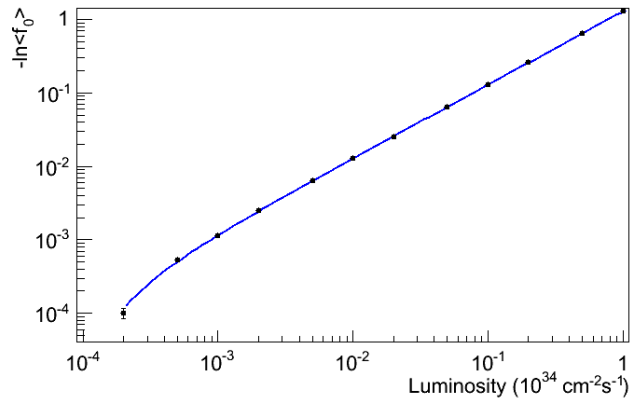
Any noise offset would be calibrated out by using the the Hlx data during the abort gap

Update for 10TeV (3_2_x)

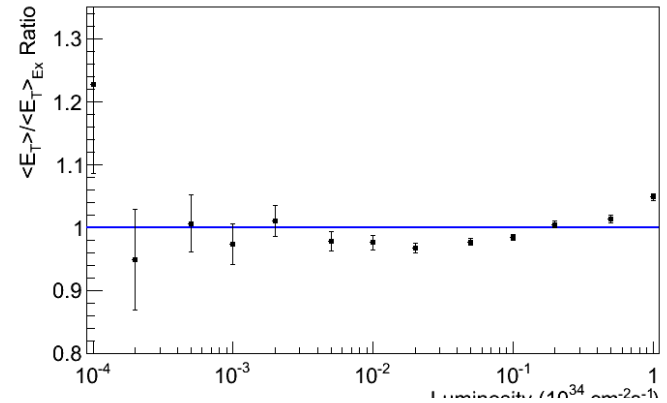
10TeV E_T Sum Linearity



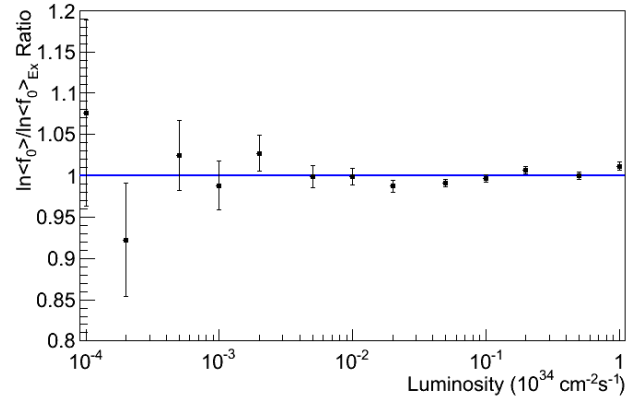
10TeV Tower Occupancy Linearity: Occupied \equiv LUT > 0



