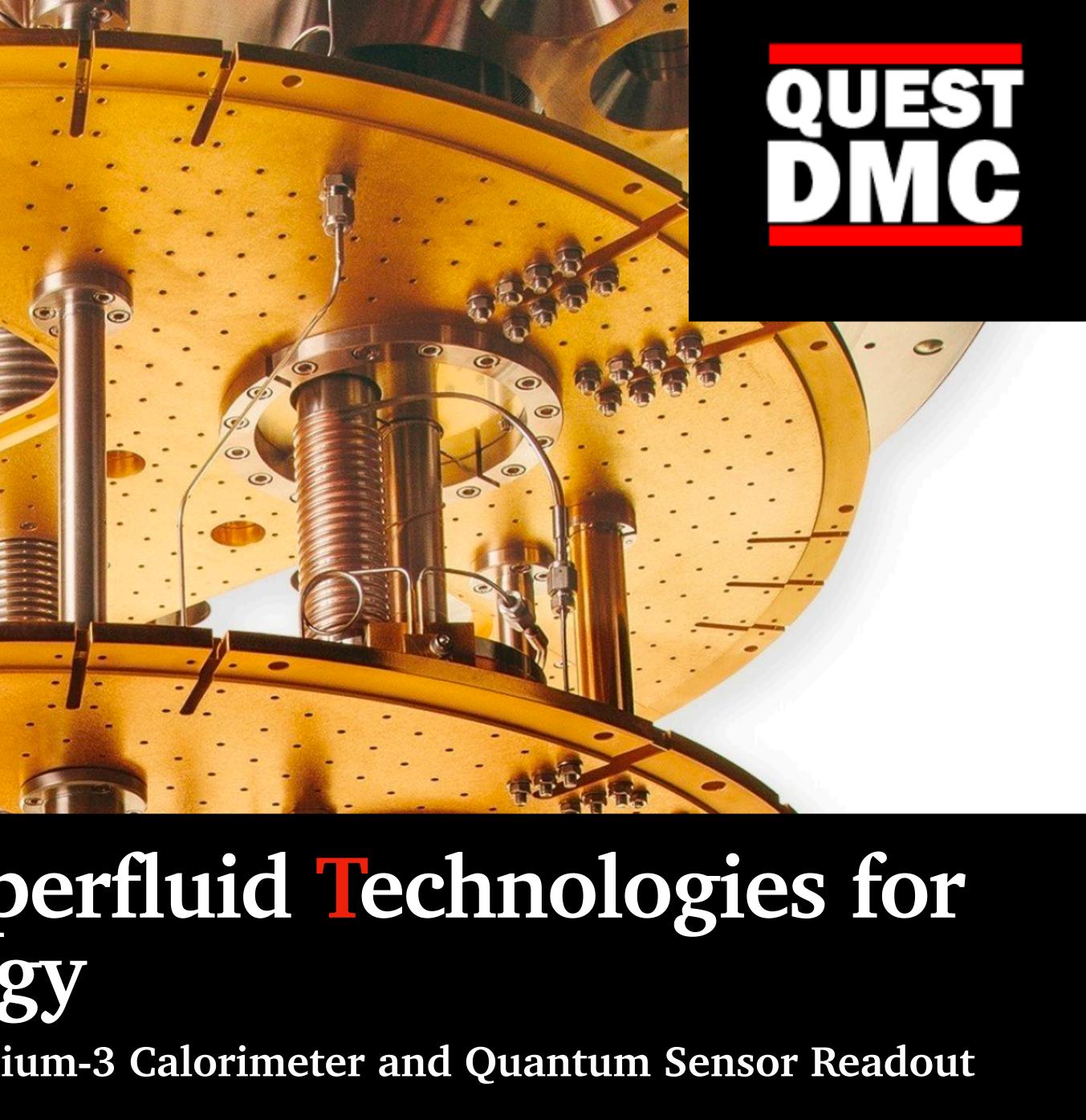
IOP APP, HEPP and 10/04/2024 Liverpool

### **QUantum Enhanced Superfluid Technologies for** Dark Matter & Cosmology Searching for Sub-GeV Dark Matter Using a Helium-3 Calorimeter and Quantum Sensor Readout

Rob Smith - Royal Holloway, University of London - rob.smith.2021@live.rhul.ac.uk/robert.smith2@physics.ox.ac.uk



# **QUEST-DMC** Collaboration











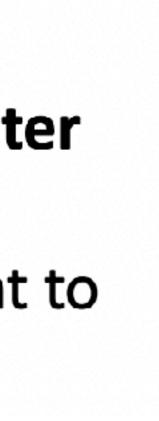
UNIVERSITY OF OXFORD



### 1. Detection of sub-GeV dark matter with a quantum-amplified superfluid <sup>3</sup>He calorimeter

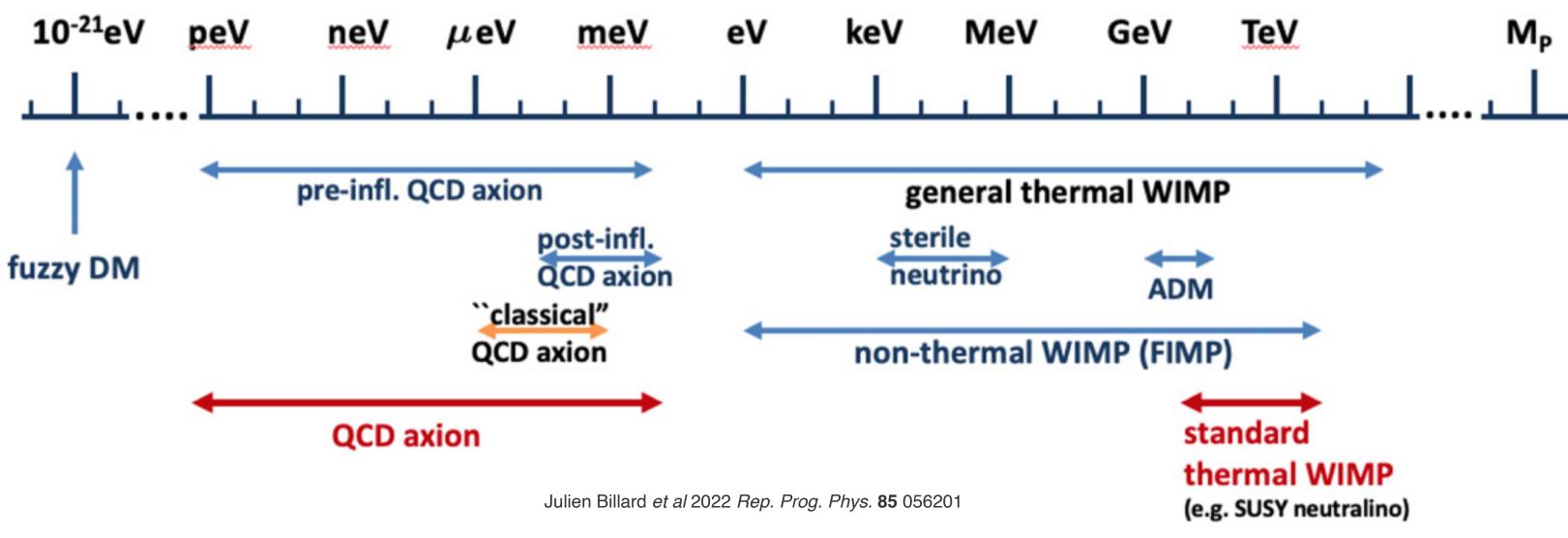
2. Phase transitions in extreme matter, relevant to cosmology and gravitational wave production





## **Motivations & Goals**

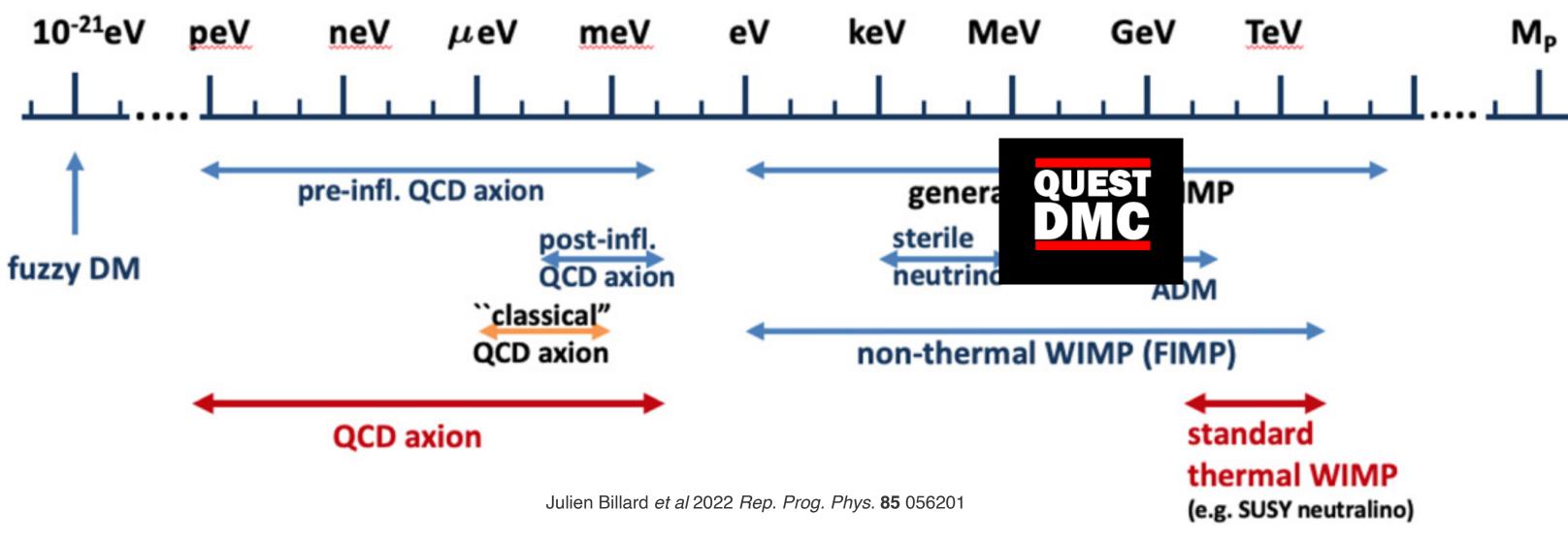
- 85% of matter in the universe is "dark"
- Theoretical motivation for sub-GeV dark matter (ADM, Hidden Sector, Freeze-in...)
- Lower threshold = lower mass DM interaction reach
- Superfluid <sup>3</sup>He target enables eV scale recoil threshold and increased spin-dependent interaction sensitivity





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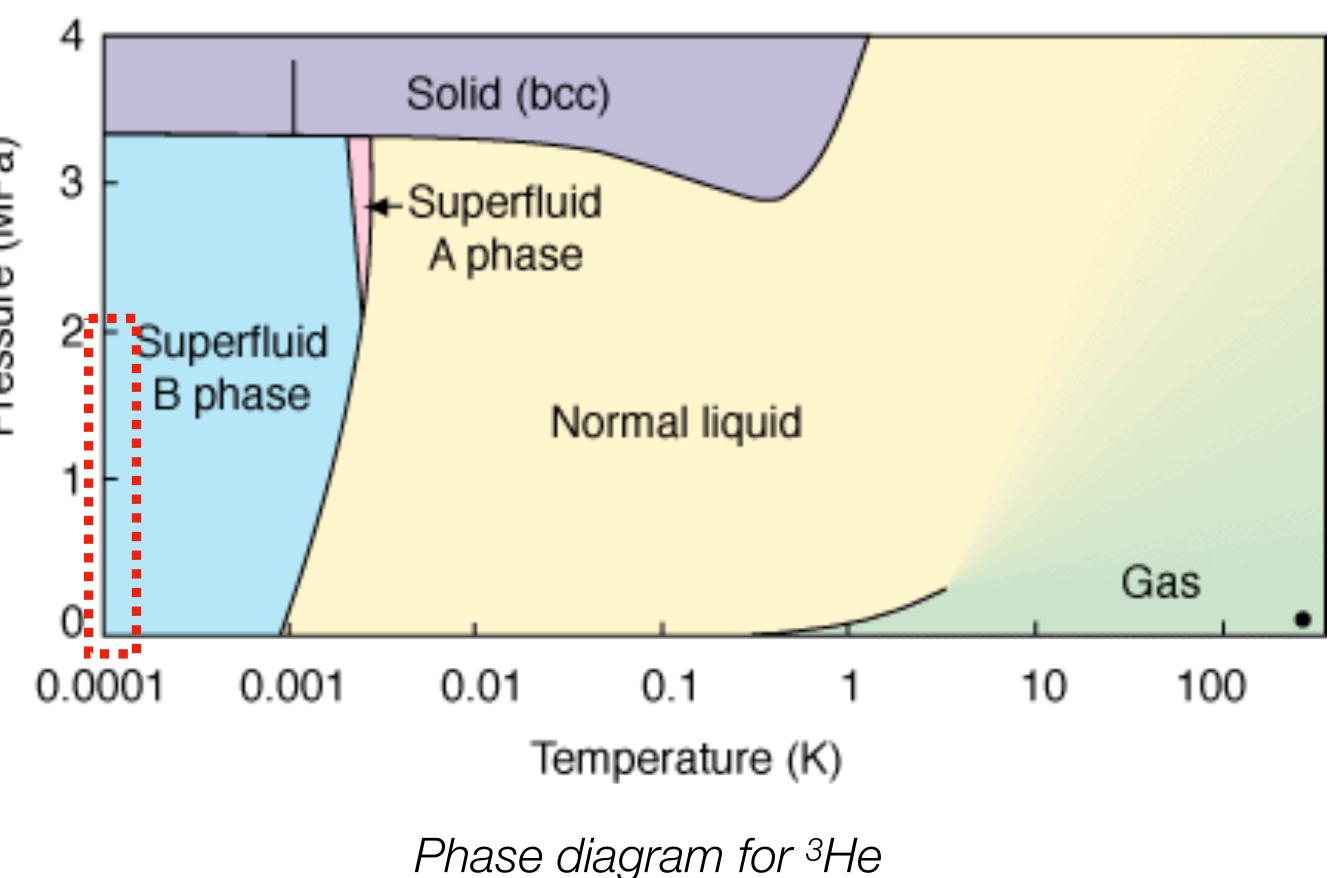




# Why Helium-3?

- Spin 1/2 nucleus sensitivity to spindependent interactions
- Superfluid below  $T_C$  of ~1 mK (0 bar)
- Form bound states analogous to Cooper pairs in superconductors
- Quanta of **10-7 eV**

