

PHYSICS OF A FUTURE MUON COLLIDER

Aspen Center for Physics March 28, 2024









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An important caveat...



The proposals are for **R&D** to see if the technology is feasible not to start construction









A muon collider is **not** at the same level of development as the HL-LHC or an e^+e^- machine

In an ideal world, multiple colliders could exist!









Why do we need another collider?





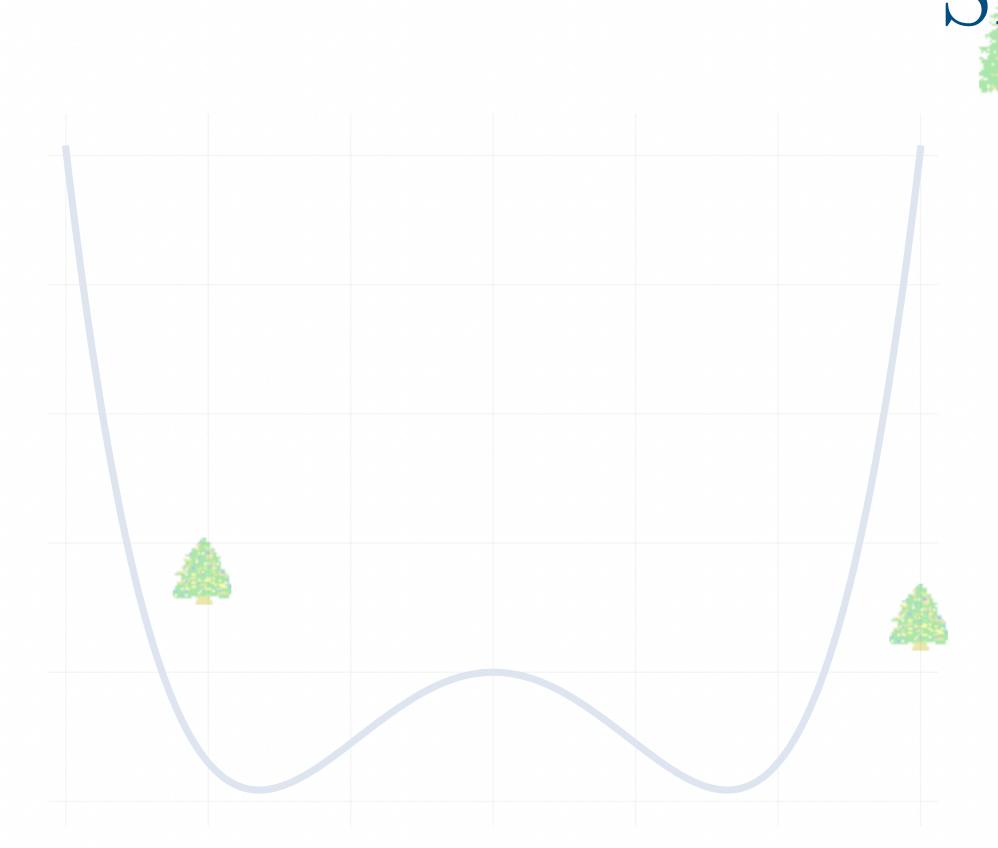








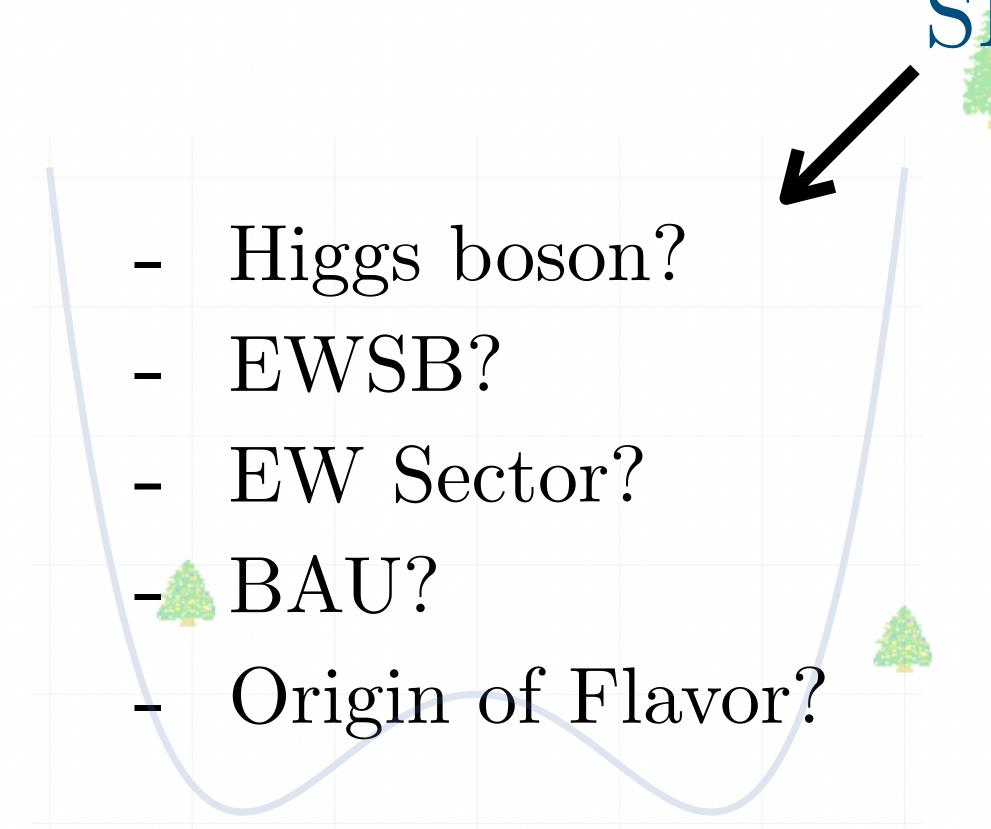






Many open questions remain in the SM & beyond

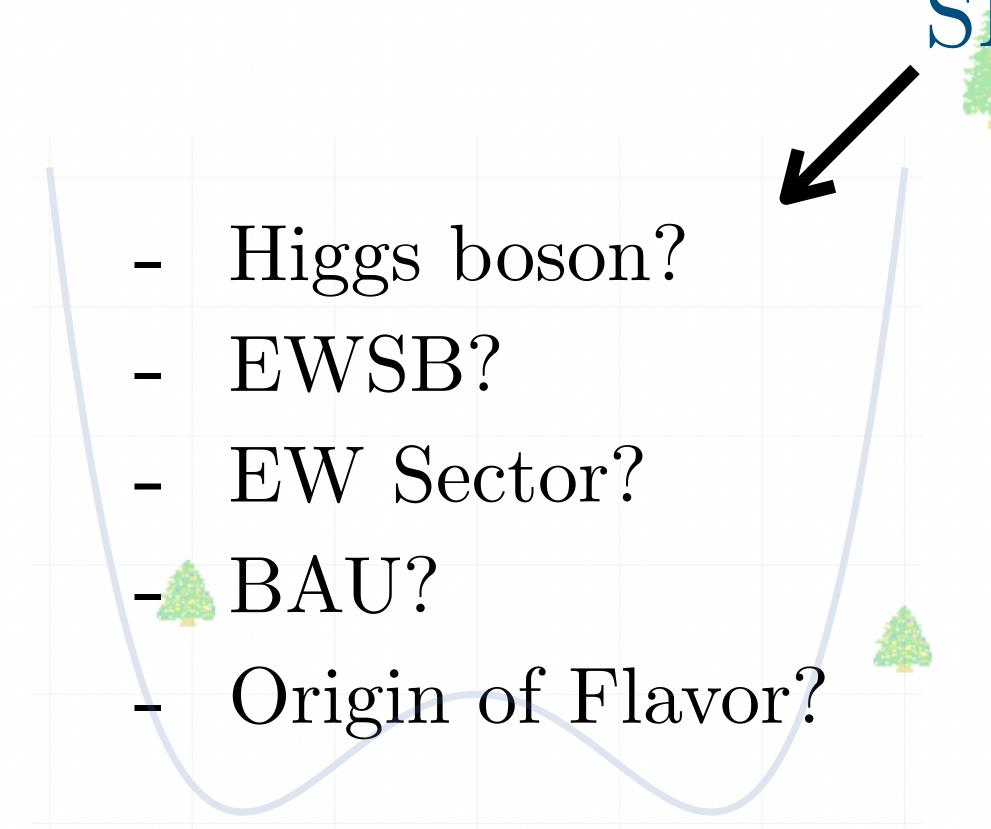






Many open questions remain in the SM & beyond



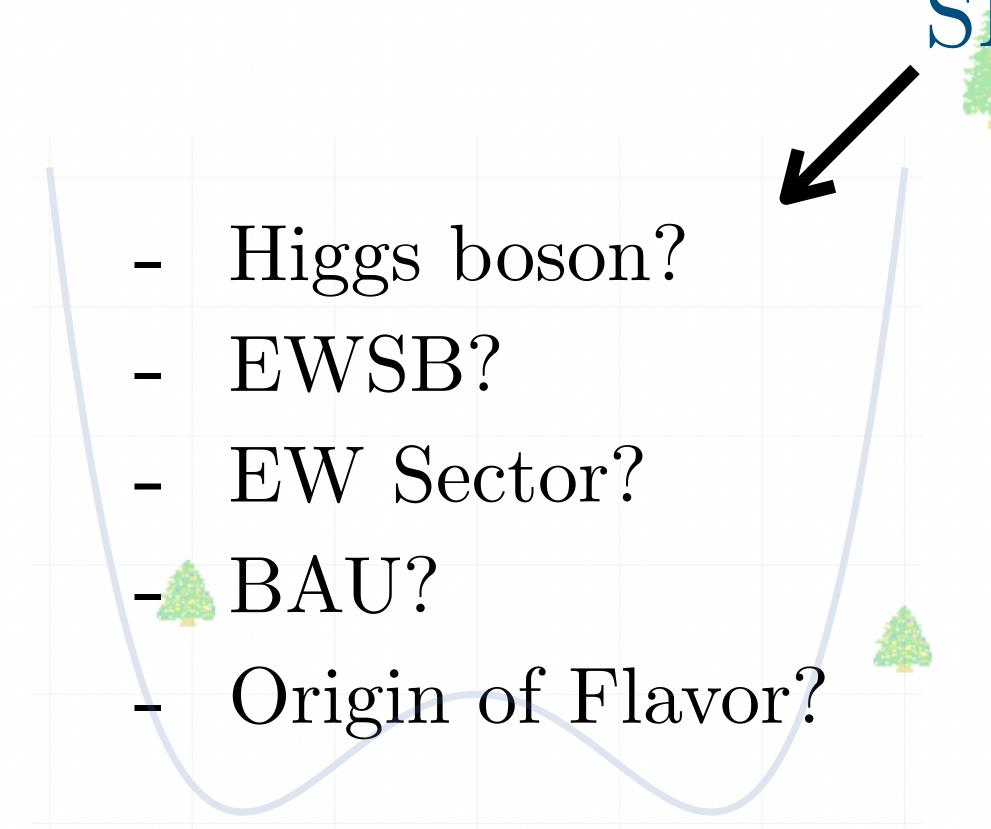




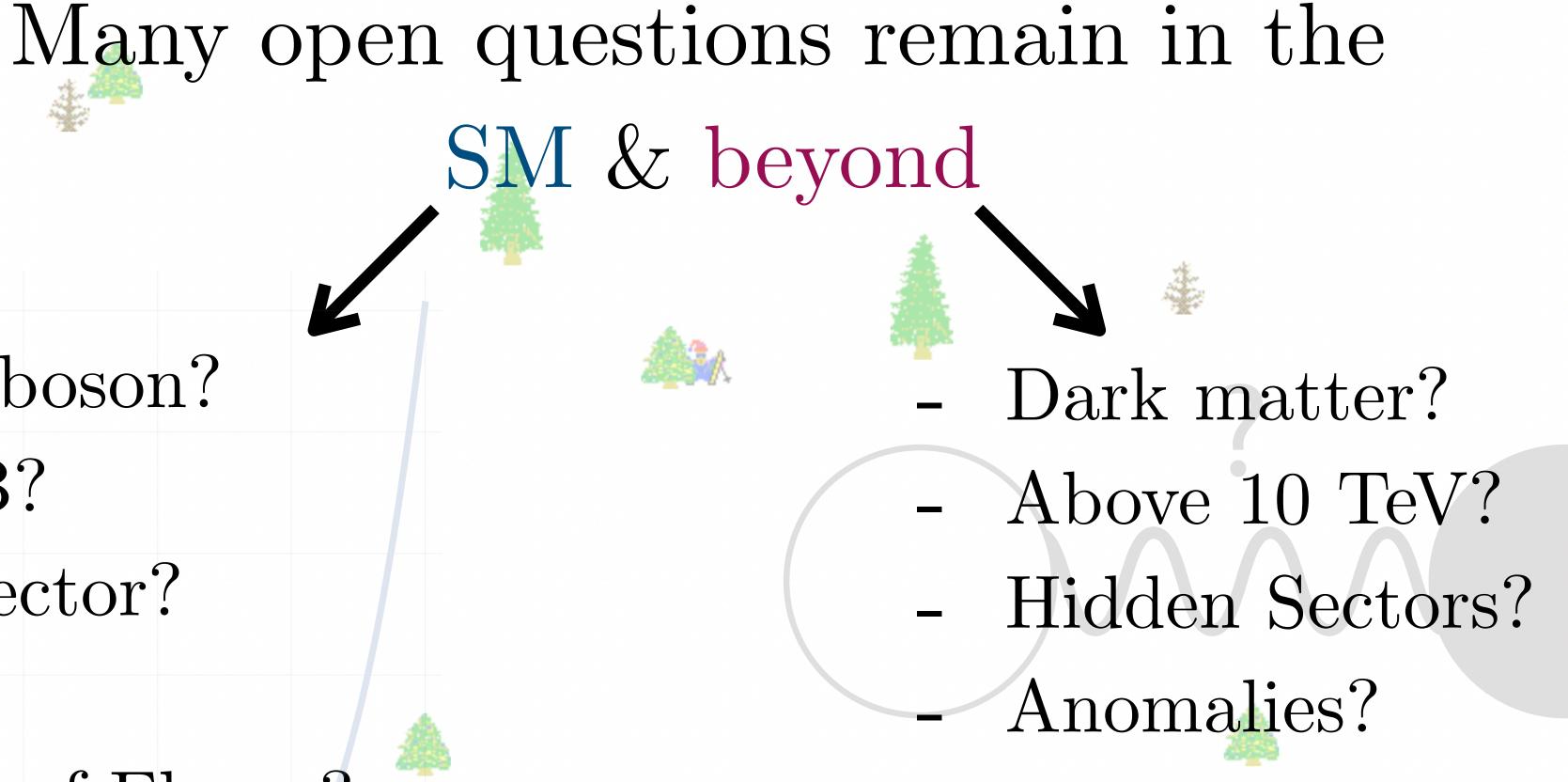


Many open questions remain in the SM & beyond - Dark matter? - Above 10 TeV? - Hidden Sectors? Anomalies?









Need empirical input





PRECISION







WHY ARE MUC WORTH THE INVESTMENT We need an experimental program that does the most "new stuff"

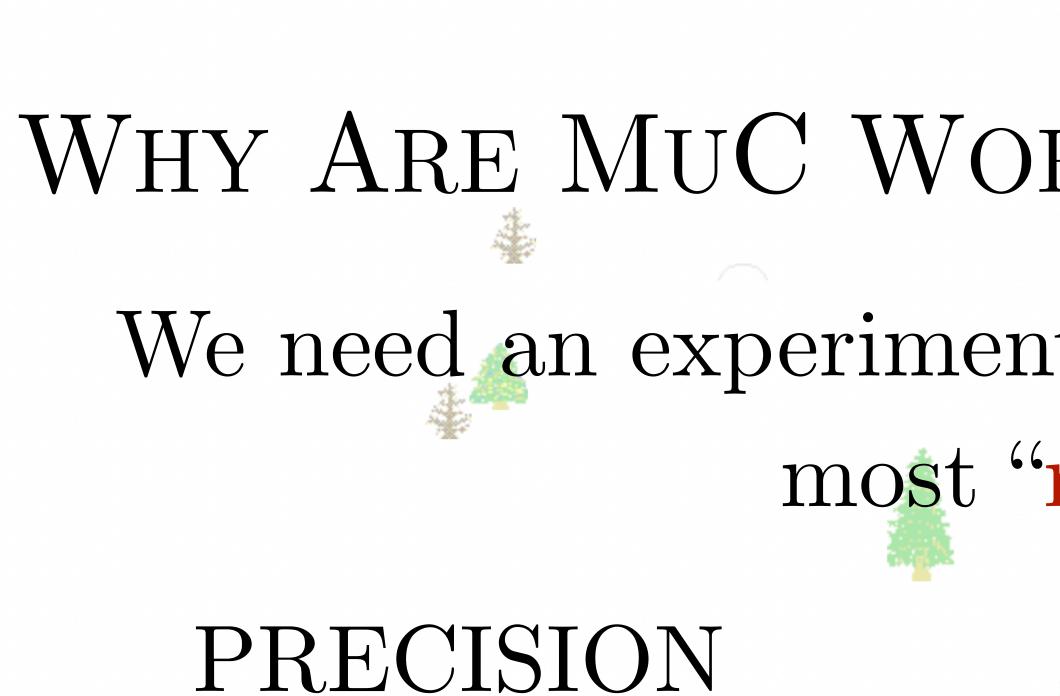












Clean Signatures **High Statistics**







 e^+e^-



WHY ARE MUC WORTH THE INVESTMENT We need an experimental program that does the most "new stuff" DISCOVERY **High Energy On-shell Production** Origin

hh

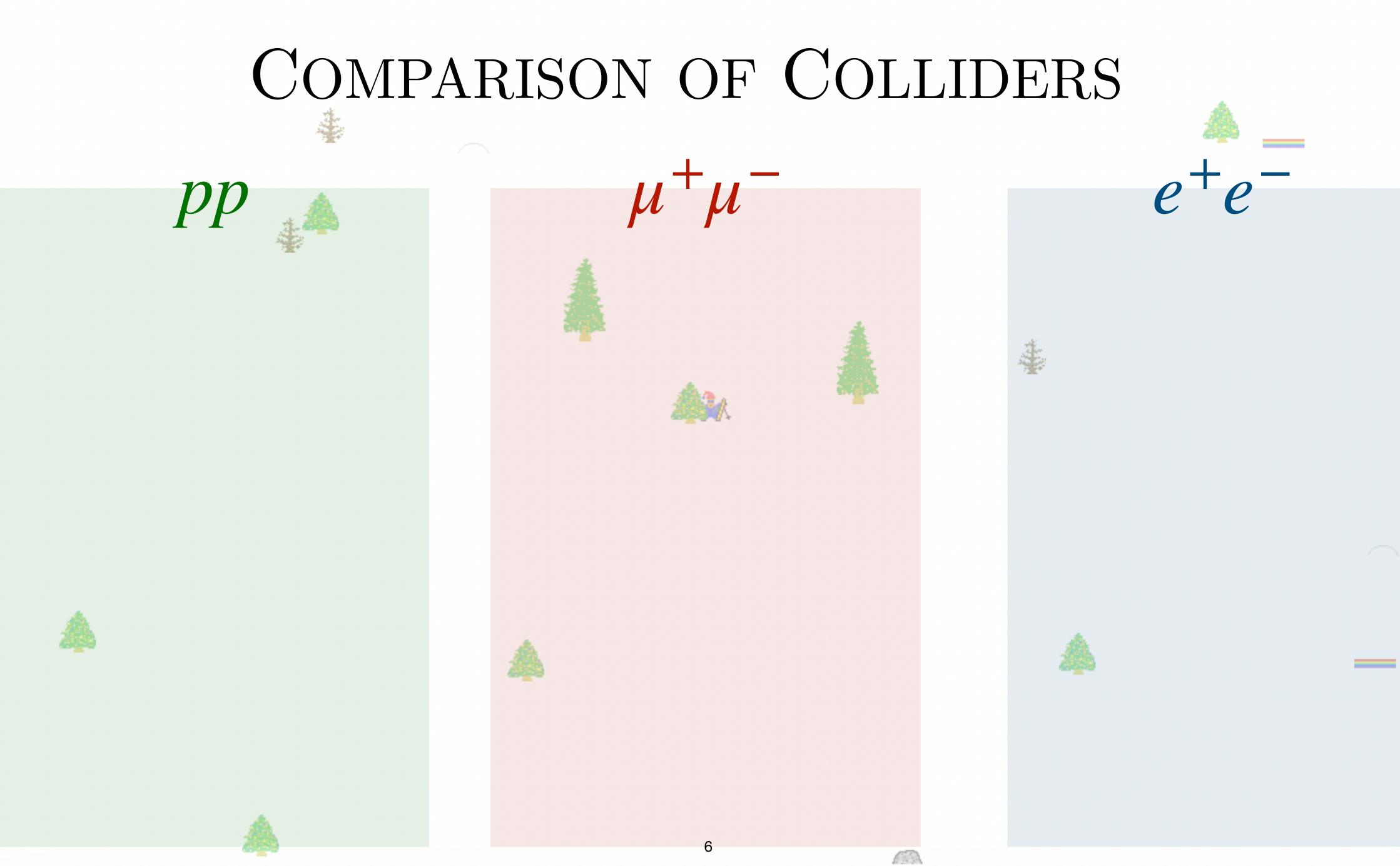




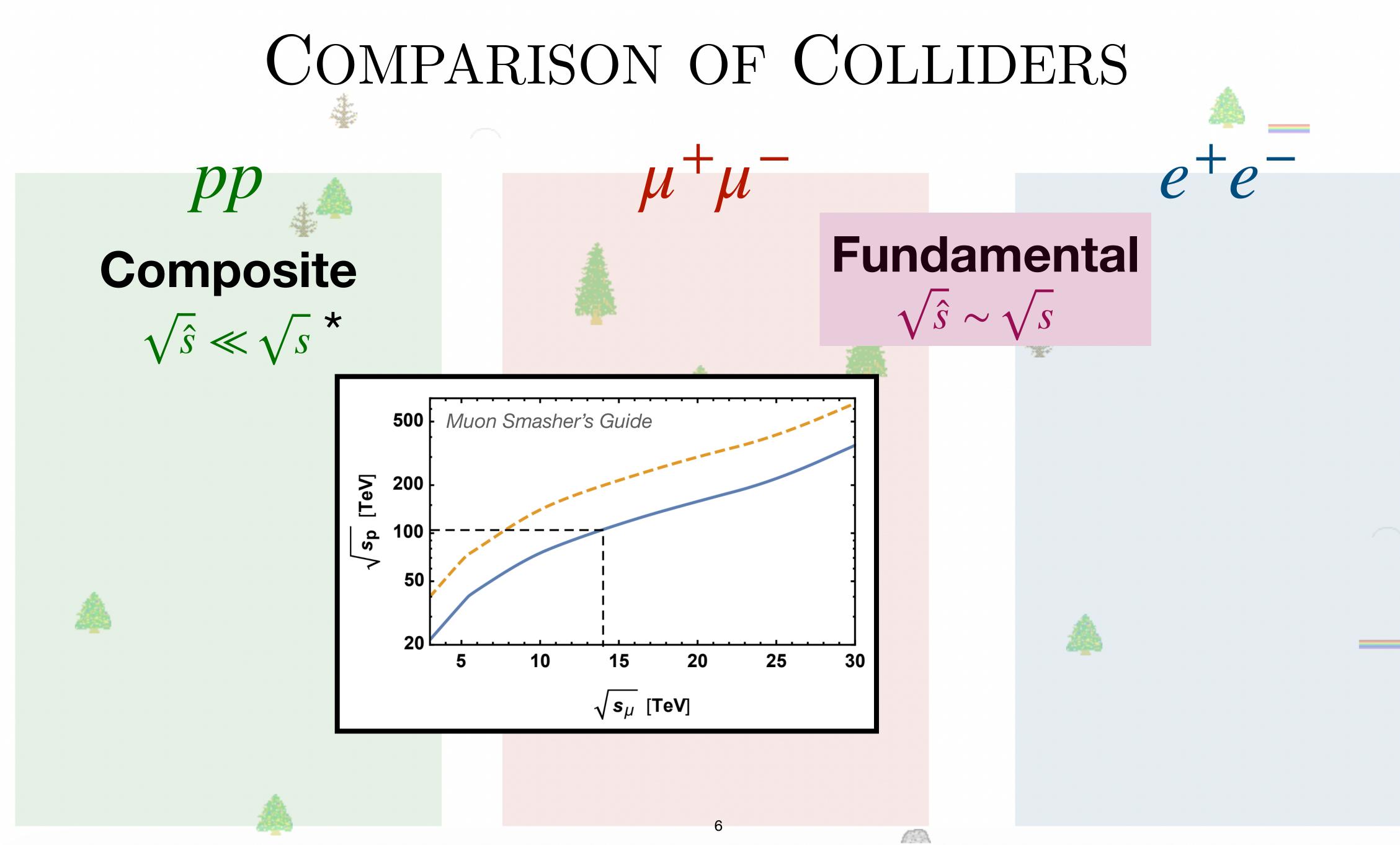
WHY ARE MUC WORTH THE INVESTMENT We need an experimental program that does the most "new stuff" DISCOVERY High Energy MuC **On-shell Production**