10TeV E_T Sum Relative Linearity

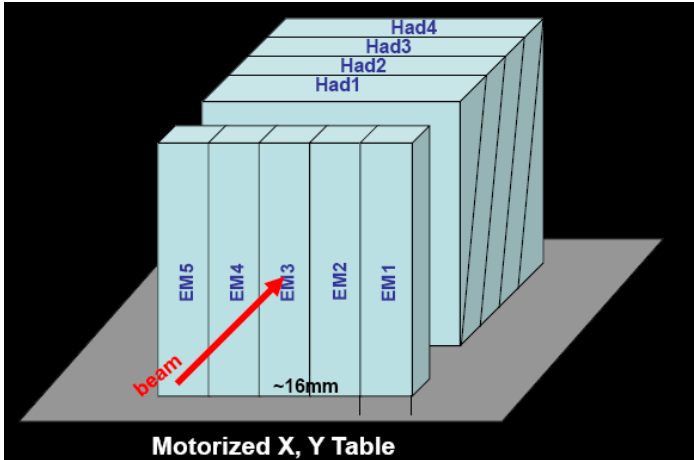


10TeV Tower Occupancy Relative Linearity: Occupied \equiv LUT > 1



Nadia Adam

Luminosity Monitors: ZDC, BRAN

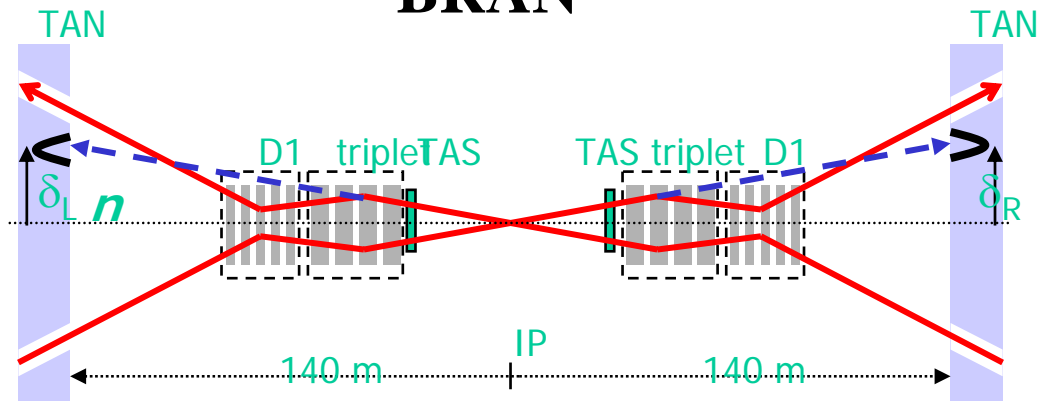


ZDC

The design of each ZDC includes (EM section) and (HAD section).

The LHC accelerator project incorporated fast ionization counters, in the TAN region, which is $\pm 140\text{m}$ from the IP

BRAN



- None of the methods discussed provides an absolute calibration for the luminosity
- Initially determine a luminosity calibration using the luminosity measurement from the LHC's measurement of beam parameters.
- Stick with that normalization until we have had a chance to study
 - CMS measurement of $\sigma_{W/Z}$. $pp \rightarrow pp\mu\mu$
 - Total cross section from TOTEM

LHC Beam Parameters

Stage	Energy[TeV]	Bunch Configuration	$\beta^*[m]$	Luminosity[$\text{cm}^{-2}\text{s}^{-1}$]	Event rate/crossing
A	0.9	2×2	11	2.6×10^{27}	0.006058
A'	10	2×2	11	2.9×10^{28}	0.0974
A'	10	2×2	2	1.6×10^{29}	0.5375
A'	10	43×43	2	6.9×10^{30}	1.0780
A'	10	156×156	2	2.49×10^{31}	1.0723
A'	10	156×156	1	1.615×10^{32}	6.9550

Integration Times based on Statistical Errors using EtSum Method

Stage	Luminosity[$\text{cm}^{-2}\text{s}^{-1}$]	Ave Time [s] 1%	Time [s] 5%	Ave Time [s] 5%	Time [s] 10%	Ave Time [s] 10%
A	2.6×10^{27}	-	-	-	425.6042	227.7675
A'	2.9×10^{28}	20.6513	1.5684	0.8261	0.3921	0.2065
A'	1.6×10^{29}	3.9007	0.3093	0.1560	0.0773	0.0390
A'	6.9×10^{30}	0.1019	0.1466	0.0041	0.0366	0.0010
A'	2.49×10^{31}	0.0407	0.1473	0.0016	0.0368	0.0004
A'	1.615×10^{32}	0.0034	0.0190	0.0001	0.0047	0.00003

Update on Lumi from Z xsec

Our recent study shows the systematic error on the Z cross section to be the following