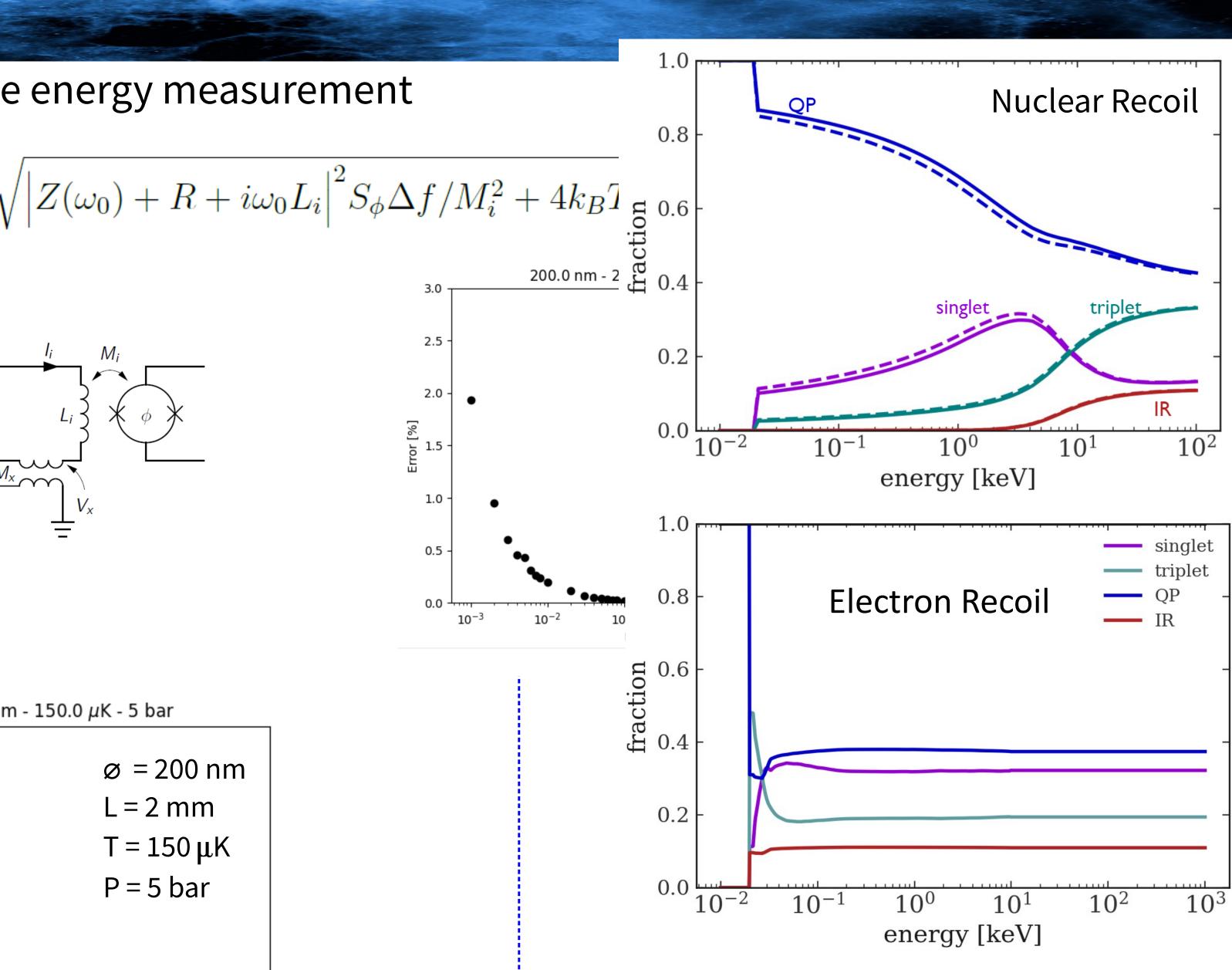
+++ $\Delta_{gap} = 10^{-7} \text{ eV}$ 

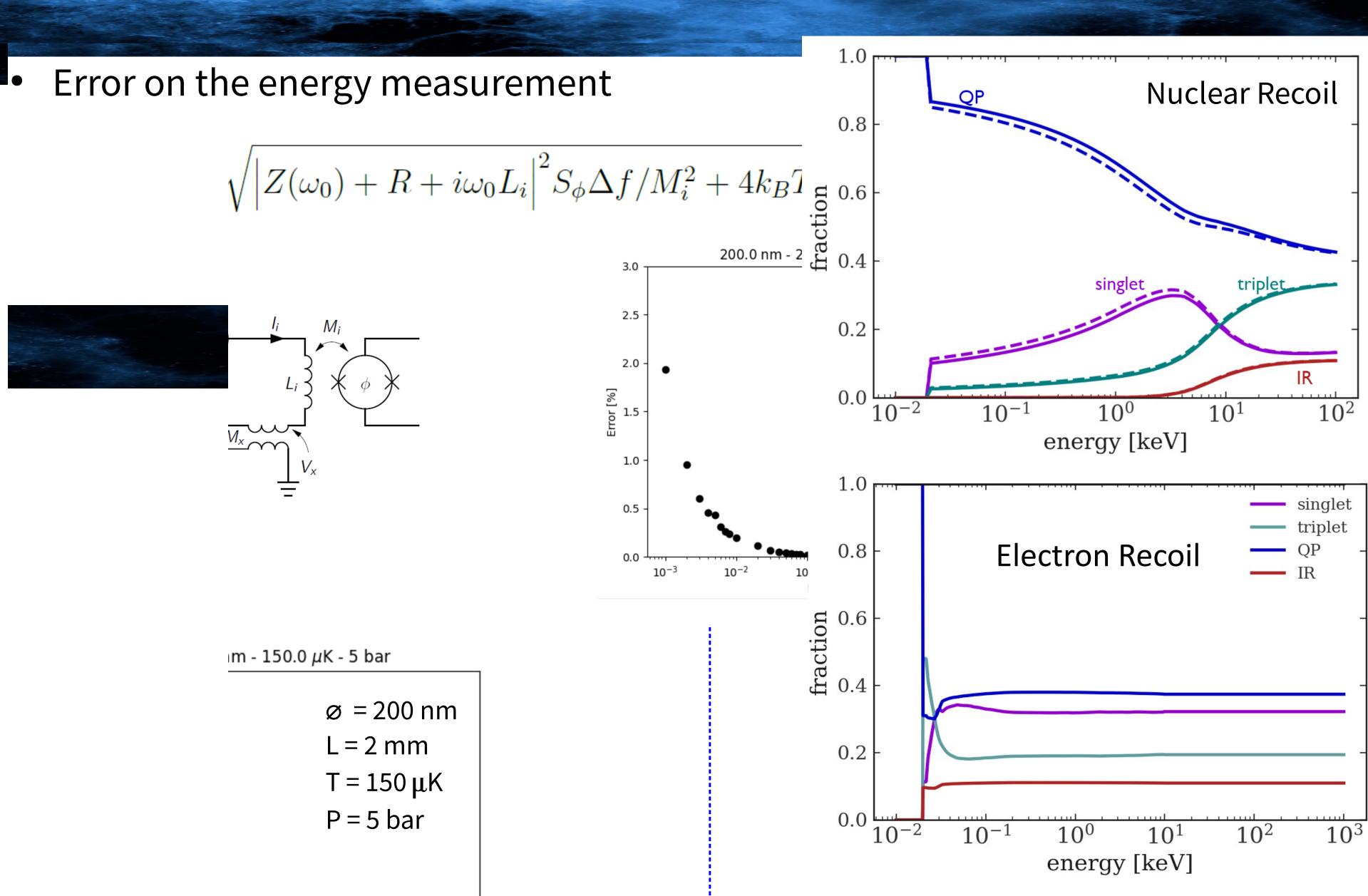




### Rolometer reconnee SOUD

### nse





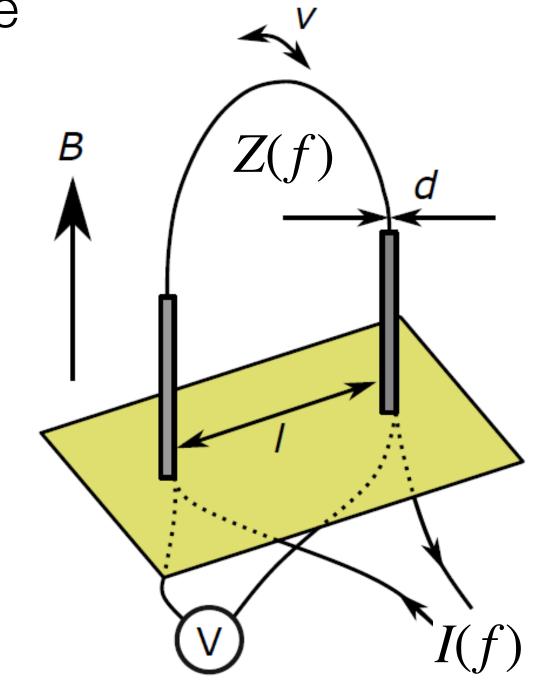
/ire diameter (0.0 bar - 150.0  $\mu$ K)

10000 12500 15000 17500 20000 7500 diameter [nm]

17

# **QUEST-DIVC Detector**

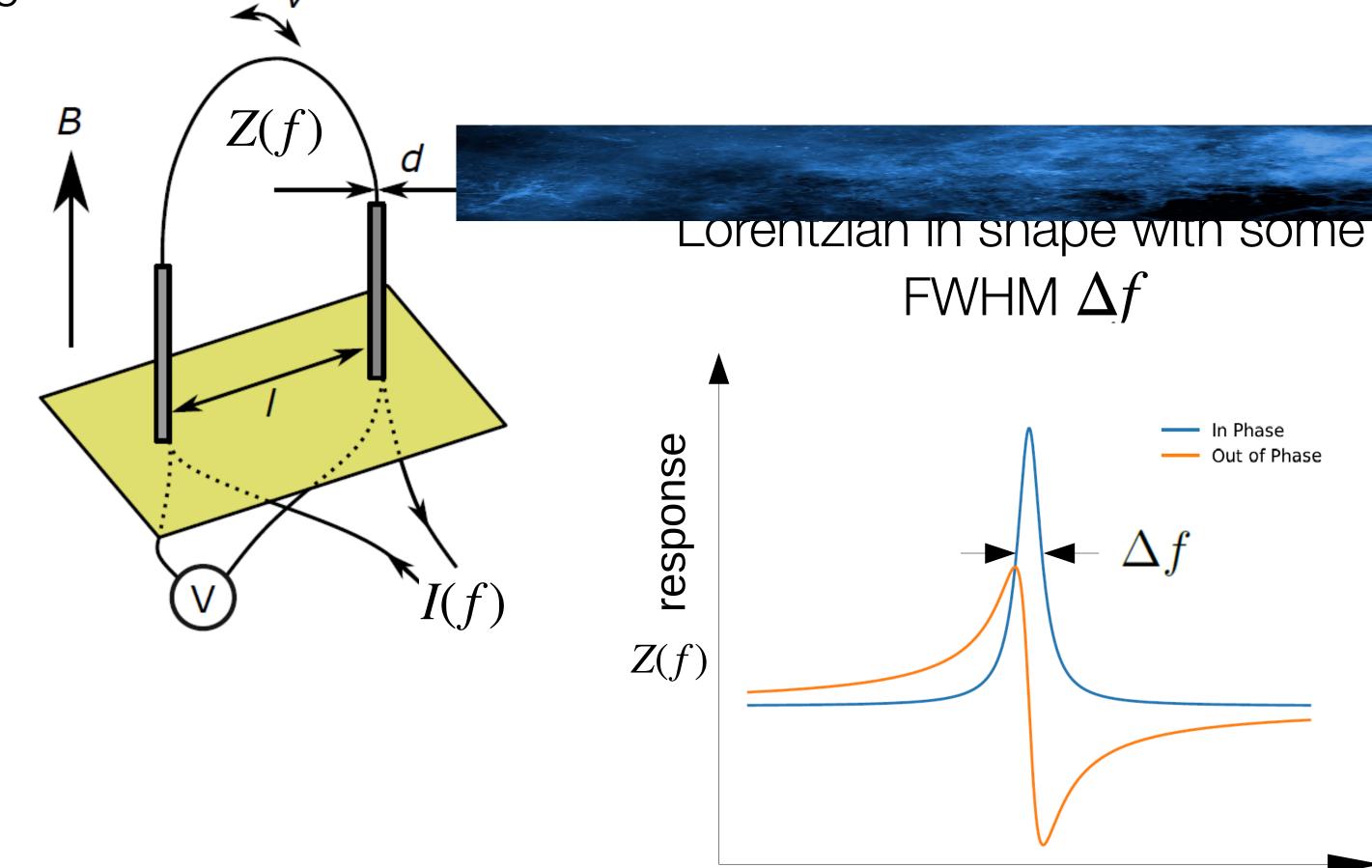
Oscillating superconducting wire in a magnetic field **B** driven at frequency **f** 



Z(f)