1 1 MuC can be a **precision** and **discovery** machine



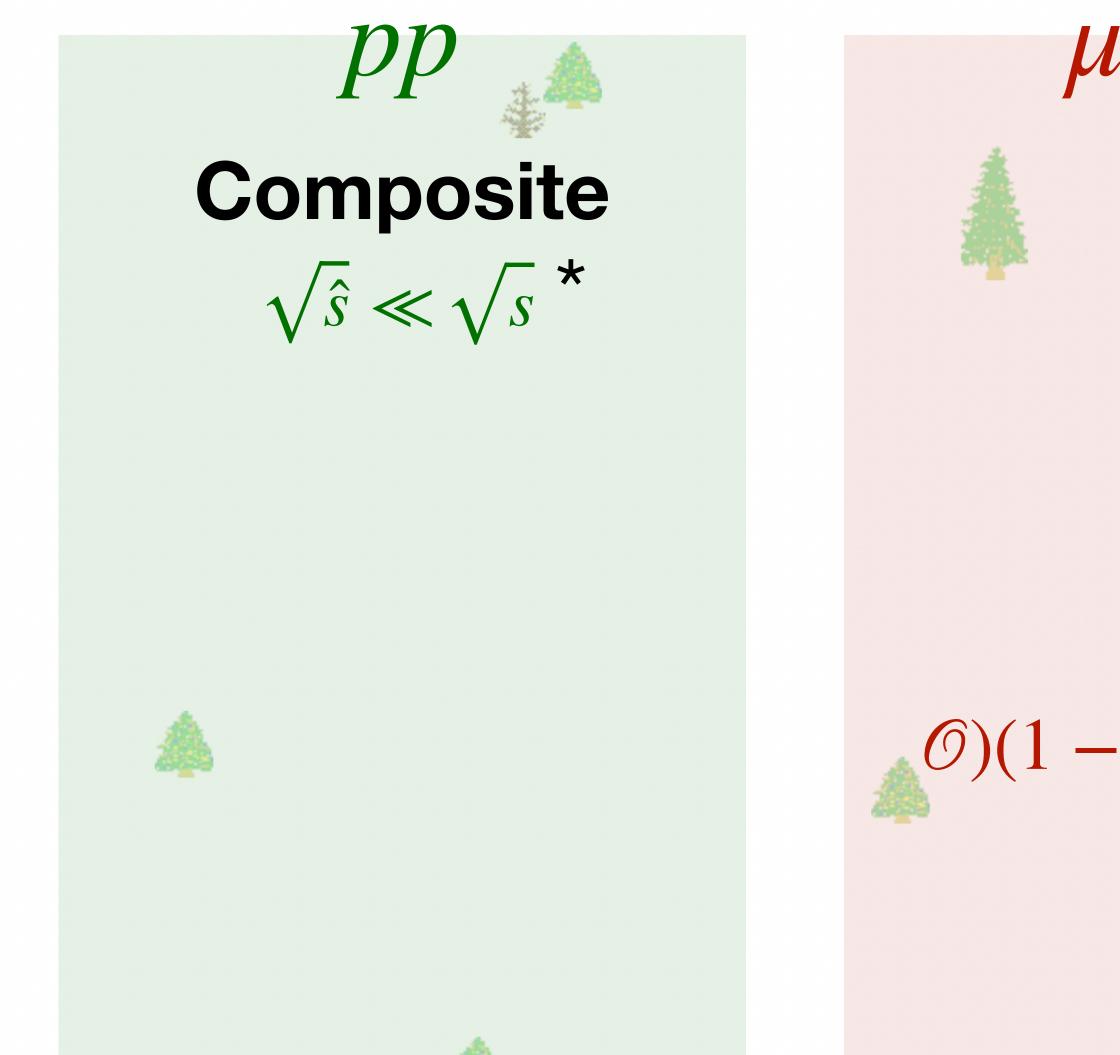








COMPARISON OF COLLIDERS $\mu^+\mu$ **Fundamental** $\sqrt{\hat{s}} \sim \sqrt{s}$ $P \propto \gamma^{4} = \left(\frac{E}{m}\right)^{4}$ $P_{\mu}/P_{e} \sim 10^{-9}$

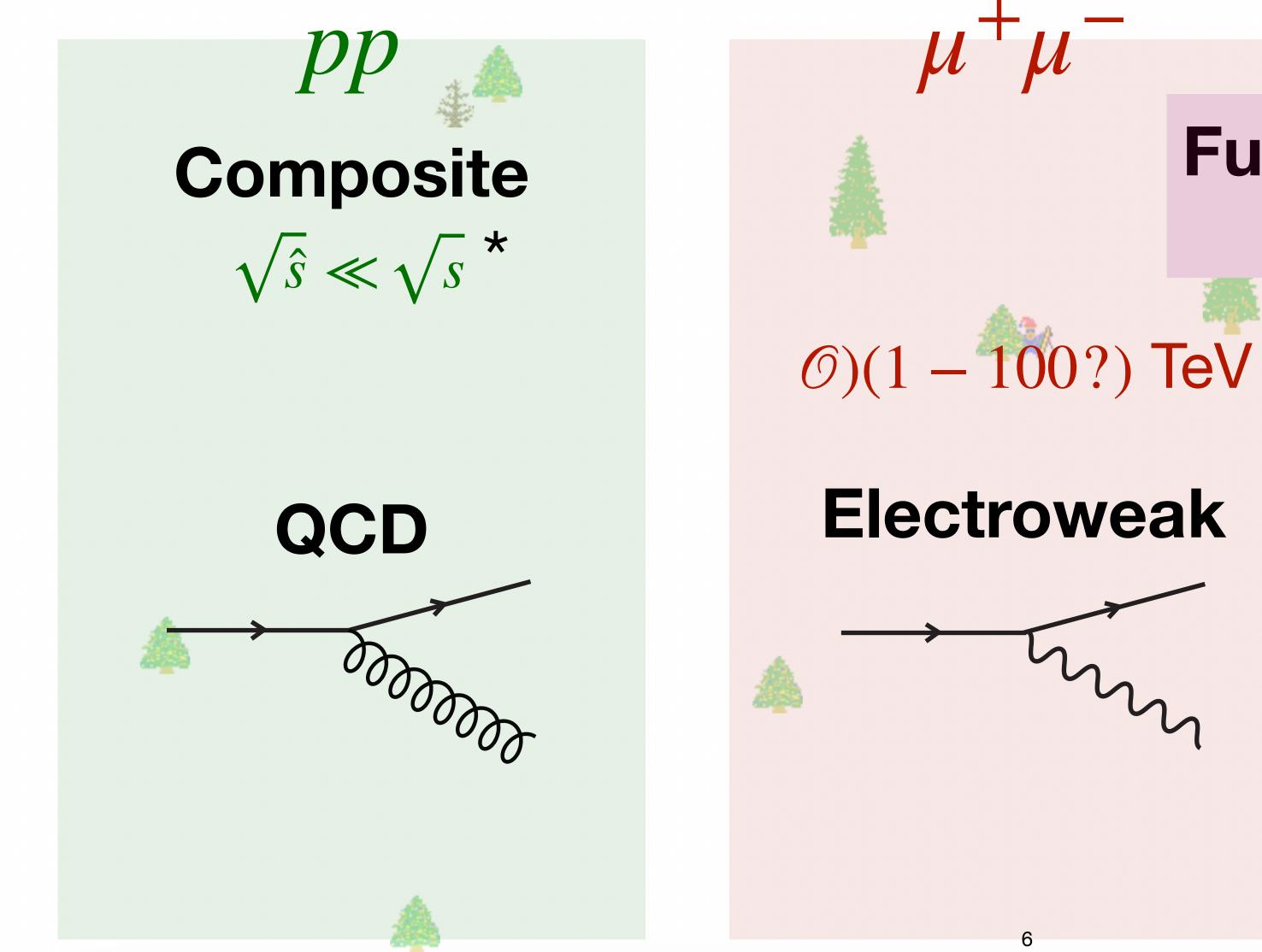


E.

Ø)(1 − 100?) TeV

Ø(100 − 300) GeV





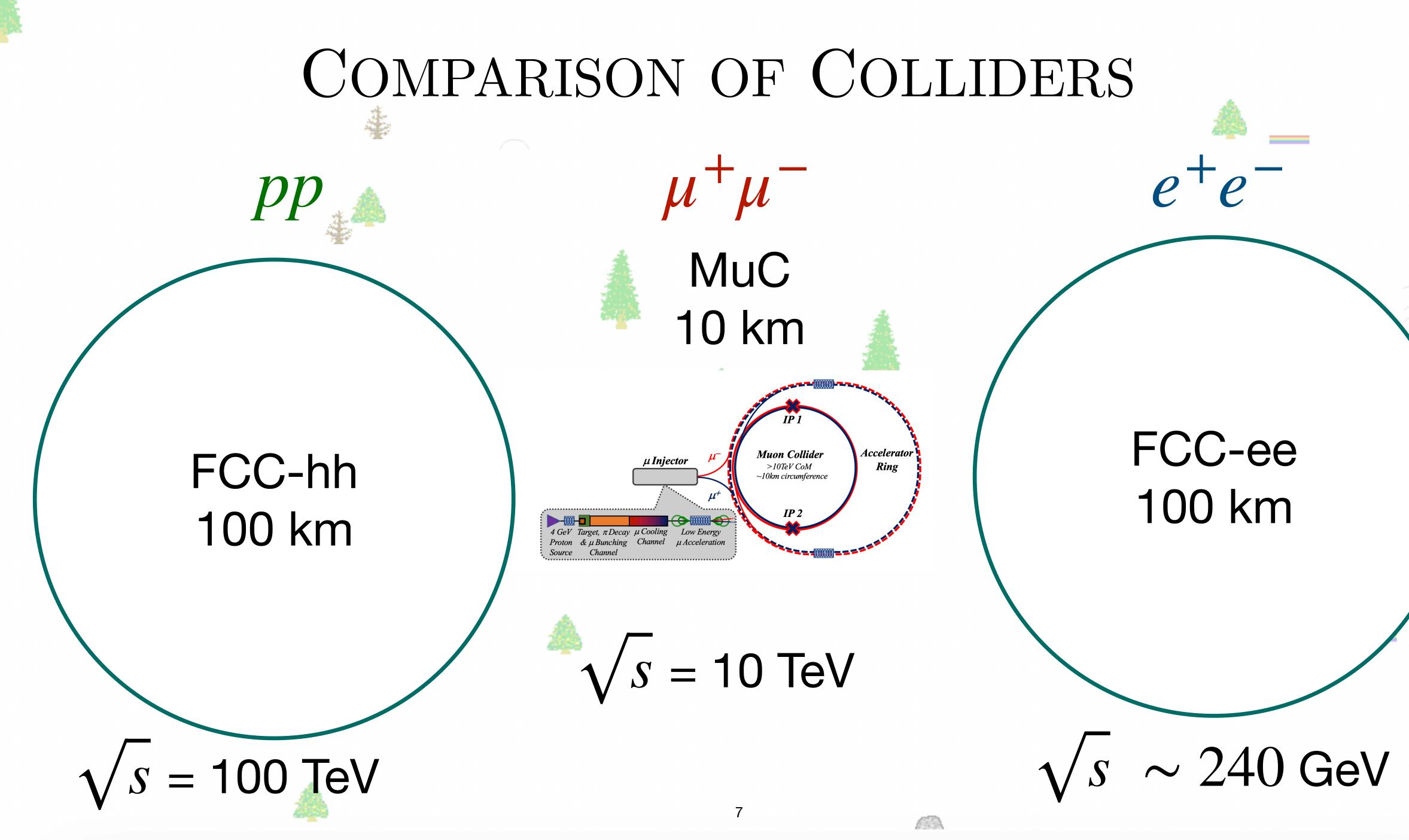
COMPARISON OF COLLIDERS

E.

Fundamental $\sqrt{\hat{s}} \sim \sqrt{s}$



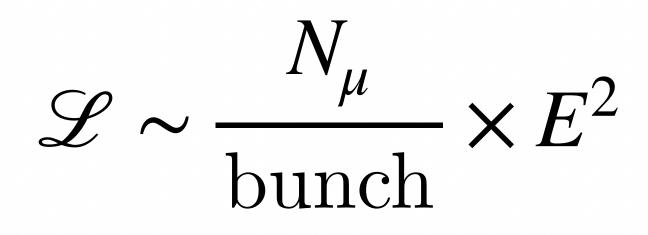


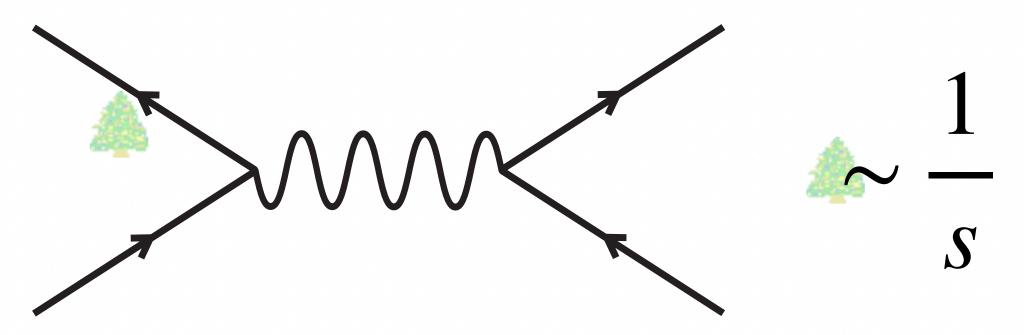




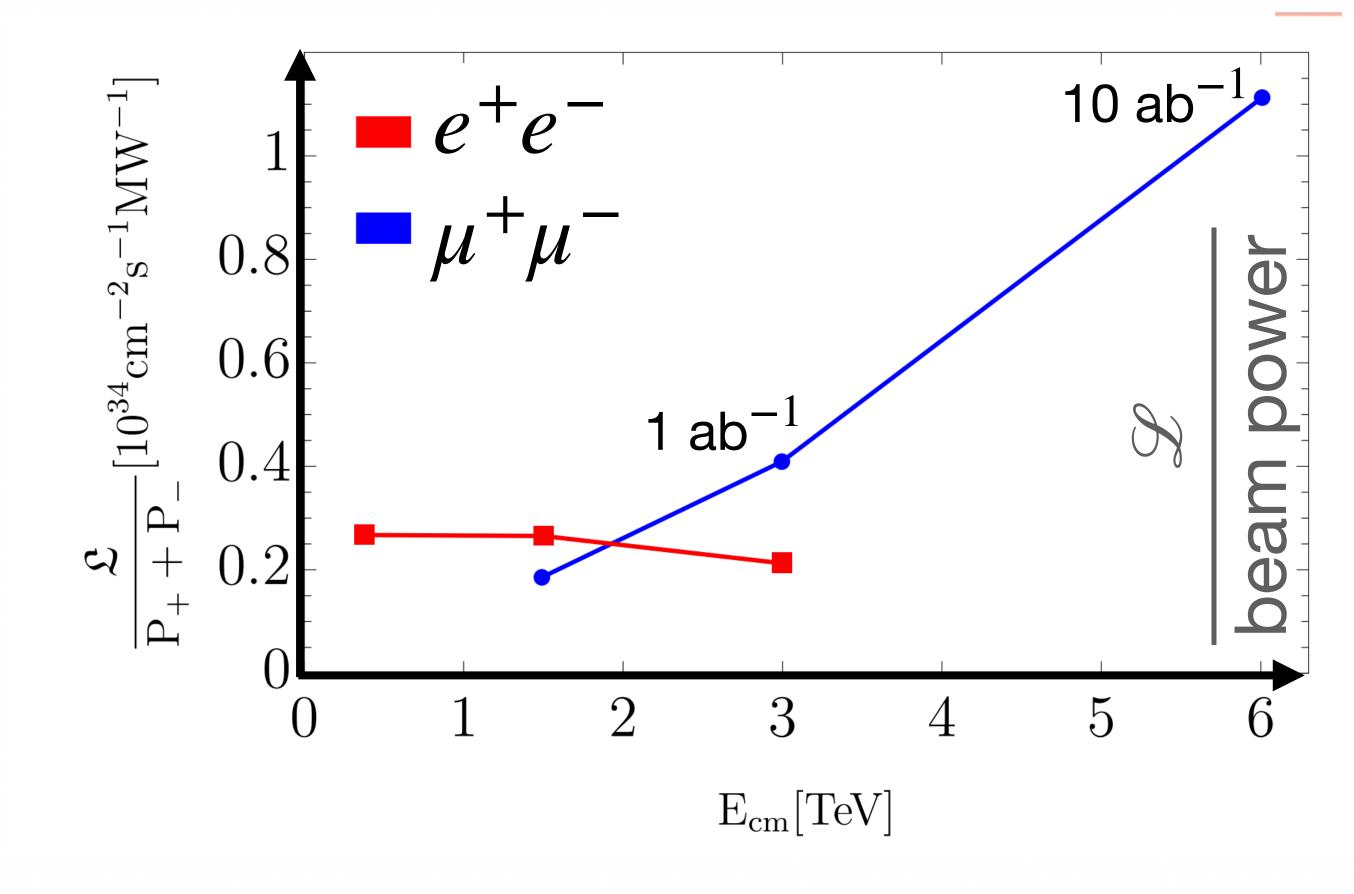
COMPARISON OF LEPTON COLLIDERS

The higher the energy, the larger the luminosity















What are the **biggest** challenges?



