



Accelerator based dark matter probes

Kate Pachal
TRIUMF

Introduction

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- How we might think about framing this to the community

Medium-mass dark matter

Today, largely talking about cases with dark matter and/or associated new particles over ~ 1 GeV in mass

Medium-mass dark matter

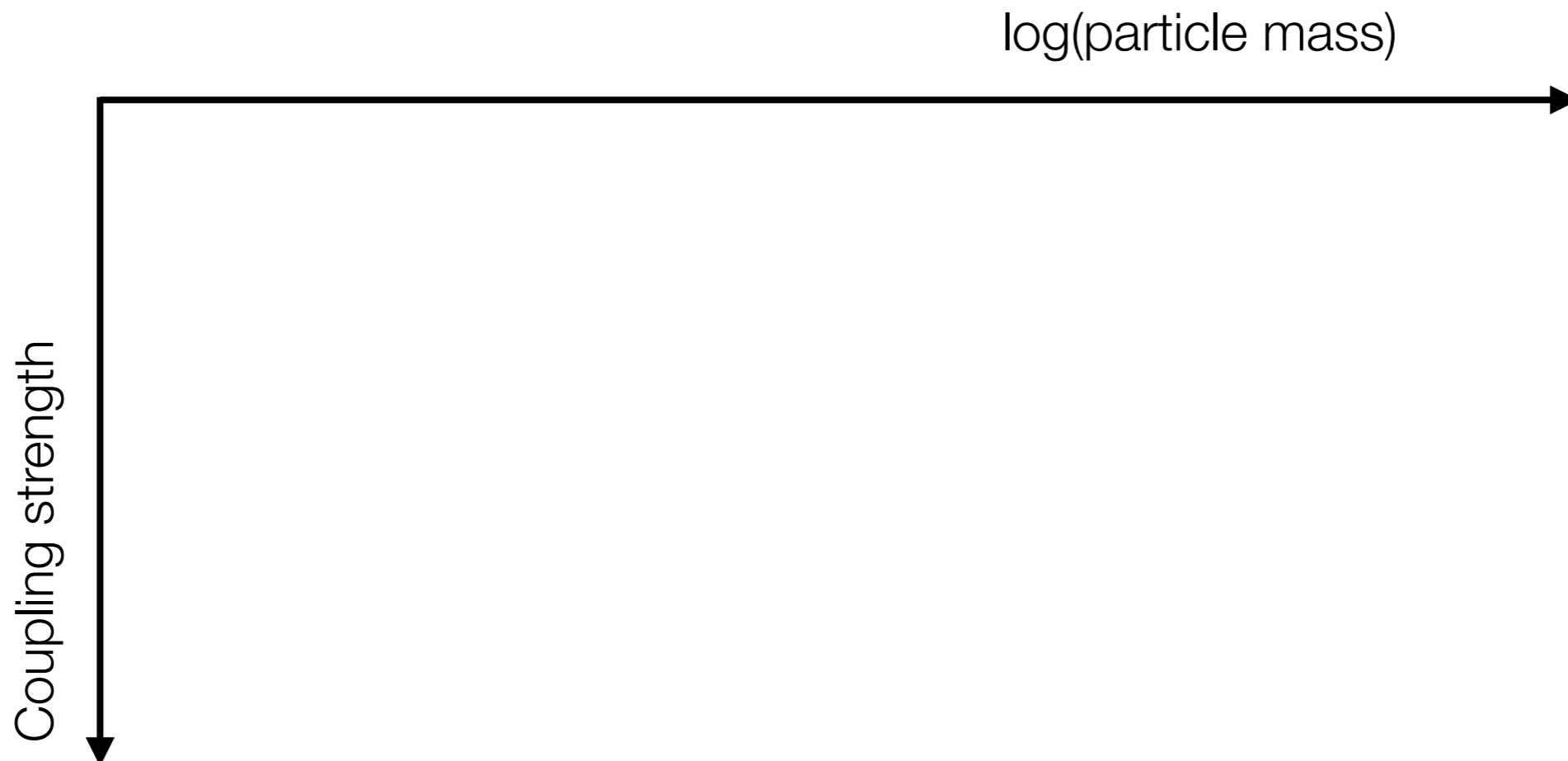
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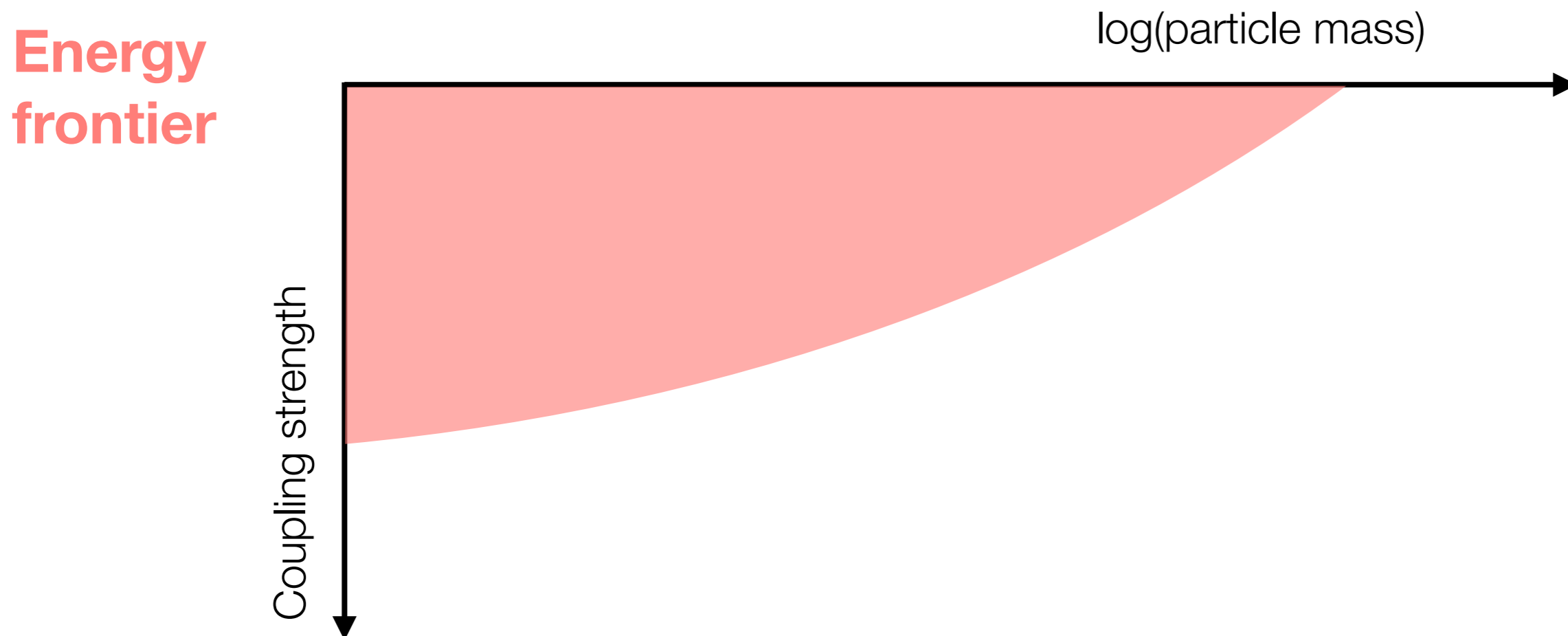
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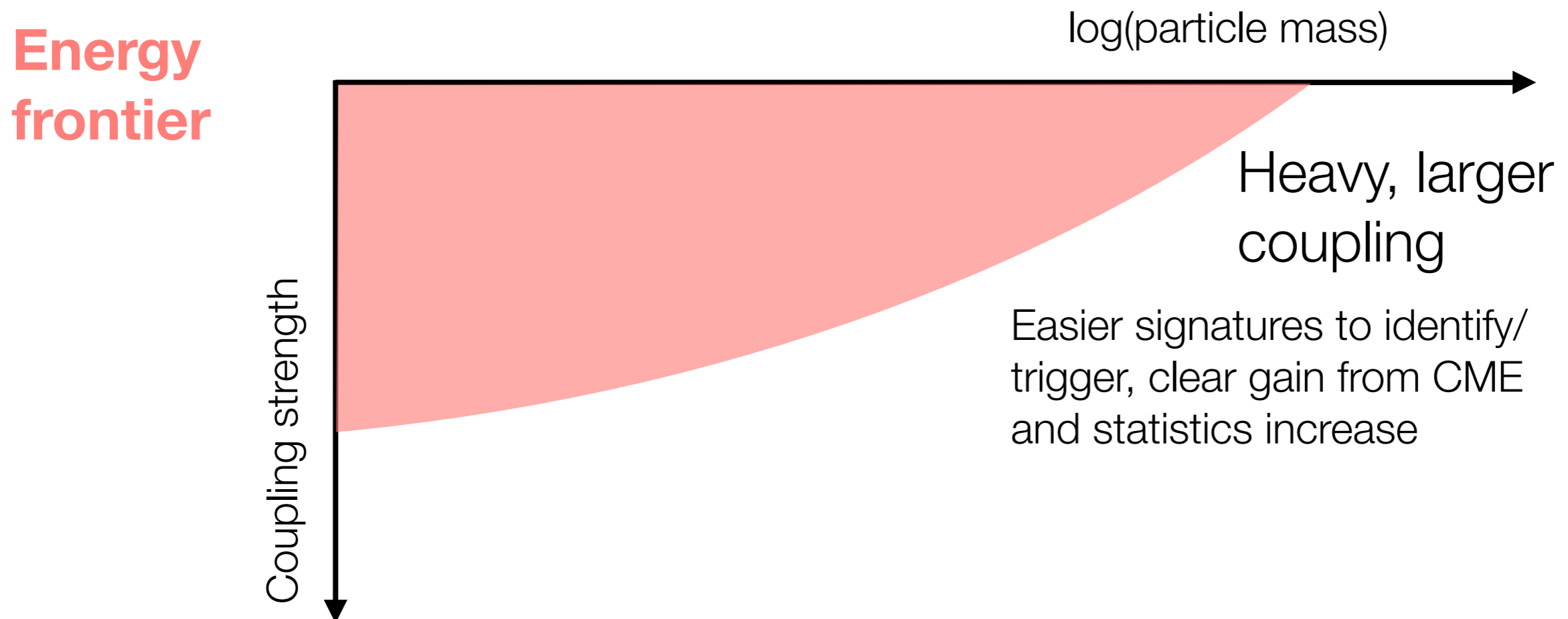
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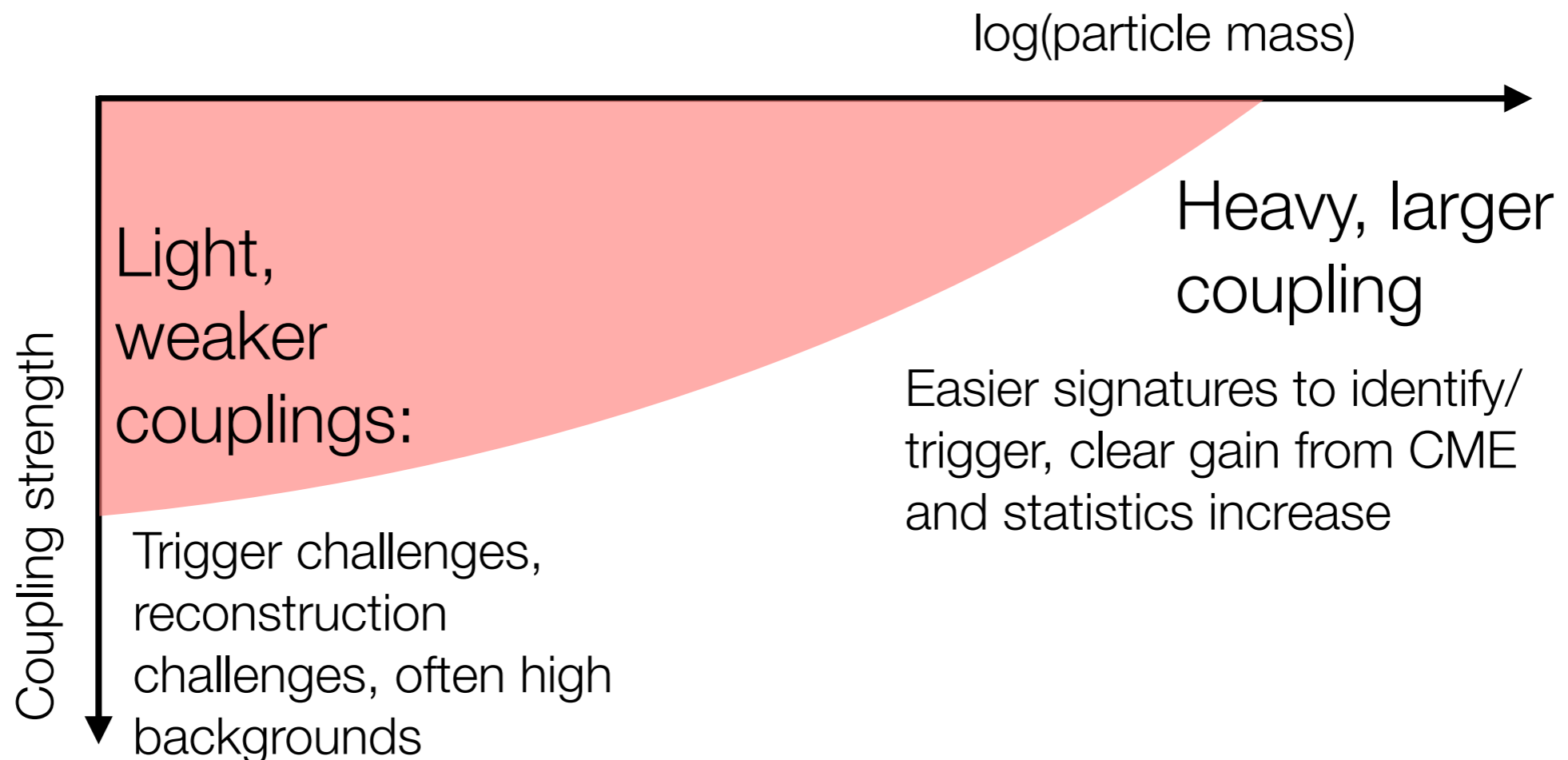