# **QUEST-DIVC Detector**

### Oscillating superconducting wire in a magnetic field **B** driven at frequency **f**

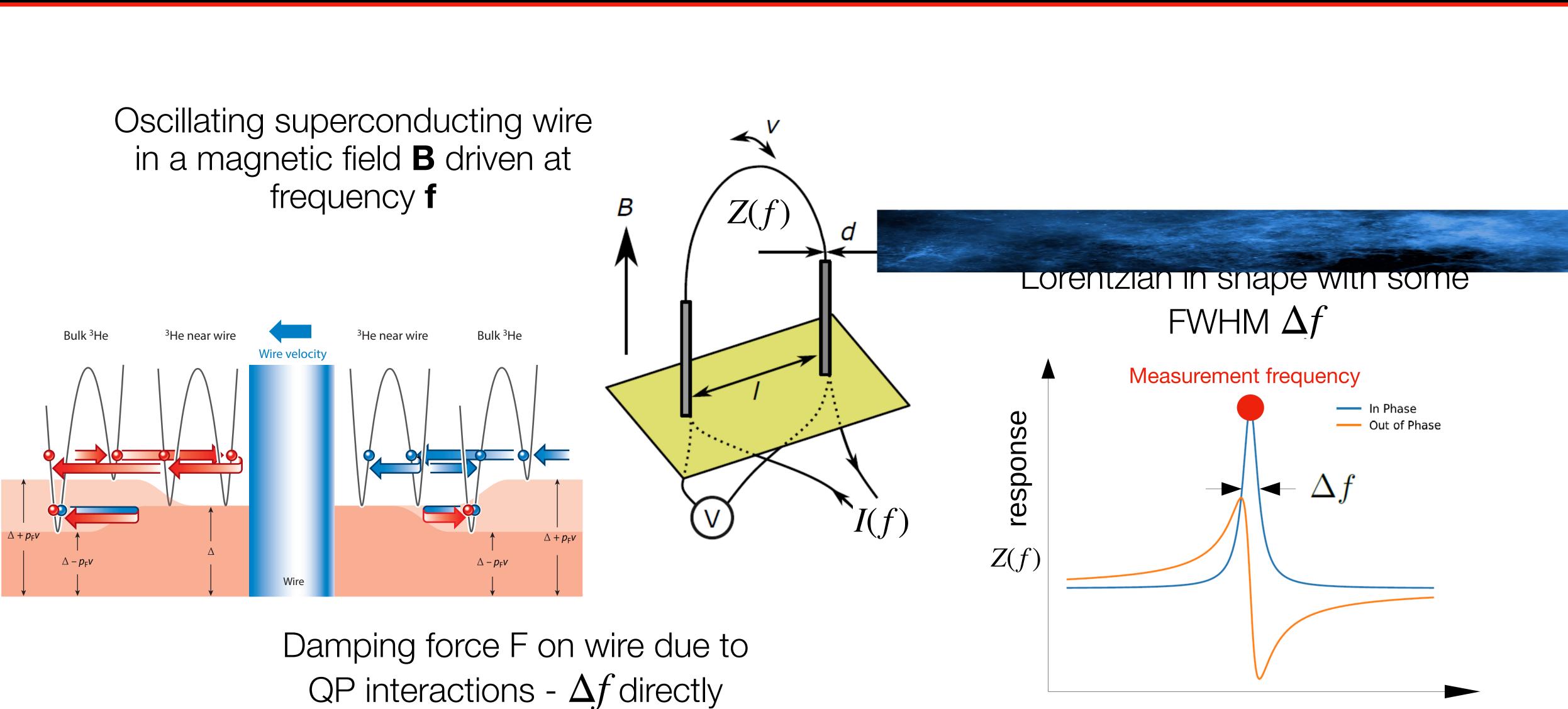


frequency



# **QUEST-DMC Detector**

### Oscillating superconducting wire in a magnetic field **B** driven at frequency **f**

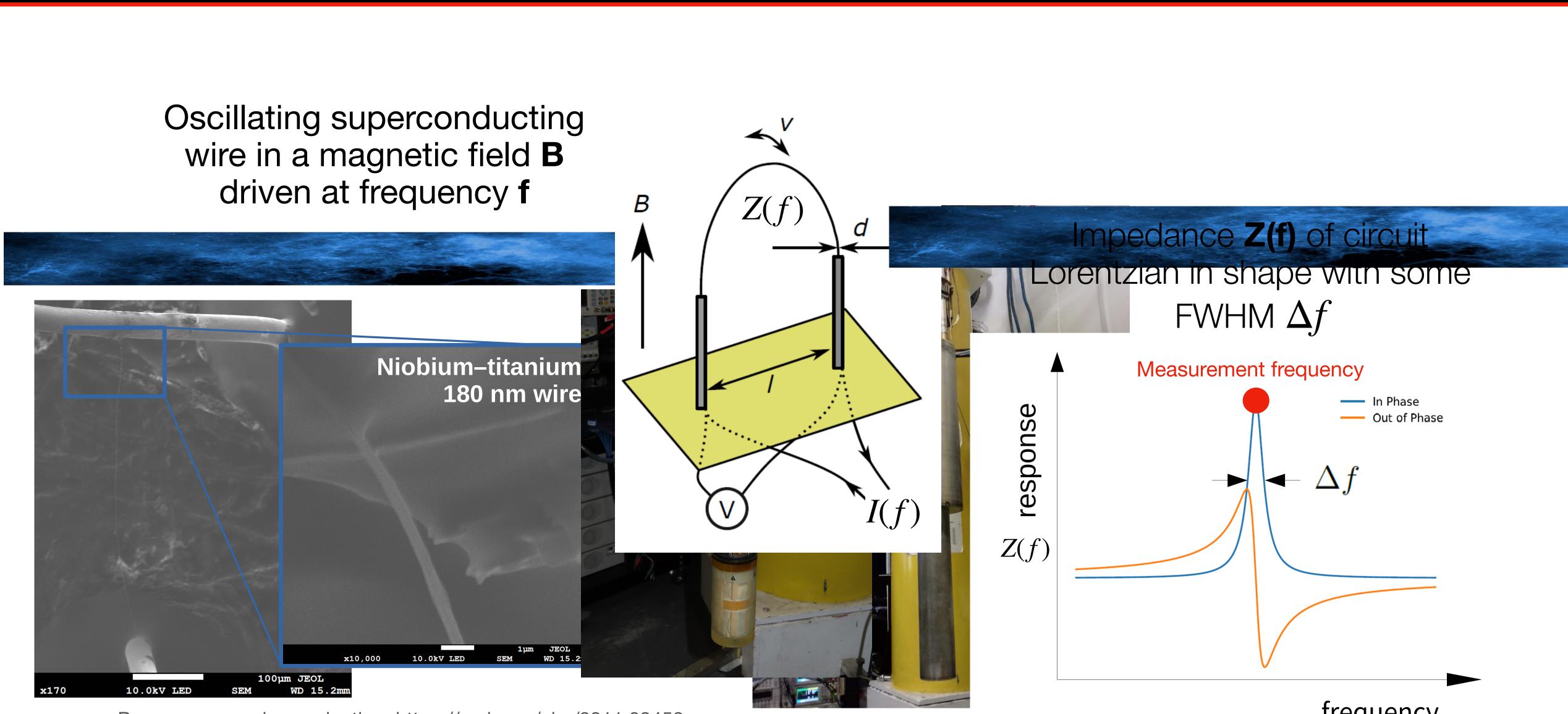


proportional to F

frequency

# **QUEST-DIVIC Detector**

### wire in a magnetic field **B** driven at frequency f



Paper on nanowire production: https://arxiv.org/abs/2311.02452

frequency

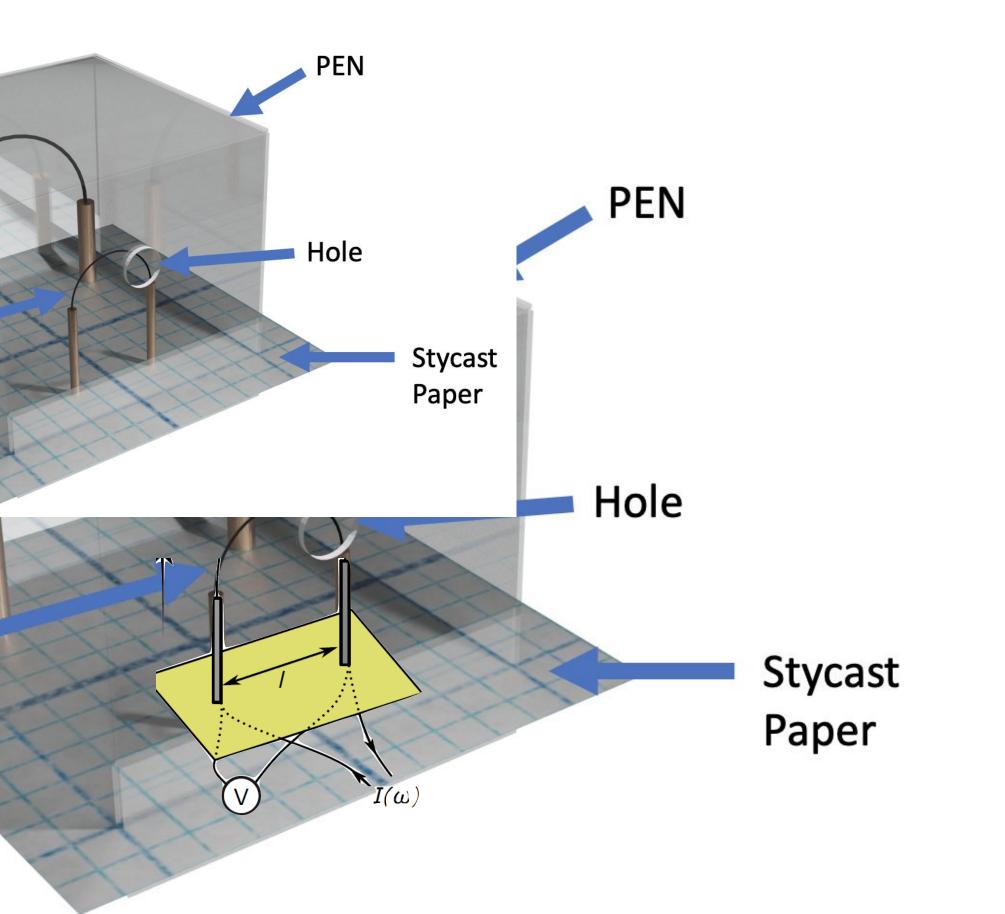
# **QUEST-DMC** Detector



Heater Wire (13.5 μm)

Thermometer Wire (4.5 μm) He; (13

Thermometer Wire (4.5 µm)



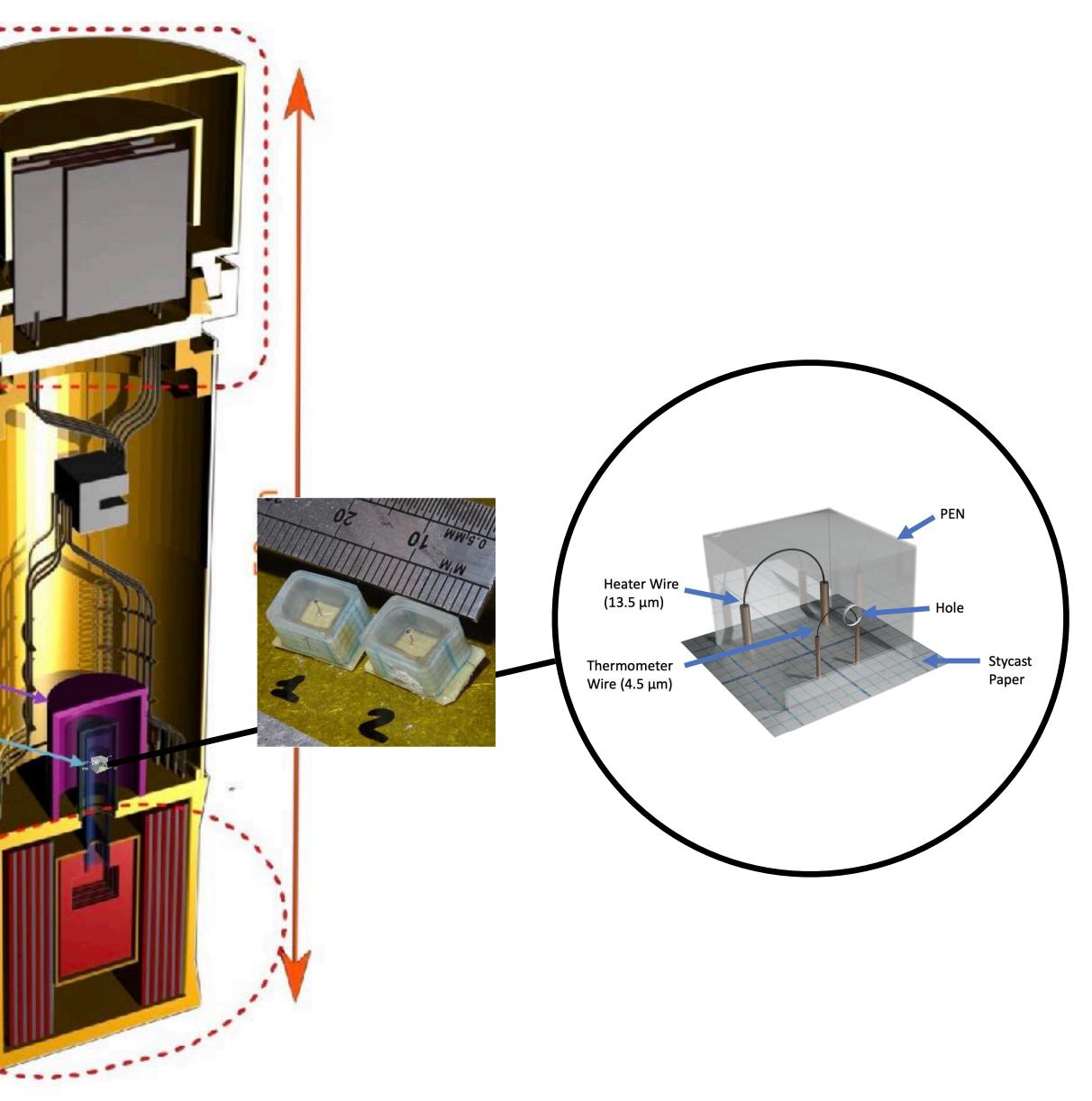
1st generation prototype cell ~ 0.33 cm<sup>3</sup>