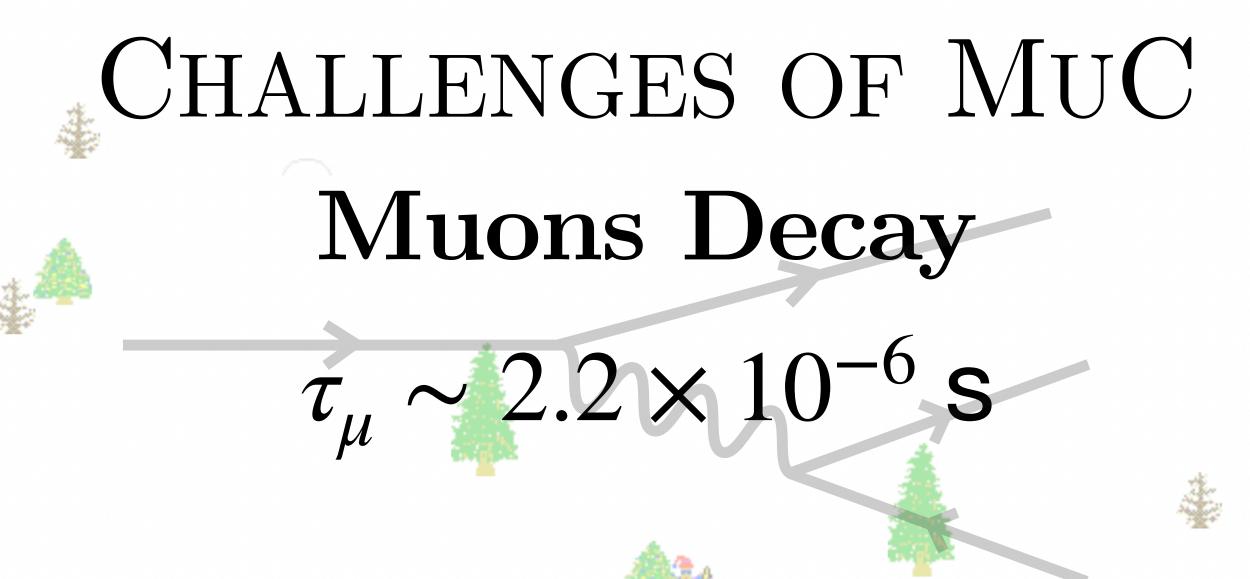


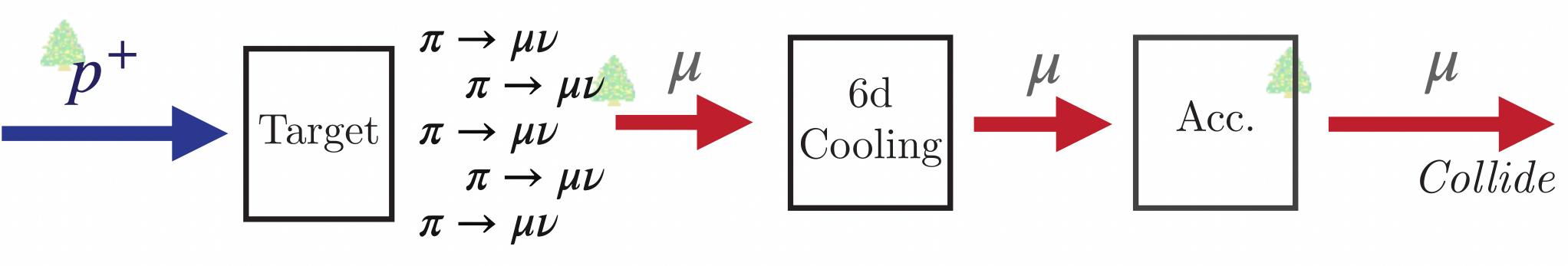








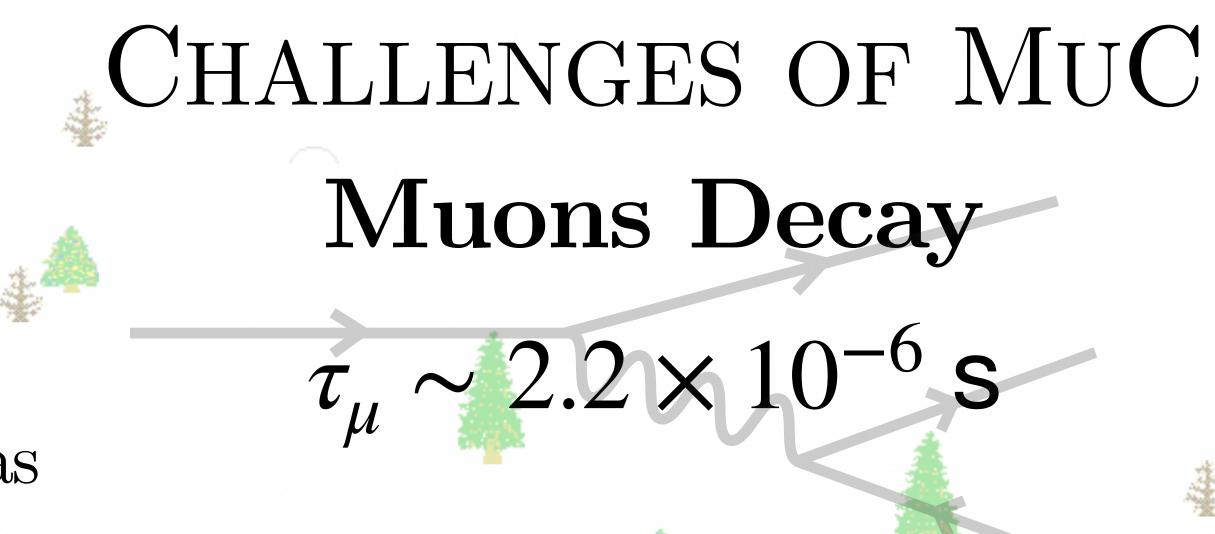


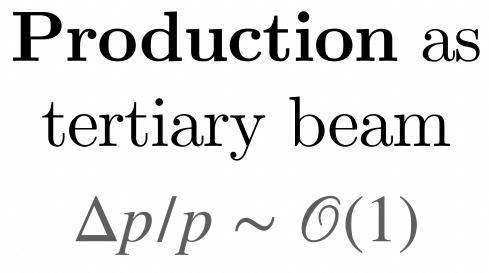


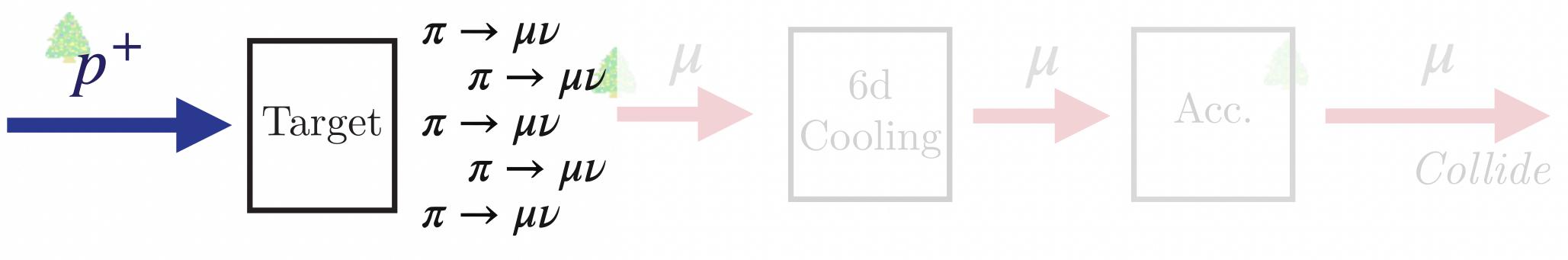














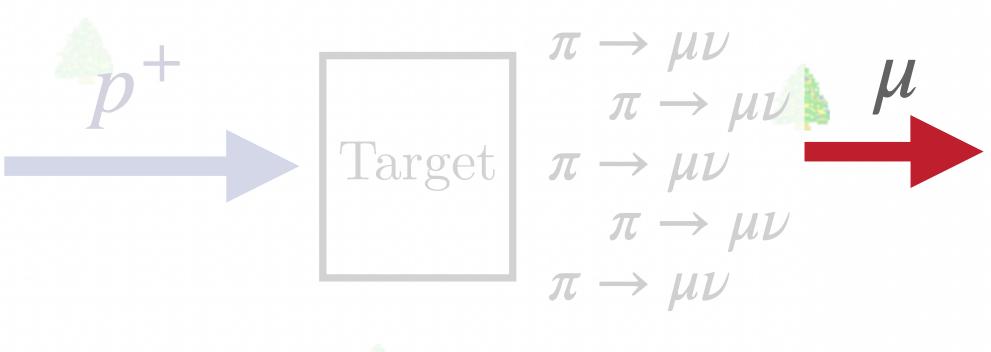






Production as tertiary beam $\Delta p/p \sim \mathcal{O}(1)$

Cooling into single collimated bunch





CHALLENGES OF MUC Muons Decay

 $\tau_{\mu} \sim 2.2 \times 10^{-6} \,\mathrm{s}$

 $0.9^{120} \sim 10^{-6}$

6d Cooling Collide

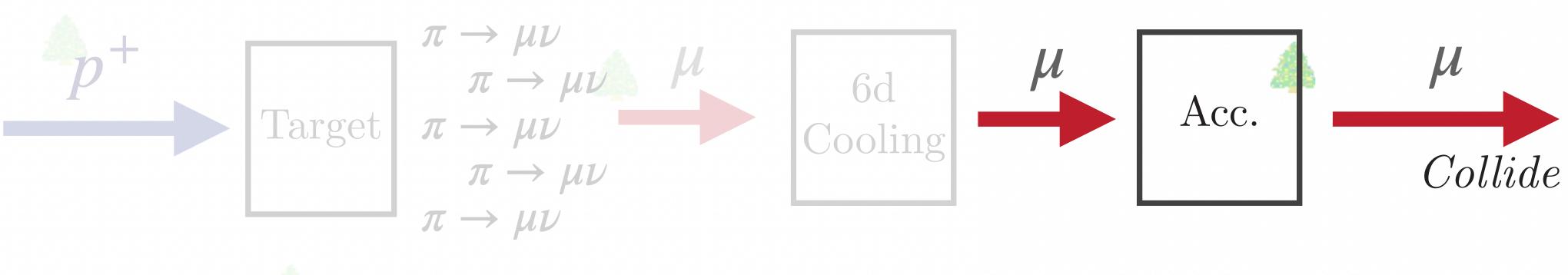






Production as tertiary beam $\Delta p/p \sim \mathcal{O}(1)$

 $0.9^{120} \sim 10^{-6}$





CHALLENGES OF MUC Muons Decay

Cooling into single collimated bunch

Acceleration and collision

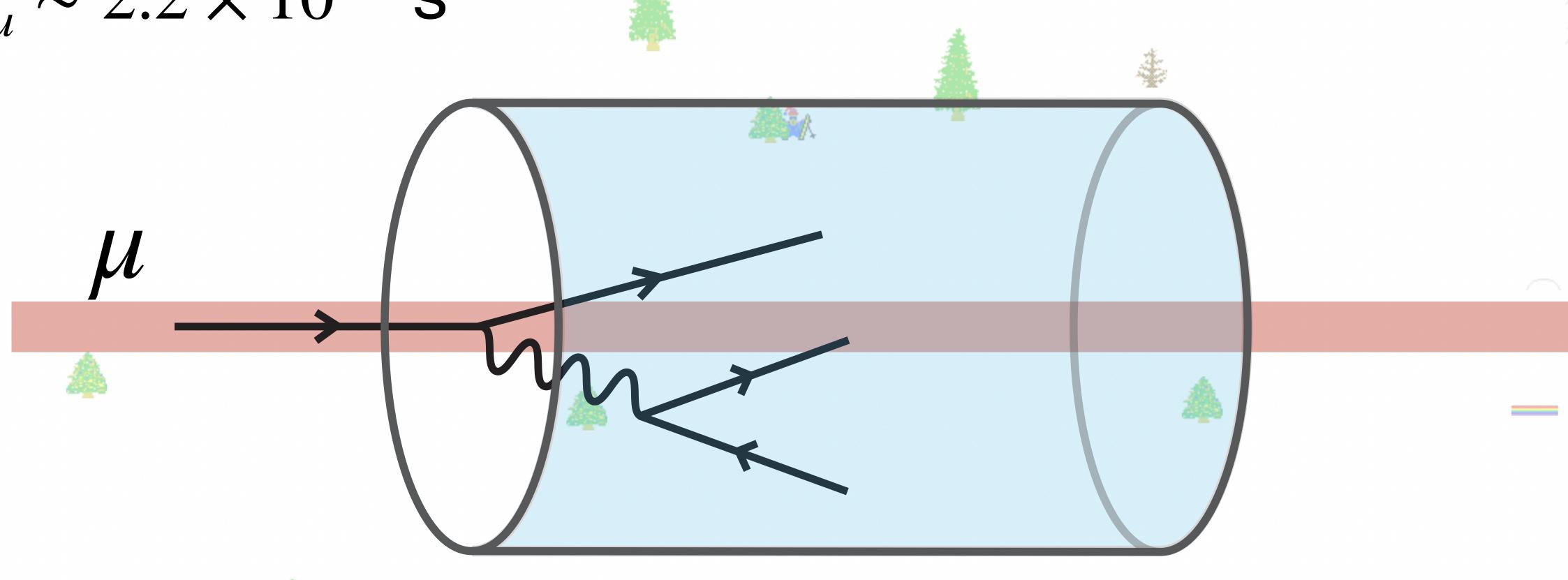
Too quick to ramp up magnets





CHALLENGES OF MUC

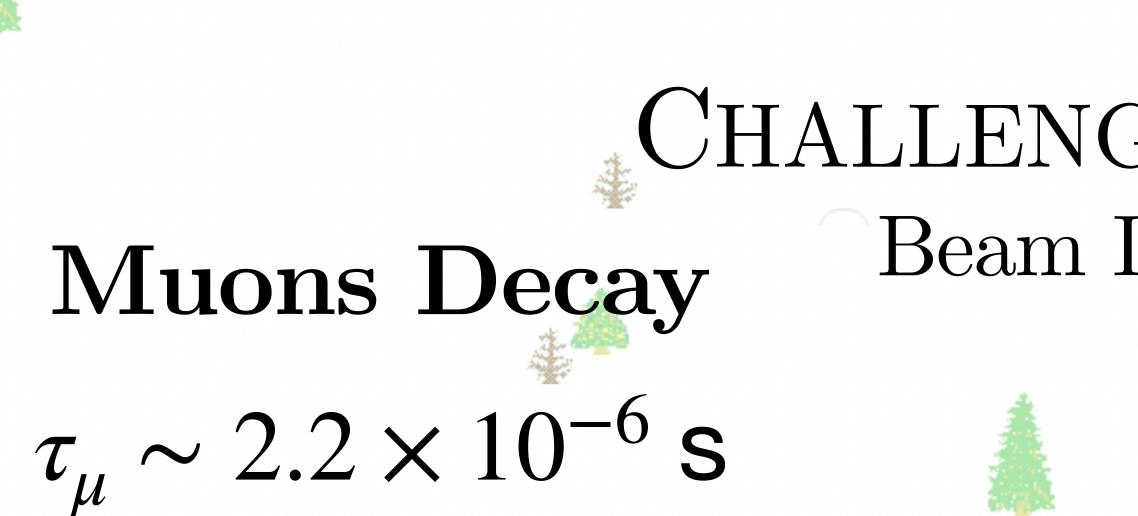
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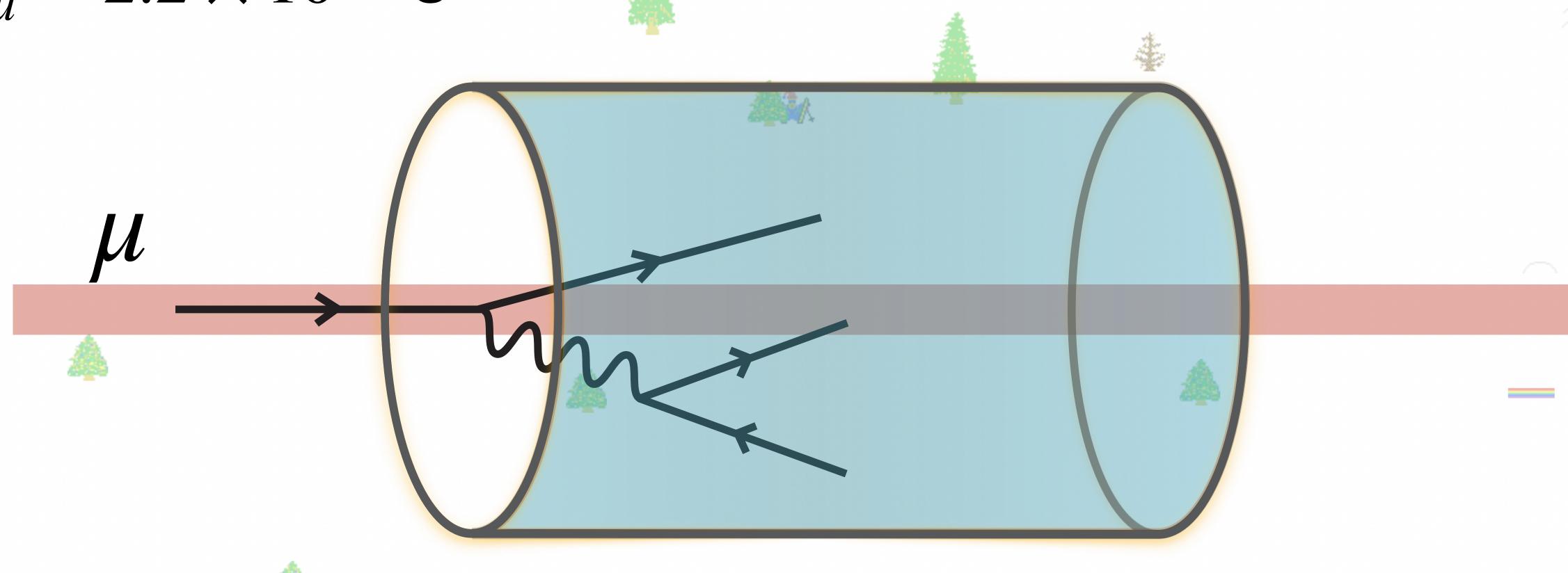








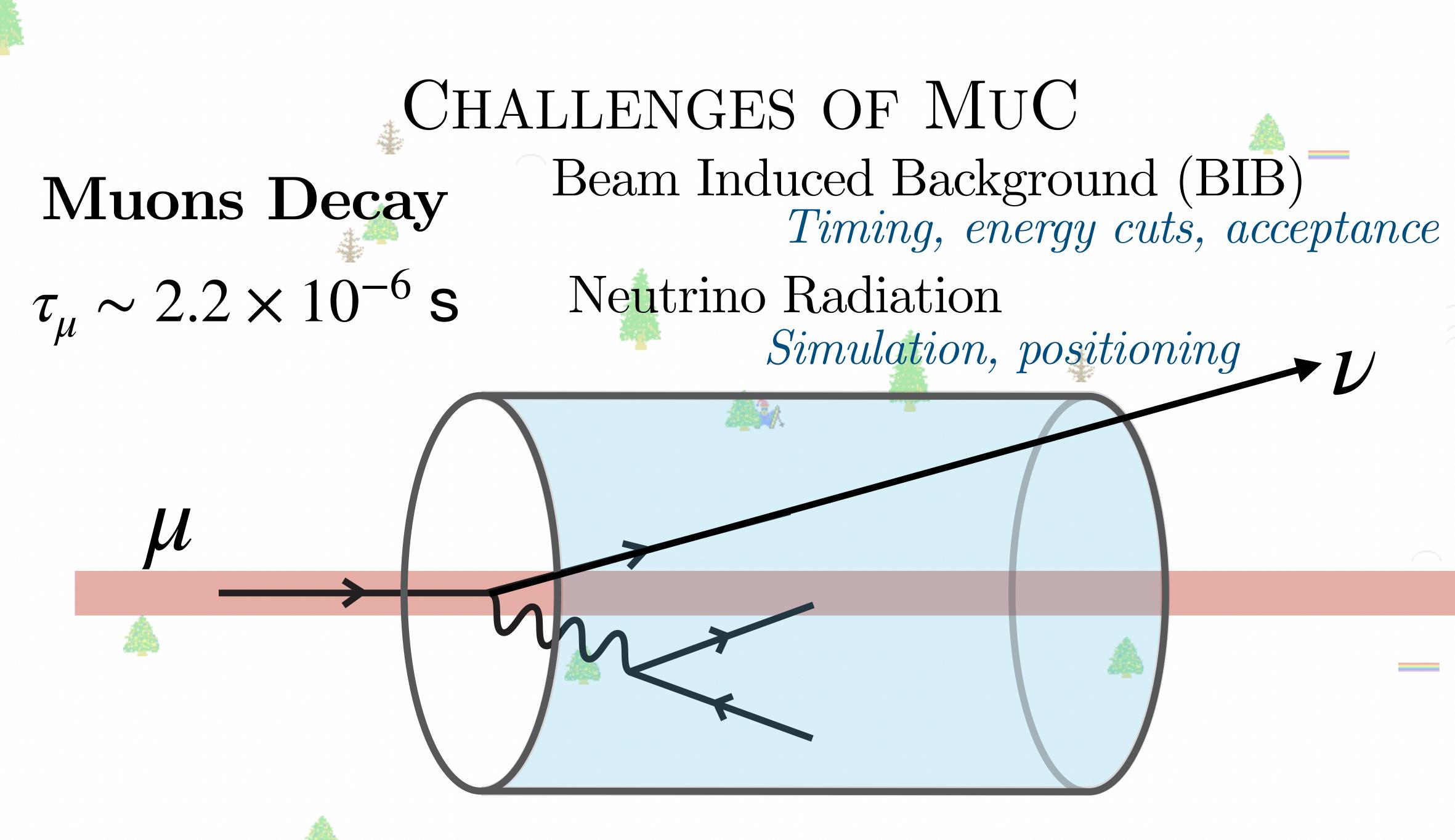






CHALLENGES OF MUC Beam Induced Background (BIB) Timing, energy cuts, acceptance







Two Paths Forward with a MuC \clubsuit













Intensity Staging





TWO PATHS FORWARD WITH A MUC





Energy Staging

3, 10, 30?, 100 TeV?







Intensity Staging

Start with smaller magnets, bigger ring

 $\mathscr{L} \to \mathscr{L} \times 10^{-1}$



TWO PATHS FORWARD WITH A MUC





3, 10, 30?, 100 TeV?



This early in the process, we cannot make fully informed decisions about what run configurations we will have

(A)



Intensity Staging

Start with smaller magnets, bigger ring

 $\mathscr{L} \to \mathscr{L} \times 10^{-1}$

84/









What is the physics reach?