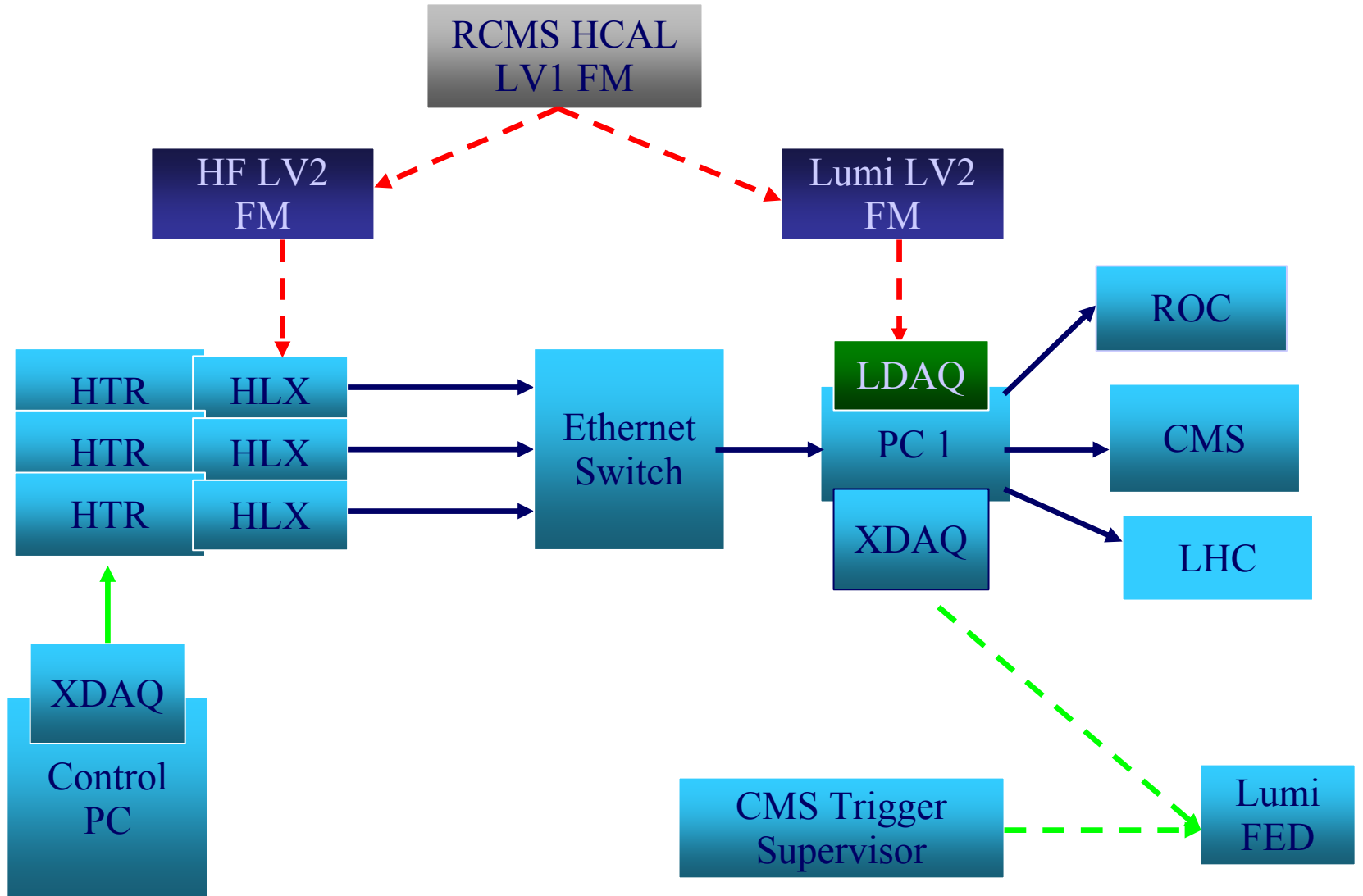
Error	Z	
	$\Delta\sigma$ (%)	ΔA (%)
Higher Order	0.23 ± 1.25	-1.97 ± 1.51
QCD Scale	0.92 ± 0.61	1.58 ± 1.35
PDF	2.75 ± 0.00	1.03 ± 0.00
Total	2.91 ± 0.22	2.73 ± 1.35