## **QUEST-DIVIC Detector**



Photon detector . Bolometer cell ~

> Copper demagnetisation stage at <100µK

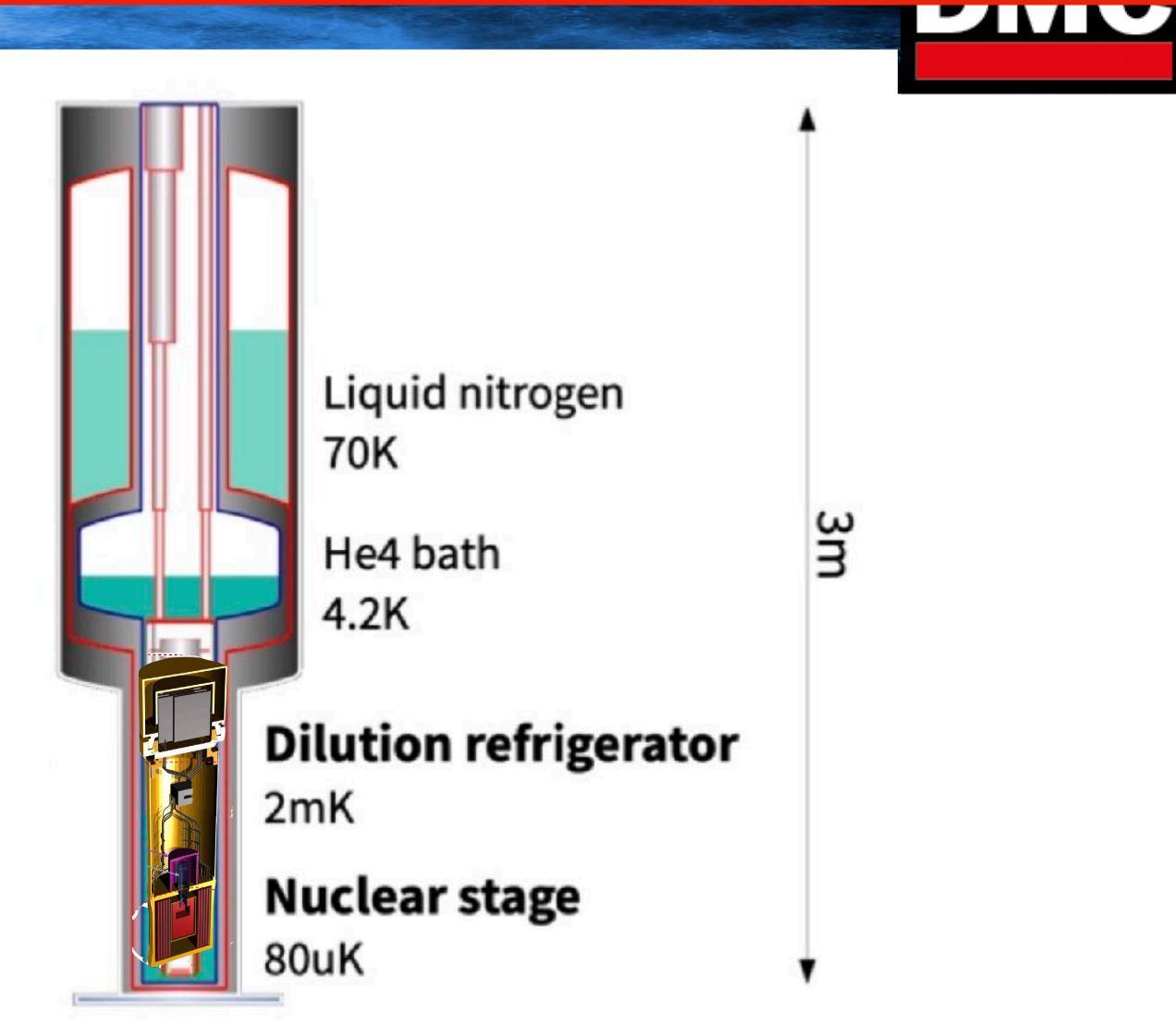


# **QUEST-DIVIC Detector**



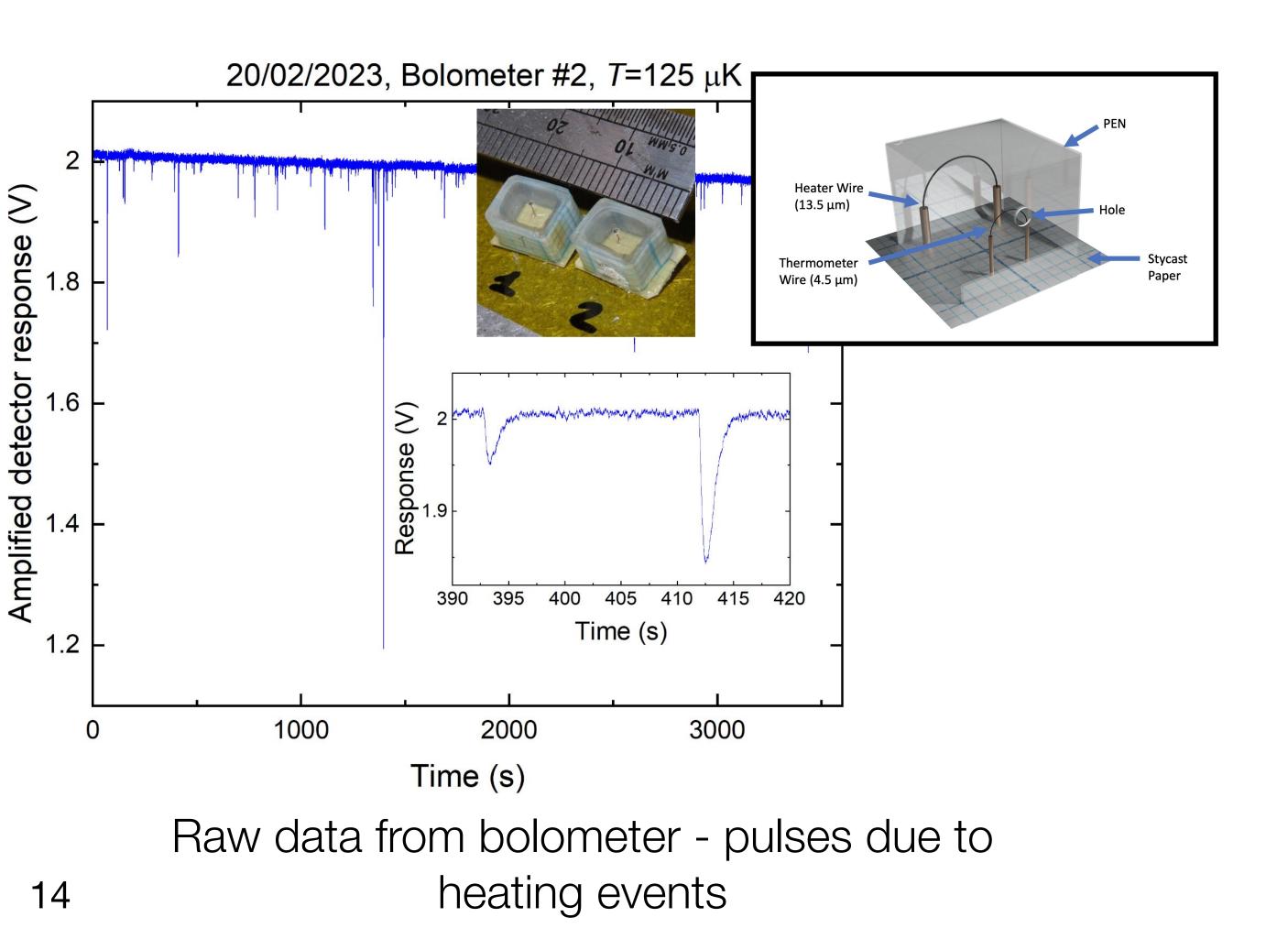






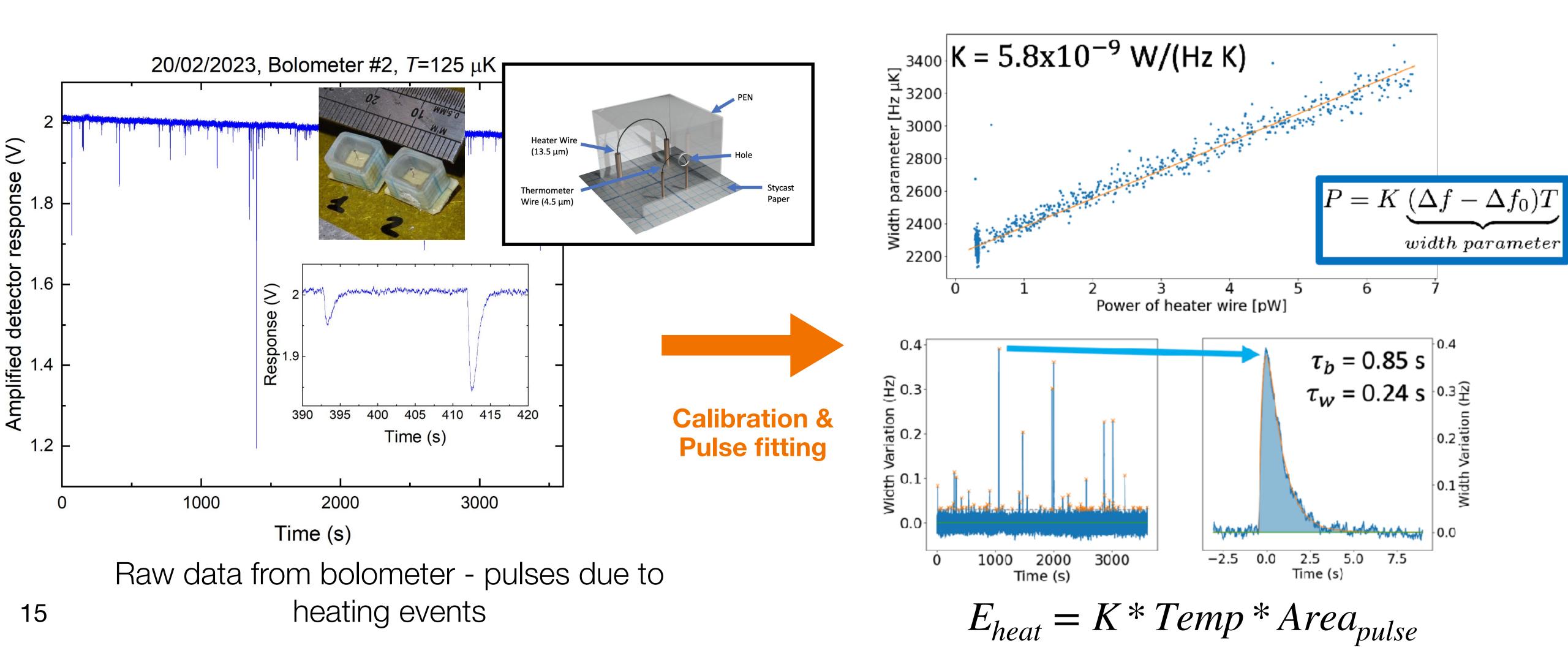


# Quasiparticle Calorimetry





# Quasiparticle Calorimetry





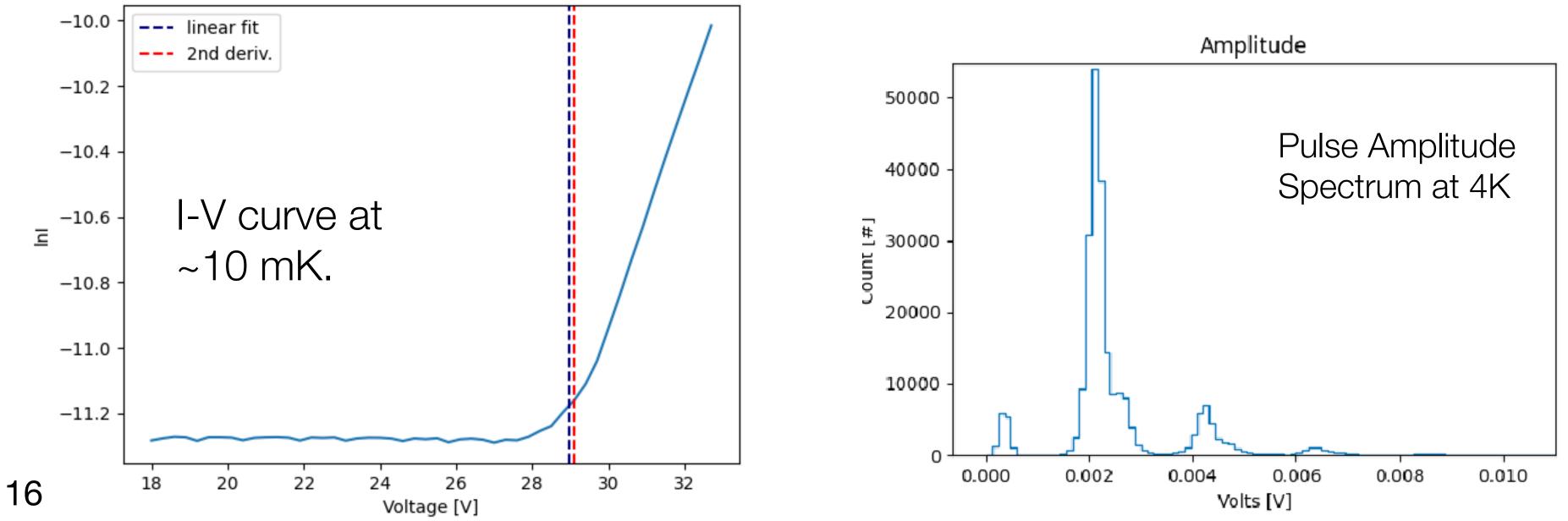
Drive heater wire past Cooper pair breaking velocity and measure thermometer wire response

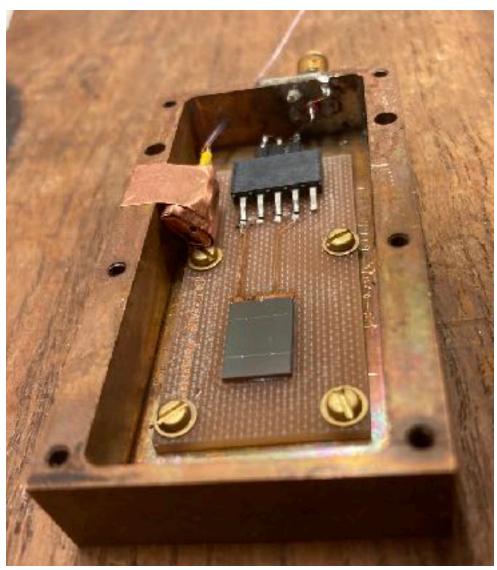


### Photon Detection

Current tests: Operating the SiPM in a dilution refrigerator at ~10 mK.

Early tests indicate the SiPM does operate at 10 mK - Work is ongoing to characterise the device at 10 mK and optimise the readout noise.

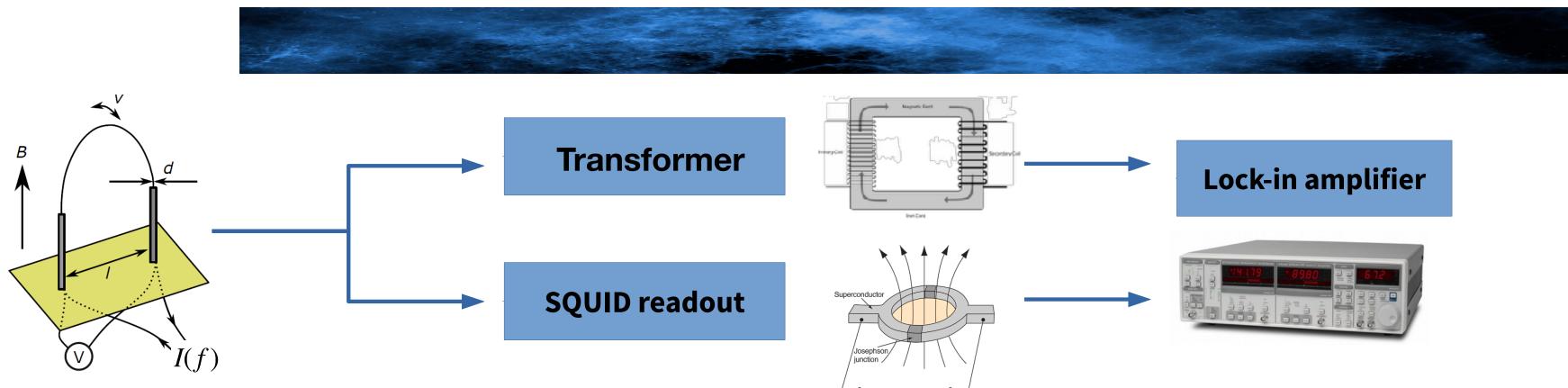








# SQUID Readout

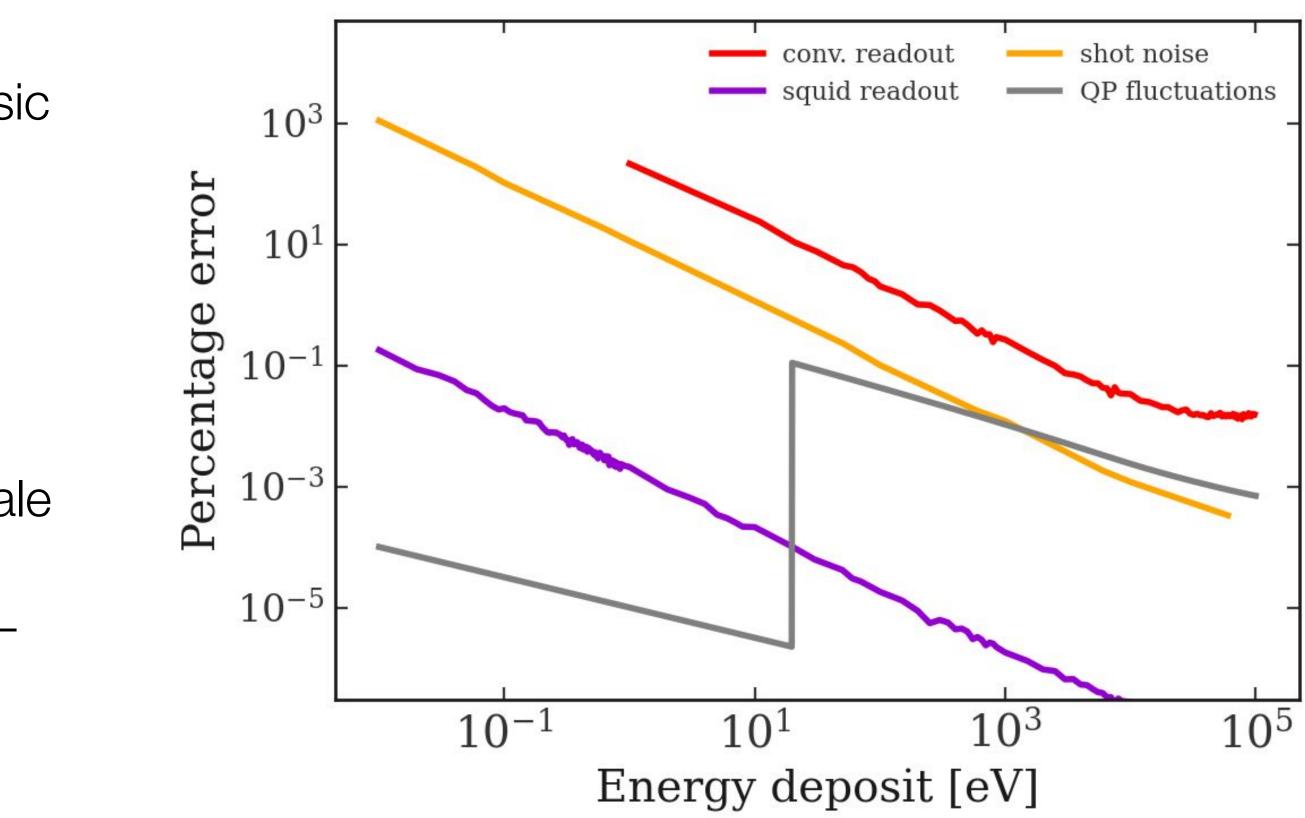


Conventional readout noise dominates over intrinsic limitation of QP noise

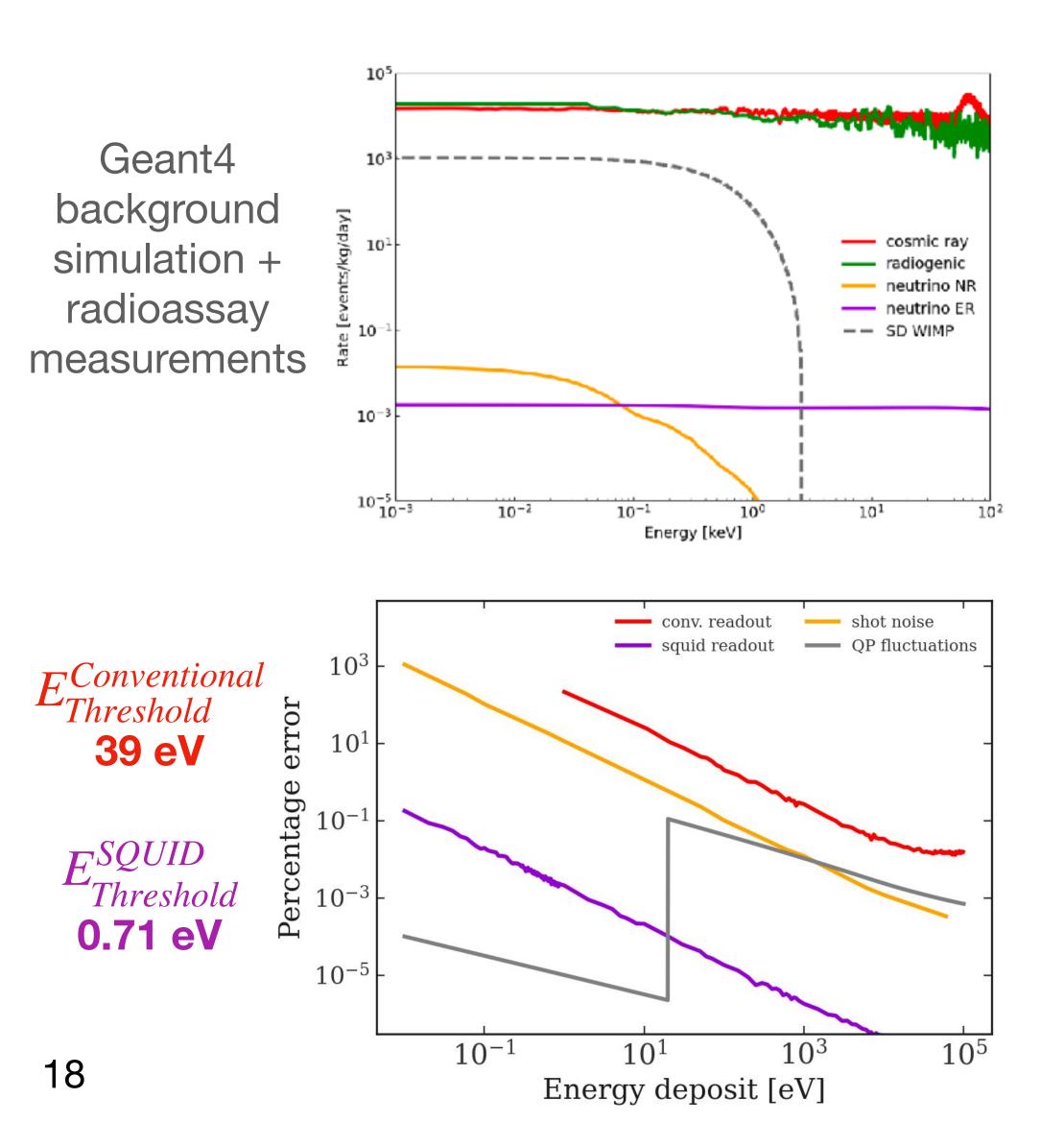
SQUID (Superconducting QUantum Interference Device) can reduce noise of readout by orders of magnitude (< pV scale)