Electroweak Precision









PHYSICS PROGRAM AT MUC Direct & Indirect Heavy States





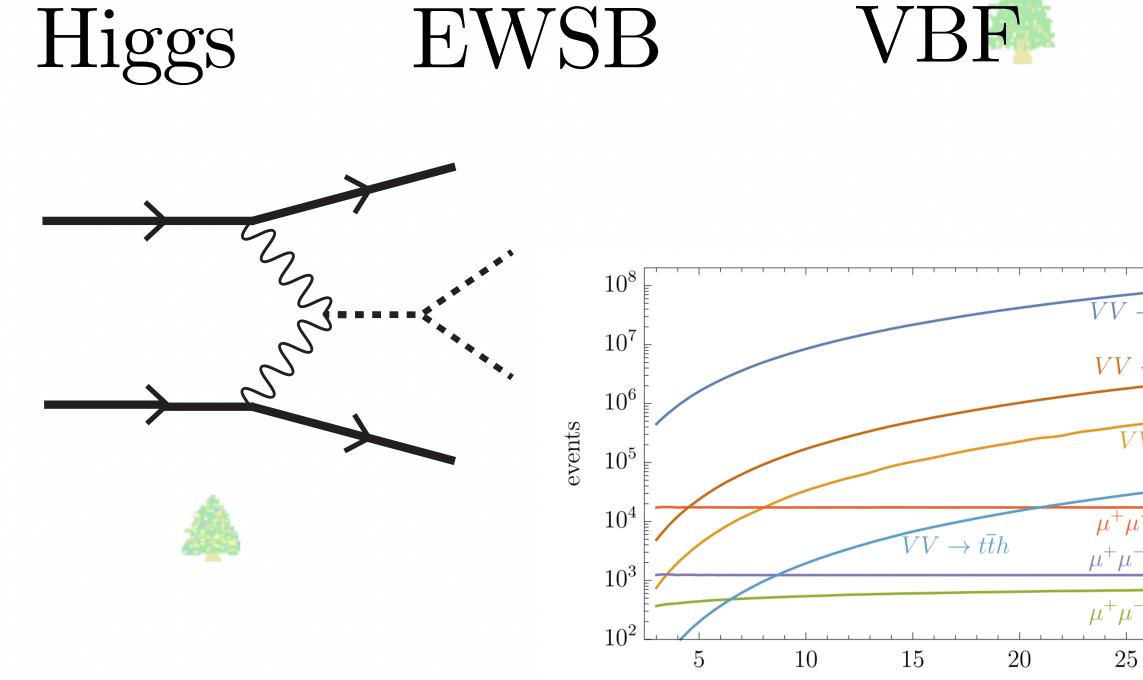






PHYSICS PROGRAM AT MUC ak Precision Direct & Indirect UDE UDE Heavy States

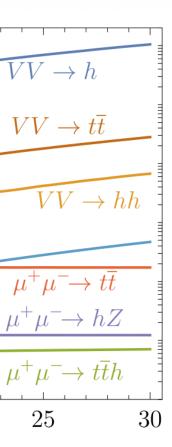
Electroweak Precision FWSB VBF



 $E_{\rm cm} \left[{\rm TeV} \right]$







14



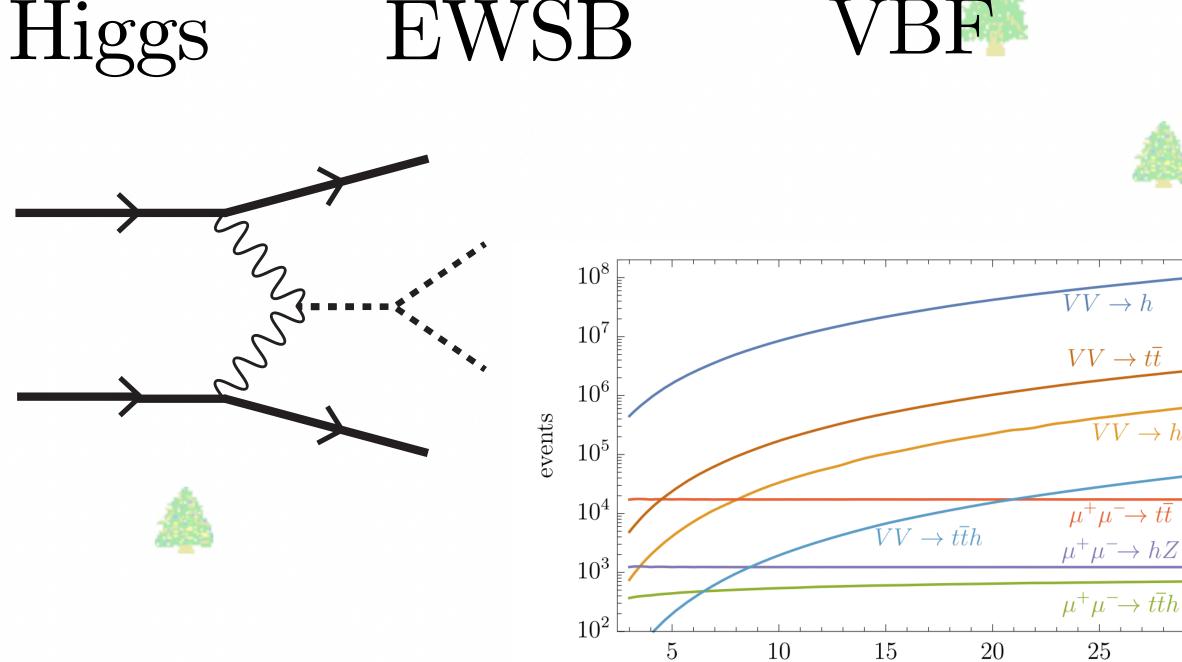






PHYSICS PROGRAM AT MUC Direct & Indirect Heavy States VBF EWSB

Electroweak Precision



 $E_{\rm cm}$ [TeV]



 $\overline{VV} \to hh$

25

New Physics Pair Production

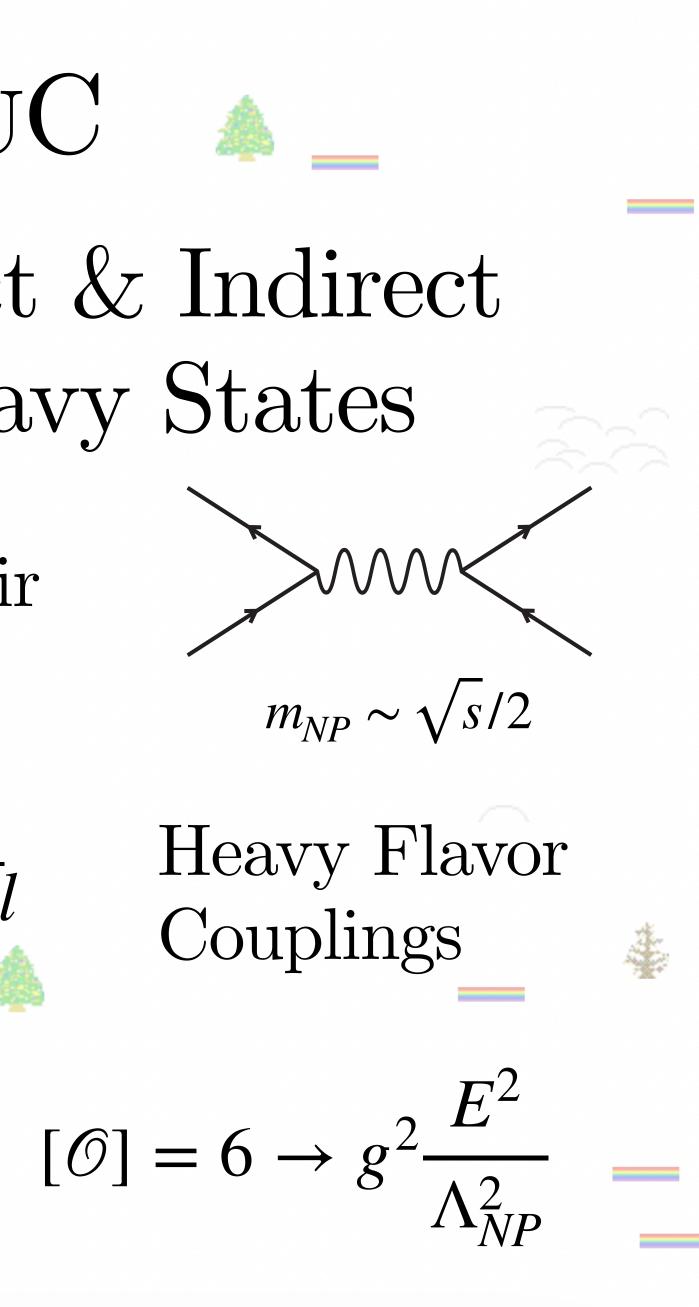
 \sim

 $m_{NP} \sim \sqrt{s/2}$

Heavy Flavor Couplings

 $\mathscr{L} \supset g \frac{m_l}{v} \phi \bar{l} l$

EFTs







































Higgs Studies





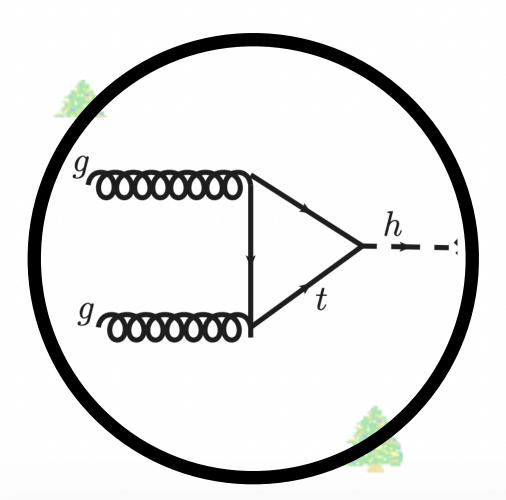


HIGGS PRODUCTION



pp Collider





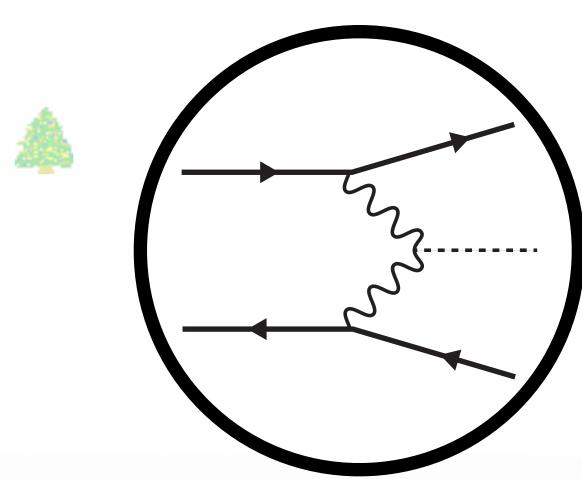


l^+l^- Collider

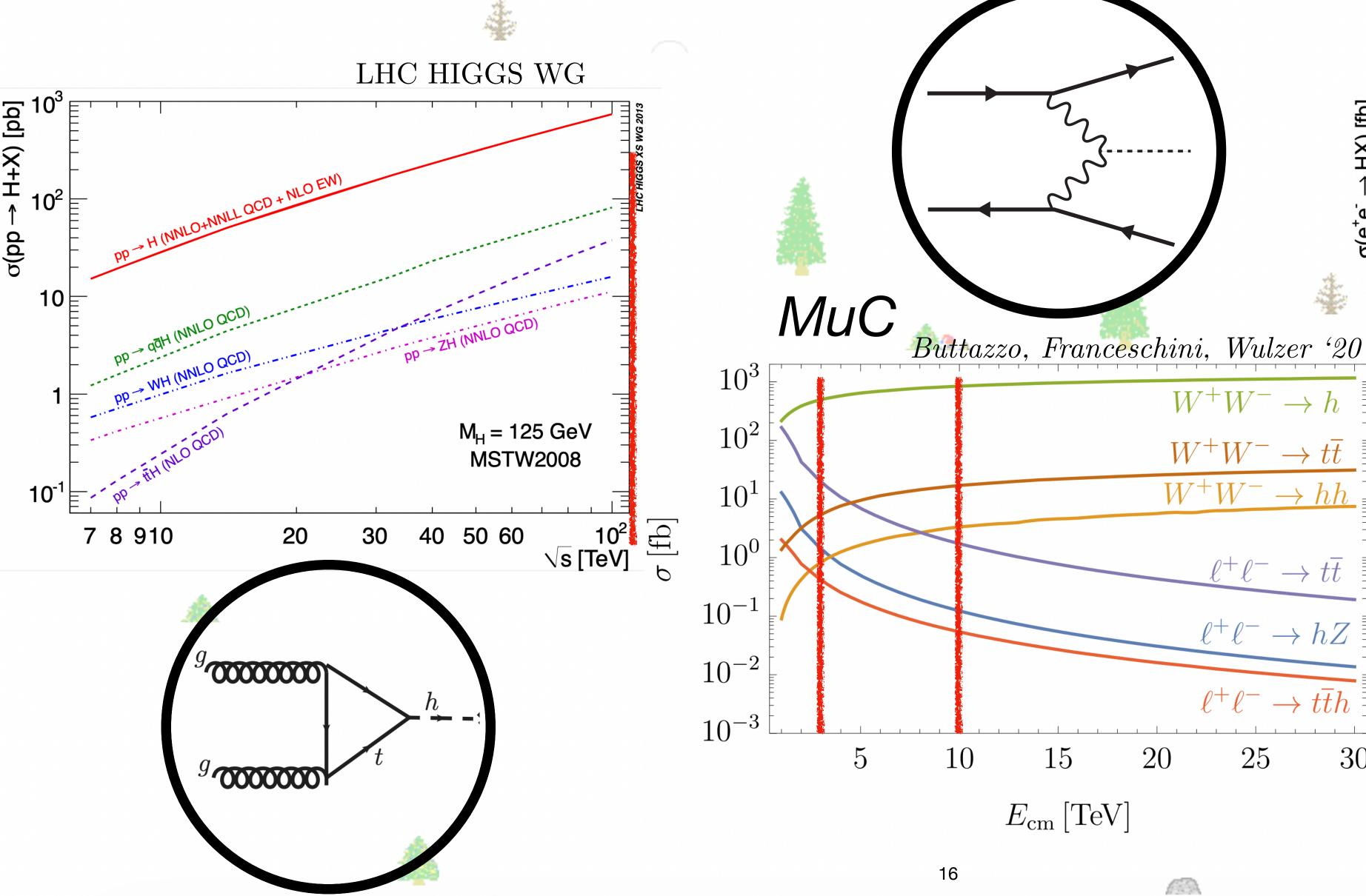










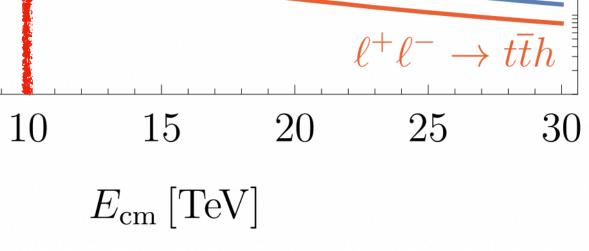


HIGGS PRODUCTION *CLIC '18* → HX) [fb] $Hv_e \overline{v}_e$ 10² ⊦ e⁺e σ(e⁺e⁻ 10

10-

10⁻²

0



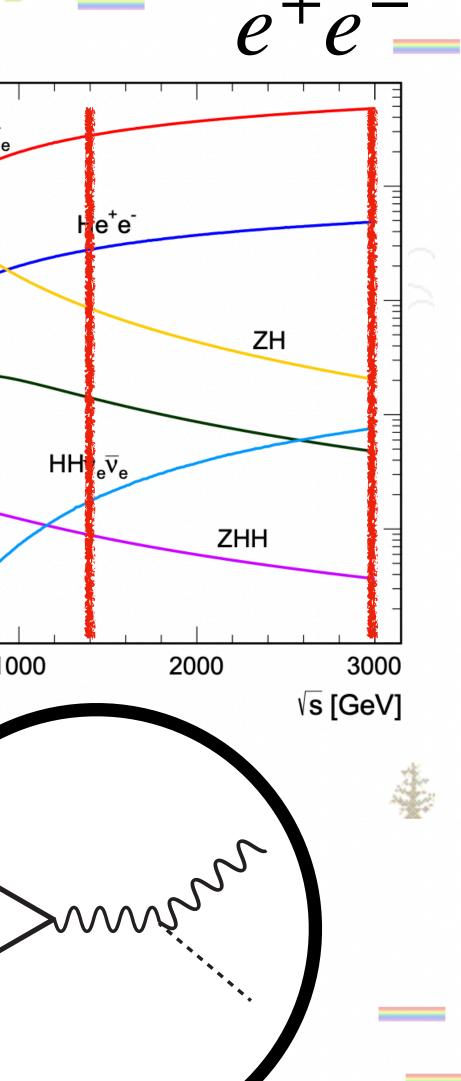
 $W^+W^- \to h$

 $W^+W^- \to t\bar{t}$

 $W^+W^- \to hh$

 $\ell^+\ell^- \to t\overline{t}$

 $\ell^+\ell^- \to hZ$



ΖH

ZHH

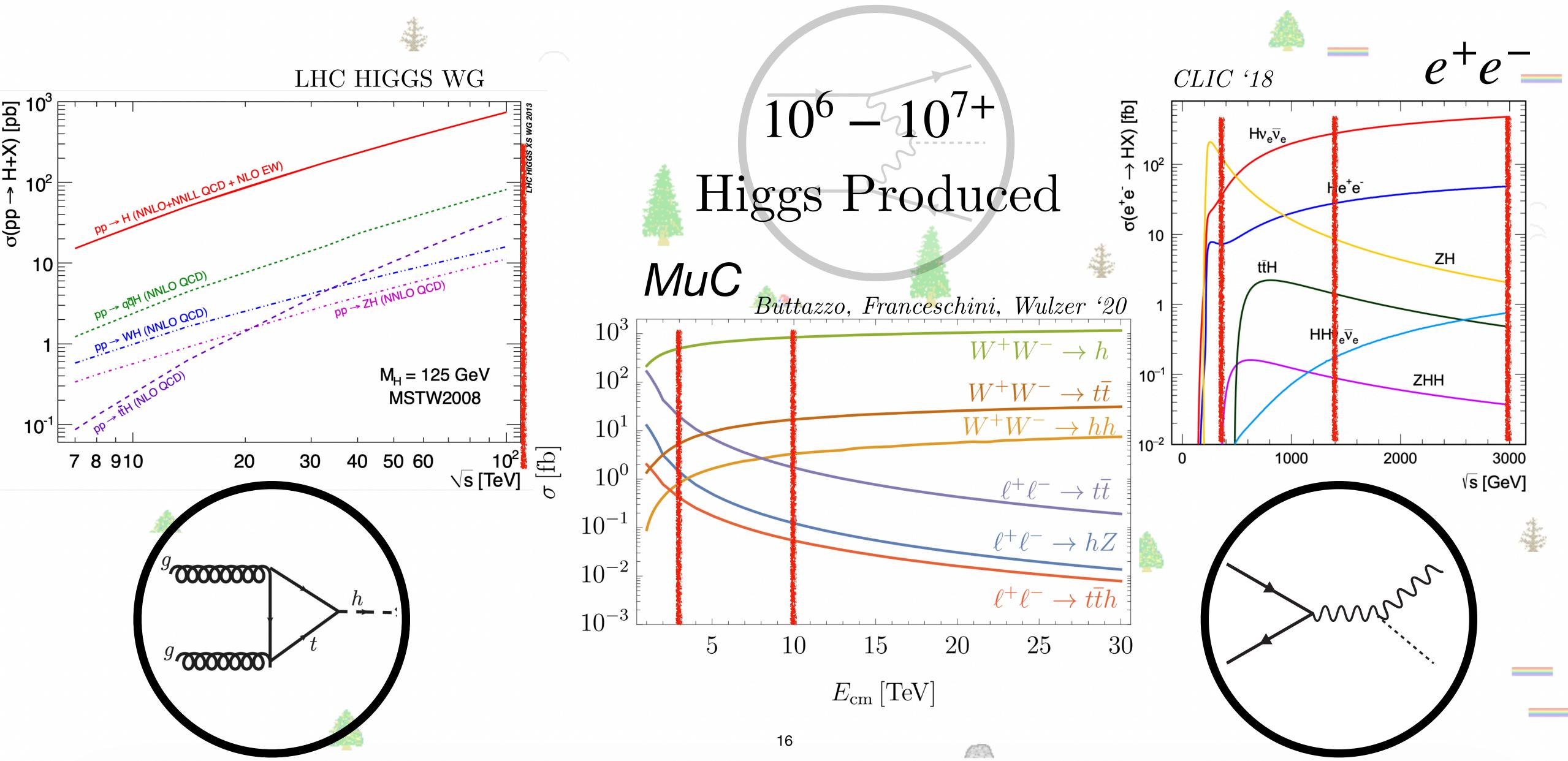
2000

tīH

1000

 $HH_{e}\overline{v}_{e}$

HIGGS PRODUCTION



HIGGS PRODUCTION

Consider precision in κ framework

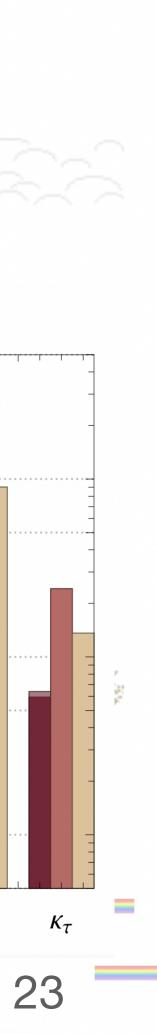
								1905.03764			230
<i>к</i> -0	HL-LHC	ILC			CLIC			CEPC	FC	C-ee	FCC-ee
fit		250	500	1000	380	1500	3000		240	365	eh/hh
κ_W [%]	1.7	1.8	0.29	0.24	0.86	0.16	0.11	1.3	1.3	0.43	0.14
κ_Z [%]	1.5	0.29	0.23	0.22	0.5	0.26	0.23	0.14	0.20	0.17	0.12
$\kappa_g \ [\%]$	2.3	2.3	0.97	0.66	2.5	1.3	0.9	1.5	1.7	1.0	0.49
$\kappa_{\gamma} \ [\%]$	1.9	6.7	3.4	1.9	98*	5.0	2.2	3.7	4.7	3.9	0.29
$\kappa_{Z\gamma}$ [%]	10.	$99\star$	$86\star$	$85\star$	$120\star$	15	6.9	8.2	$81\star$	$75\star$	0.69
$\kappa_c \ [\%]$	-	2.5	1.3	0.9	4.3	1.8	1.4	2.2	1.8	1.3	0.95
$\kappa_t \ [\%]$	3.3	_	6.9	1.6	_	_	2.7	—	—	_	1.0
$\kappa_b \ [\%]$	3.6	1.8	0.58	0.48	1.9	0.46	0.37	1.2	1.3	0.67	0.43
κ_{μ} [%]	4.6	15	9.4	6.2	$320\star$	13	5.8	8.9	10	8.9	0.41
κ_{τ} [%]	1.9	1.9	0.70	0.57	3.0	1.3	0.88	1.3	1.4	0.73	0.44



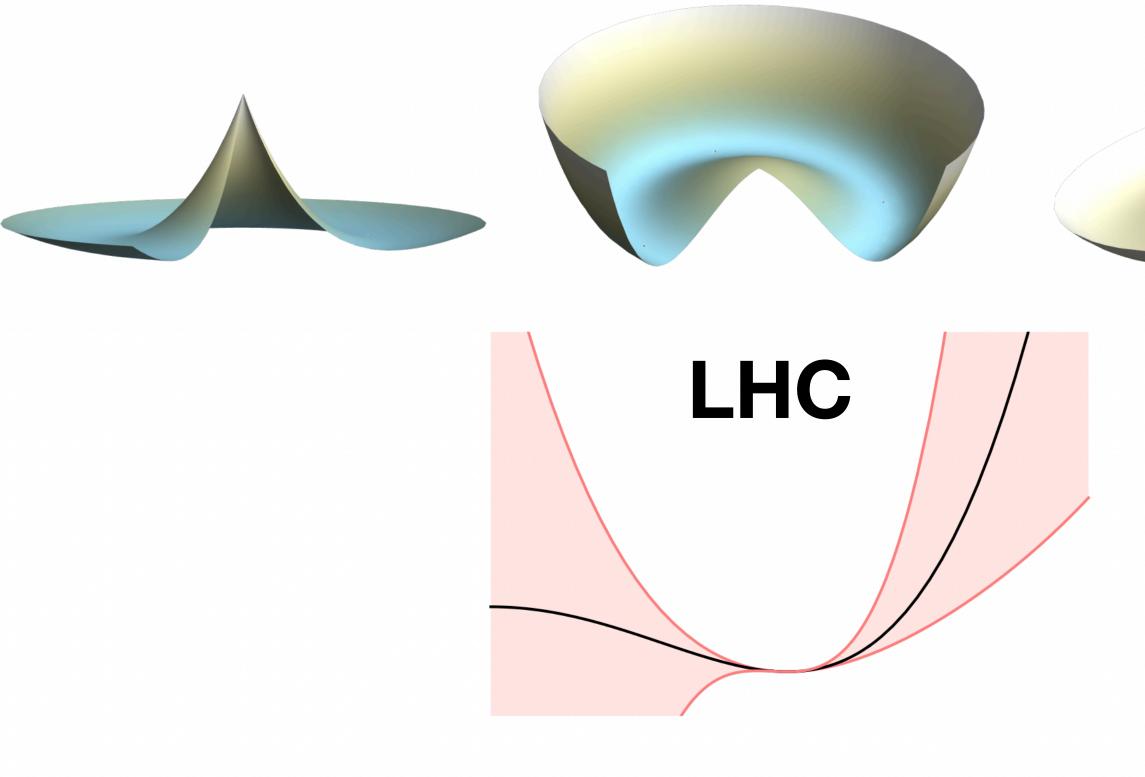
10⁶ – 10⁷⁺ Higgs Produced



08.02633 $BR_{BSM}=0$ Fit Comparisons $\mu^{+}\mu^{-}$ h **10000** \blacksquare 10 TeV $\mu^+\mu^-$ @ 10/ab ■ HL–LHC 0.0610 $\square 250 \text{ GeV } e^+e^-$ 0.23Precision [%] 0.150.641.00.500.896.00.16 0.10^{-1} 2.00.05 κ_Z κ_g $\kappa_{Z\gamma}$ κ_t κ_b κ_W Kγ κ_c κ_{μ} 0.31Scenario **B**R_{inv} Meade & Forslund 23 fixed at 0 kappa-0 16 11 12 12