Hence the absolute luminosity measurement will be dominated by statistic at startup

CMS AN-2009/088

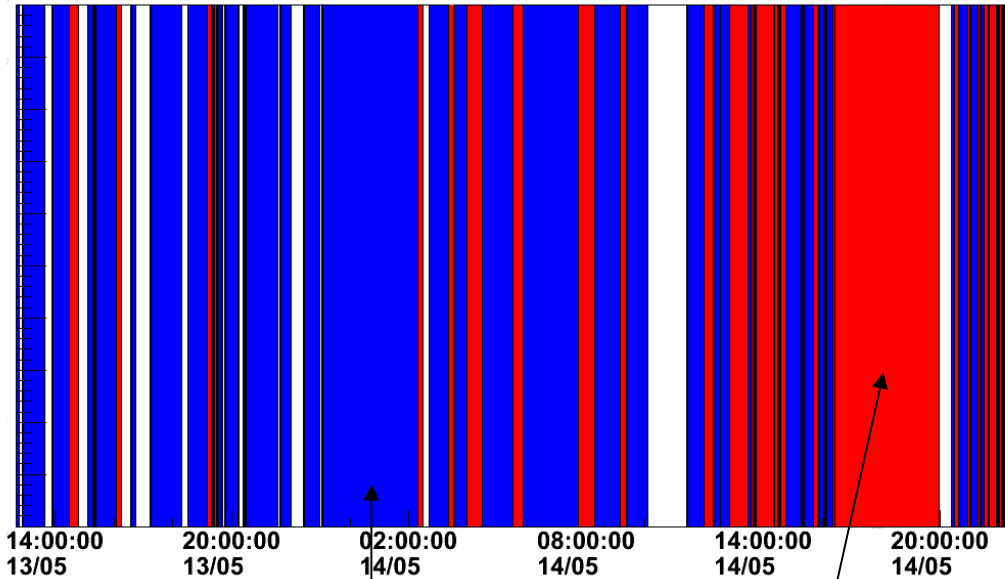
JHEP 09 2008 133 , JHEP 05 (2008) 062

