







# SoLAr detector prototyping

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### The SoLAr detector



#### Challenges

- Achieve an excellent energy resolution
- Low-energy background mitigation strategy
- Neutrino flavour tagging
- Identify neutrino direction (angular resolution)
- Calibration at MeV energies across the detector
- An efficient event reconstruction for online triggering



#### **Physics motivation**

- Detecting the Solar <sup>8</sup>B and hep neutrino fluxes via both CC and ES reactions
- Detecting Supernova neutrino bursts
- Detecting other processes in the MeV scale



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#### **Novel detector concept**

SolAr

- Pixelated readout providing true 3D reconstruction from both charge and light
- Integrated array of VUV SiPMs on anode plane
- Easily scalable for a kiloton-scale LAr-TPC
- Online localized triggering for dealing with high data rates





### Existing SoLAr prototypes



#### 1) Small scale v1 - October 2022 @ The University of Bern

- $7x7 \text{ cm}^2$  anode plane, 5 cm drift 3 stacked PCBs
- 16 Hamamatsu ceramic packaged VUV SiPMs with connector pins
- 4 LArPix v2a chips
- Observed cosmic muon tracks

#### 2) Small scale v2 - July 2023 @ The University of Bern

- 30x30 cm<sup>2</sup> anode plane, 30 cm drift single PCB
- 64 Hamamatsu SMD packaged VUV SiPMs
- 20 LArPix v2b chips slots for 64 chips
- Observed cosmic muon tracks and <sup>60</sup>Co gamma source







### SoLAr V2 prototype

- 30x30 cm<sup>2</sup> anode plane
- 64 SMD packaged VUV SiPMs
- 6840 3*x*3 mm<sup>2</sup> pixels

Solar

• 500 V/cm electric field

SMD SiPMs are V2's biggest advantage over V1

- No need for stacked PCBs to fit connector pins
- No ceramic mount distorting electric field lines and taking up pixel space





### SoLAr V2 run

#### Readout diagram

SolAr





### SoLAr V2 run

Overview

Solar

- 1. Achieved good LAr purity
- 2. Achieved very low charge hit threshold  $\sim 3.8 \text{ ke}^- \rightarrow 100 \text{ keV}$
- 3. 85.7% of the charge events found a corresponding light event match within a search window of 10  $\mu s$
- Two days of cosmic run with nominal 15 kV HV
- Special <sup>60</sup>Co source run
- Special runs with 7.5 kV and 3.75 kV HV
- Special runs with varied SiPM bias over voltage







### Data Processing

Reconstruction methods

#### **Charge tracks fitting**

#### 1. DBSCAN clustering

SolAr

- L) Cluster hits in the *xy*-plane
- 2) Determine the intervals in *z* between clusters
- 3) Generate fake data filling dead areas withing the z intervals
- ) Cluster hits + fake data in the xy-plane
- 5) Cluster hit labels from first stage with hit *z*-coordinates
- 6) Remove fake data

#### 2. RANSAC regression

- Fit line to clustered hits
- Use hit charge as weight for line fit
- Optionally re-cluster and fit outliers to find secondary tracks

dQ/dx is obtained by defining same sized stacked cylinders along the fitted tracks



#### Light "track" fitting

- 1. Obtain light signal *xy*-coordinates
  - Select top 5 SiPMs with largest light signal
  - Minimum of 3 SiPMs with non-zero signal
  - *x* and *y* are determined by SiPM's coordinates
- 2. Estimate light signal *z*-coordinate
  - Average the *z*-coordinate of the charge hits within the SiPM's quadrant
  - Use total average z-coordinate if no hits in the SiPM's quadrant
  - Average is weighted by hits charge

#### 3. RANSAC regression

- Fit line to light coordinates
- Fit quality depends strongly on angle of incidence to anode

Light fit is a coarse approximation to the charge track and is expected to be approximately parallel to it



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### Data Processing

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Single cosmic ray event

#### Integrated charge and light event display



Single track dQ/dx plot showing gap in dead area



Crosses indicate quadrants with dead chips. 1 chip = 1 quadrant with 60 pixels



#### 2D event charge vs. light Both axis in log scale







Analysis

#### ~80 minutes of cosmics – all single-track events



- Score indicates "goodness of fit". Shorter tracks are not well fit
- Fit quality doesn't affect total event light and charge