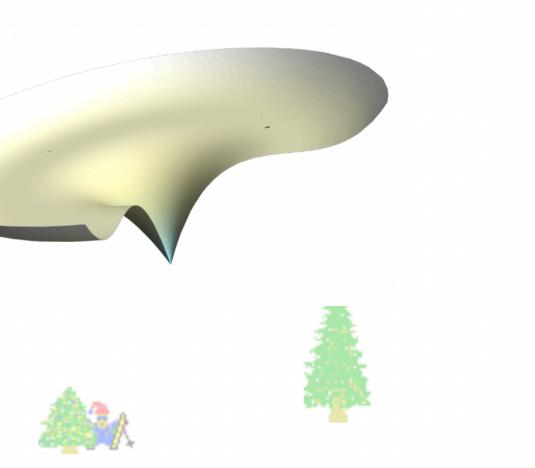
Slide Credit: P Meade, N Craig, R. Petrossian Byrne HIGGS POTENTIAL

























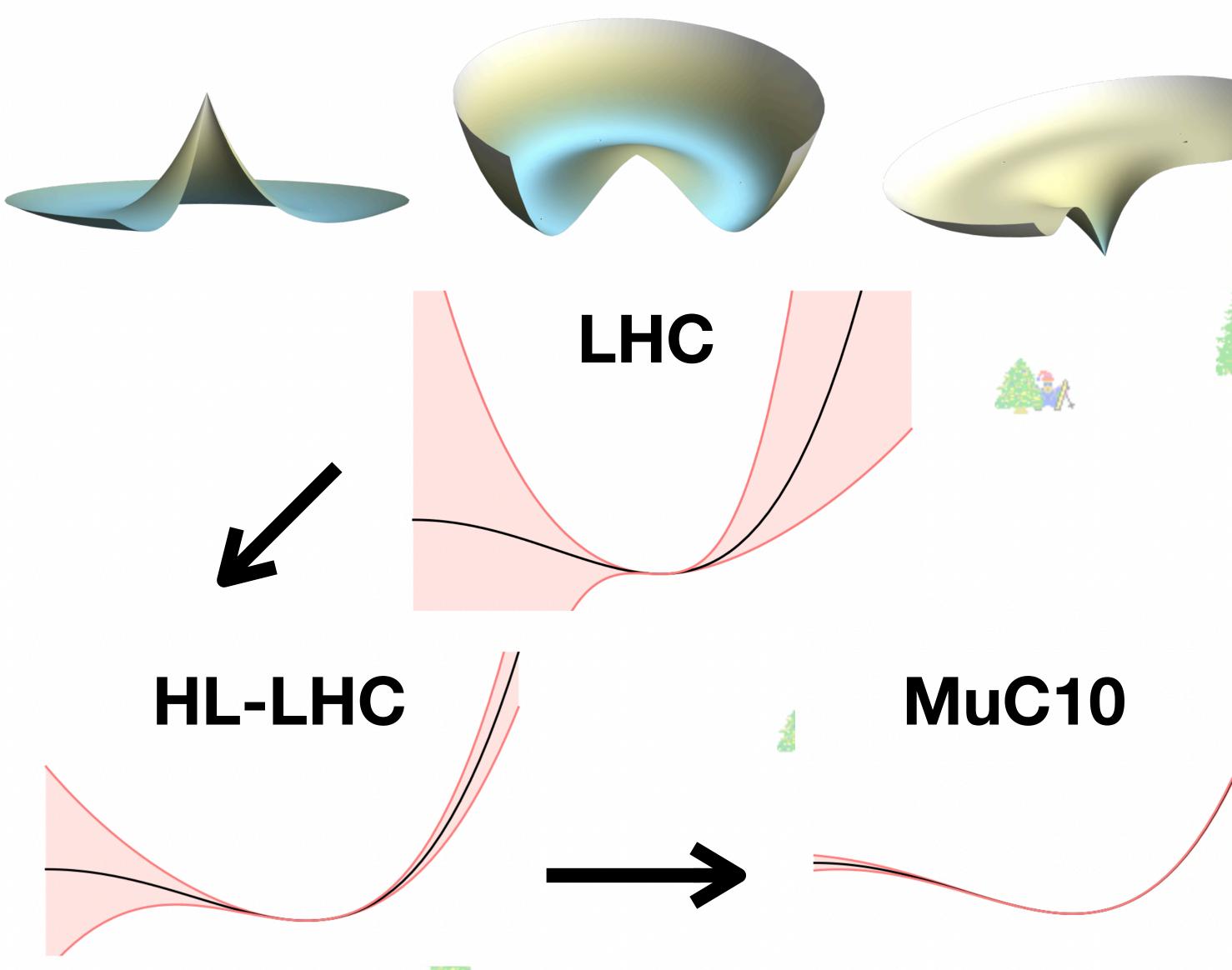








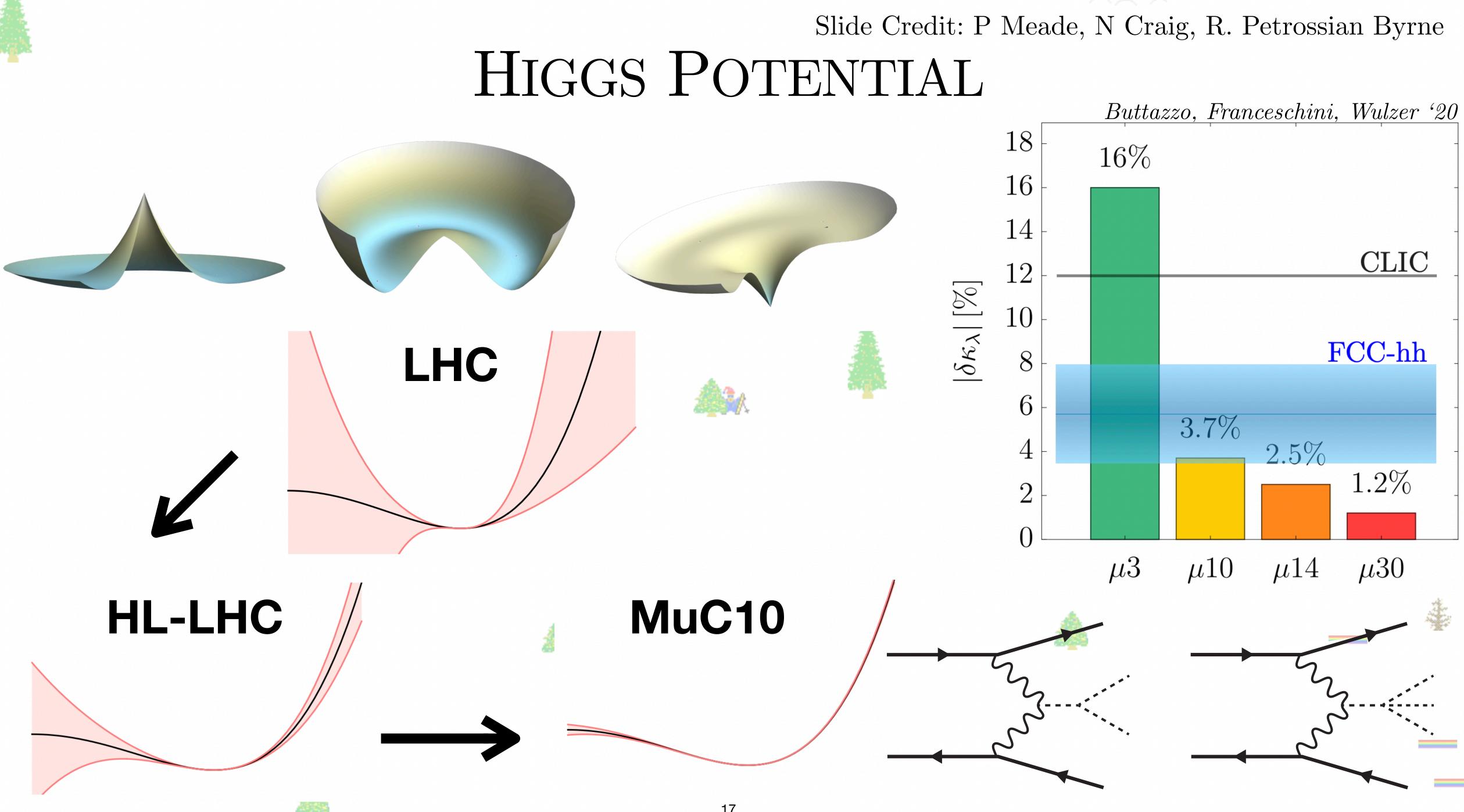
Slide Credit: P Meade, N Craig, R. Petrossian Byrne HIGGS POTENTIAL























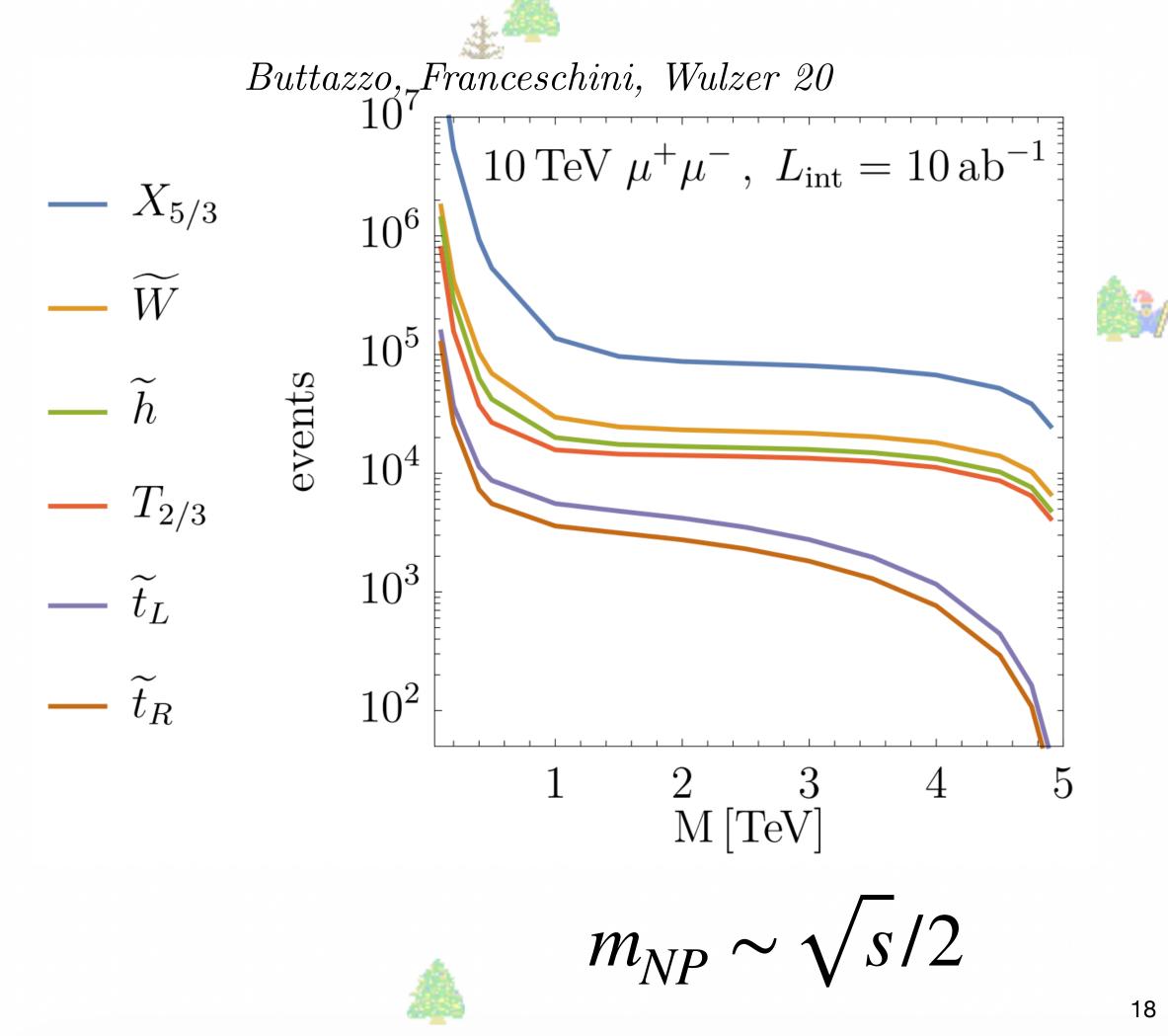


Examples of Direct Production of Heavy New Physics

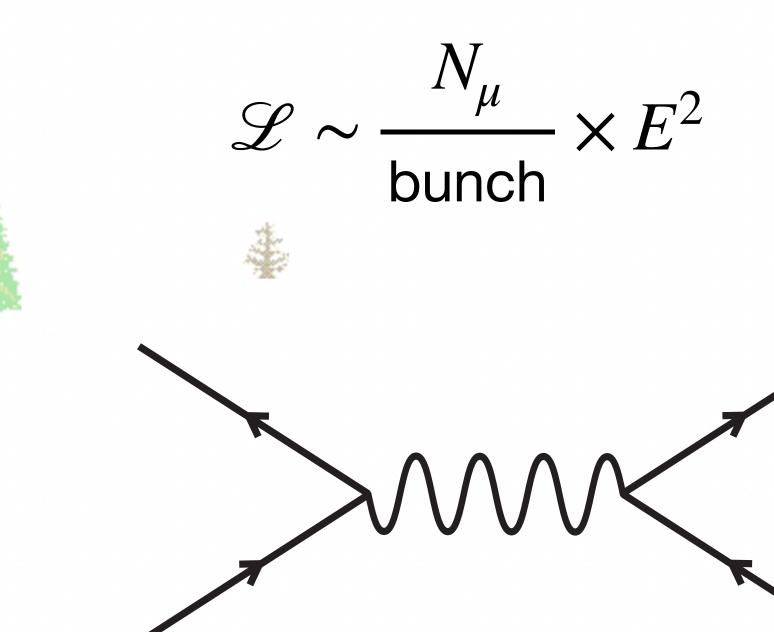




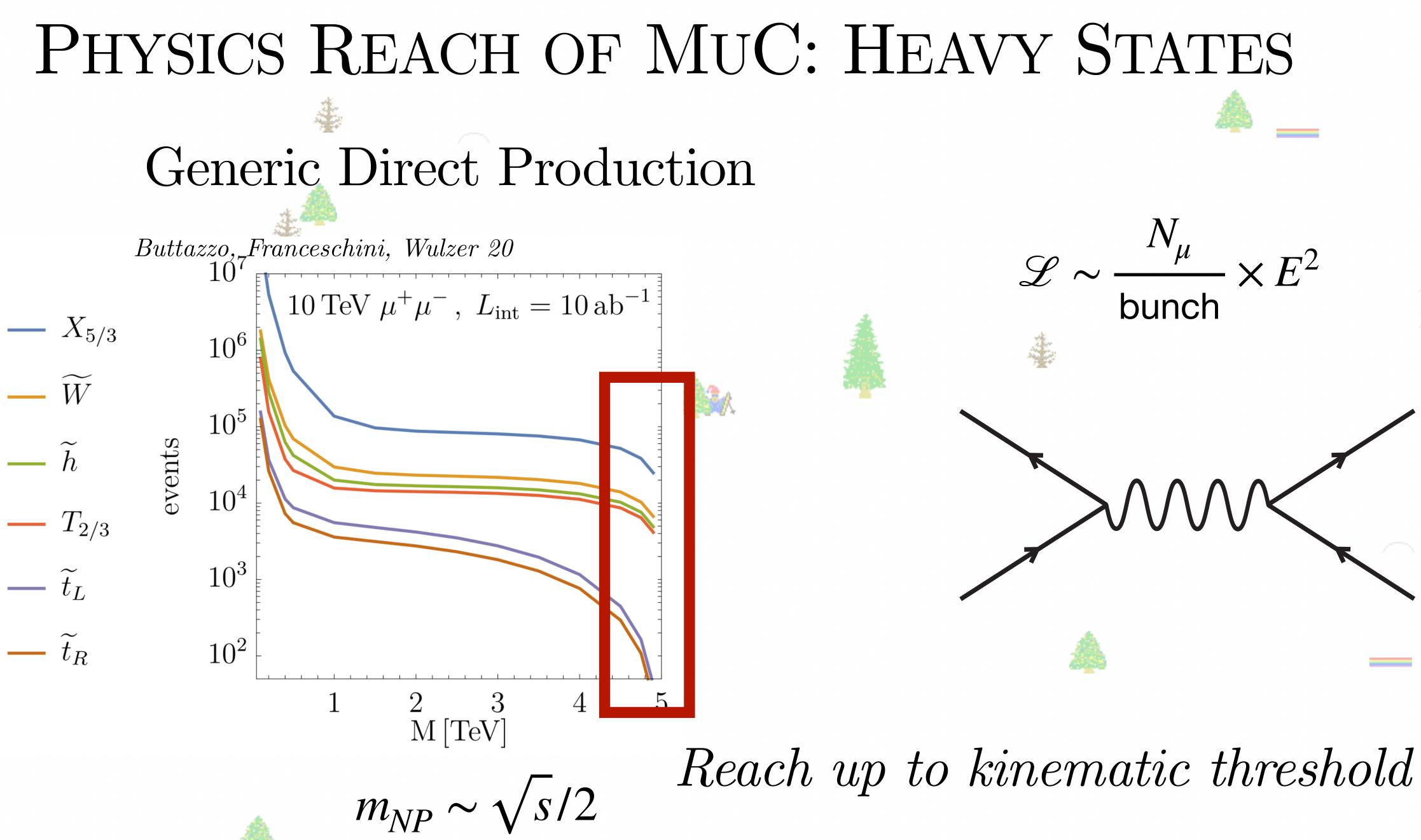
Generic Direct Production



PHYSICS REACH OF MUC: HEAVY STATES



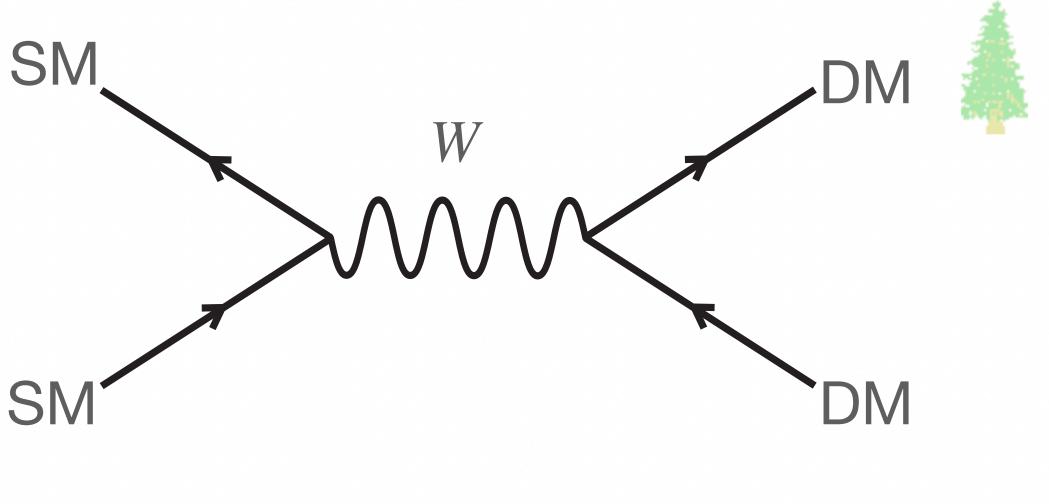




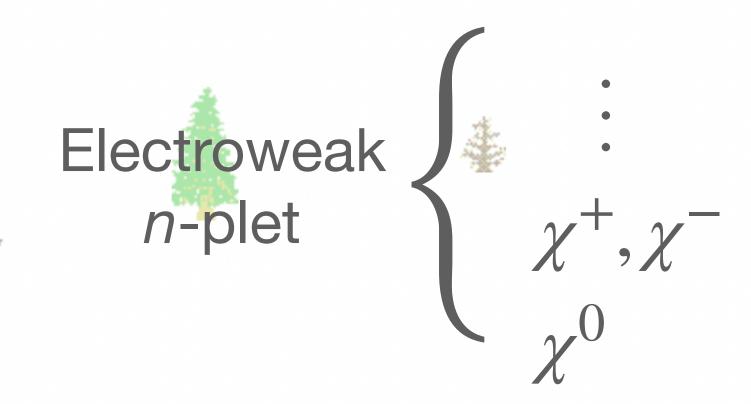




PHYSICS REACH OF MUC: WIMPS For dark matter models coupling to EW bosons, MuC is an ideal place for searches



Bottaro, Buttazzo, Costa, Franceschini, Panci, Redigolo, Vittorio 21, 22



Mass fixed by freeze-out abundance

EW n-plet	Mass [TeV]		
2 _{1/2}	1.08		1
3 0	2.86		
4 _{1/2}	4.8	freeze out →	1
5 0	13.6	Freeze out →	→Ω _x h ² -
5 1	9.9		Y 1
6 1/2	31.8		-
7 0	48.8	[\ ←	-Y _{EQ}
9 0	113	$-20\frac{1}{1} \qquad 10^{1} \qquad \qquad$	10" 10





SM

PHYSICS REACH OF MUC: WIMPS For dark matter models coupling to EW bosons, MuC is an ideal place for searches

 W, Z, γ

DM

mono-W (incl.)

mono-W (lep.)

mono-y

mono-Z

di-W(SS)

MIM (comb.)