SQUID readout of nanowire operated at < mK scale

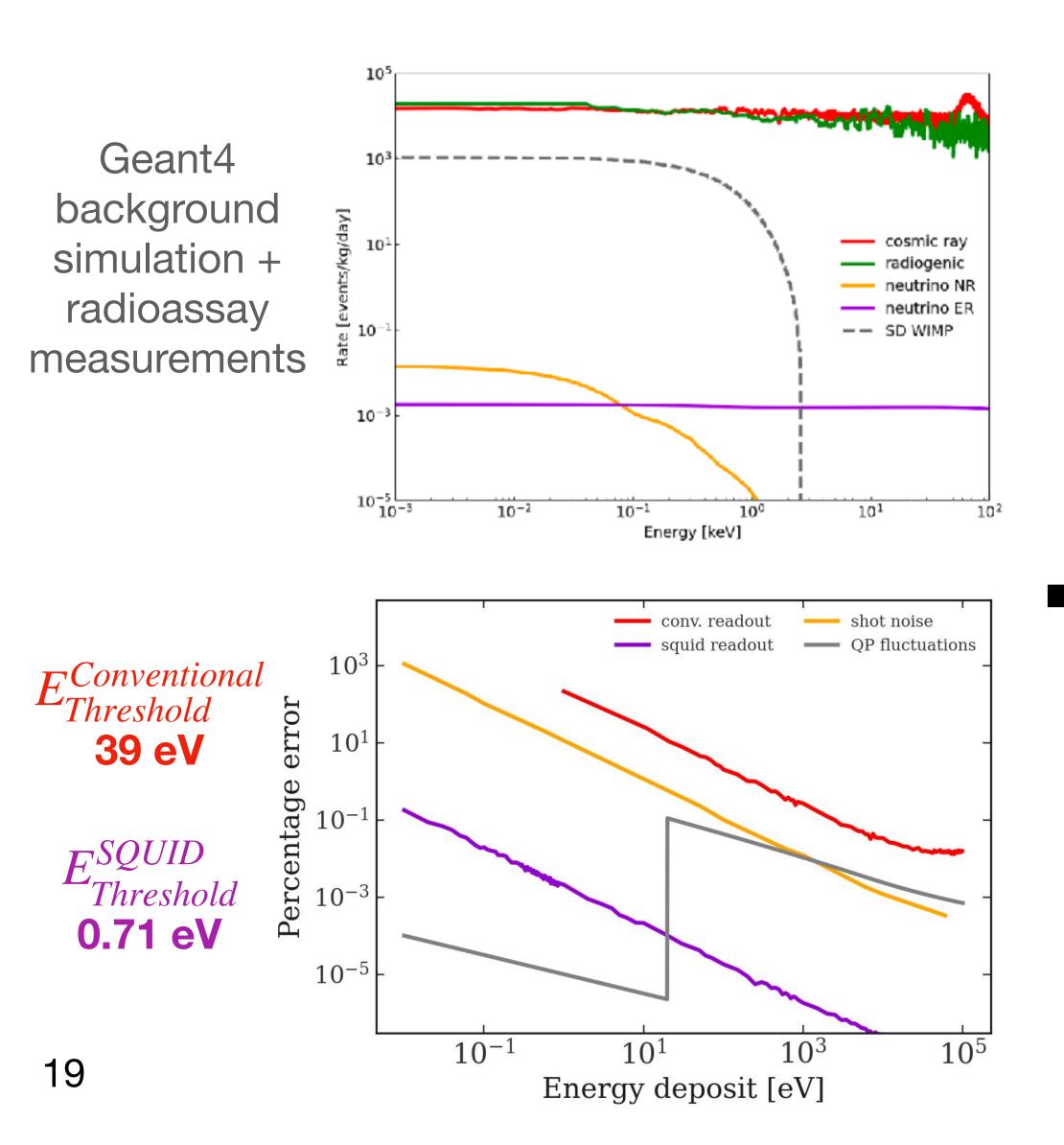
Ongoing work to measure QP shot noise at RHUL



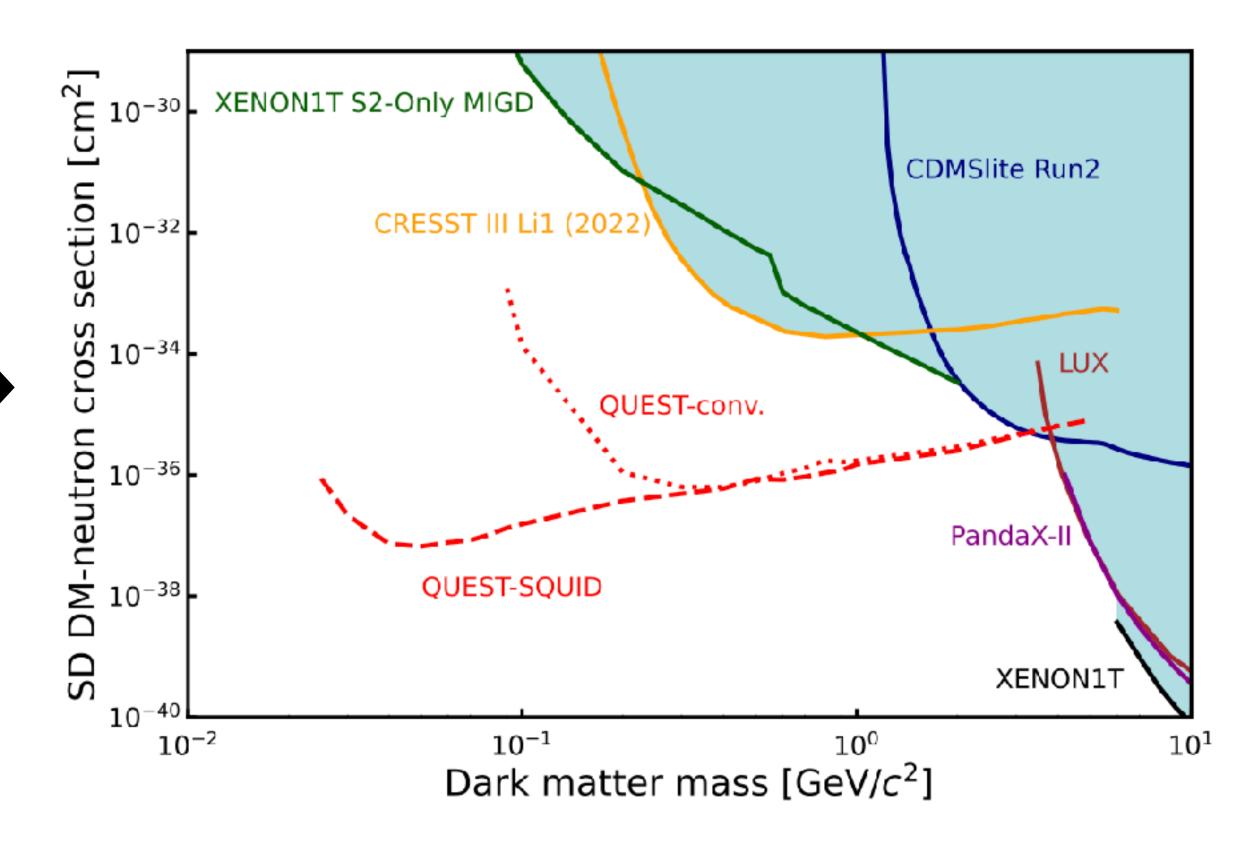
## **Estimated Sensitivity**



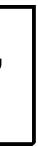
## **Estimated Sensitivity**



### 6 month run, 50% livetime, 200nm wire, 5 x 1cm<sup>3</sup> cells (0.1 g / cm<sup>3</sup>)



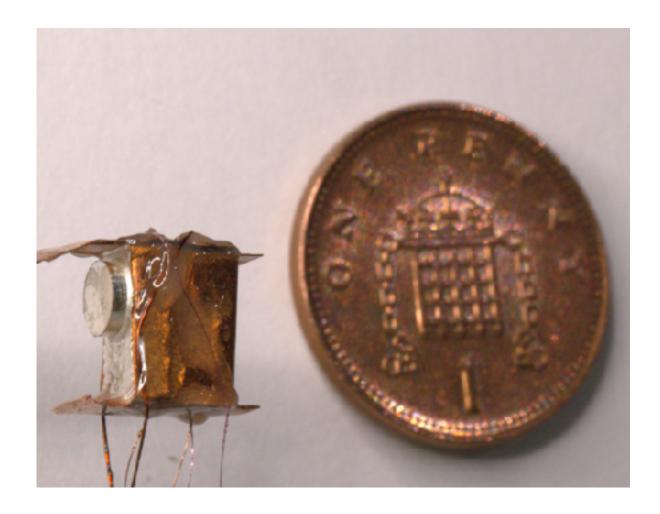
https://doi.org/10.1140/epjc/s10052-024-12410-8

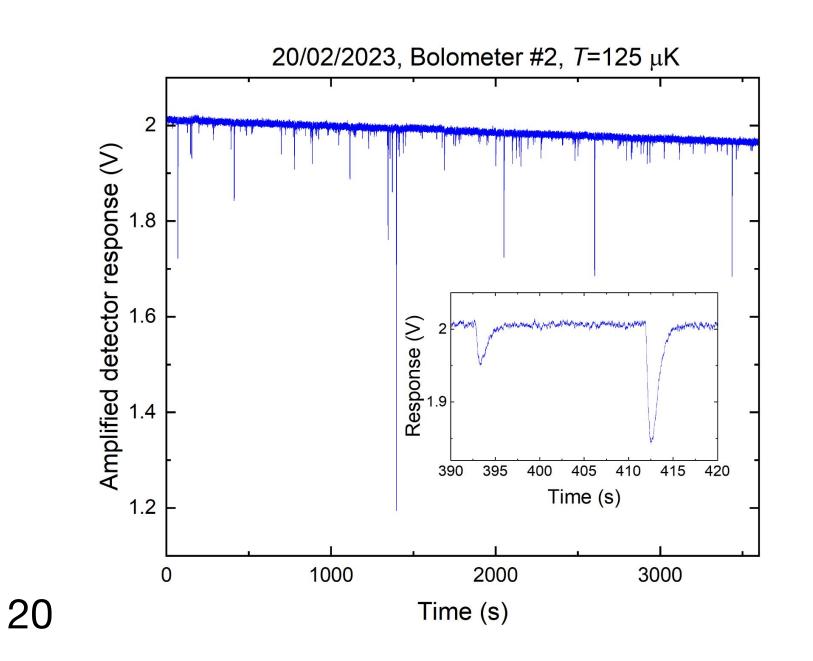


## Conclusions

Proof of concept for eV scale threshold DM detector

Operated bolometers with nanoelectromechanical resonator (NEMs) detector





Next generation bolometer (see above) to be installed at Lancaster with full SQUID readout

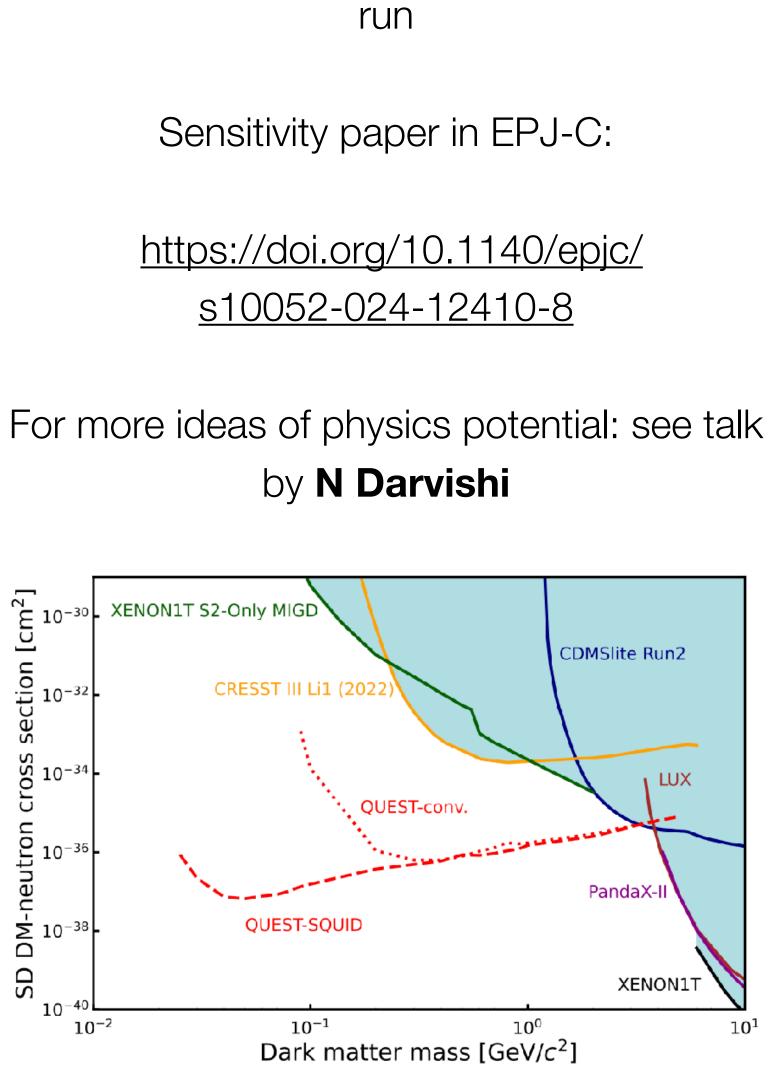
mK scale photon detection, shot noise measurement, monoenergetic source calibration

**Aim:** 6 month dark matter search data taking run

<u>s10052-024-12410-8</u>

### **Ongoing work:**

### by **N Darvishi**



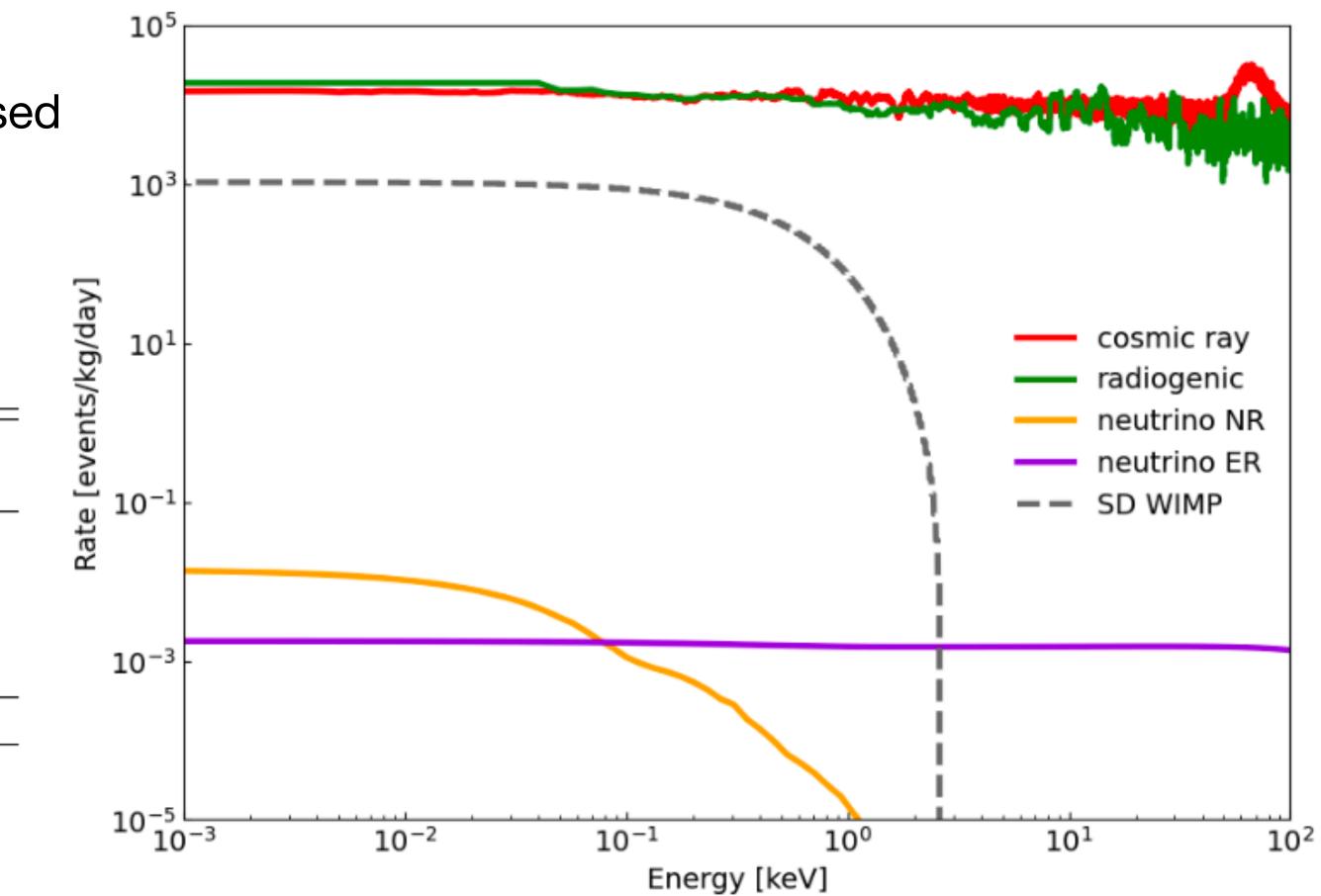
# Backup Slides

# Background

- G4 simulation done for cosmic ray and radiogenic event rates
- BUGS (Boulby Underground Germanium Suite) used for extensive radioassay of detector and cryostat materials