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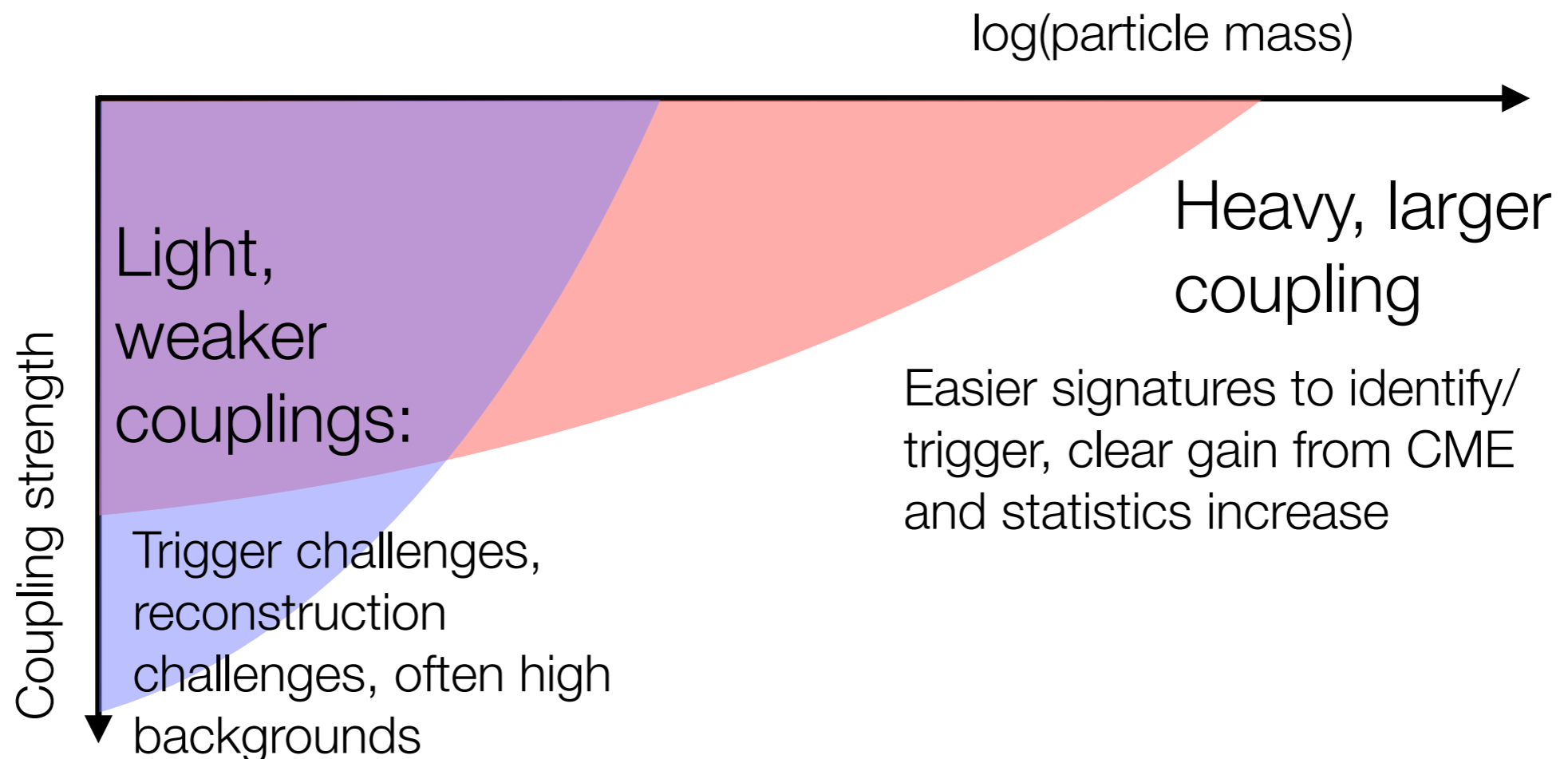


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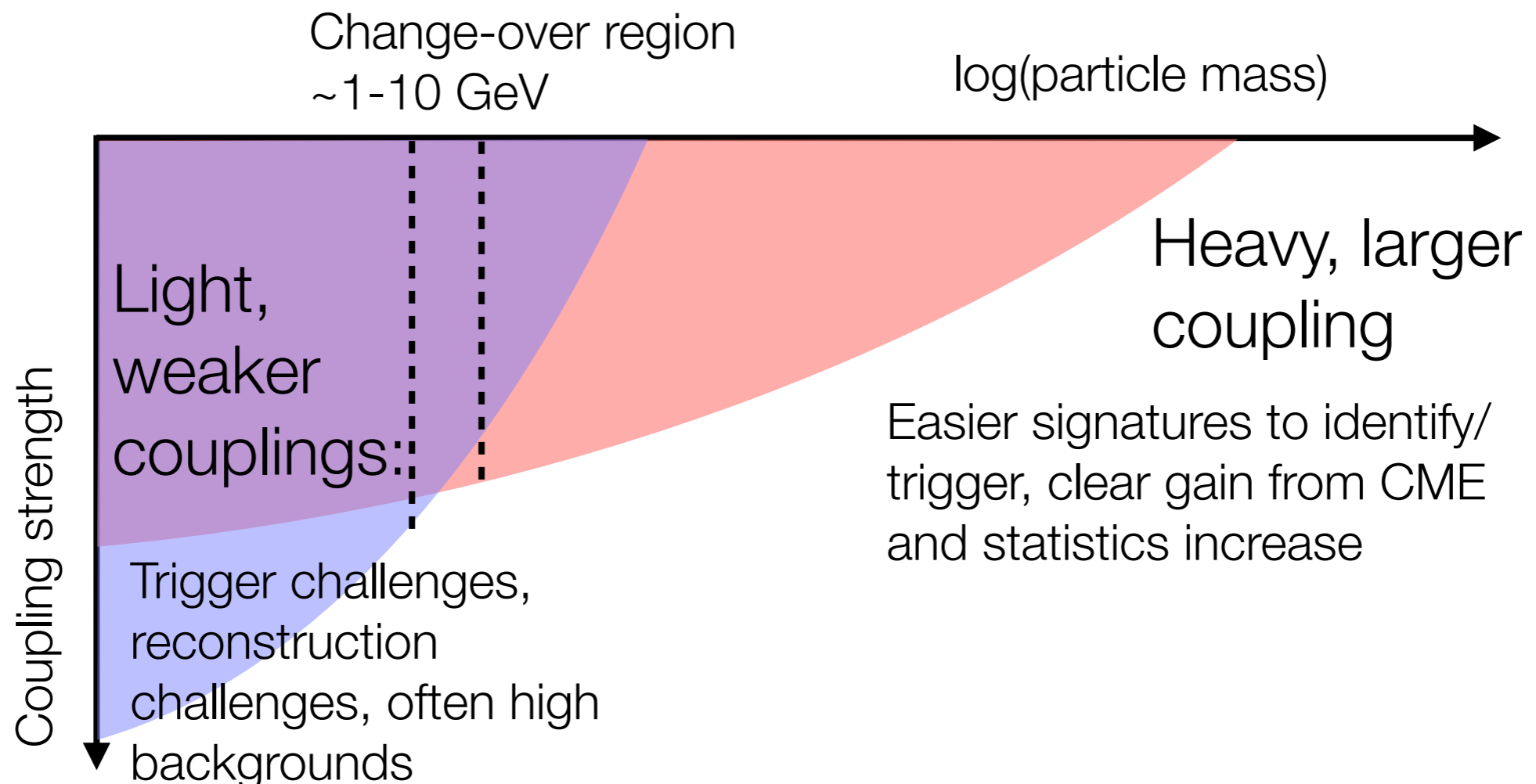