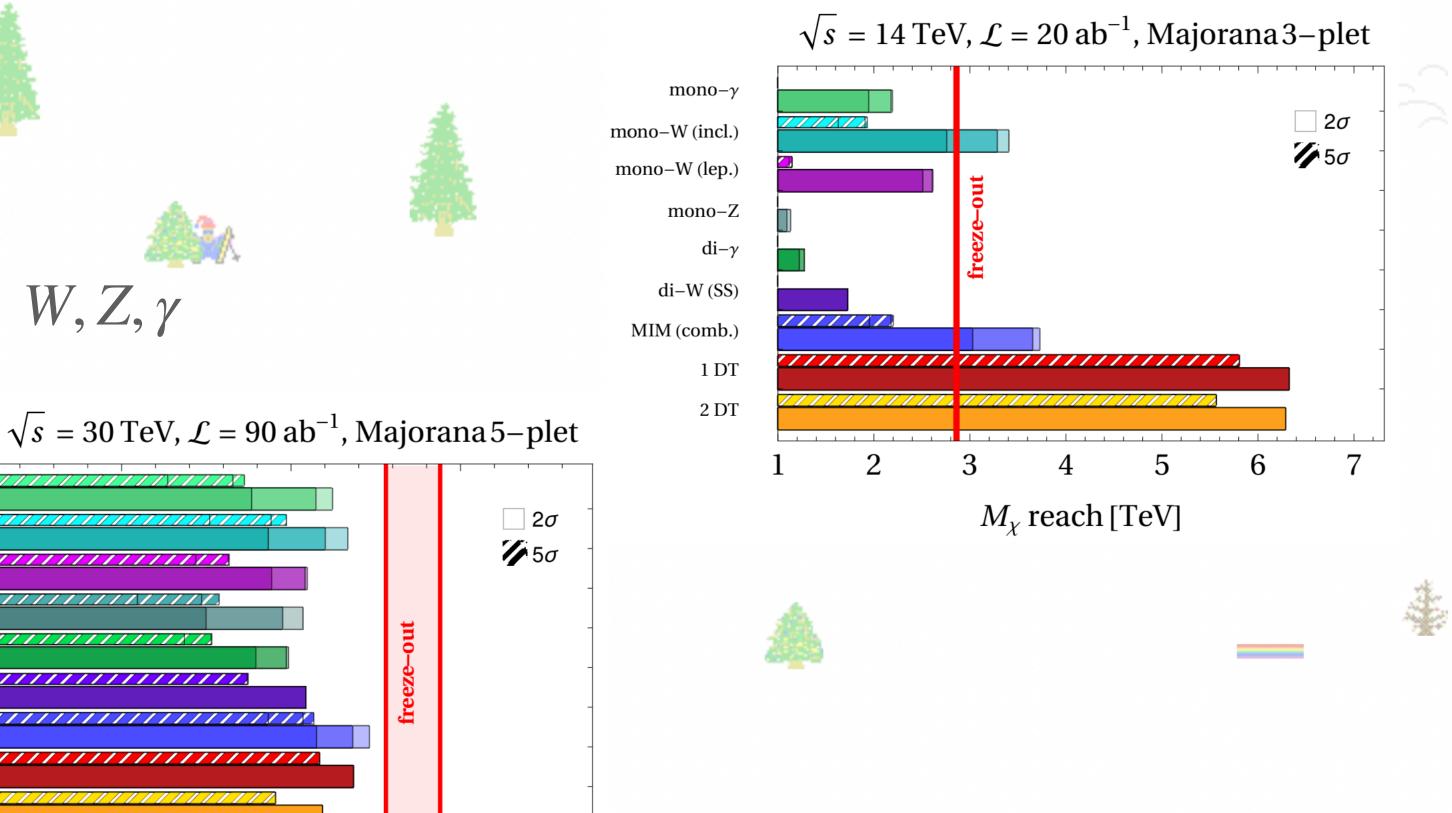
 $di-\gamma$

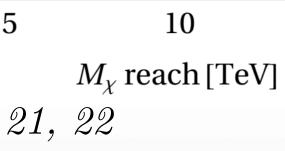
1 DT

2 DT

 $\mu^+\mu^- \to \chi\bar{\chi} + X$

Bottaro, Buttazzo, Costa, Franceschini, Panci, Redigolo, Vittorio 21, 22





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Examples of Indirect Production of Heavy New Physics







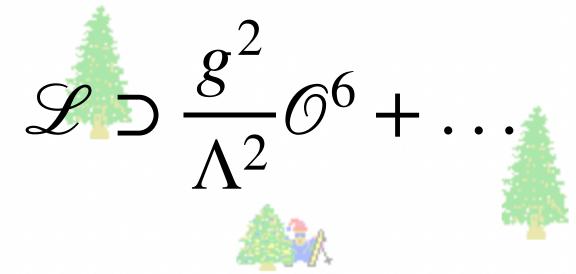






PHYSICS REACH OF MUC: INDIRECT

EFT Approach for Energy \leftrightarrow Precision





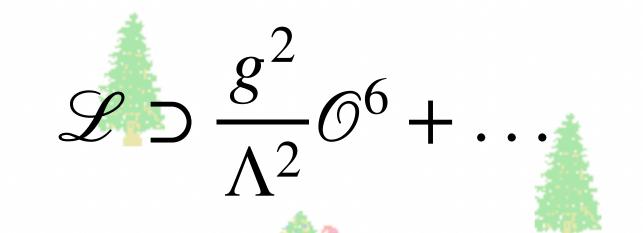






PHYSICS REACH OF MUC: INDIRECT

EFT Approach for Energy \leftrightarrow Precision



Say you can measure something to 1% precision









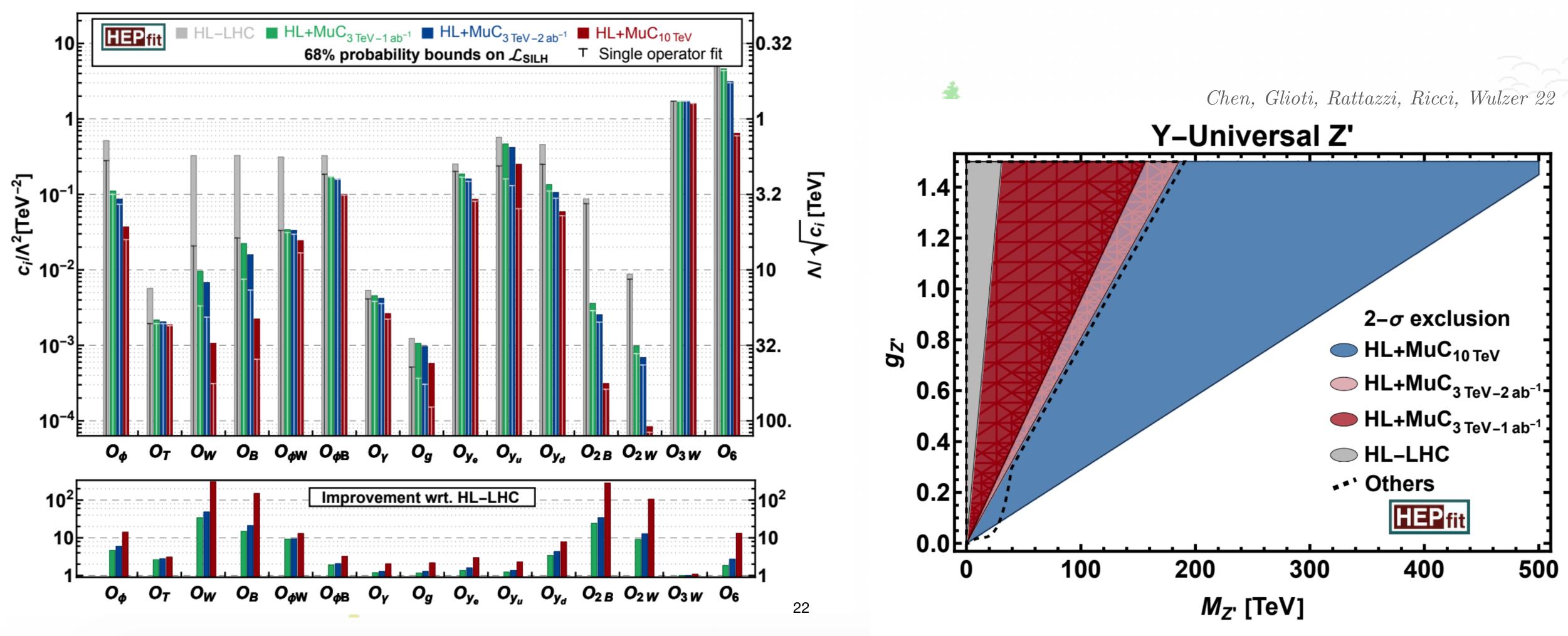
$\frac{\Delta \mathcal{O}}{\mathcal{O}} = 0.01 \approx \frac{E^2}{\Lambda^2}$ $E \sim 10 \text{ TeV}$ $\Lambda \sim 100 \text{ TeV}$

Can still be probing new physics at much higher scales!



PHYSICS REACH OF MUC: INDIRECT

EFT Approach for Energy \leftrightarrow Precision







Any thing else we could do with a muon collider?

















÷.

Any thing else we could do with a muon collider?











Yes! See Zahra's Talk Today



- experiment ready to build
- The run parameters of a MuC should be flexible depending on the status of other colliders • We can explore new physics at the 10 TeV scale *soon* and in a
- clean environment











• A muon collider is a technology concept, not a fully on-shell







- experiment ready to build
- The run parameters of a MuC should be flexible depending on the status of other colliders • We can explore new physics at the 10 TeV scale *soon* and in a
- clean environment
- $(\checkmark \checkmark)$ What makes a MuC special is our ability to sprint to the 10+ TeV frontier. Why should we do luminosity staging?











- experiment ready to build
- The run parameters of a MuC should be flexible depending on the status of other colliders • We can explore new physics at the 10 TeV scale *soon* and in a
- clean environment
- $(\checkmark \checkmark)$ What makes a MuC special is our ability to sprint to the 10+ TeV frontier. Why should we do luminosity staging?









(A)



THANKS!



































Challenges of MuC

Recent Improvements

Cooling

Magnets

Community Interest "Need N Miracles"

MICE, Simulation, Timescales

20 T Dipoles 30 T Solenoids 10³ T/s Ramping

IMCC + US R&D

"No Showstoppers Identified"

ESPPU + P5 \implies Time is now!







Demonstrator Facilities & Beam Dump

Before construction (or approval) of a full collider we need to demonstrate technologies



Staging options are important to ensure a muon collider can collect data even if upgrades need more time for R&D

3 TeV

10 TeV

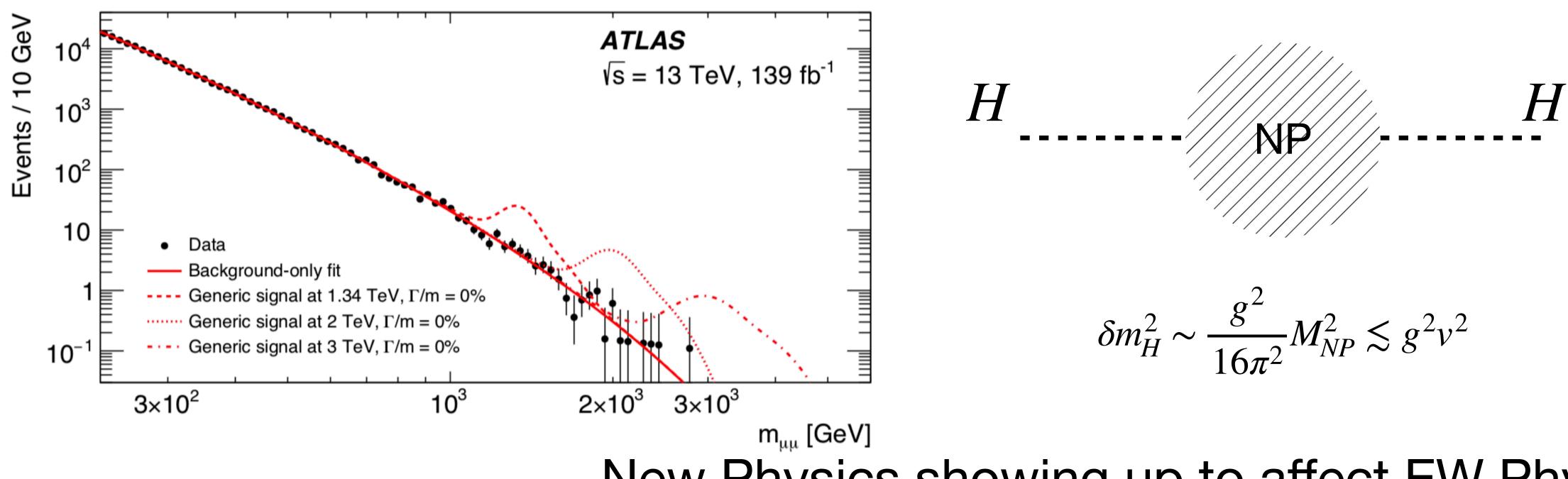
30 TeV

100 TeV

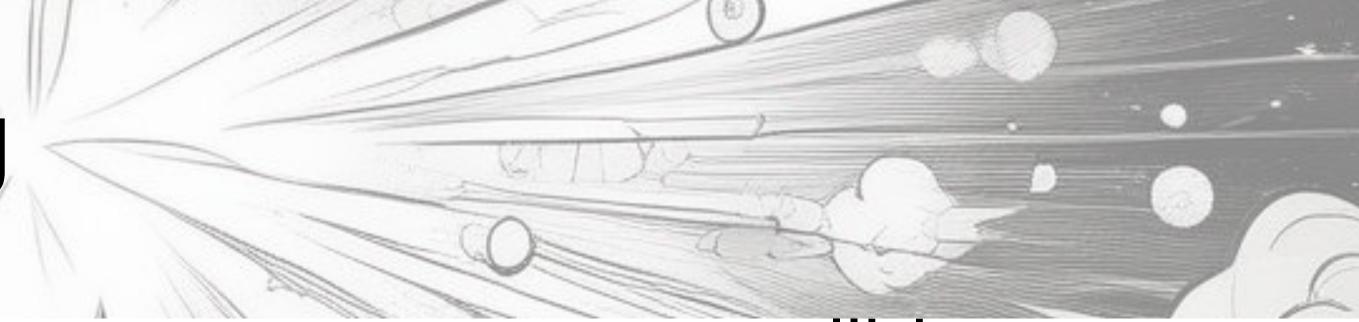


Staging options are important to ensure a muon collider can collect data even if upgrades need more time for R&D

3 TeV



Slide Credit: D. Buttazzo



Beyond typical LHC $\sqrt{\hat{s}}$

New Physics showing up to affect EW Physics?



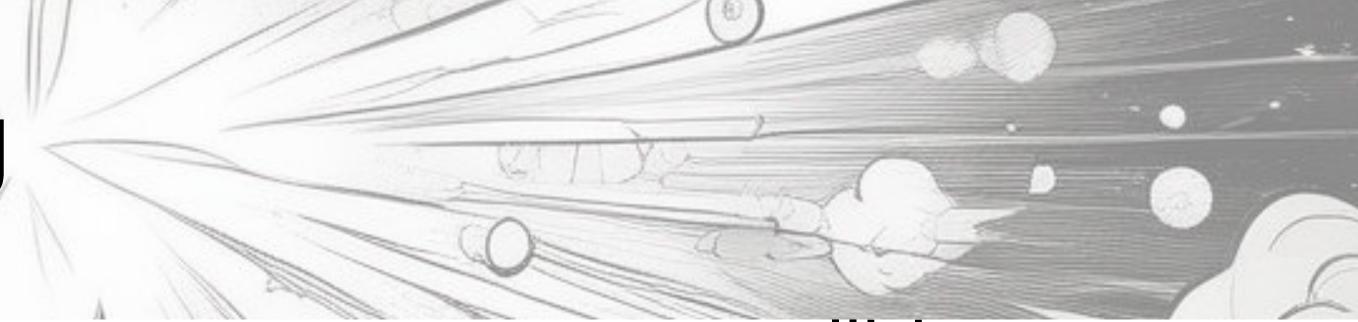
Staging options are important to ensure a muon collider can collect data even if upgrades need more time for R&D

3 TeV

10 TeV See I

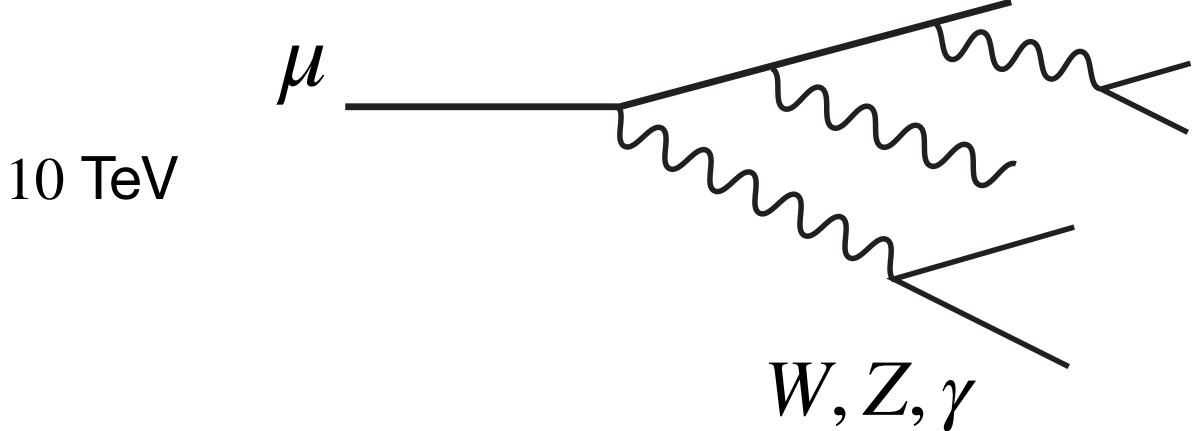
$$\frac{\alpha}{4\pi} \log^2 \left(\frac{E^2}{m_W^2} \right) \times \text{Casimir} \sim 1 \text{ for E } \sim$$

Chen, Glioti, Rattazzi, Ricci, Wulzer 2202.10509

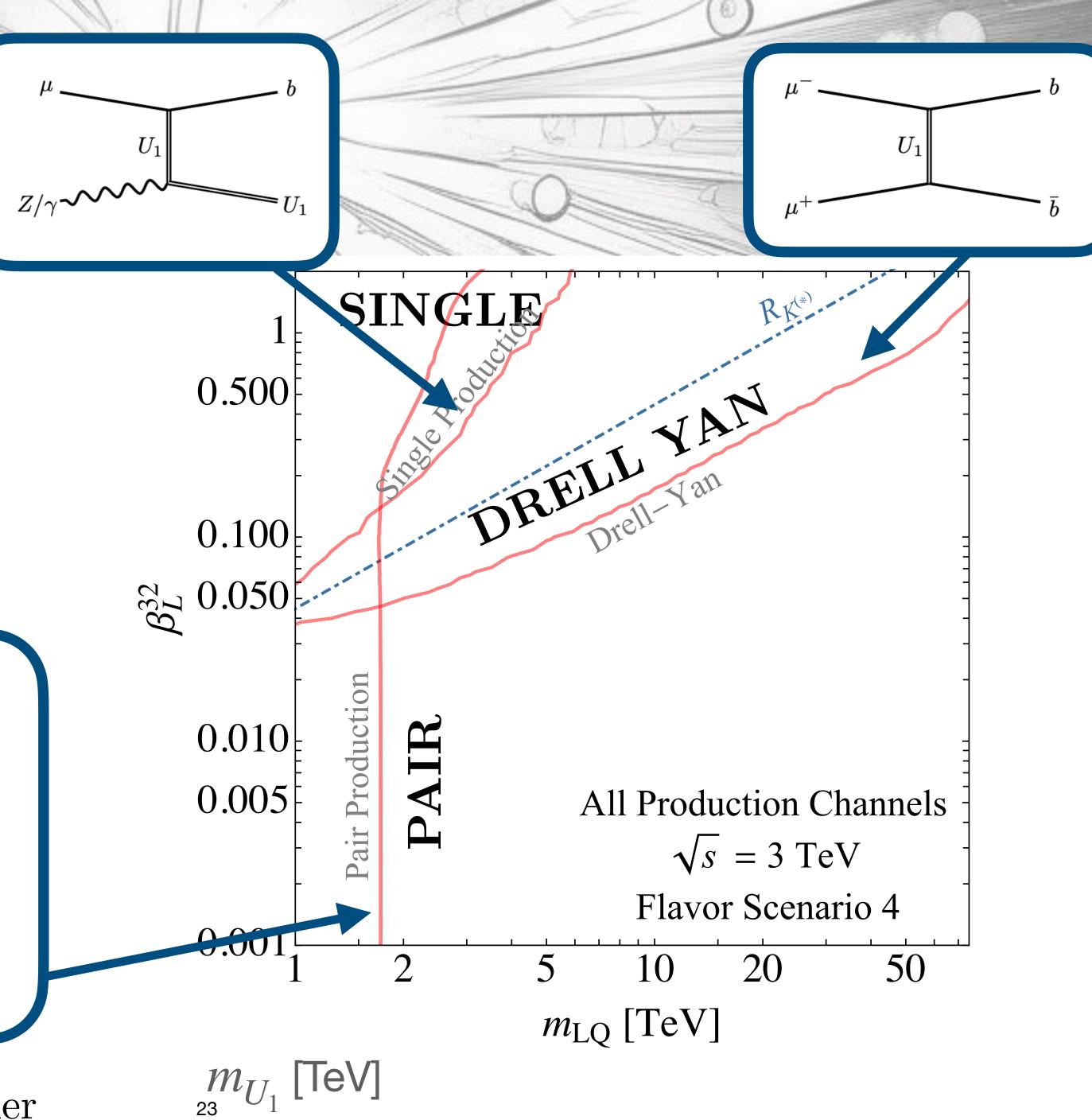


Beyond typical LHC $\sqrt{\hat{s}}$

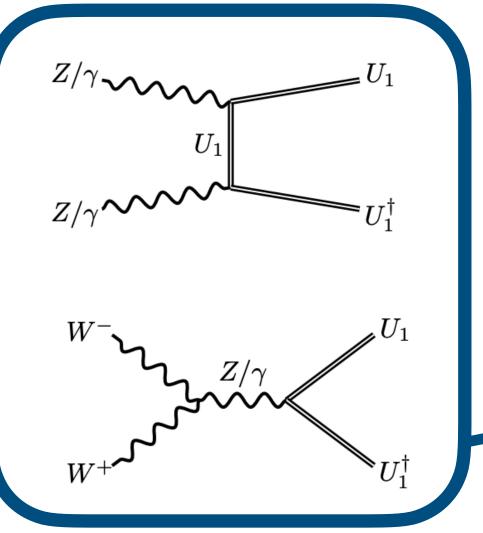
See new SM phenomena (e.g. EW jets)



Ex: Leptoquarks



Minimal U_1 Leptoquark EFT $U_1 = (3,1)_{2/3}$ 5σ confidence limits $3 \text{ TeV } \mu C$