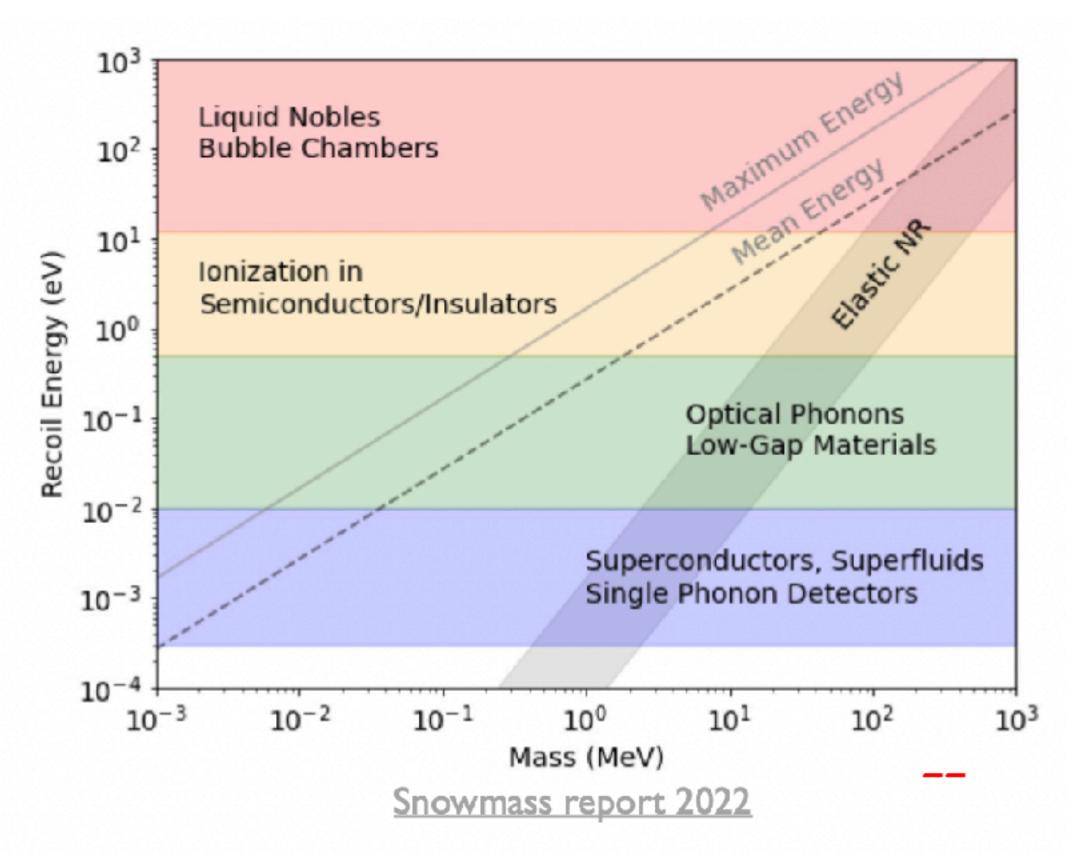
Component	Expected counts [0-10 keV]		Uncertainty
	/kg/day	/cell/day	
Cosmic ray	$1.05 \times 10^{5}$	3.31	$11 \ \%$
Radiogenic ER	$8.31 \times 10^4$	2.61	14~%
Solar $\nu$ ER	$1.51 \times 10^{-2}$	$4.76  imes 10^{-7}$	2 %
Solar $\nu$ NR	$6.37 \times 10^{-4}$	$2.01 \times 10^{-9}$	2 %
TOTAL	$1.88 \times 10^{5}$	5.92	

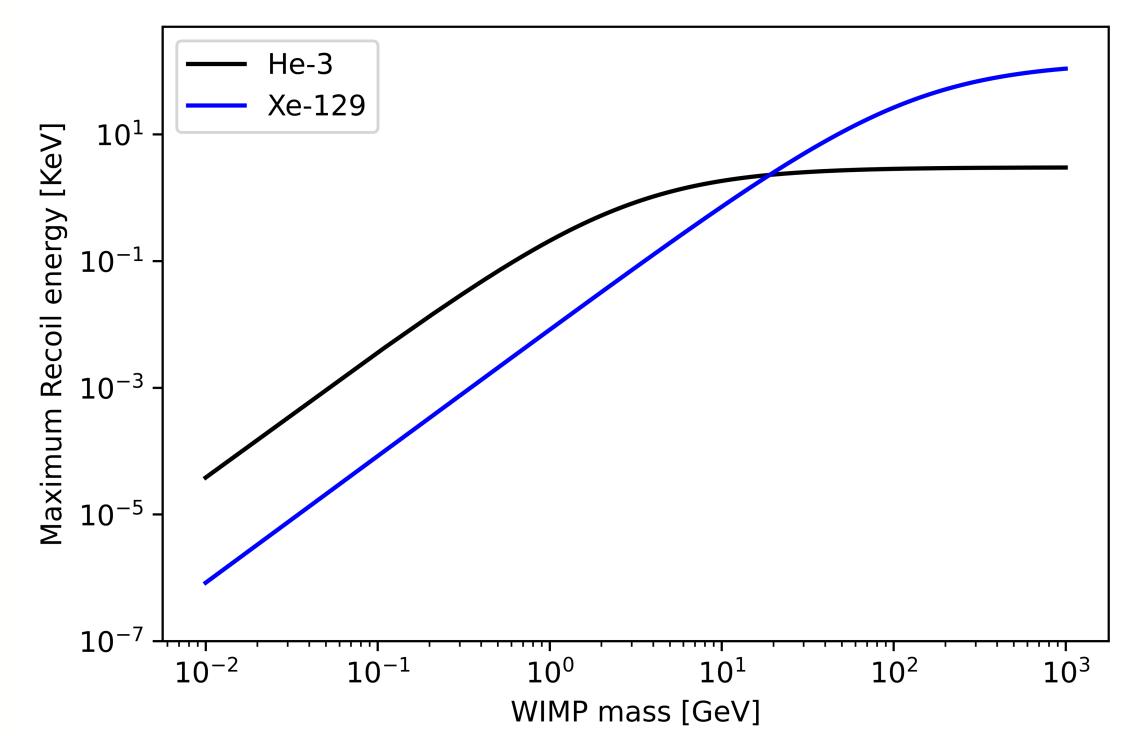
Expected background rates in ROI (0-10 keV) per kg and per cell (0.033 g) assuming 90% CR veto



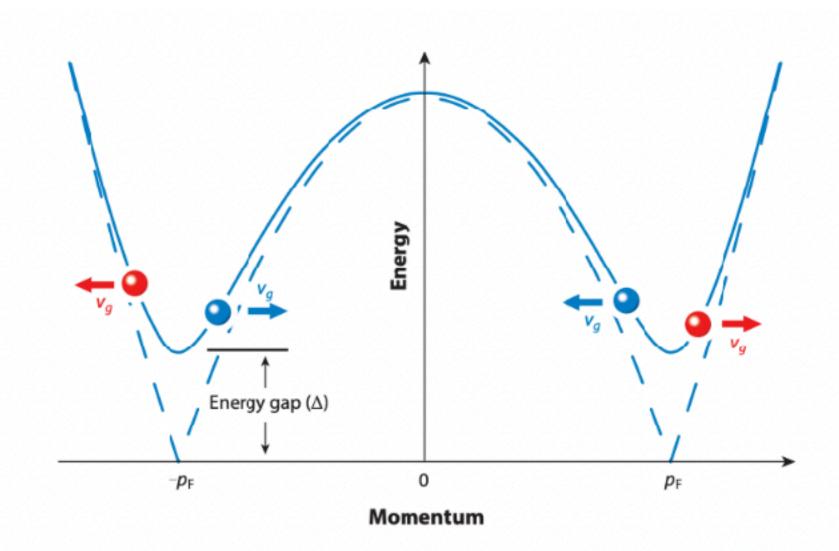
# Recoil Energies

- Lower mass nucleus = lower maximum recoil energies with lighter DM  $\bullet$
- Superfluids able to reach low recoil energy/low mass parameter space lacksquare





# Andreev Scattering



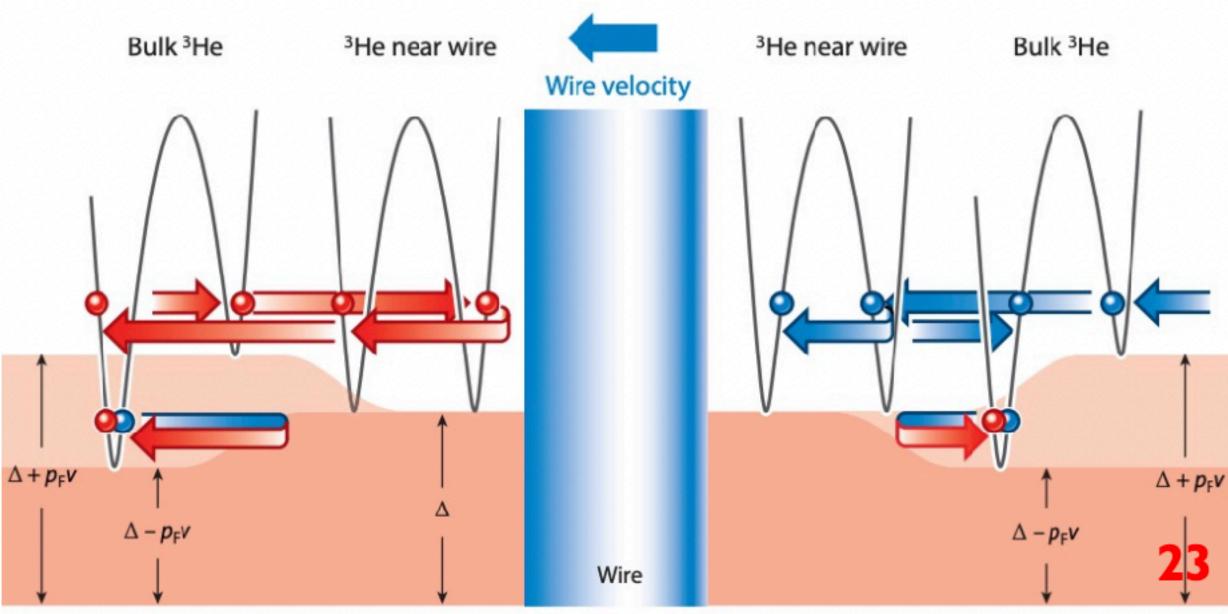
- momentum.

- Fluid flow and relative motion of an object ٠ can increase/decrease the gap.
- Only quasiparticles from in front and ٠ quasiholes from behind can transfer momentum [2pF], increasing the damping.

Ref https://www.annualreviews.org/doi/pdf/10.1146/annurev-conmatphys-031016-025411

Quasiparticle dispersion curve, with energy minima at the Fermi

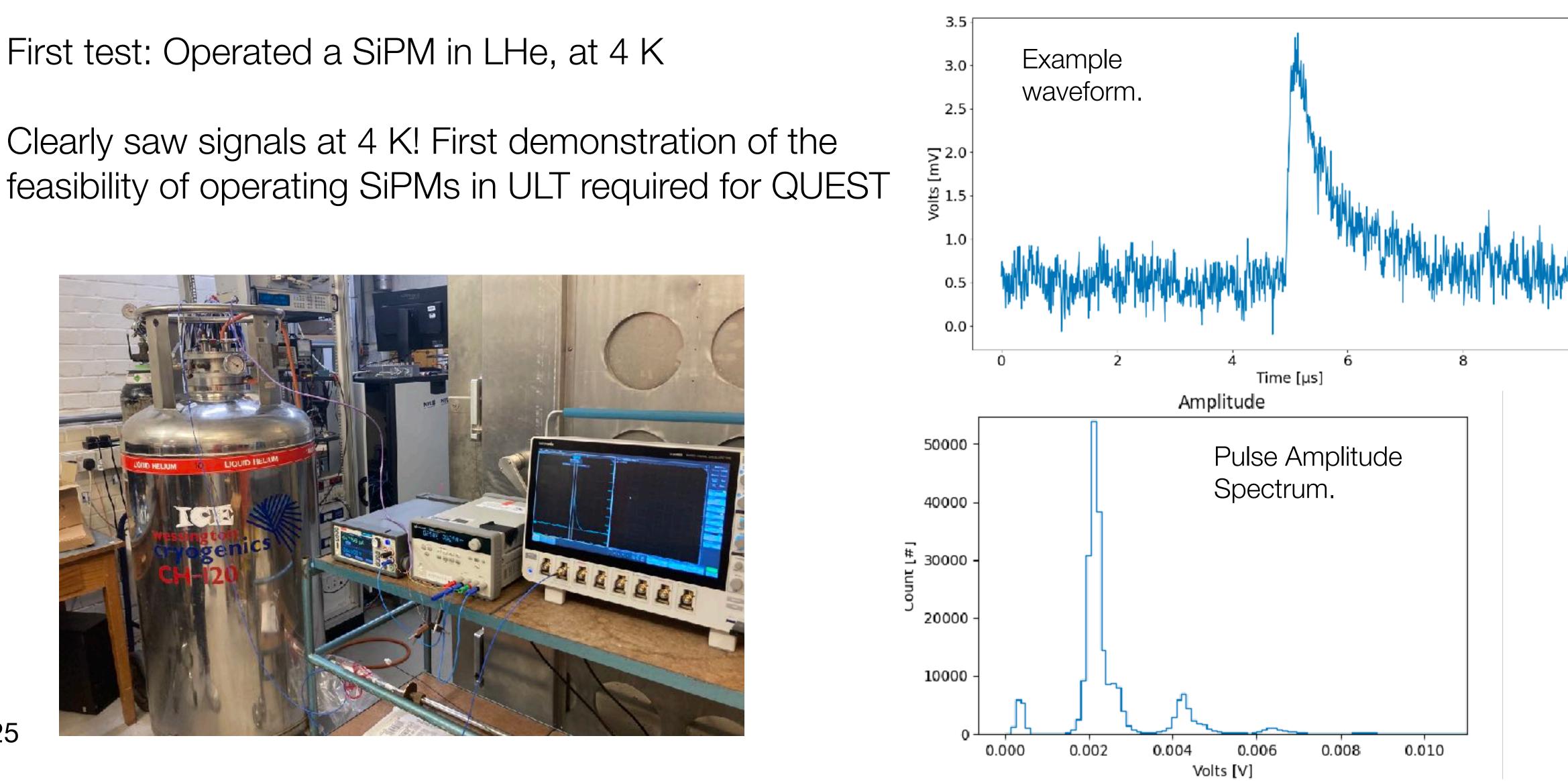
Group velocity (slope) parallel to momentum for particles and antiparallel for holes. Becomes zero at pF, so in some scattering process particle drops to min then moves up other side of curve as a hole, with velocity reversed but momentum same.