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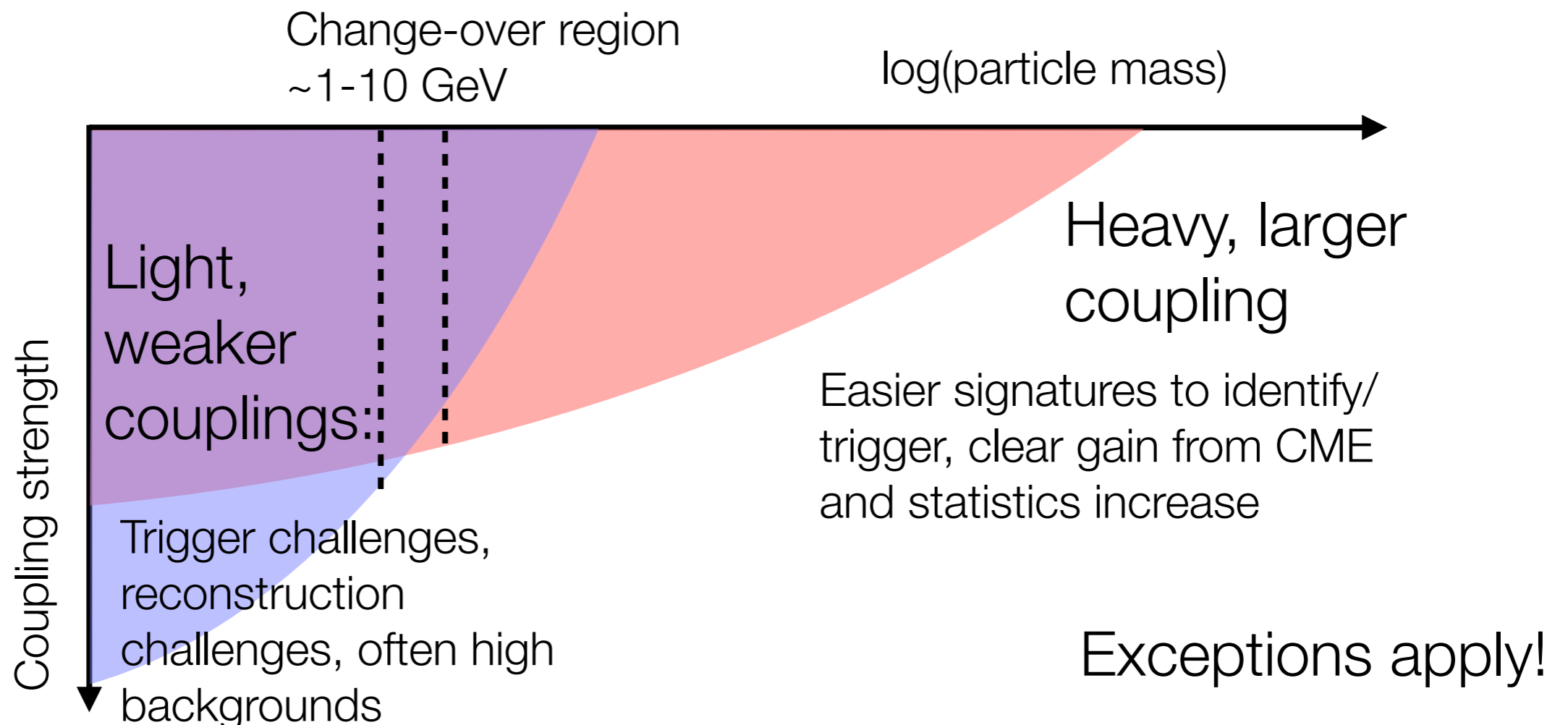