HF Luminosity Readout Path



Lumi Continuous Running Mode

Adam Hunt



Lumi as LV2 FM
Communicates to
HCAL L1FM

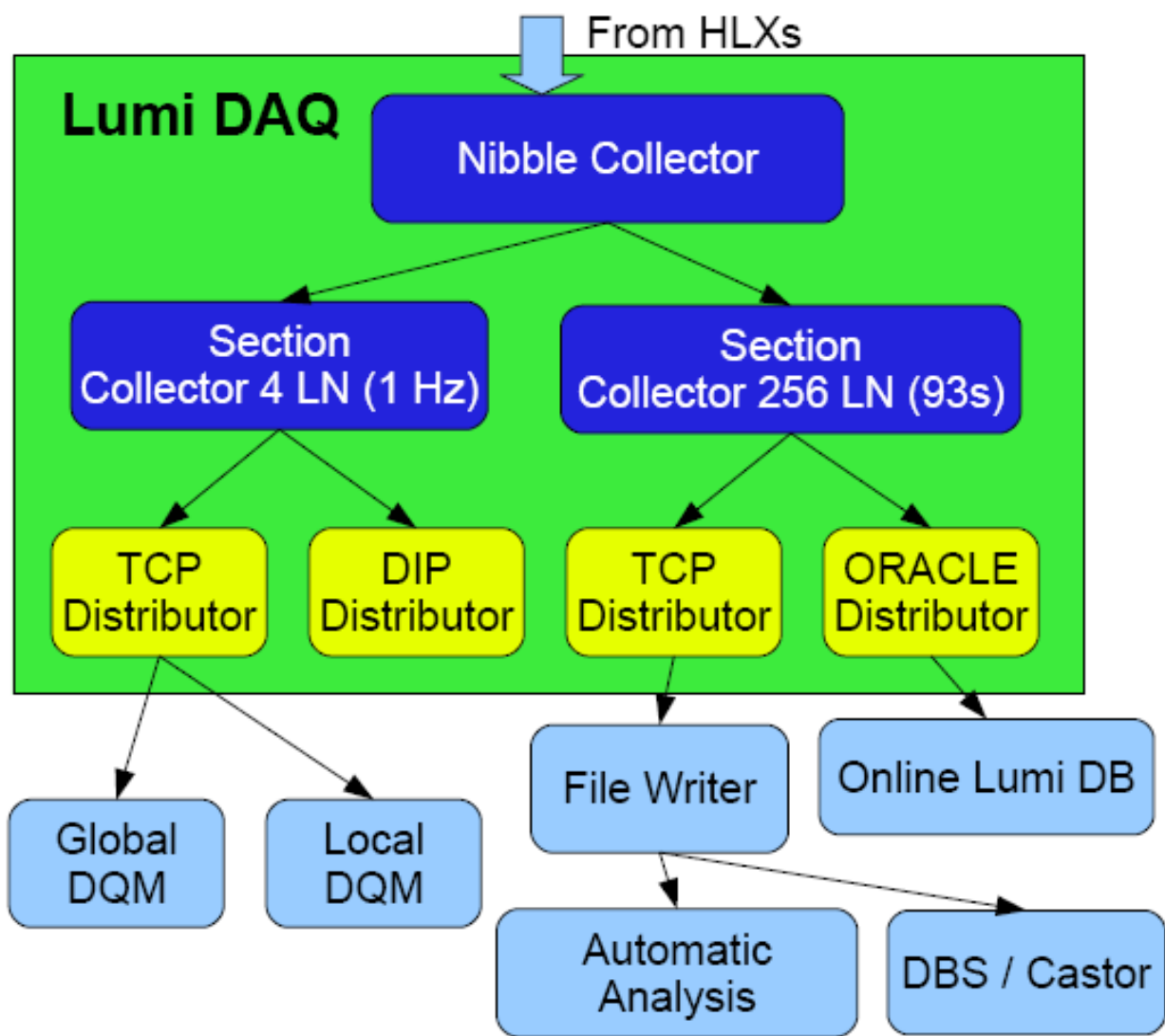
Lumi start with CMS
And continuous until
the next 0C0