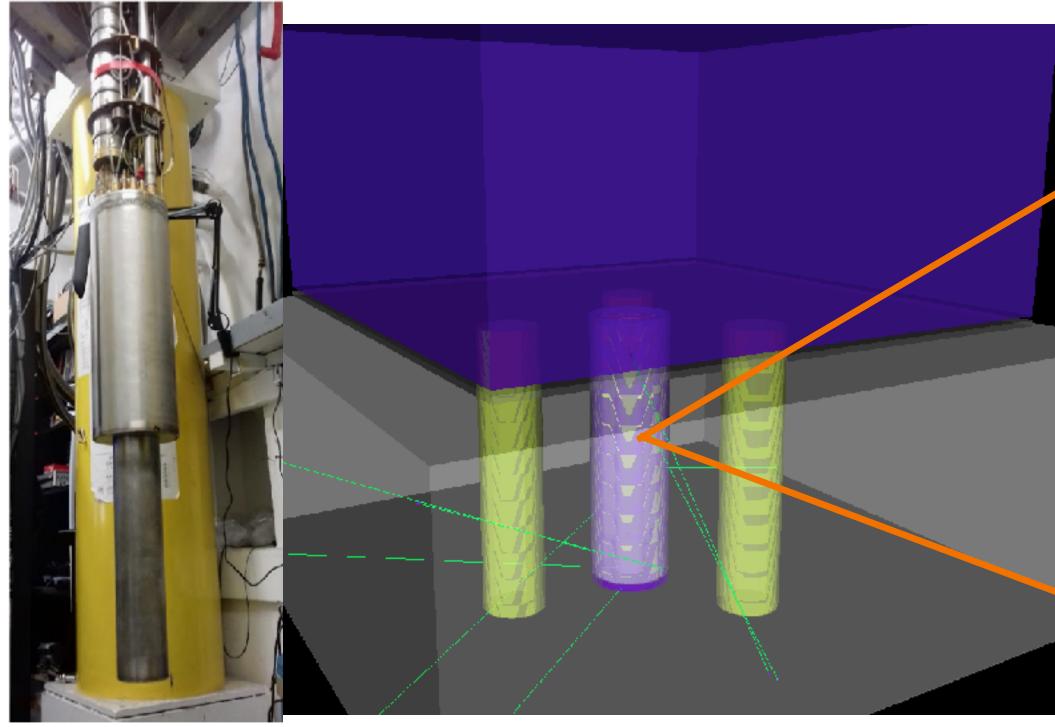
### **Detector - Scintillation**

### First test: Operated a SiPM in LHe, at 4 K





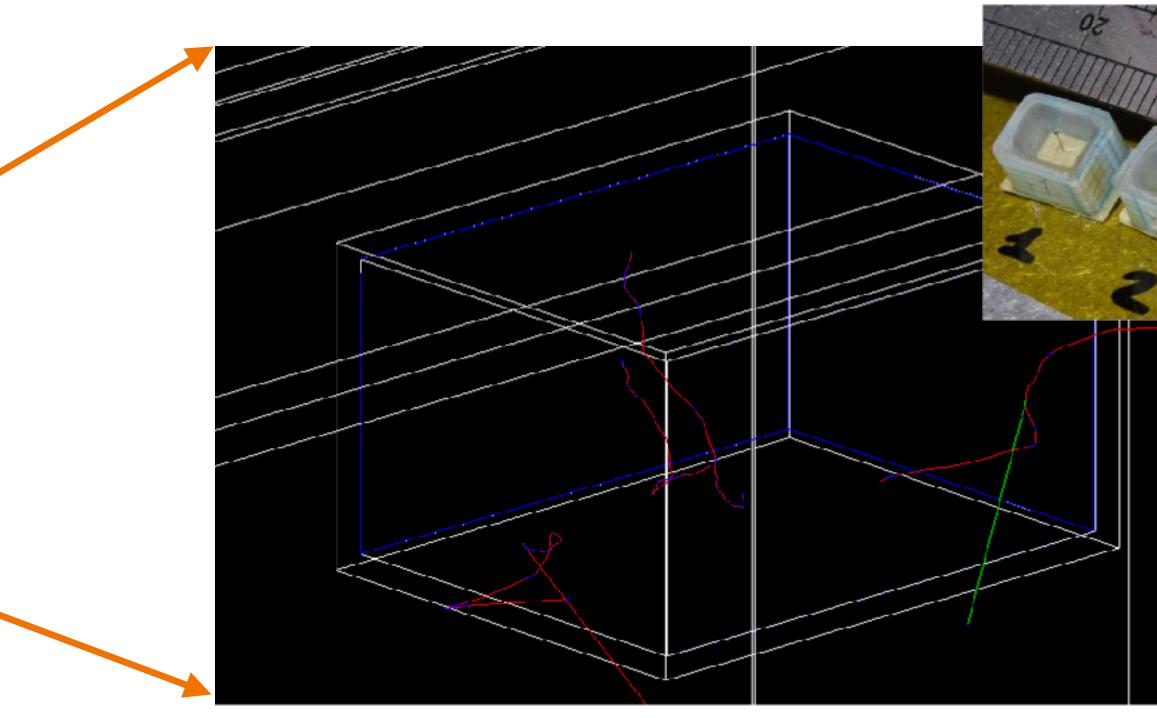
# G4 geometry of cryostat and surrounding materials



Credit: P Franchini

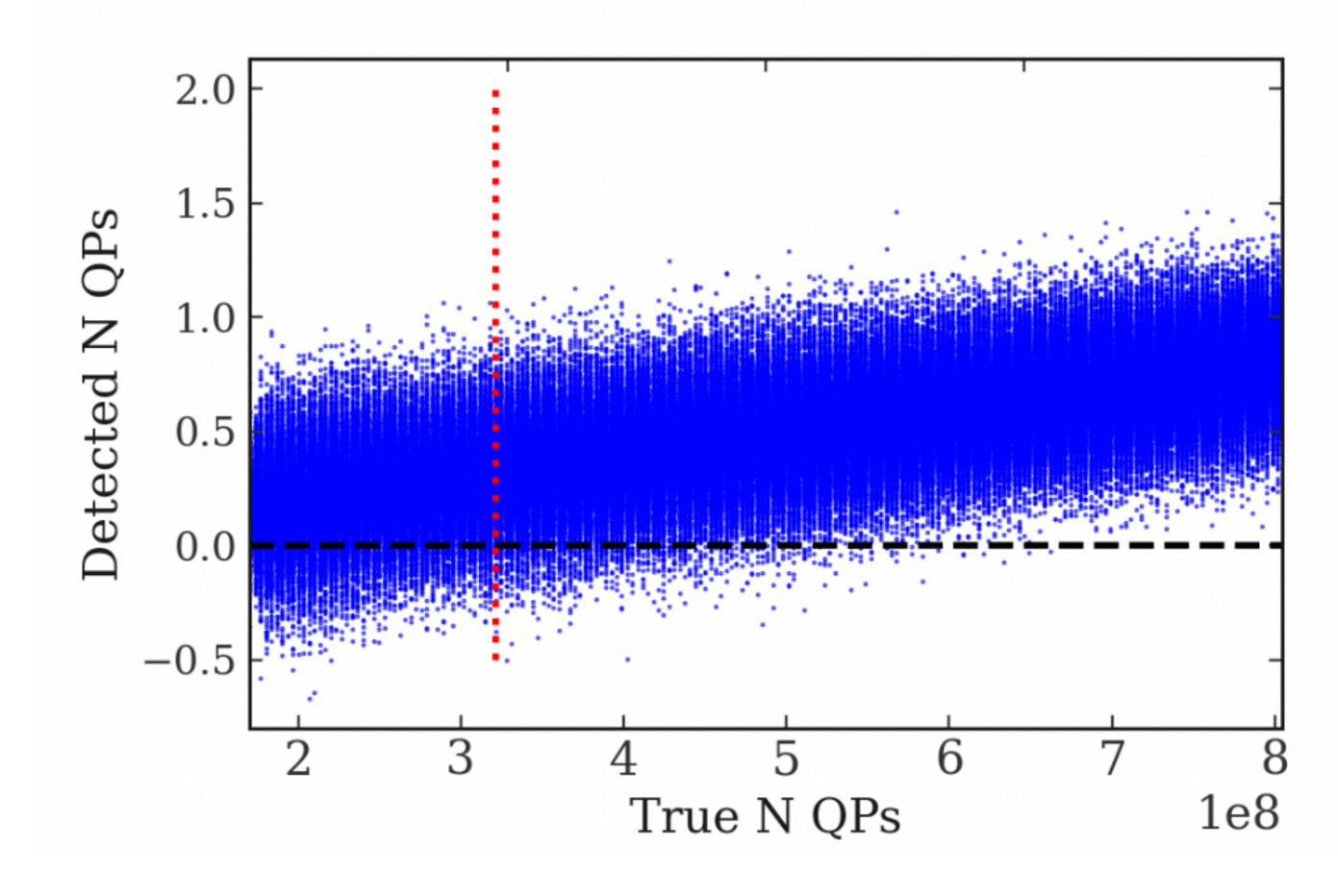
# MC event rates produced within cell of <sup>3</sup>He

Normalise with reference values to produce expected bg event rates





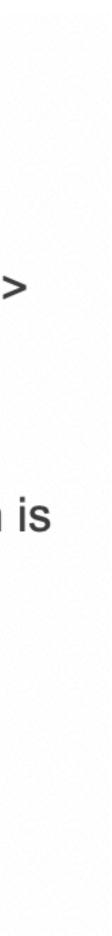
## Energy Threshold



Resolution at threshold – 95% confidence energy > zero.

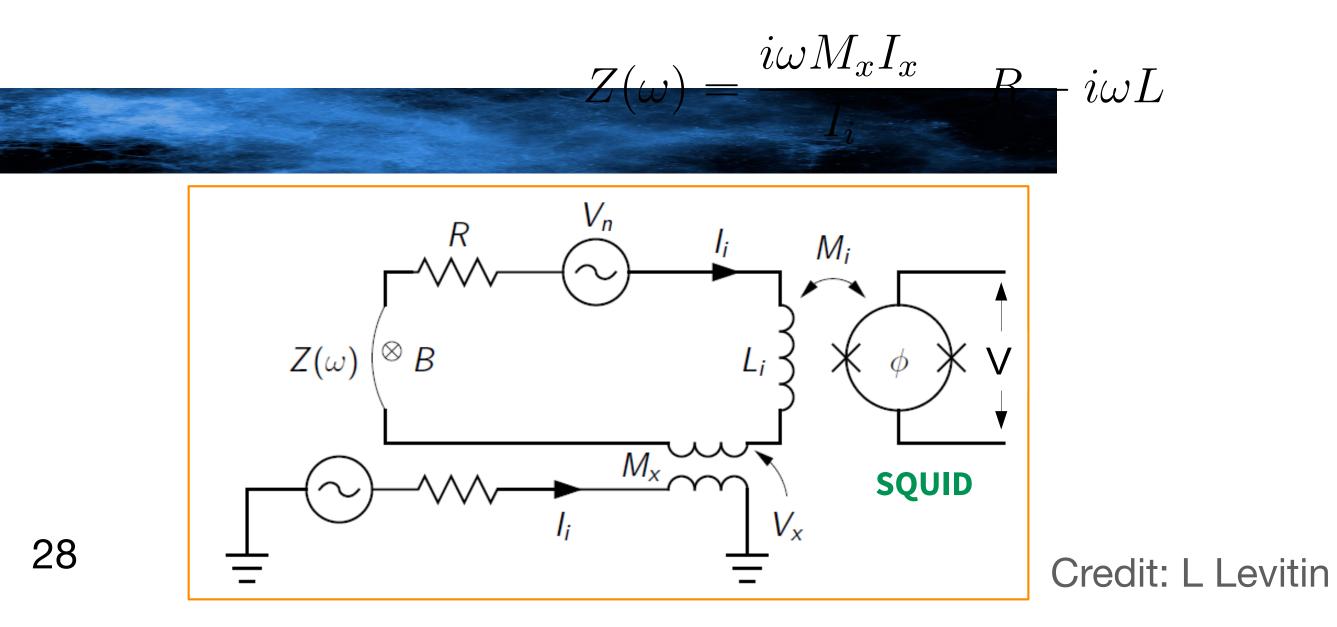
- Conventional readout: 39 eV
- Squid readout reduces noise, so resolution is dominated by shot noise.
- Squid readout: 0.71 eV

Credit: E Leason

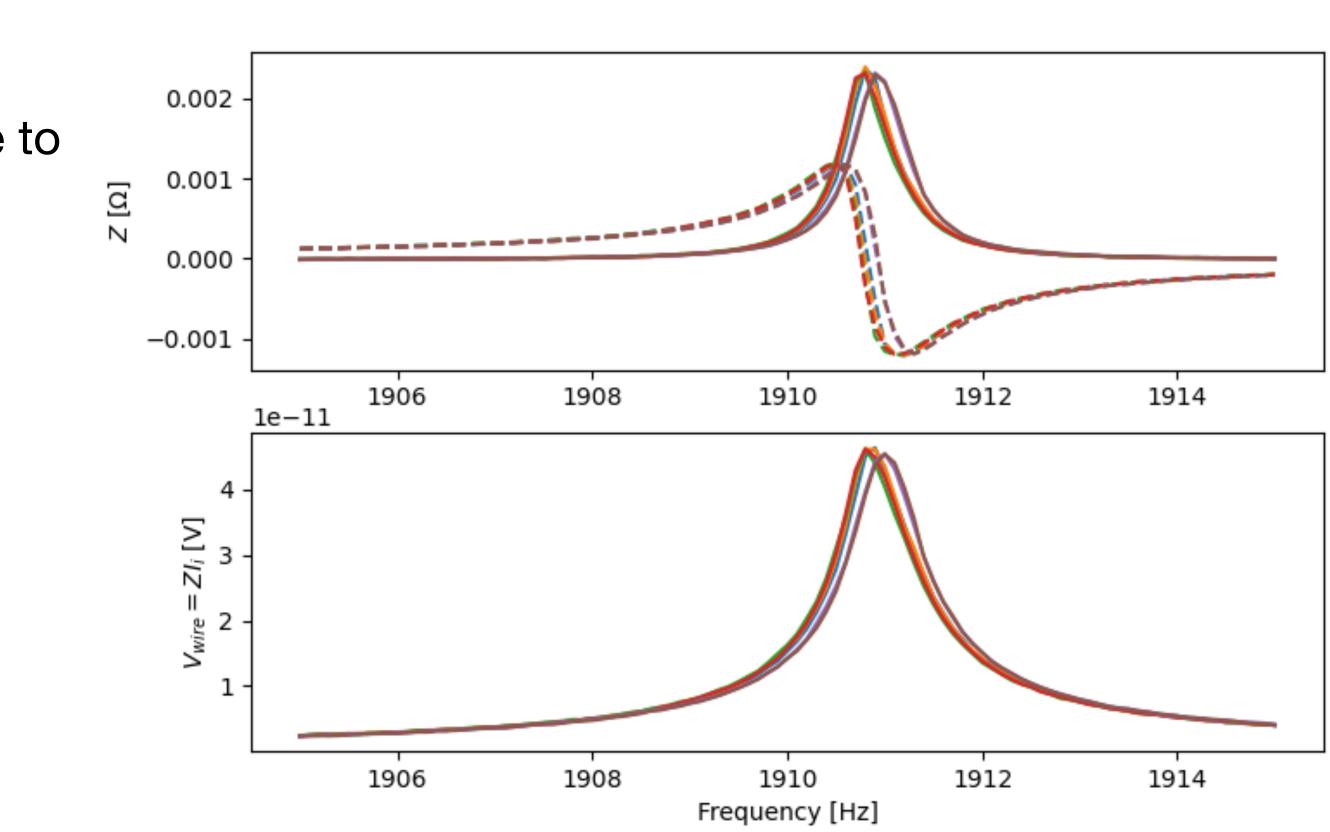


# DC SQUID

- Quantum interference between the junctions leads to  $\bullet$ extreme sensitivity to changes in magnetic flux
- Detected as change in IV characteristics sensitive to 1/2 integer values of applied flux
- Phase sensitive measurement of Ii/Ix is used to  $\bullet$ extract impedance of the wire