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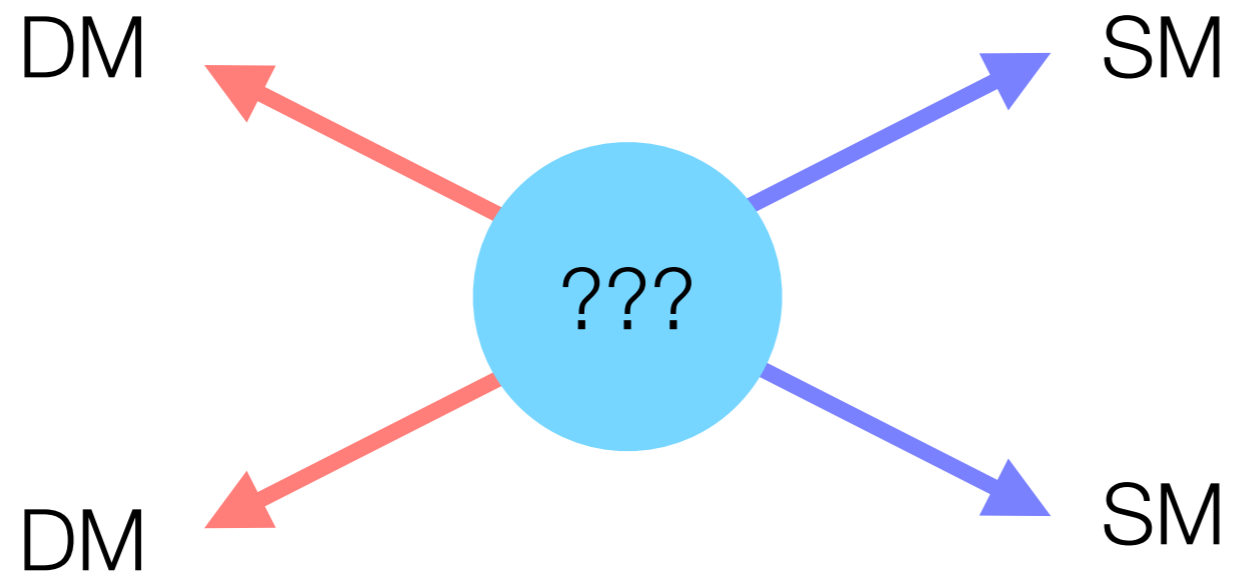
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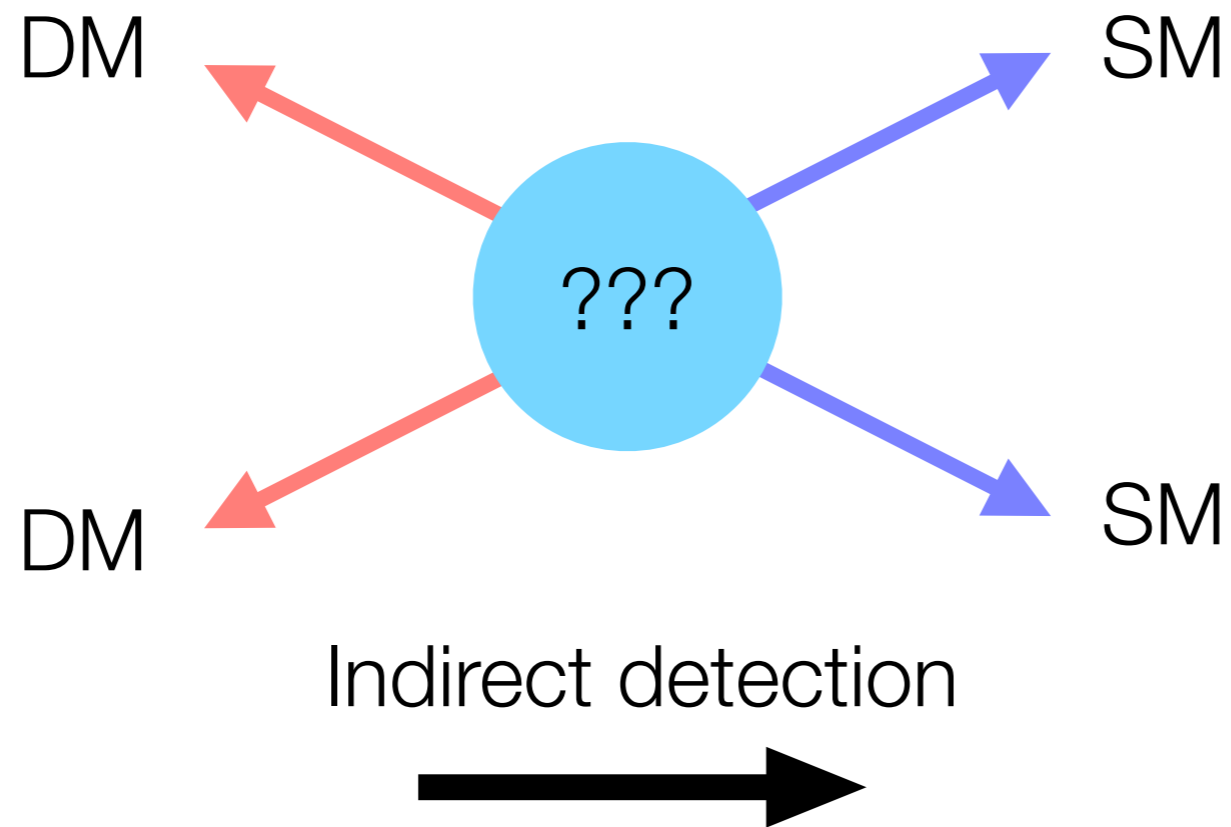
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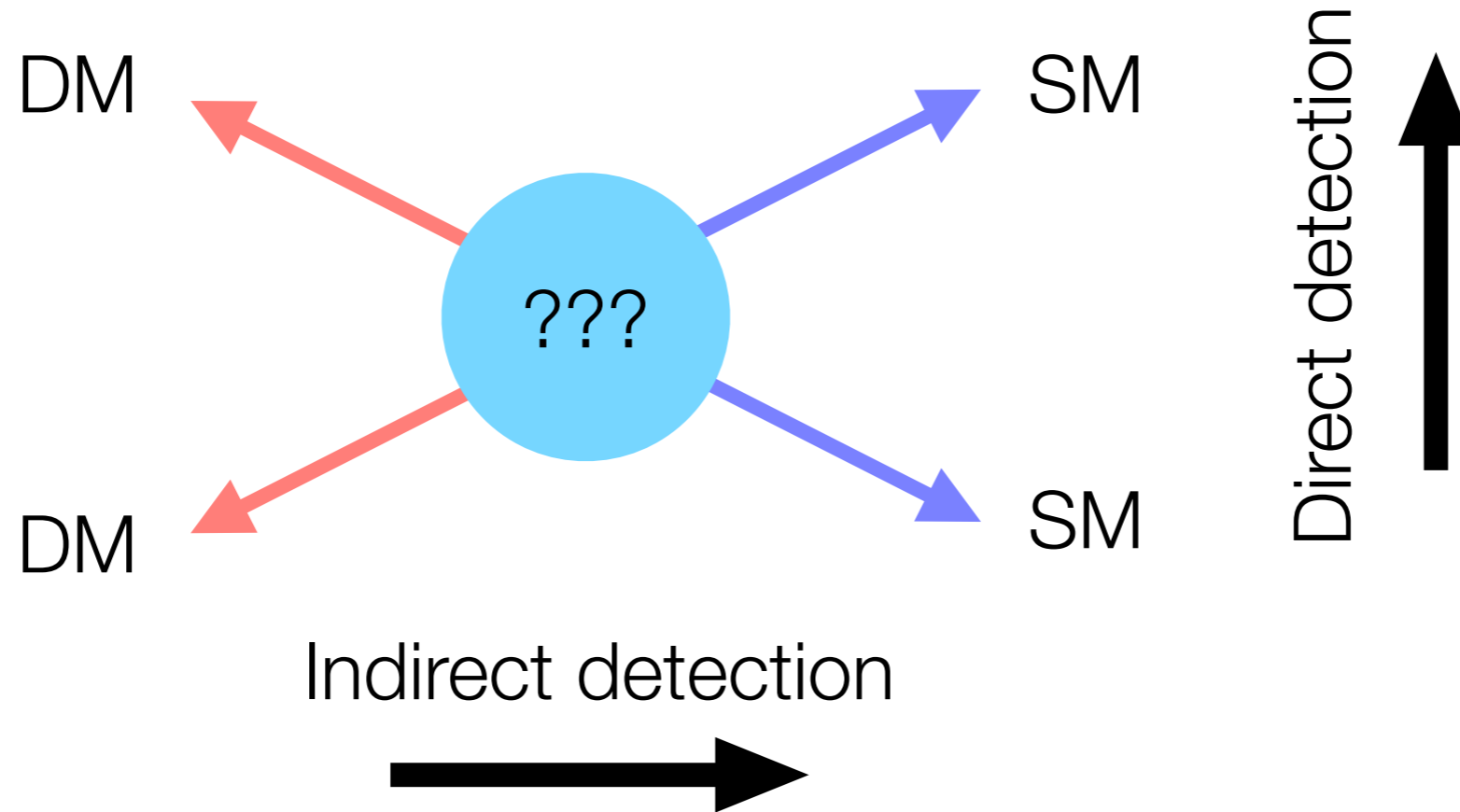