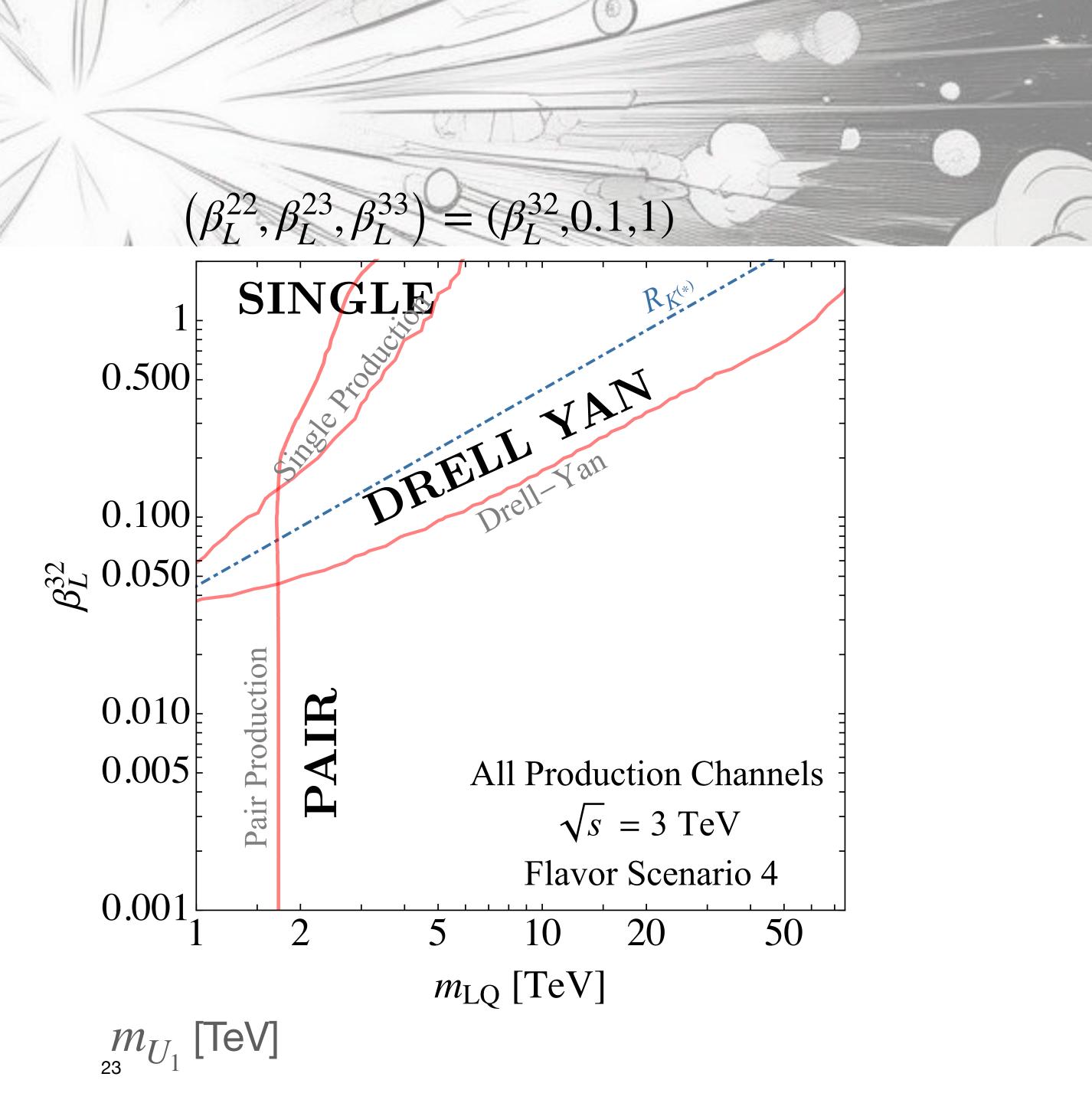
2104.05720 ${\bf CC},$ Asadi, Capdevilla, Homiller



Ex: Leptoquarks

Minimal U_1 Leptoquark EFT $U_1 = (3,1)_{2/3}$ 5σ confidence limits $3 \text{ TeV } \mu C$

2104.05720 ${\bf CC},$ Asadi, Capdevilla, Homiller



Ex: New Physics with Higgs Mixing

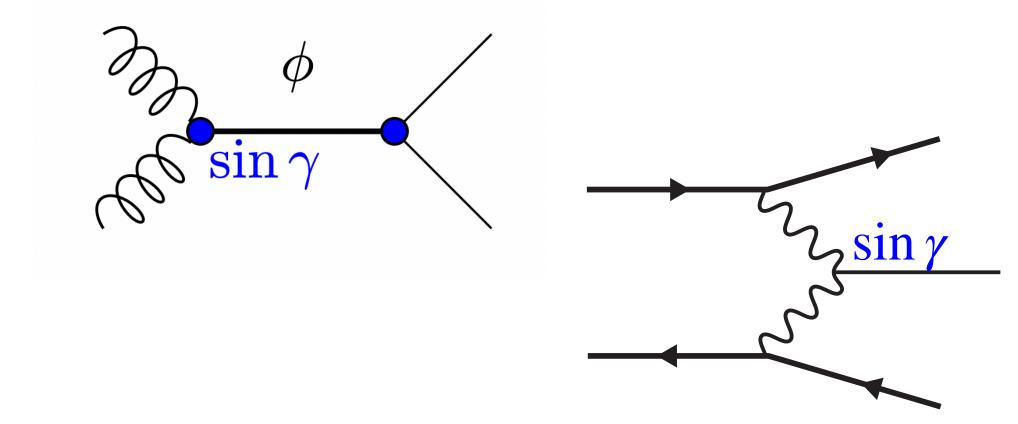
Benchmark model: New singlet S mixes with Higgs

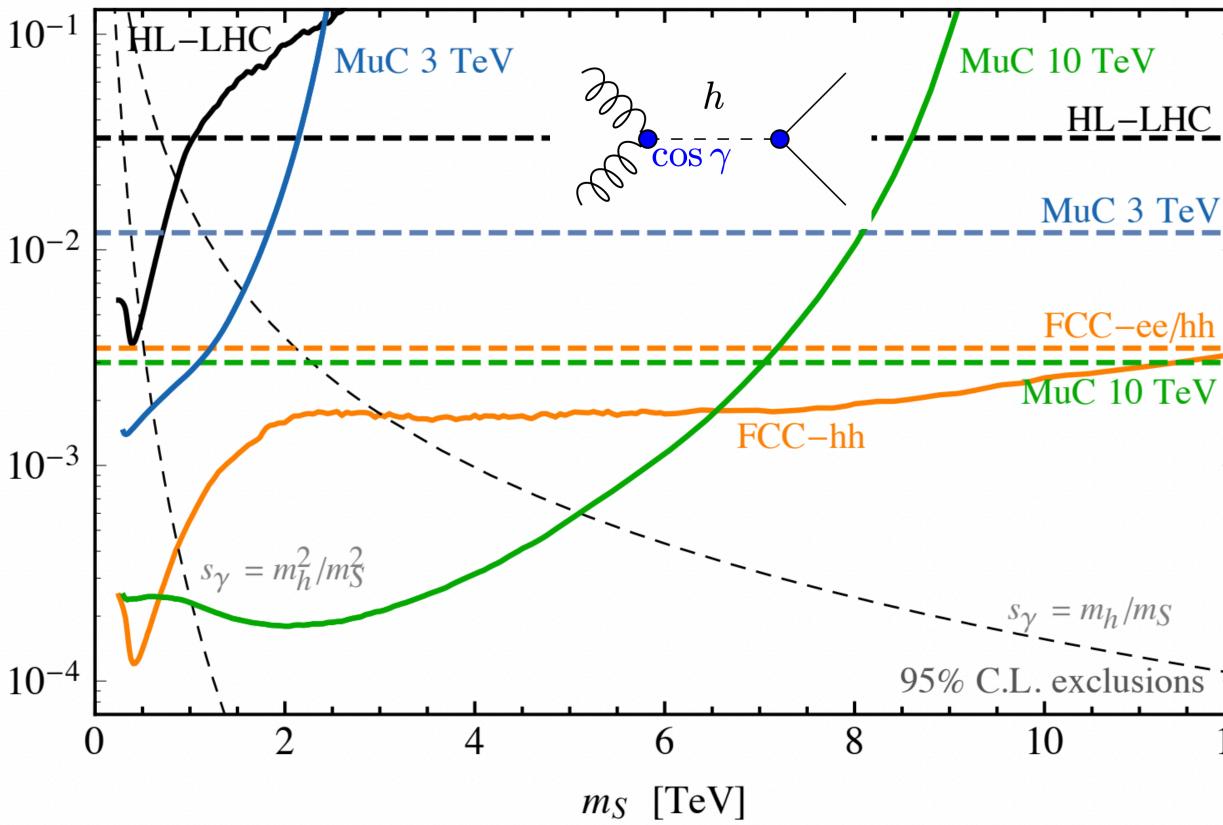
 $\sin^2 \gamma$

$$h = h_0 \cos \gamma + S \sin \gamma$$

$$\phi = S\cos\gamma - h_0\sin\gamma$$

$$\phi \rightarrow hh, ZZ, WW$$





Buttazzo, Redigolo, Sala, Tesi 1807.04743







Ex: New Physics with Higgs Mixing

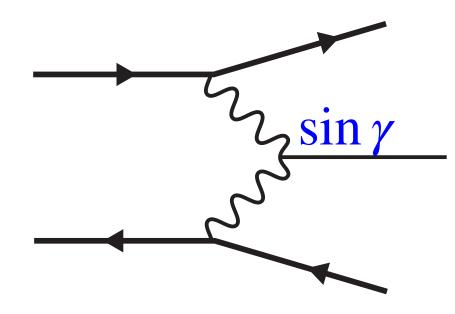
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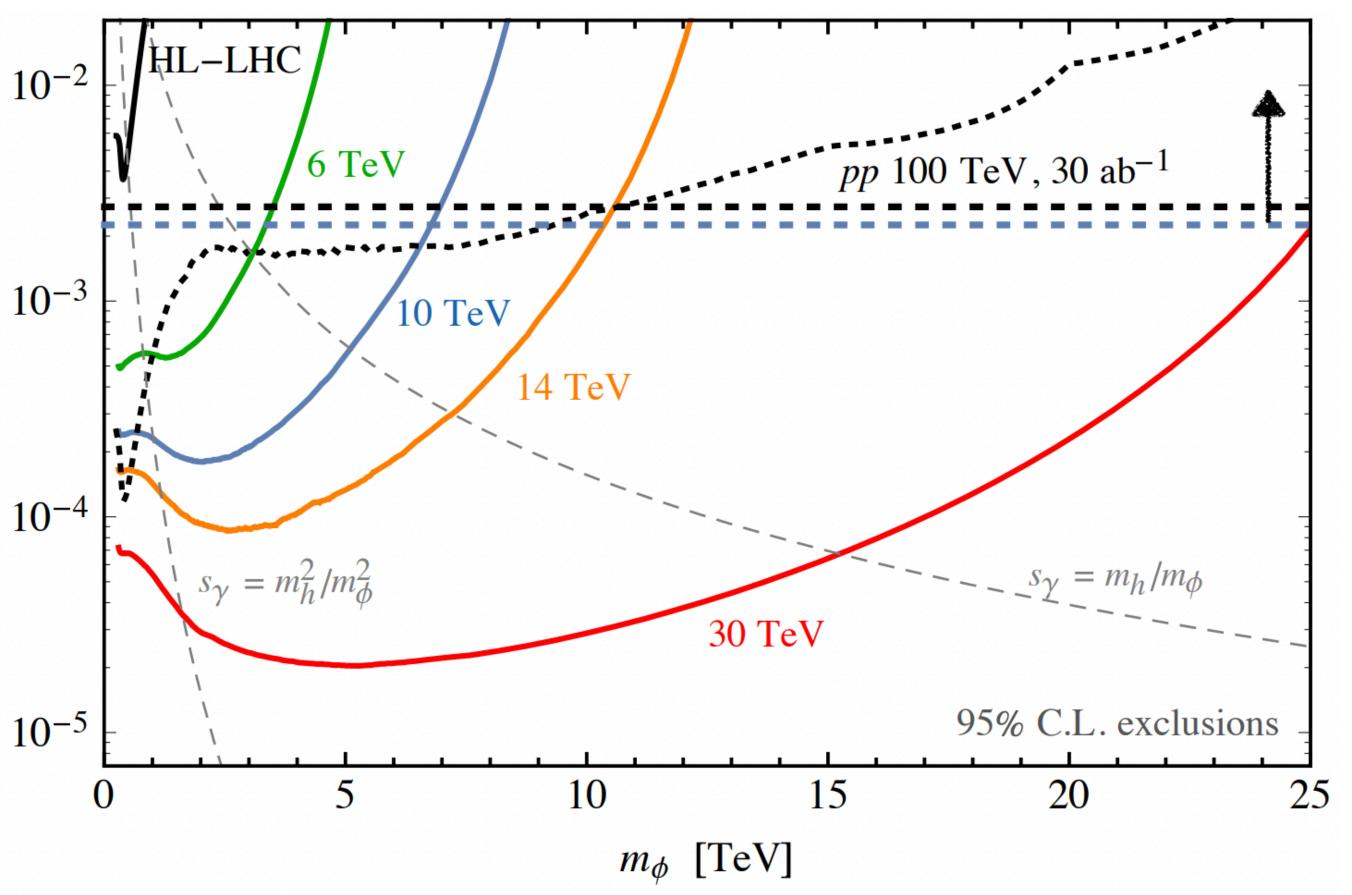
 $\sin^2 \gamma$

$$h = h_0 \cos \gamma + S \sin \gamma$$
¹⁰⁻

$$\phi = S\cos\gamma - h_0\sin\gamma$$

 $\phi \rightarrow hh, ZZ, WW$





Buttazzo, Redigolo, Sala, Tesi 1807.04743





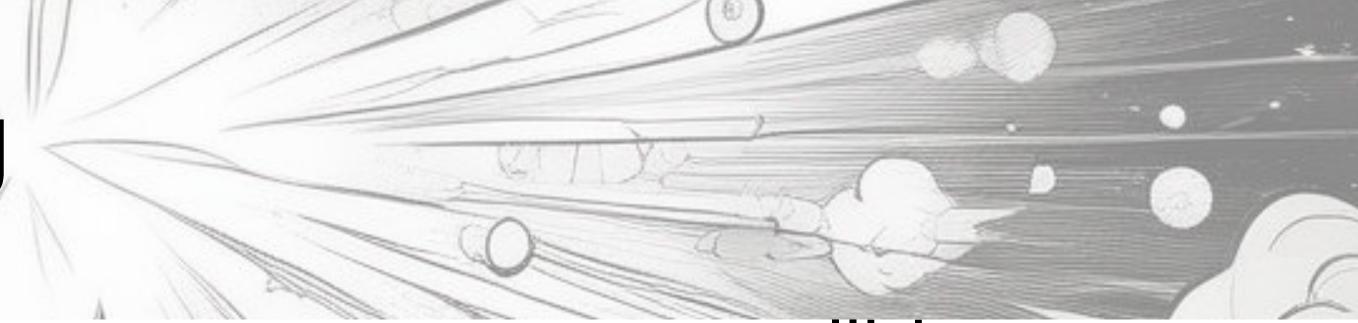
Staging options are important to ensure a muon collider can collect data even if upgrades need more time for R&D

3 TeV

10 TeV

30 TeV

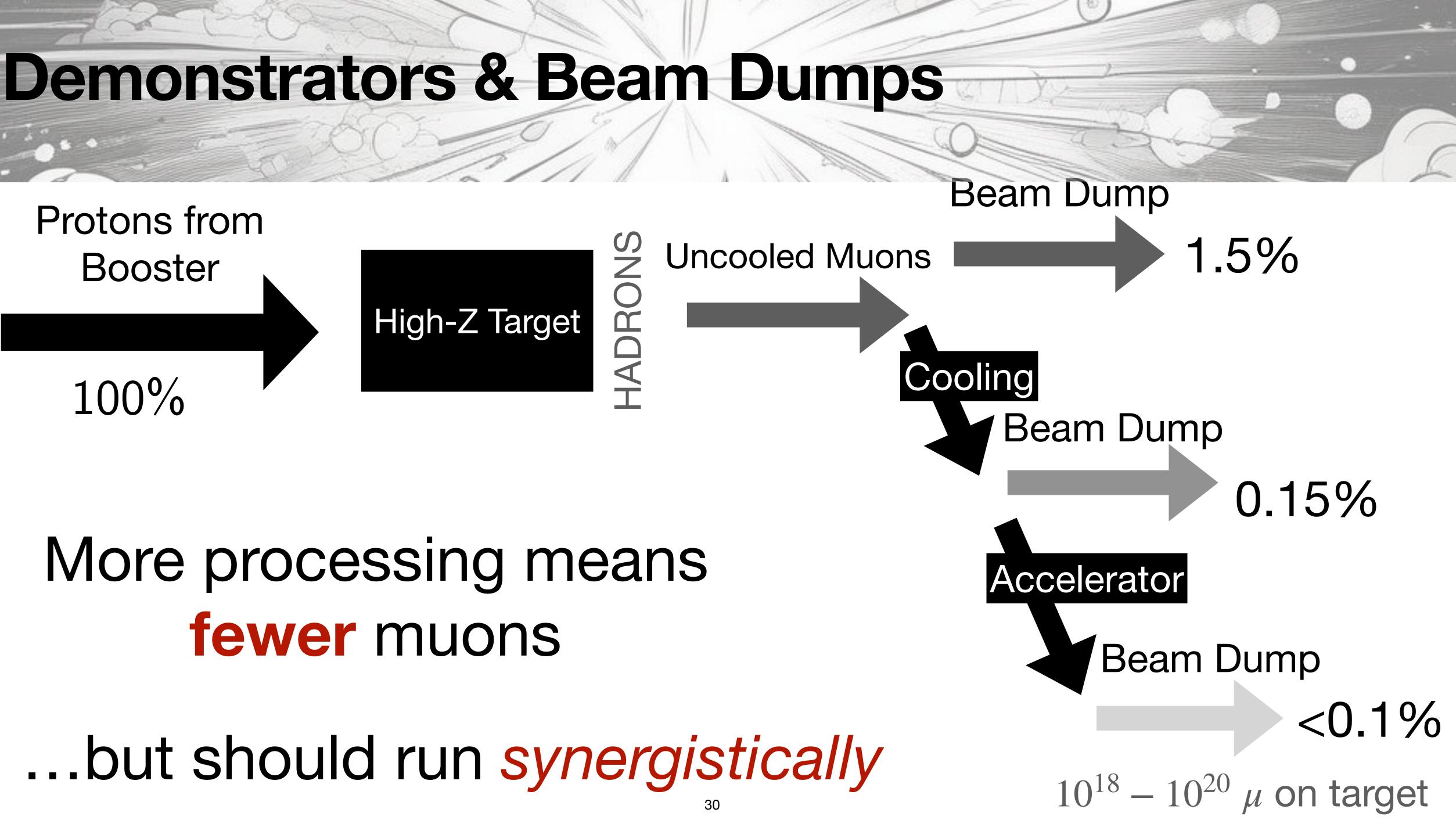
100 TeV



Beyond typical LHC $\sqrt{\hat{s}}$

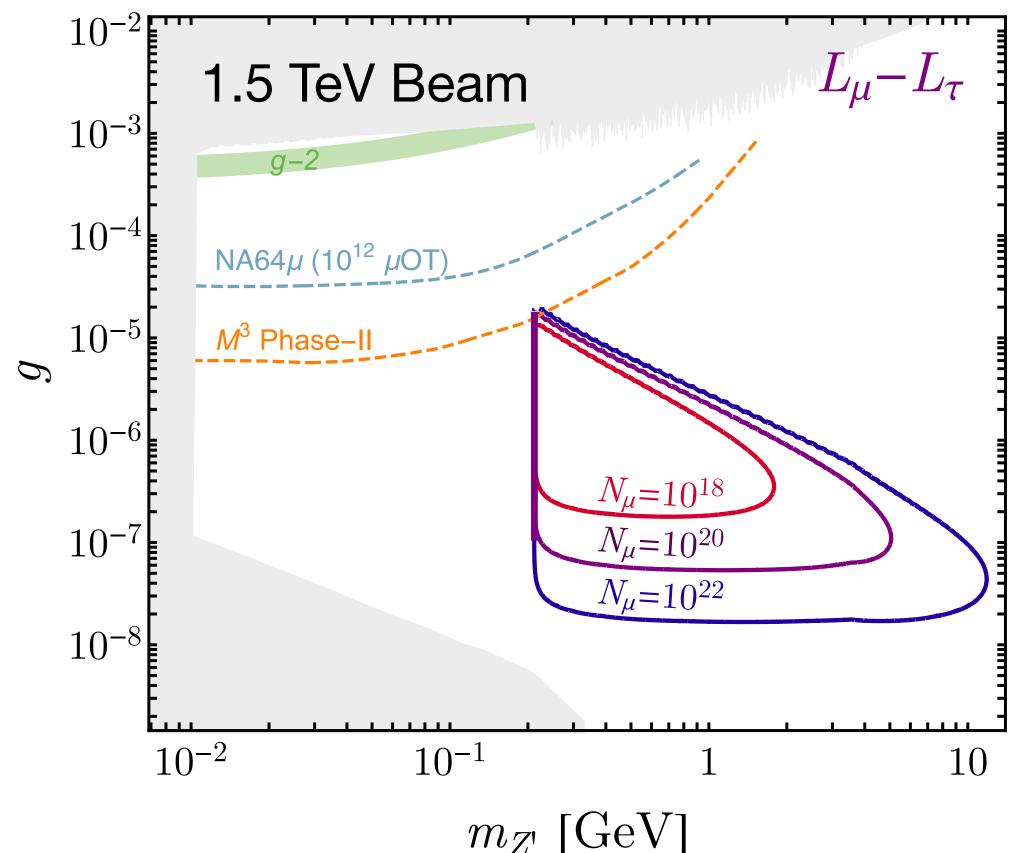
See new SM phenomena (e.g. EW jets)

Reach Thermal Target DM Candidates Dramatically new Energy Frontier Limits of reasonable projections?