Extra Lumi Run

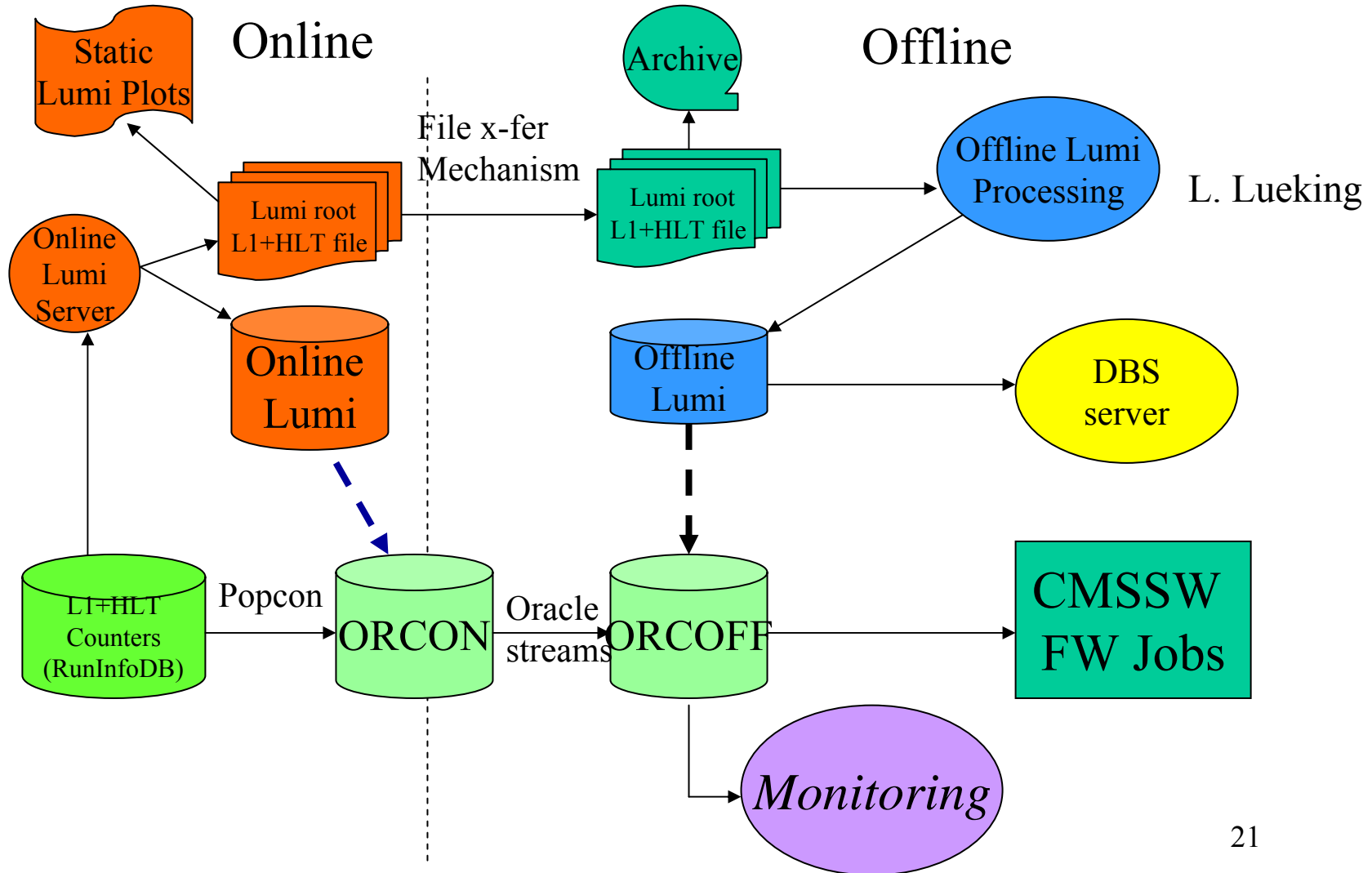
Both HCAL + Lumi run

Lumi run 87% of time
CMS was on 57% of the time

Software Architecture



Schematic diagram of lumi DataFlow



- Summary and details luminosity data would be loaded into the event during prompt reconstruction.
- The loading happens ~ 24 h from the end of the run
- Lumi data is defined in CMSSW/DataFormat/Luminosity

Lumi Summary

Inst Lumi/LS

Error Lumi

Quality of Lumi

Scalers

deadTime

Lumi Details

EtSum+Occ

Bunch by Bunch Lumi

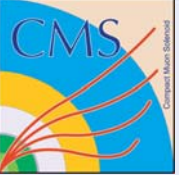
Error on Lumi

Lumi Quality

- Lumi ROOT file is accessible in DBS (Local DBS at the moment)
 - “find file where file like *CMS_LUMI%
 - In DBS Prod Tier0
 - Lumi data structure in ROOT file described in LumiWiki_ROOTFileStruct (update in progress)

- Once DAS + Lumi integration is complete
- A query page would be available to calculate the lumi for on demand
- The page would interface Lumi+DBS+Quality+Config DB and allow the user to correlate the different inputs

- Can be accessed from **LumiWiki_DataQualityMonitoring**
- Online:
 - DQM
 - Static plots from our raw data
 - LHC monitoring accessed from WBM
 - OMDS browser a week of lumi data
- Offline
 - Offline plots of processed lumi data (under development)