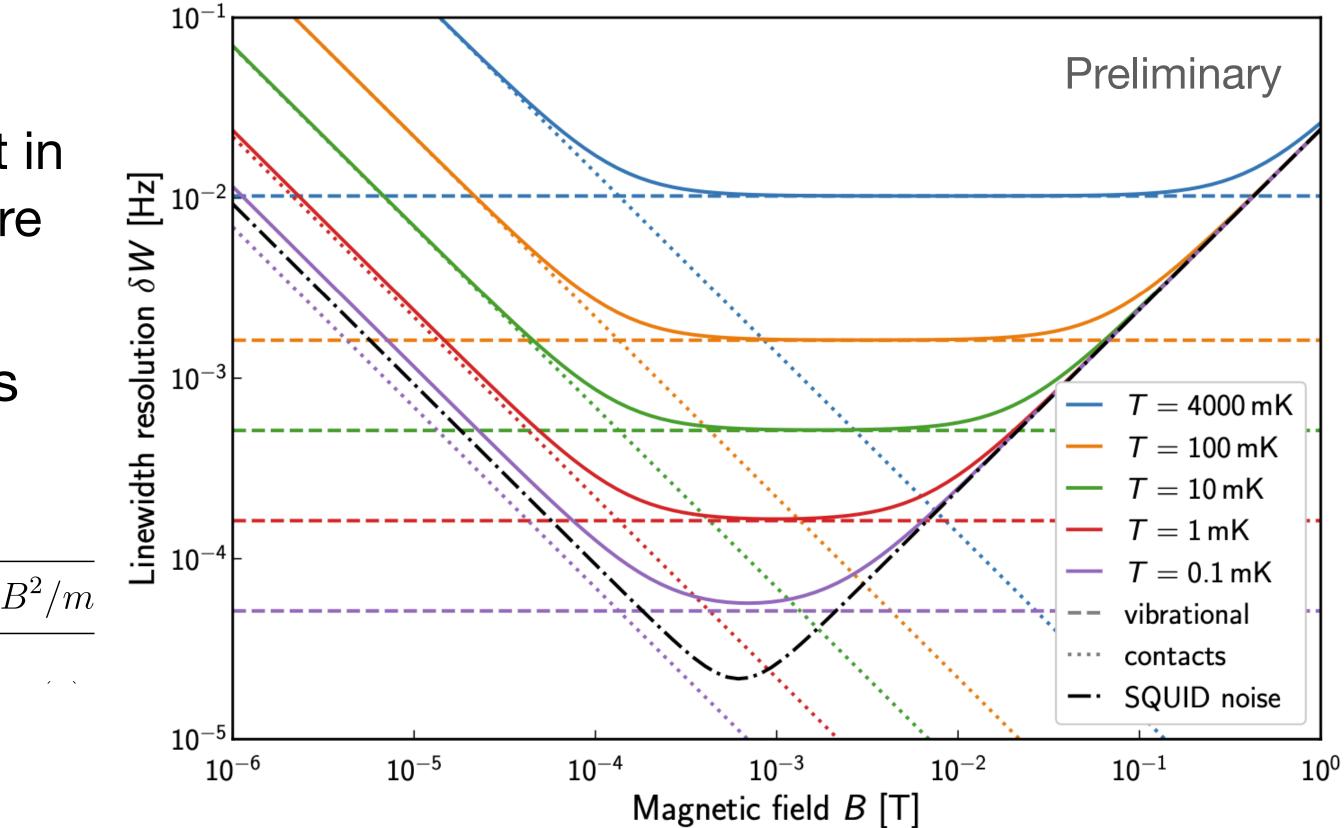
Frequency sweeps of 315nm wire taken at 4.2K

# SQUID Resolution Model

# Expected SQUID noise in width measurement in ${}^{3}$ He bolometer at $130\mu$ K and 5 bar of pressure

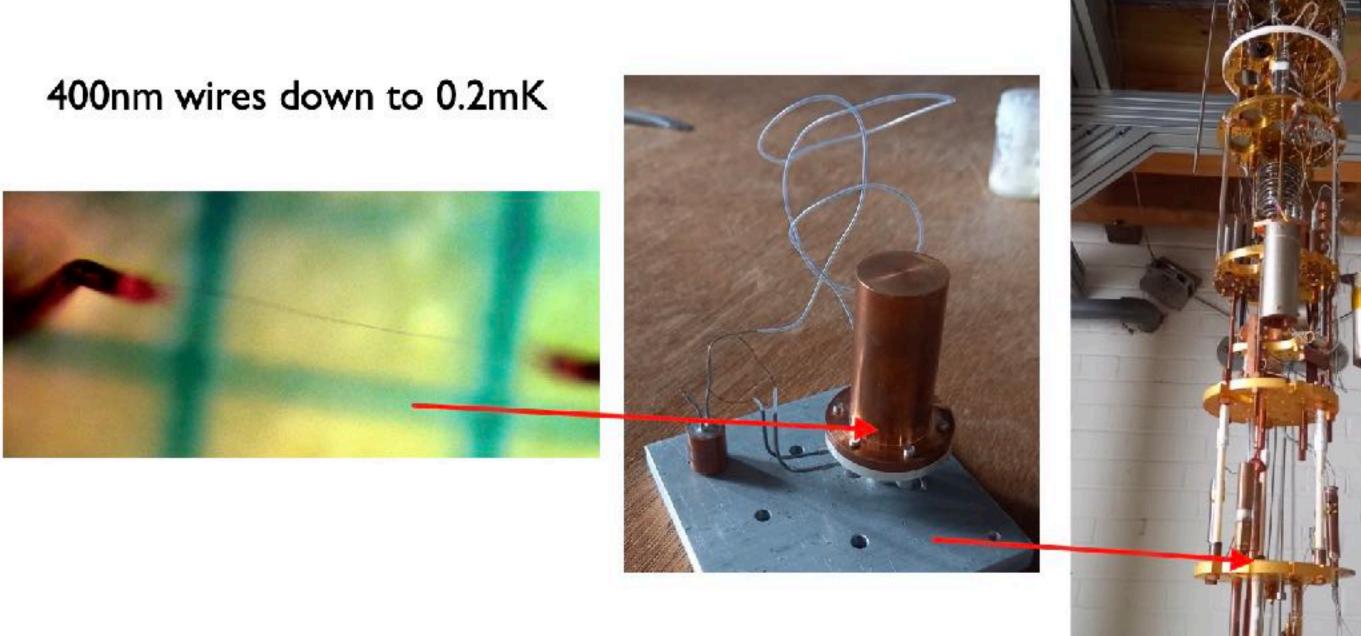
# Ratio between sum of noise contributions and maximum voltage across the wire

$$\frac{\delta W}{W} = \frac{\sqrt{\left|Z(\omega_0) + R + i\omega(1 - \alpha^2\eta)L_i\right|^2 S_\phi \Delta f / M_i^2 + 4k_B T R \Delta f + k_B T l}}{V_v^{\text{max}}}$$

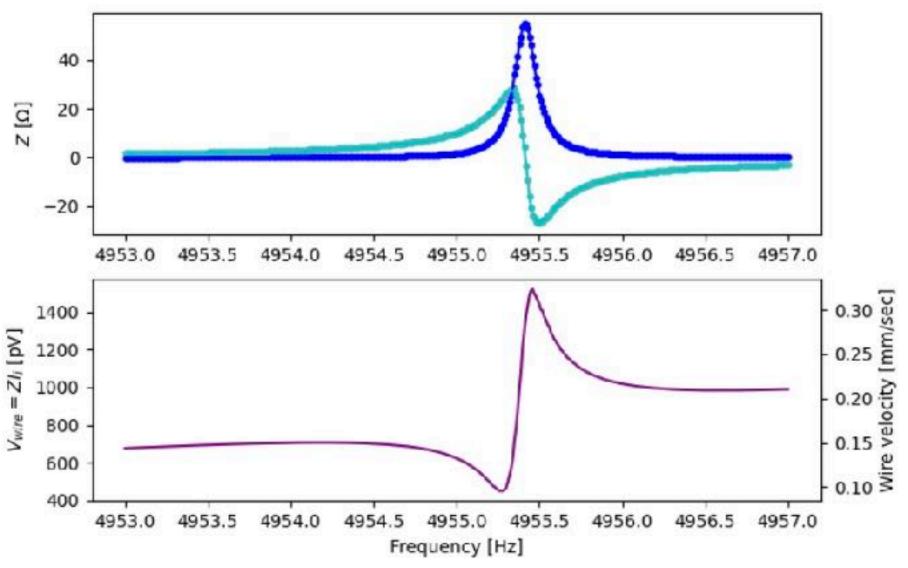


Credit: L Levitin

## SQUID Readout



### SQUID readout of VWR operated at sub-mK scale

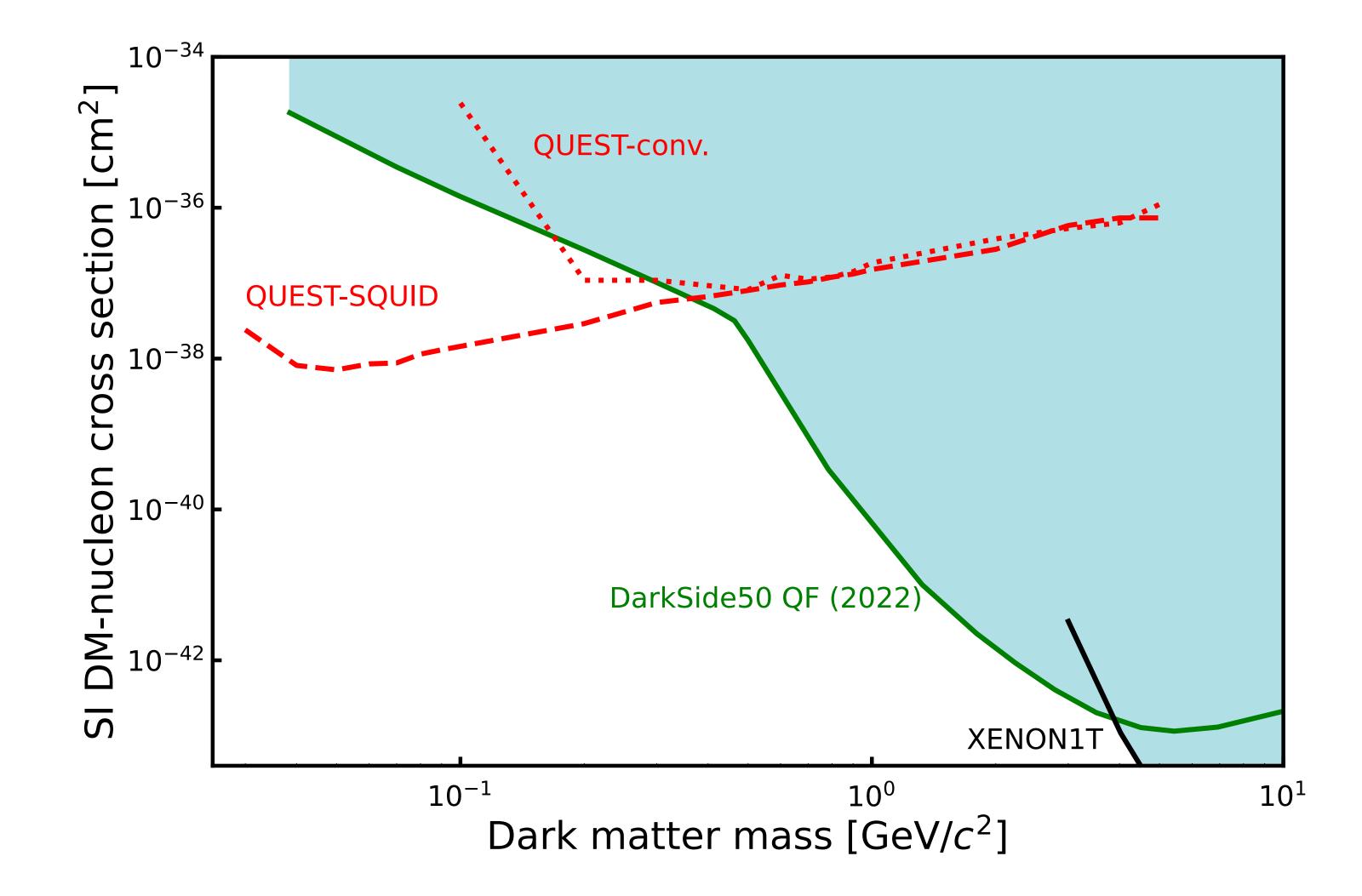


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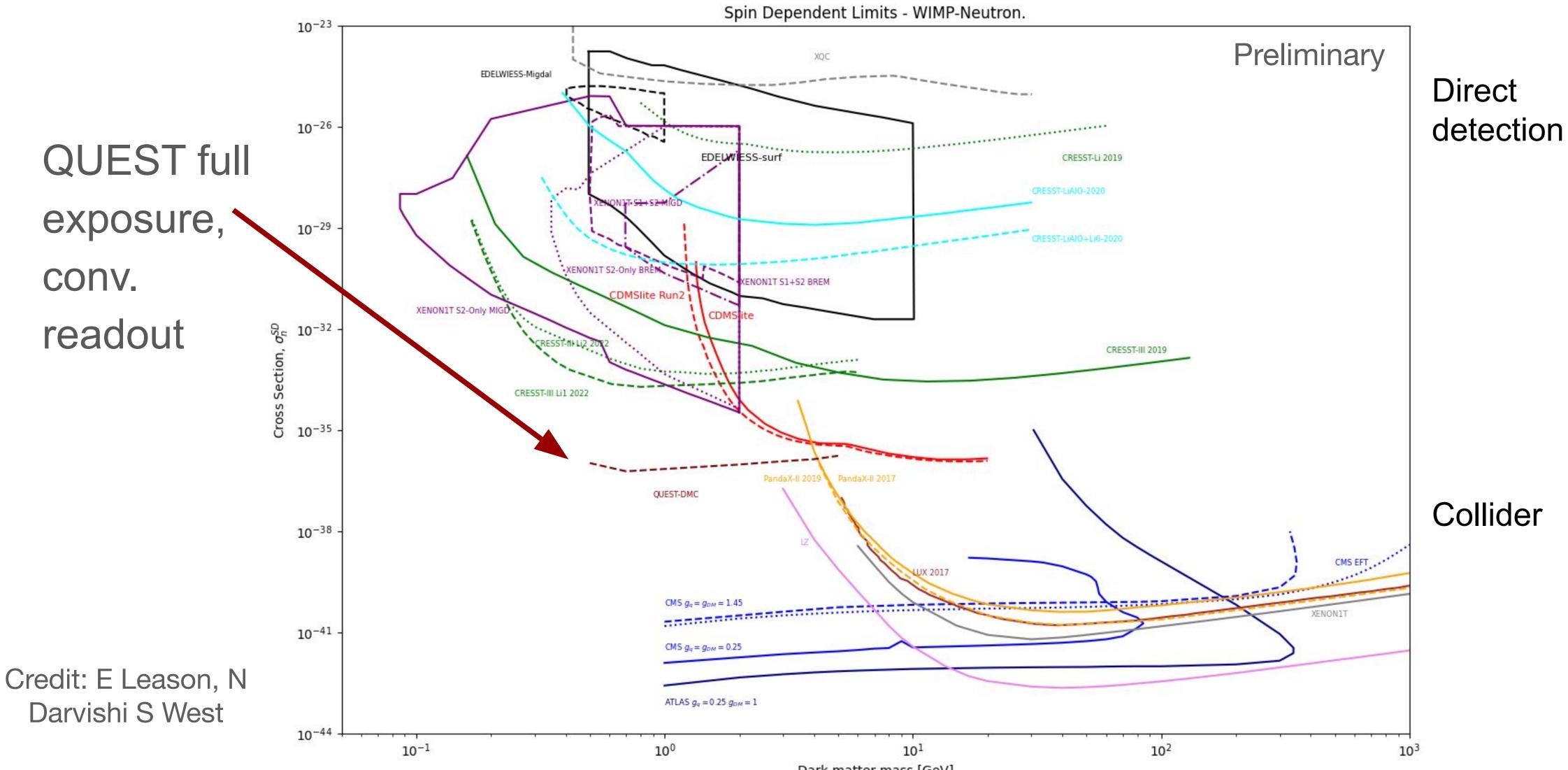


# Spin-Independent Sensitivity



Credit: E Leason, N Darvishi S West

# Sensitivity Wider Context





Dark matter mass [GeV]