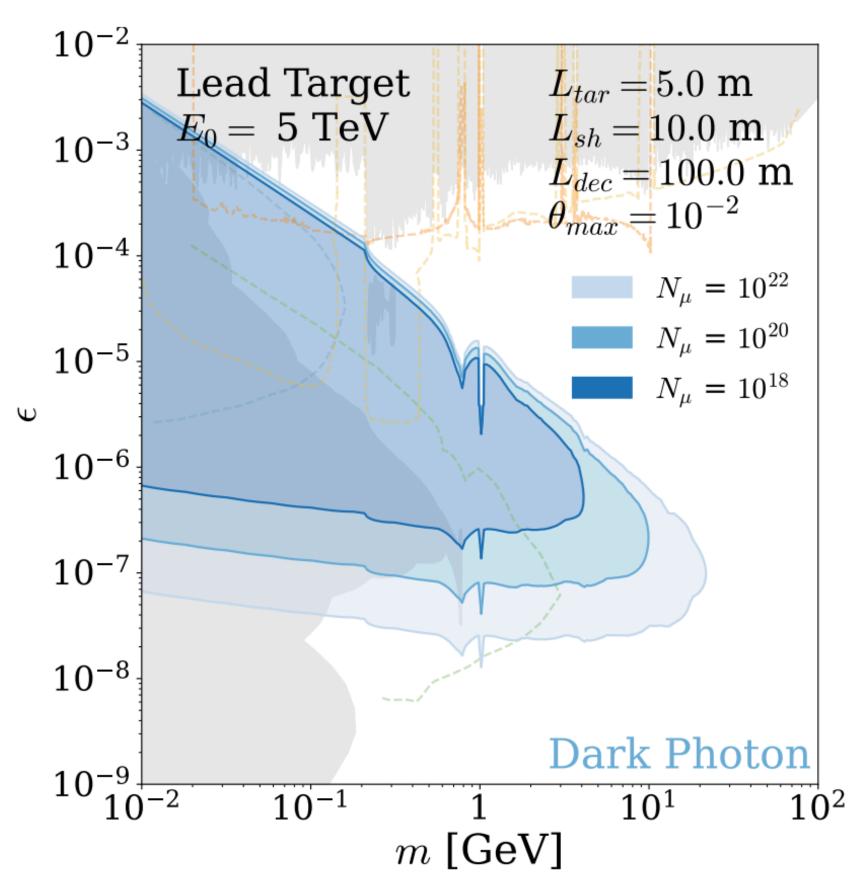


Muon Collider Beam Dump 2050ish 3, 10 TeV MuC **Dark Photon**

2202.12302



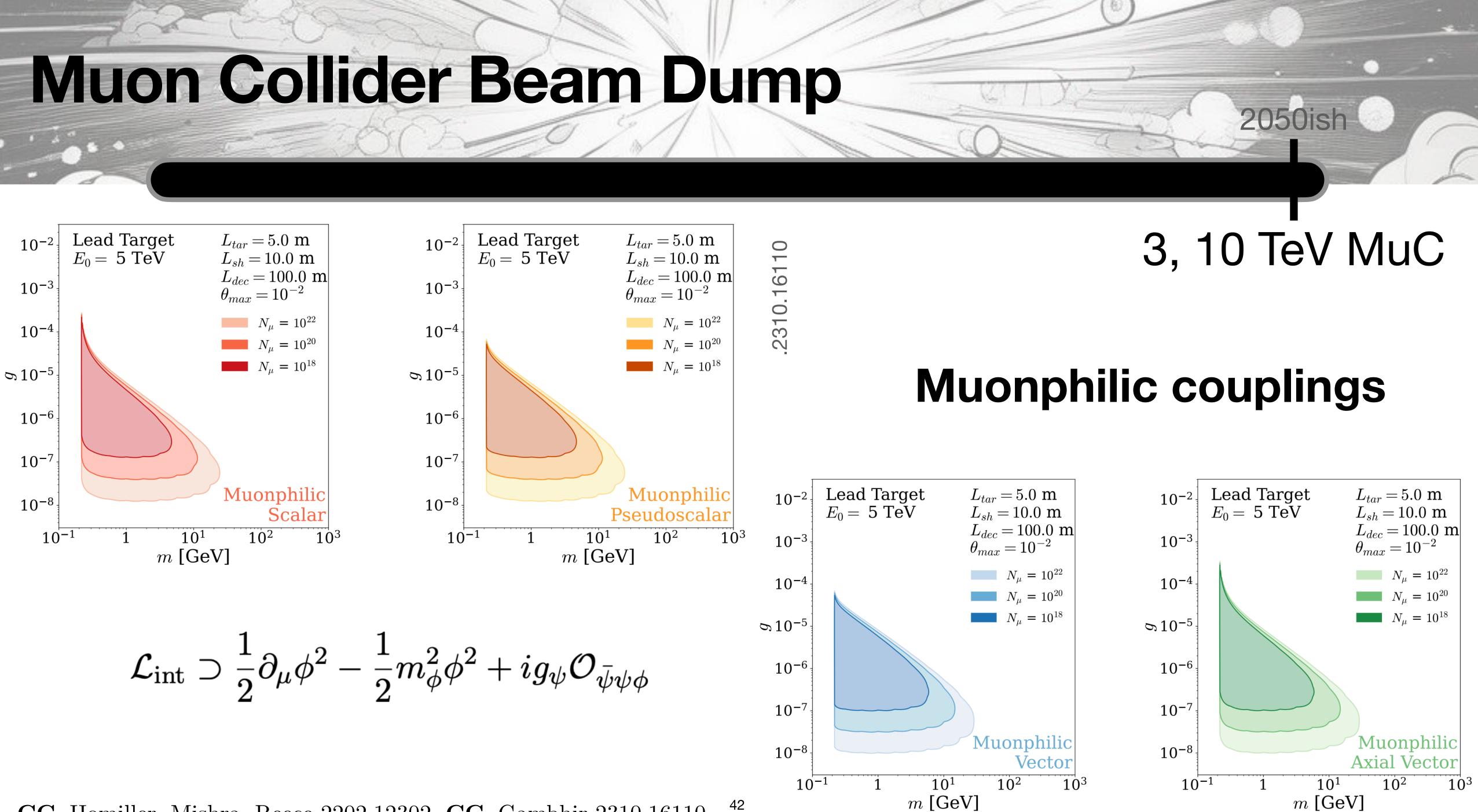
CC, Homiller, Mishra, Reece 2202.12302, CC, Gambhir 2310.16110



2310.16110

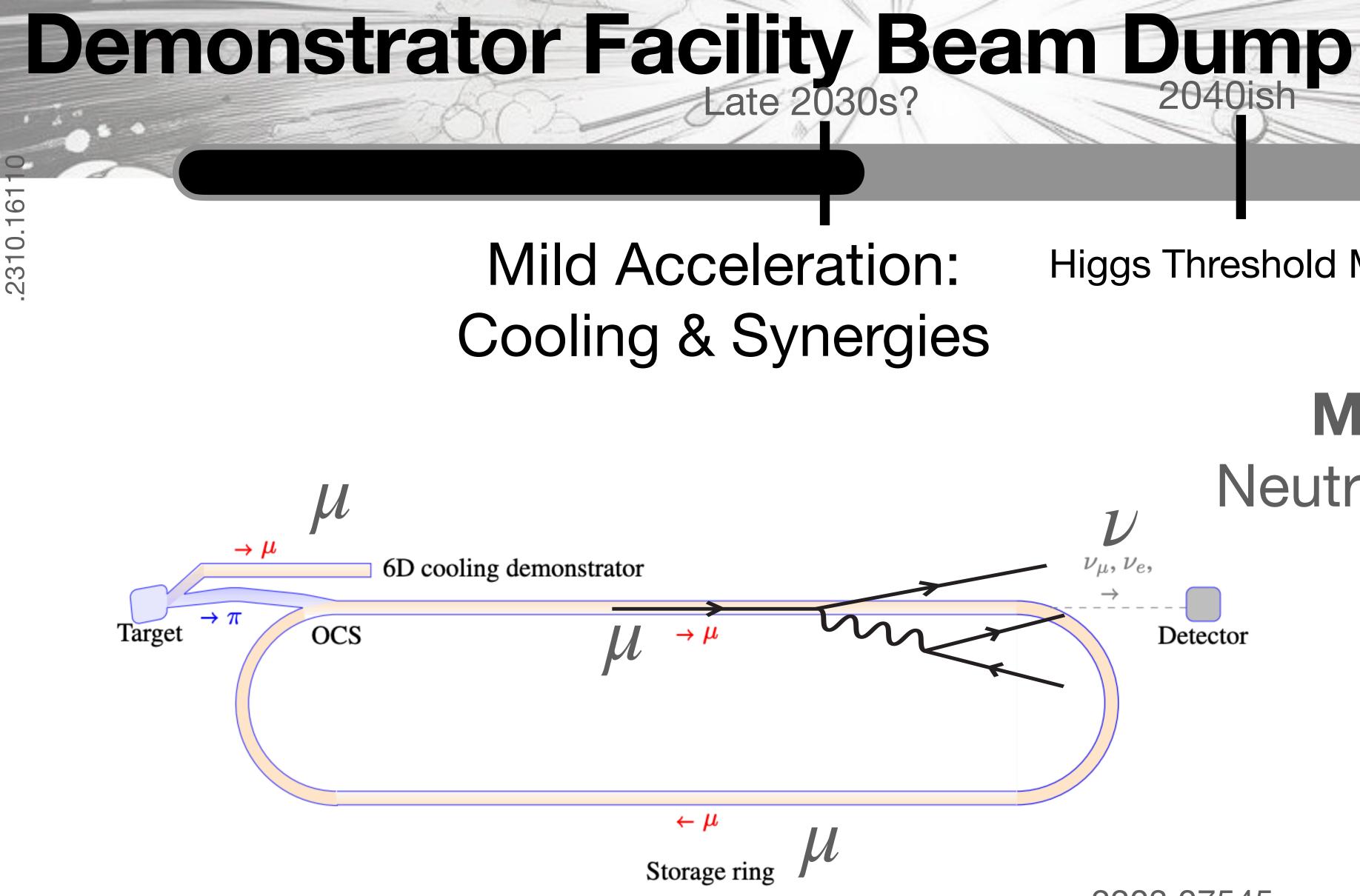






$$\mathcal{L}_{\rm int} \supset \frac{1}{2} \partial_{\mu} \phi^2 - \frac{1}{2} m_{\phi}^2 \phi^2 + i g_{\psi} \mathcal{O}_{\bar{\psi}\psi\phi}$$

CC, Homiller, Mishra, Reece 2202.12302, CC, Gambhir 2310.16110



2040ish

3, 10 TeV MuC Higgs Threshold MuC

Motivation: **Neutrino Facilities?**

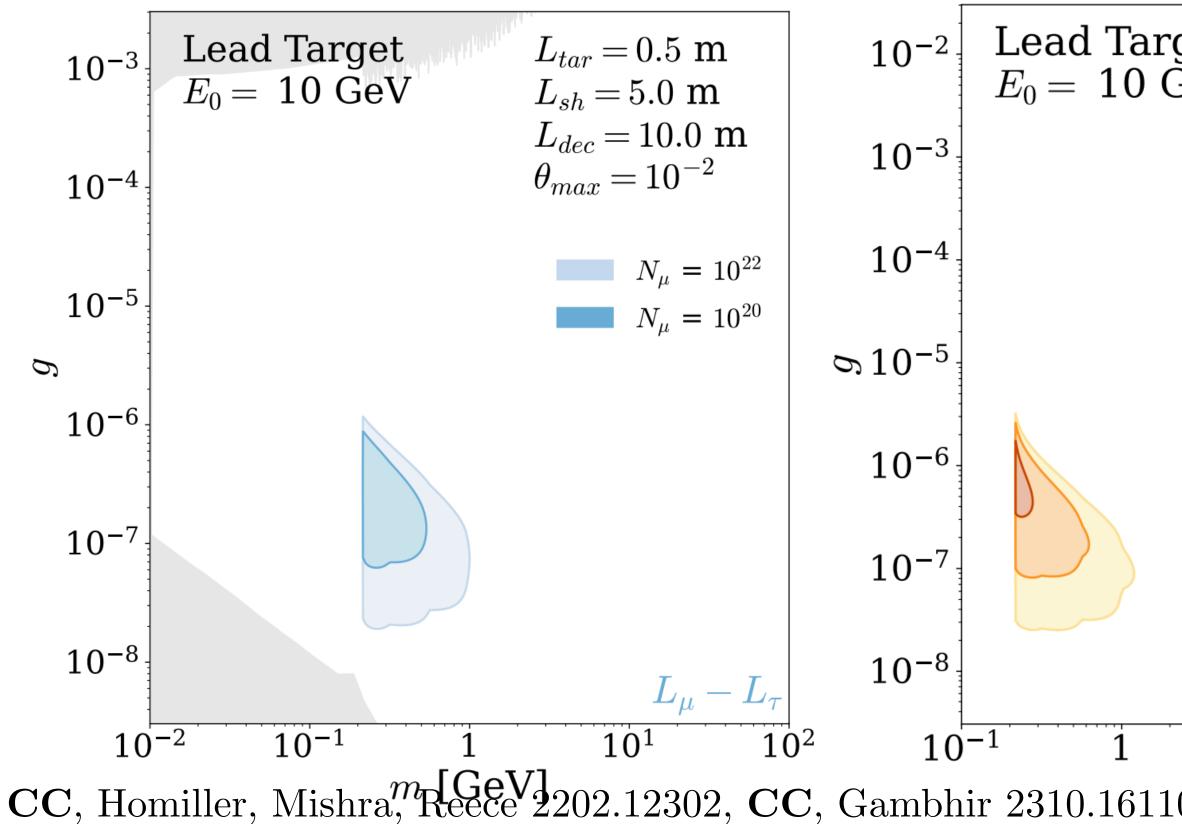
2203.07545



2050ish

Demonstrator Facility Beam Dump Late 2030s? 2040ish

Mild Acceleration: Cooling & Synergies



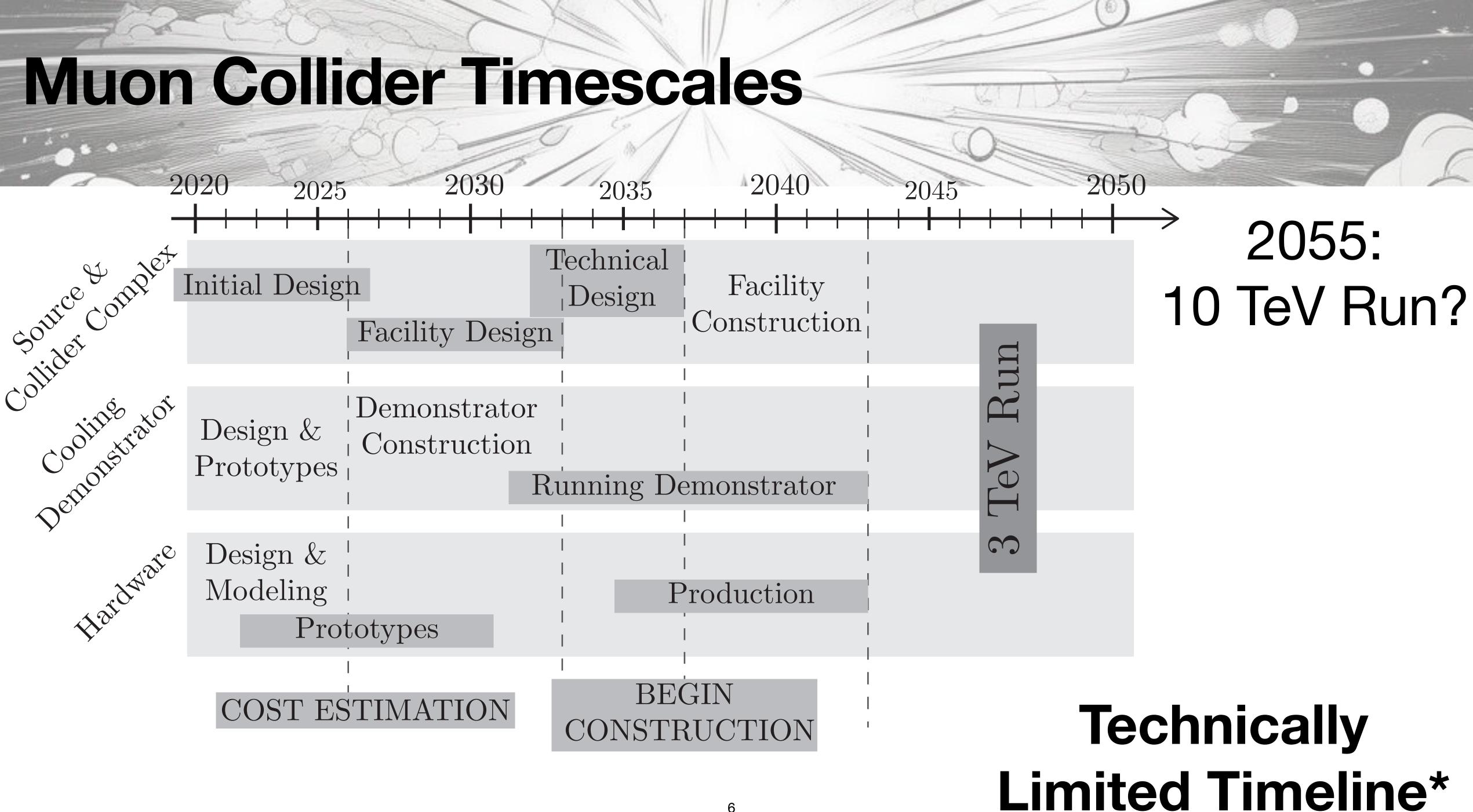
ion: Higgs Threshold MuC 3, 10 TeV MuC rgies

rget GeV	$L_{tar}{=}0.5$ 1 $L_{sh}{=}5.0$ n	n 10 n	$b^{-2} \begin{bmatrix} L \\ E \end{bmatrix}$	ead Tar $G_0 = 10$ C	get GeV	$L_{tar} = 0.5 \text{ m} \ L_{sh} = 5.0 \text{ m}$
	$L_{dec} = 10.0$ $ heta_{max} = 10^-$	m)-3			$L_{dec} = 10.0 \text{ m}$ $ heta_{max} = 10^{-2}$
	$N_{\mu} = 1$ $N_{\mu} = 1$)-4			$N_{\mu} = 10^{22}$ $N_{\mu} = 10^{20}$
	$N_{\mu} = 1$	0^{18} 510	-5			$N_{\mu} = 10^{18}$
		10) -6			
		10)-7			
	Muonph Pseudosca)-8			Muonphili Scala
10 10m@C		103	10^{-1}	1	10^1 m [GeV	10 ²]

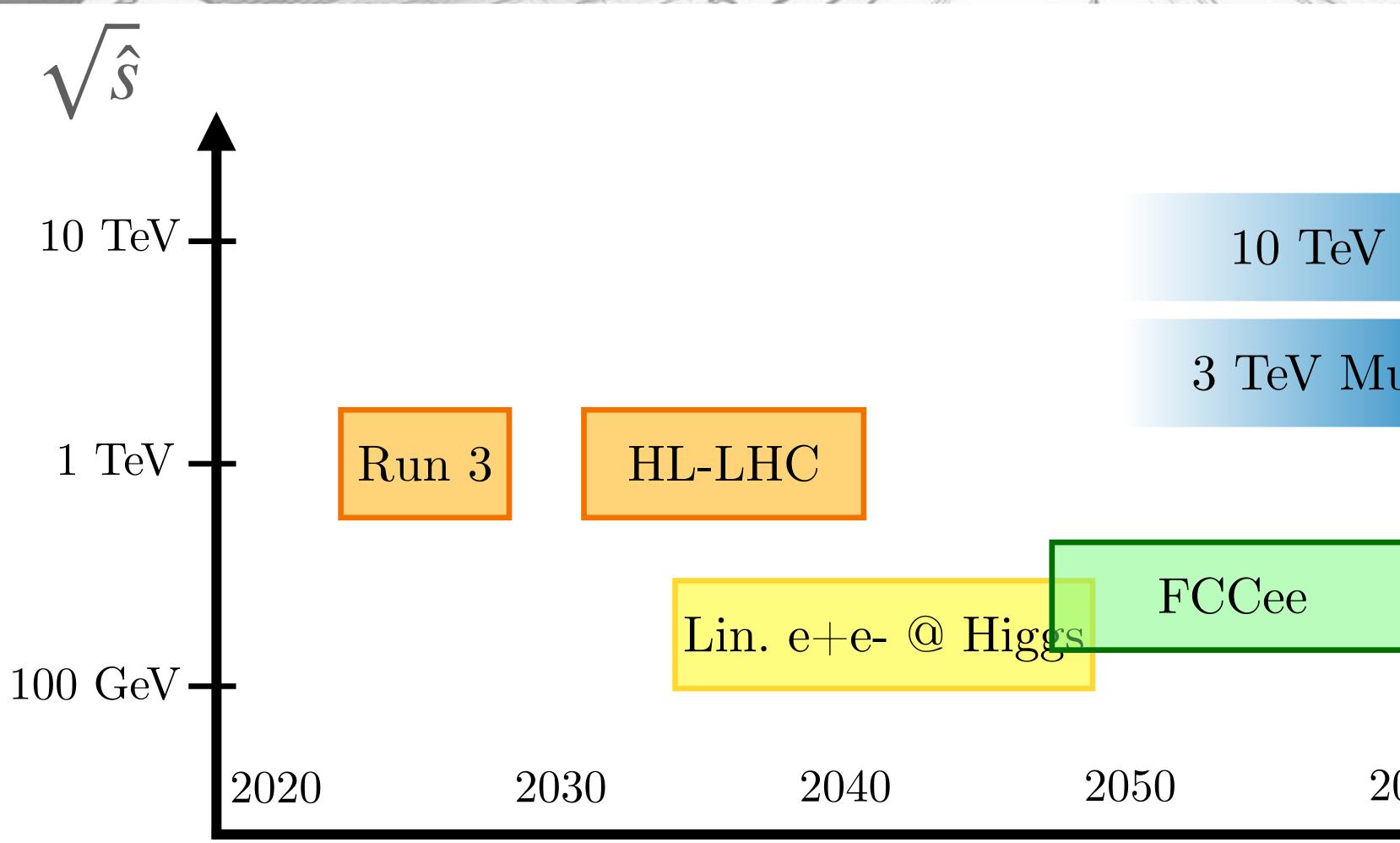


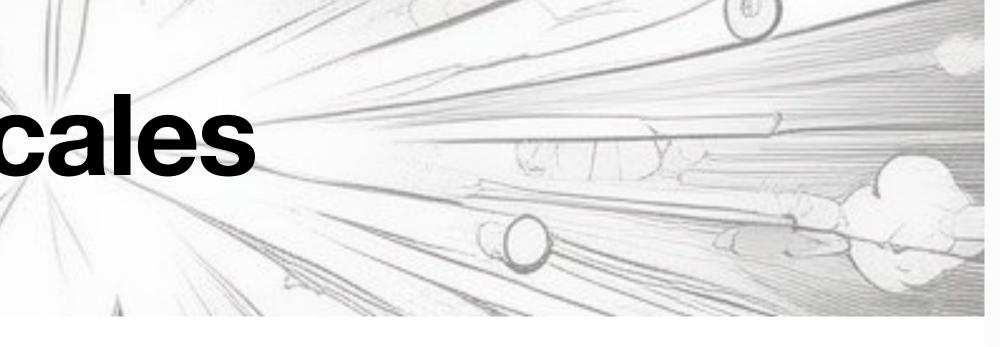
2050ish

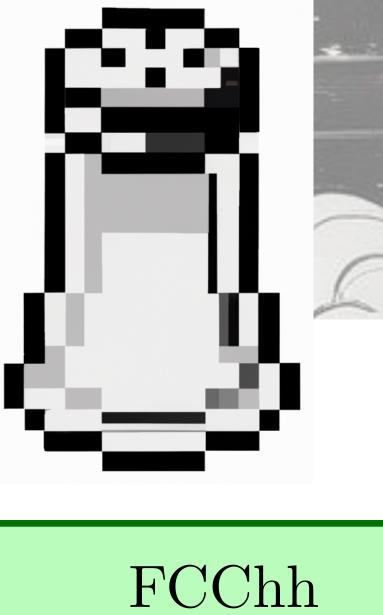




Future Collider Timescales







10 TeV MuC

3 TeV MuC