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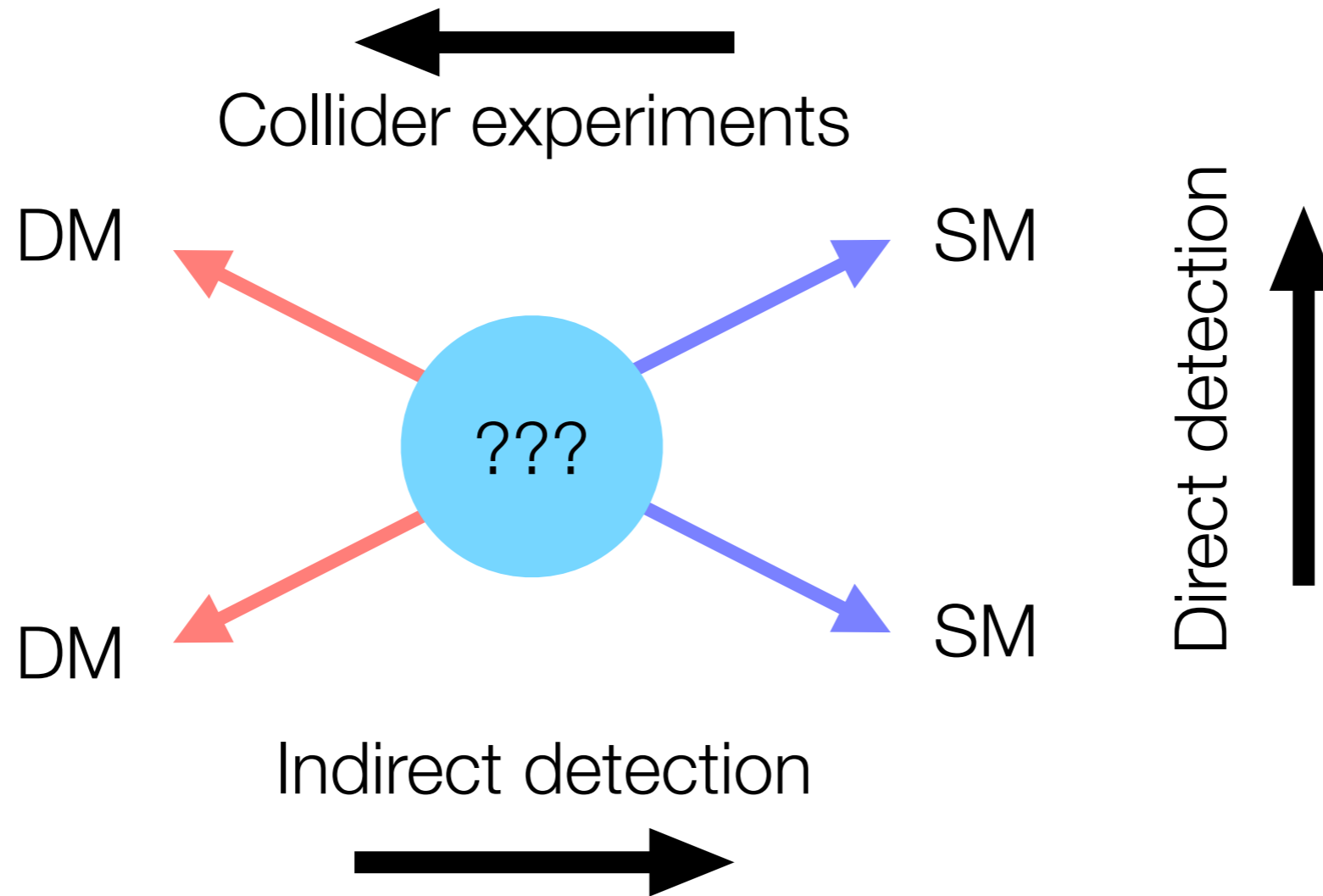
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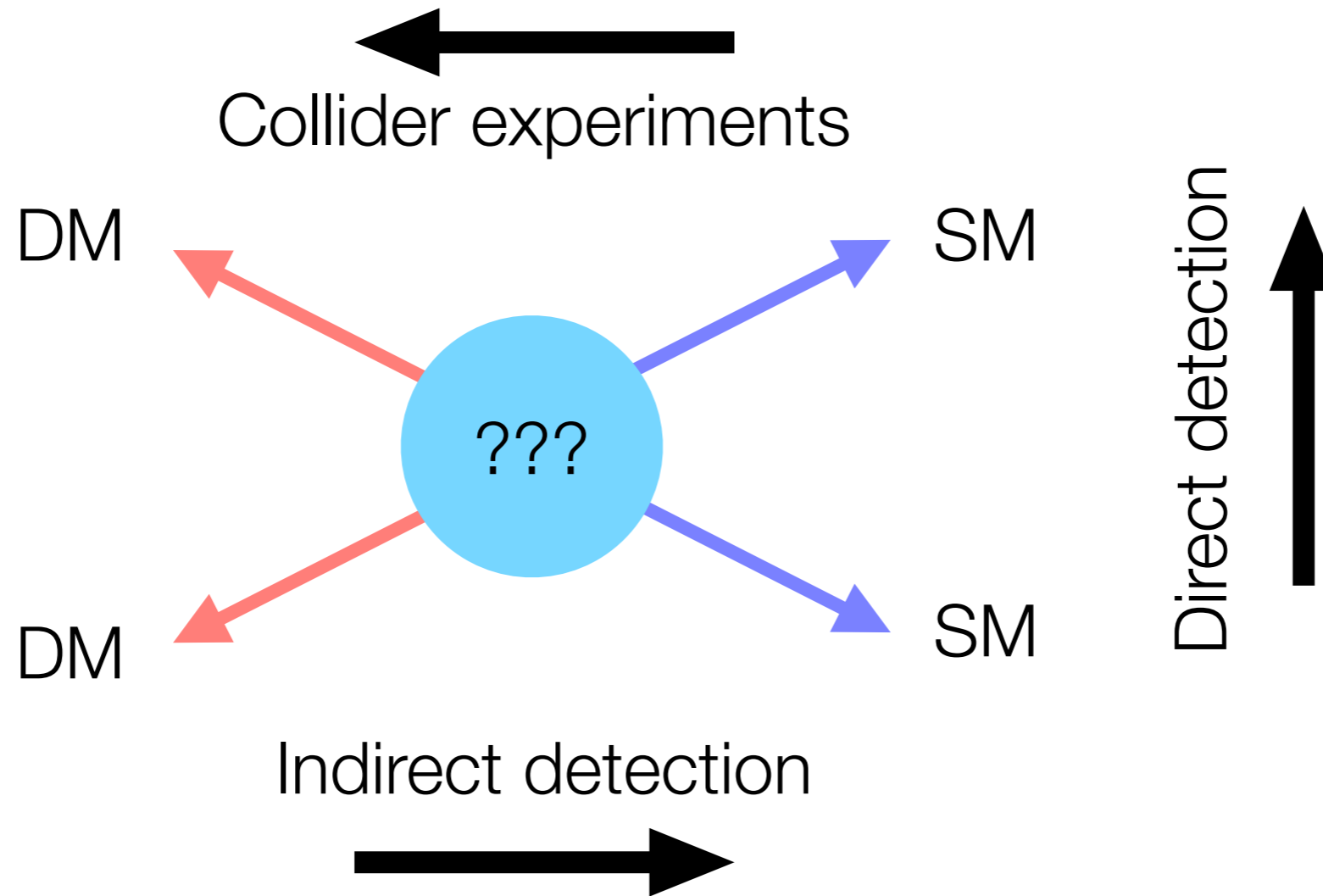
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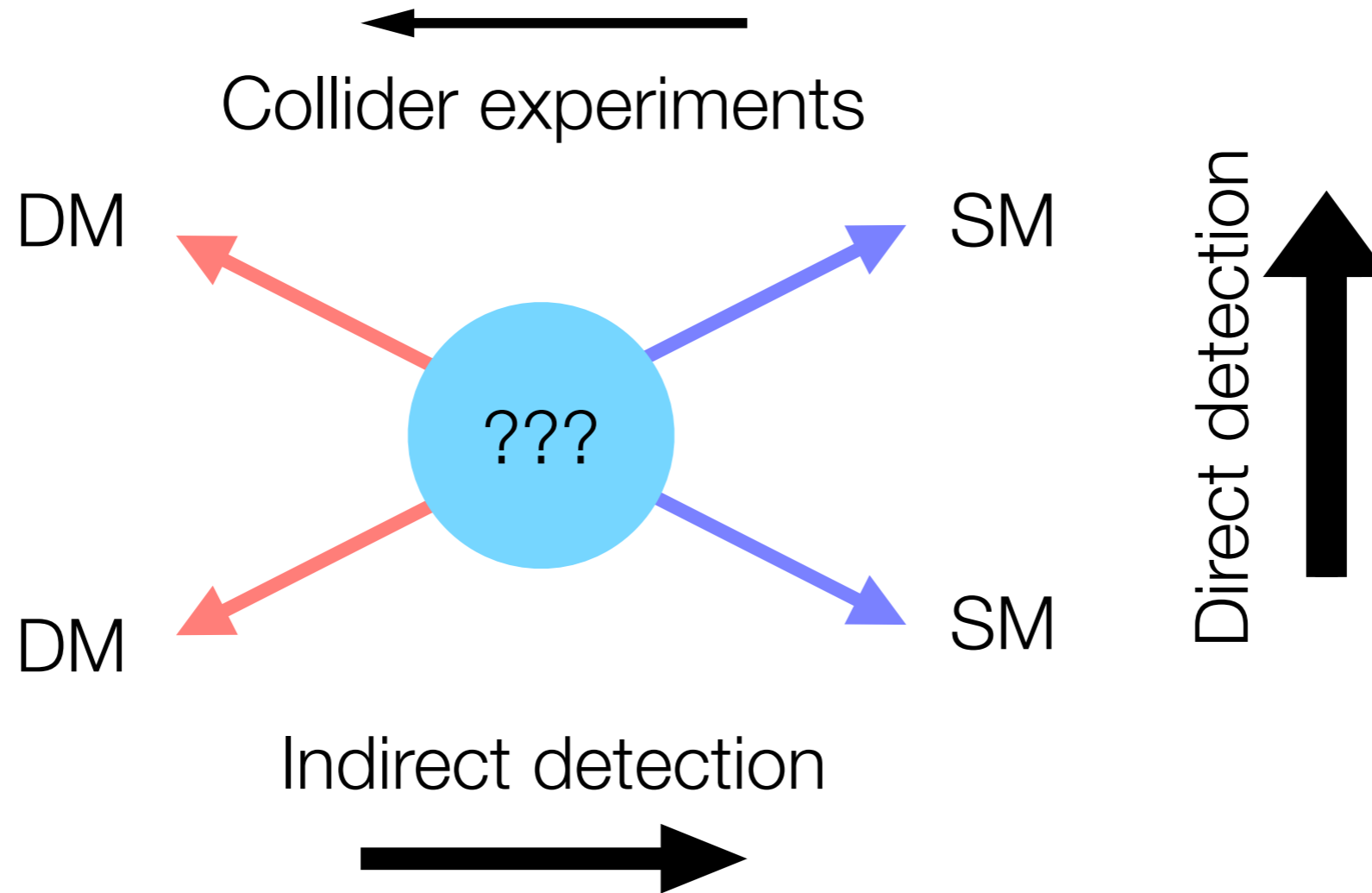