New Offline Monitoring Tool



Web CondDB - Mozilla Firefox

File Edit View Go Bookmarks Tools Help

https://cmsweb.cern.ch/conddb/lumi/index

Princeton SPIRES arXiv CDS Wikipedia Gmail PDG CMS Lumi Page Private EVO DiLepton Lumi LumiSystemOperations MyQED IM

https://svn.prince...tion_VH_OJ108.tex Web CondDB

Dashboard DBS Discovery ProdRequest PhEEx SiteDB **CondDB** Support valerieh » logout

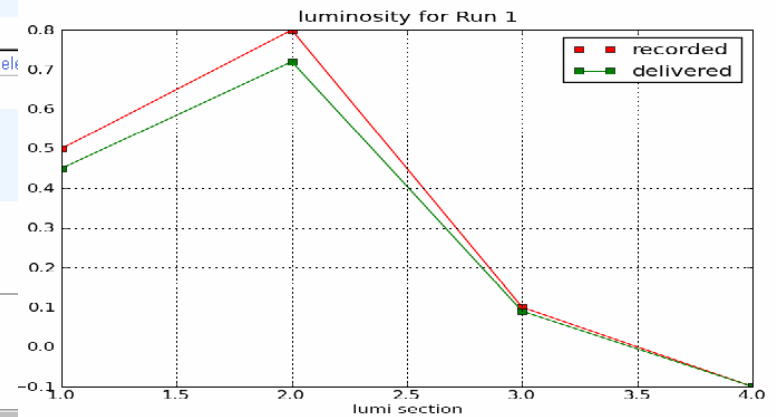
CondDB Navigation :: Overview - IOV Management - IOVOld - Global Tag - Luminosity

[dummy](#) | [cms_orcoff_prod](#)

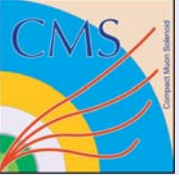
dummy

tag name	class/record	start run	end run	revision	operationlog	plot
chi	luminosity	1	4	0	USER=futyand HOSTNAME=srv-C2C04-21 PWD=/nfshome0/futyand/CMSSW_3_1_0_pre9/src	--select a run--
cruzet	luminosity	80	84	0	USER=futyand HOSTNAME=srv-C2C04-21 PWD=/nfshome0/futyand/CMSSW_3_1_0_pre9/src	--select a run--
craft	luminosity	9,080	9,084	0	USER=futyand HOSTNAME=srv-C2C04-21 PWD=/nfshome0/futyand/CMSSW_3_1_0_pre9/src	--select a run--
crame	hltinfo	9,080	9,090	0	USER=futyand HOSTNAME=srv-C2C04-21 PWD=/nfshome0/futyand/CMSSW_3_1_0_pre9/src	--select a run--
cream	hltinfo	10	90	0	USER=futyand HOSTNAME=srv-C2C04-21 PWD=/nfshome0/futyand/CMSSW_3_1_0_pre9/src	--select a run--