Solar

#### ~80 minutes of cosmics – all single-track events

300 - 250 integral [p.e. - Log] 10 00 Track length [mm]  $10^{2}$ Total Light 100 - 50 105 10<sup>6</sup> Total charge [e - Log]

Mean Light vs. Charge per 1mm track length step

For every 1mm step in track length, plot mean light vs. charge

• Scattering on large tracks due to low statistics



Peaks in track length correspond to peaks in light vs. charge

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### Analysis

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#### Dead area map and angle of incidence





- Peaks in track length appear to be due to dead areas in the anode plane
- Short tracks could be due to tracks that cross only a corner of the live area







### Analysis Single tracks fully crossing the TPC



• Mean  $dQ/dx \approx 4.6 \text{ ke/mm}$ 

Solar

• Measured electron lifetime from cathode/anode crossers: 1.828ms



2D event charge vs. light

Both axis in log scale





### Analysis

Solar

#### Track reconstruction from light

- A cut at 30 mm eliminates short tracks, likely confined to a single SiPM quadrant.
- The light fit depends strongly on the angle of incidence of the track to the anode.
  - Luckily most tracks have a shallow angle to the anode
- Light "tracks" have lower granularity than charge tracks but result in a similar direction
- Increasing the light threshold removes events without improving the cosine distribution



#### The direction of the light fits align well with their charge track counterparts!



### Proposed SoLAr prototypes

#### Small scale prototype

• Custom-made SiPMs with charge pads mounted on top of photosensitive cell



Solar



 Test of alternative readout chips when available, i.e.
LightPix
Q-Pix

#### Mid scale Demonstrator

- 2025-2028 prospect at Boulby Underground Laboratory
- Few-ton scale LAr detector underground
  - $\circ$  1100 m overburden
- 1.6 x 2.6 x 2 m<sup>3</sup>
  - 1 m drift distance
- 30x30 cm<sup>2</sup> readout anode tiles
  ≈ 6400 pixels per tile
- First measurement of flavor tagged solar neutrinos in LAr

#### Boulby Underground Laboratory





### Summary

- $\,\circ\,$  SoLAr is the first integrated light and charge detector
- $\,\circ\,$  V2 prototype produced essential data to support the concept
- V2 data analysis is still ongoing but progressing fast
  - Most charge events are correctly matched to their light counterparts
  - Reconstructing events required refined methods due to large areas without charge pixel coverage
  - Track reconstruction from charge is performing remarkably well
  - Light "tracks" are in reasonable agreement with their respective charge counterparts
- SoLAr is a rapidly growing project with much more to come!





# Thank you!

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### Backup Slides

SoLAr



### SoLAr V1 prototype

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### ArgonCube cryostat @LHEP - University of Bern

Solar





#### V2 run used the fully instrumented and automated ArgonCube cryostat, which can cool and filter the Argon!



## Light readout diagram

SolAr







# Charge readout diagram





Solar

LArPix chips form a "Hydra" network

- Only root chips communicate with PACMAN
- Root chips must not be neighbors
- Remaining chips connect through one or more neighbor chips
- All connected chips have a single root chip



### FAQ

Solar

- DBSCAN parameters:
  - xy\_epsilon = 8 mm (2 pixels)
  - Z\_epsilon = 8 mm
- RANSAC parameters:
  - Residual\_threshold = 6
  - Max\_trials = 1000
- RANSAC has random component and may not always return the same fit
- Cylinder parameters are determined dynamically but bound by DBSCAN parameters



### Cylinder parameters

dh estimation

SoLAr



 $dl \hat{u} = [xy_epsilon (\hat{x} + \hat{y}) + z_epsilon \hat{z}]$ 

 $dh = dl (\hat{u} \cdot \hat{v})$ 



### Cylinder parameters

dr estimation

SoLAr





Minimum dr is bound to  $\frac{dl}{4}$ 





### Light to charge z-coordinate estimation



