## Modeling Breakdown and Gradient Limits

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# We study rf gradient limits at the Fermilab Muon Test Area.



### Many people have contributed to this work.

Normal Conducting A. Hassanein A. Moretti A. Bross Y. Torun D. Huang R. Rimmer D.Li, M. Zisman D.N. Seidman S. Veitzer Superconducting M. Pellin G. Elam A. Gurevich J. Zasadzinski Th. Proslier L. Cooley G. Wu

Plasma Phys RF RF, instrumentation RF, instrumentation RF, Instrumentation cavity design, expts. cavity design, expts. Expt design High E / materials Plasma modeling

ALD, expts ALD, expts. SCRF theory SC theory and exp SC theory and exp SCRF SCRF Purdue FNAL FNAL IIT JLab LBL LBL Northwestern U Tech-X

ANL/MSD ANL/ES NHMFL IIT IIT FNAL FNAL

## What determines the operational rf gradient limits (NC & SF)?

- Accelerator performance is limited by arcing.
- The arcing problem is very old and not adequately described anywhere. (even after ~110 years, - A "breakdown" of the scientific method?) Data is sparse and clustered, hard to compare.
- Our basic assumption is that all arcs have a lot in common: Warm accelerator, SRF, Tokamak, laser ablation, cathodic arcs, large/small gap, lightswitches, micrometeorites, +/-, e-beam welding, high pressure, cavities, RF to DC,
- We want a model that:

   is simple,
   can explain all features of the discharge in detail,
   including accelerator gradient limits,
   in all environments,
   and can point the way to a solution.

## The breakdown model.

- Coulomb explosions trigger breakdown fatigue (creep) and Joule heating help.
- Breakdown arcs are initiated by FE ionization of fracture fragments.
- The arcs produced are small, very dense, cold, and charged +(50-100) V to surface.
- Small Debye lengths,  $\lambda_D = \sqrt{rac{\epsilon_0 KT}{n_e q_e^2}} = \sim nm$ , produce fields,  $E = \phi / \lambda_D \sim GV/m$ .
- High electric fields produce micron-sized unipolar arcs.
- Unipolar arc energy produces craters and surface roughness.







Emax - damage equilibrium

Arc electrons to wall

### OOPIC Pro modeling shows us how the arc starts.



# What is a Unipolar Arc?

- A unipolar arc is an inertially confined plasma on an equipotential surface. ٠
- The literature is not very descriptive, neither is the name. It is very bipolar. ٠
- Unipolar arc parameters: • The arc is dense. Electron motion Ion cloud Electrons diffuse away The plasma is charged to ~50 V. FE electrons maintain the plasma. Potential, a. u. Electron Cloud Ions heat the surface. 1.0 FE, ion currents can be large. 0.8 Radius, microns MG Magnetic fields possible. 0.6 0.4 Arc energy goes into craters. 0.2 0.0 -0.2 -0.4 0.0 3.0 5.0 5.0 5.0 7 5.0 7 In our case: ٠ Things are very bipolar. 8.0 Electrons return elsewhere. 7.5 13.0 Arc energy goes into craters. 6.0 length, microns 4.5 3.0

0.0

22.0 23.0

# Where does the unipolar arc fit in plasma physics?

•

1

X

•

X

1

25 • The unipolar arc is <u>not</u> a "plasma". 50% ionization hydrogen plasma "Plasmas" are defined by:  $(4\pi/3)n\lambda_{n}^{3} < 1$ Laser Focus Plasm 20  $\lambda_D < L$ (size)  $\lambda_{\rm D} = 1 \, \rm nm$ > 1 cm Z-pinches  $N_D \gg 1$  (screening) tubes High ωτ > 1 (collisionality) pressure Fusion = 10 nmarcs reactor 15 The Debye length is too short  $\log_{10} n \ (\mathrm{cm}^{-3})$ Fusion screening is marginal (!?) experiments Alkali Low pressure metal plasma 10 Solar corona Glow discharge Flames  $\lambda_n > 1 \text{ cm}$ Earth iono-5 sphere Traditional plasma methods Numerical & atomistic methods Solar wind Earth (1 AU) plasma sheet 0 -2 -10 1 2 3 4 5 log<sub>10</sub> T (ev)

## Unipolar arcs attack surfaces,

- ... and they do it very efficiently.
- The interactions of high density, low temperature plasmas with materials was studied actively in the fusion community until about 1990.
- Numerical modeling of self-sputtering at high fields and high temperatures shows high secondary atom yields, but codes give surface temperatures of ~10000 degC so the surface could not survive.
- Erosion rates on the order of,  $r = n_{\rm I} v_{\rm I} Y(\lambda_{\rm D}, \phi, T_{\rm surf}) / V_{\rm A}$  are ~ 1 m/s.