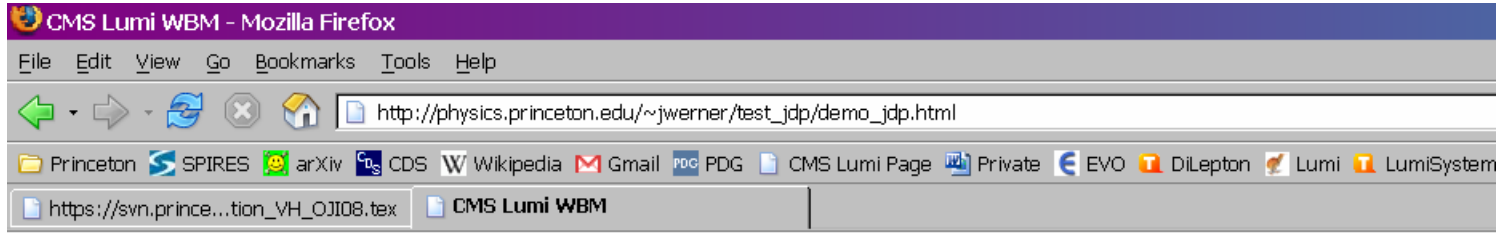
Zhen Xie



Find: rutgers Find Next Find Previous Highlight Match case

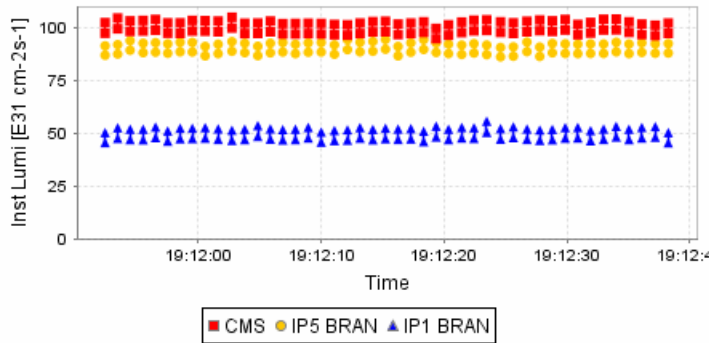


LHC Monitoring

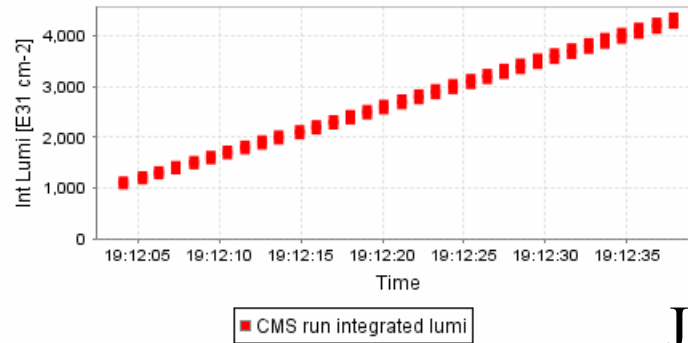


CMS Lumi WBM

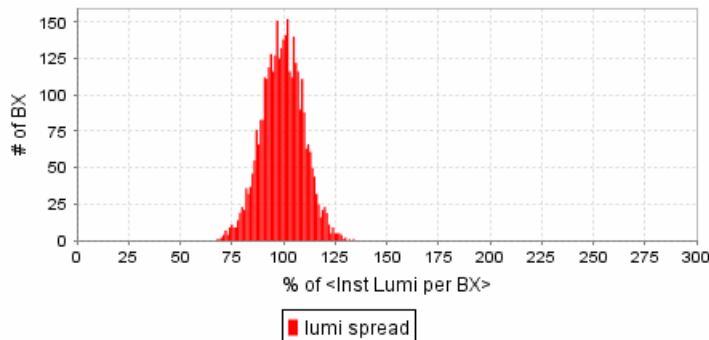
Log Scale Inst Lumi



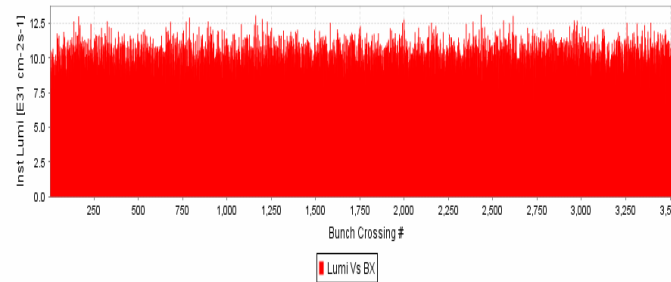
Log Scale Int Lumi



Log Scale Lumi Spread



Log Scale Lumi Vs BX



J. Werner

- CMS will use multiple relative luminosity monitors
- The Calibration procedure is well planned
- Several studies on W/Z and $pp \rightarrow pp\mu\mu$ to calibrate lumi exist
- Lumi is operating in “Continuous mode”
- Lumi is planned to be loaded into the event during CRAFT
- ROOT files should be accessible via DBS
- Discovery page depends on DAS + lumi integration
 - Estimate (D. Evans) multi variable query capabilities might be accessible