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Tired: all three approaches are probing the same thing (interchangeable)

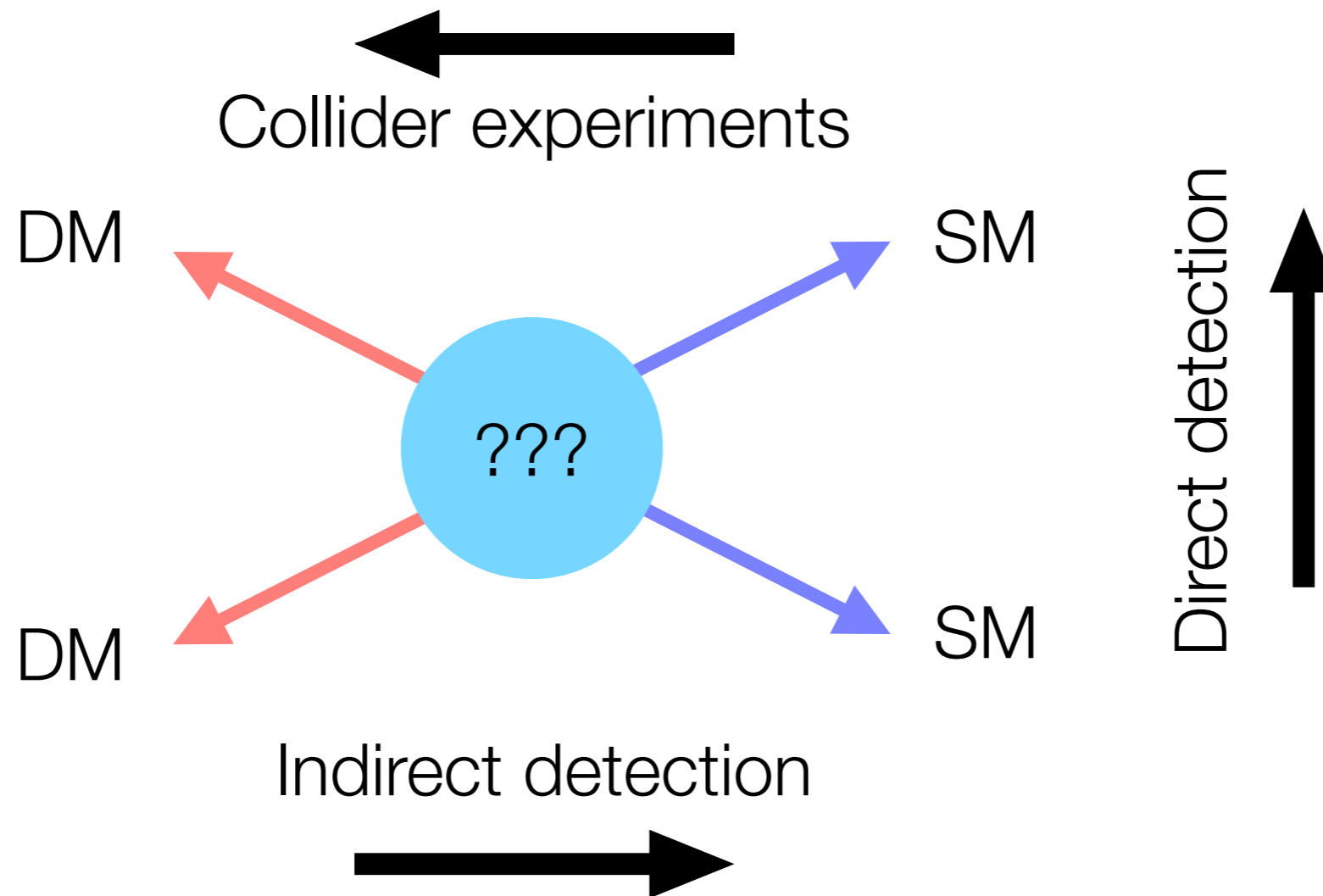
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Wired: different DM scenarios may be accessible to only one or two of the three approaches