### The unipolar arc is complex.



## Much of the arc is experimentally accessible.

We are continuing to model the arc with OOPIC Pro and VORPAL.

#### Trigger

We can measure  $E_{local}$ , emitter size, and density of breakdown sites,  $n(\beta)$ , n(r),  $n(E_{local})$  of sites What is the material and magnetic field dependence ?

#### Ionization

Optical radiation(t) describes the arc (core or edge?), degree of ionization? X rays give time development, power.

#### Unipolar arc

Basic dimensions and parameters could be measured better.

 $E_{surface}$ 

Dependence of damage parameters on power (or anything else)

# What happens to the cavity energy?

- X ray data show how energy leaves the cavity. Relativistic electrons take it.



At the MTA our 805 MHz pillbox has:

- An easily measured risetime ~ 4 20 ns
- Stored Energy ~ 1 J
- Electron energy ~ 4 MeV
- Electron current ~ 4 A, (40,000 (?!) times the field emitted currents)

### We can compare measured and predicted rise times.

We can look at rise times of the shorting current pulse.

- The initial few ns have been modeled in detail in OOPIC Pro.
- The end of the breakdown event was measured with x rays.



### There is a spectrum of enhancement factors.

• Everyone sees roughly the same thing.



## The properties of breakdown sites have been measured.

	E <sub>local</sub> V/m
Lord Kelvin, ('04)	9.6E9
Alpert et al, JVST ('64)	8e9
KEK ('09)	8E9
CERN ('09)	10.8E9
Us ('03)	8E9
Cox ('74)	~7E9

CERN data seems to show deformation of emitter tips at high fields ('09).



radius, m	
	theory
3E-8 to 8E-8	exp
	w
2E-8 to 4E-8	w
~5E-8	w
< 5E-8	w



Cox ('74) measured emitter area vs  $E_{local}$ .

# What is the surface field in the unipolar arc?

- Electrohydrodynamic spinodal decomposition gives a reasonable result.
- $E_{surf} \sim 1 \, GV/m$
- Wavelength ~ 2  $\mu$ .
- Enhancements seem to come from fracture, if dimensions ~ 10 nm.

Laser material interiactions by Getvilas. et al. (2009) and J. Wang and Guo (2005)





and surface waves in CLIC prototype Cu cavities (Izquierdo, 2008) . . . .



## Breakdown events damage the surface

- More energy => more damage
- More damage => Higher enhancement factors => Lower operating fields
- Exponential damage spectrum => logarithmic dependence of operating field.



# We can calculate all aspects normal rf operation.

• Emax vs. Pulse Len. • Emax vs. f 1000 Local E field Tensile stress ~ tensile strength Gradient (MV/m) NLC prototype E (MV/m) 100 Waveguide SS Cu CLIC Au 100 urface field 10 ⊾ 0.1 10 L 1000 10 Frequency (GHz) • BD rate vs. Pulse len. DC breakdown Fermilab linac Local E field, eta E (GV/m) Local Electric Field (µs time scale) rate (arb. units) SLAC / NLC prototype (ns time scale) 0.1 ukdown 0.01 0.01 0.0001 10.5 0.0001 0.001 0.01 0.1 Gap (m) Pulse length (µs or ns) • BD rate vs. E • Emax vs. T 100 rate or probability (arb. units) 9 10 SLAC/NLC prototype CERN/CLIC prote =IELD 0.1  $\geq$ umop 0.01 ITICAL D MD simulation Ecr vs T R 0.001 100 1000 250 450 650 850 1050 1250 1450 Gradient (MV/m) SURFACE TEMPERATURE (K) • Material dep. E<sub>max</sub> vs. pressure • 1000 Maximum Gradient (MV/m) 00 field (MV/m) nospheric DC, and high pressure rf Lab G data 0.0001 0.0 10 ∟ 100 Pressure (Torr) 1000 Tancila etra al AD

10

### Summary

- We can calculate all aspects of arcing.
- Unipolar arcs seem to be the key.
- All data is relevant and explainable.
- There are many applications: Tokamaks, SRF, small gap, laser ablation, micrometeorites, e-beam welding, . .
- Our immediate interest is understanding effects of B fields.

We have a movie you can look at through the CLIC 09 website. We are planning a meeting on Unipolar arcs at Argonne in January