Inspired: the future of the field needs all three to ensure success



Benchmarks

Dark matter models at the LHC



No EFTs

Mediator masses
around energy
scale of collider

Dark matter models at the LHC

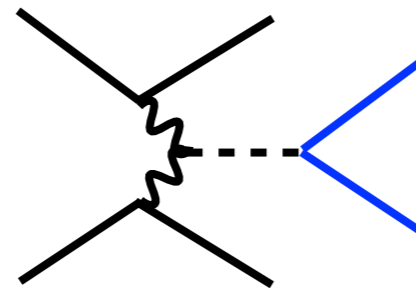
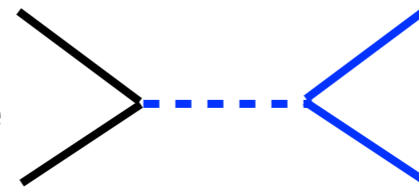
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Simplified models

Spin-1
mediator, one
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Simplified
Higgs portal

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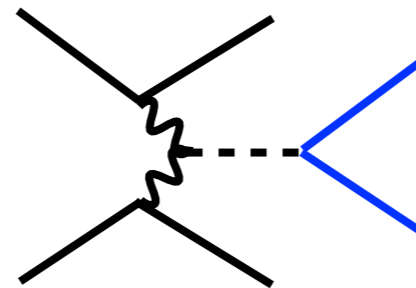
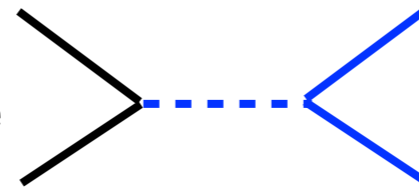
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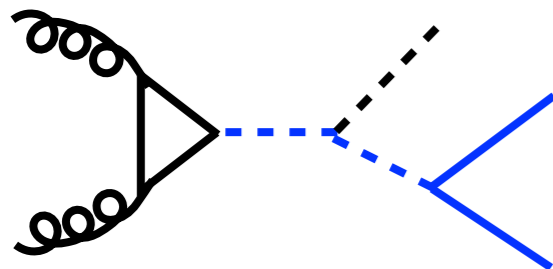
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2HDM + pseudoscalar



Still simple, UV-
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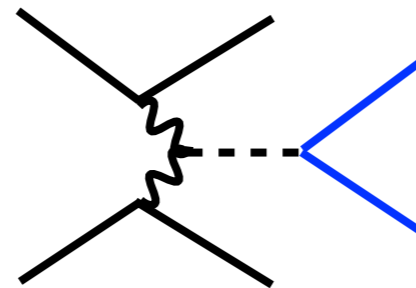
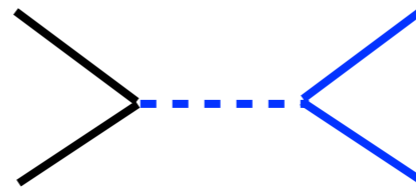
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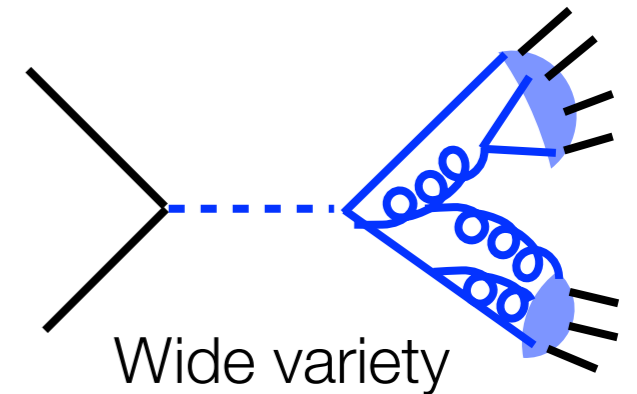
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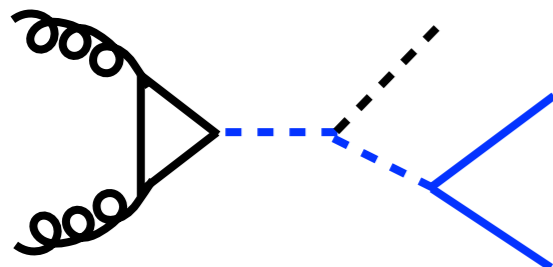


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Extended dark sectors



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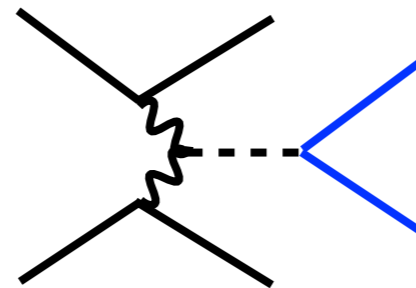
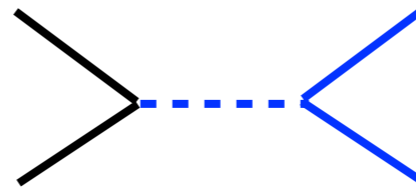
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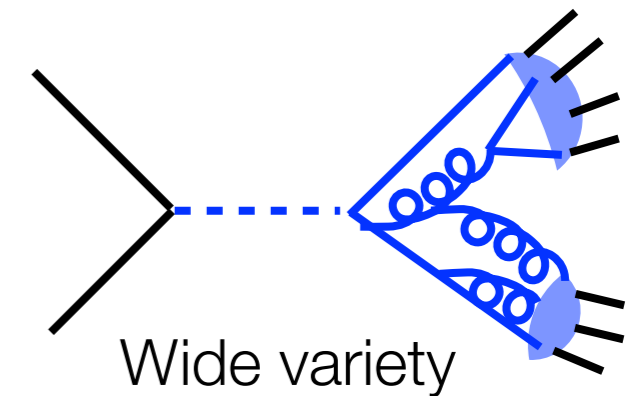
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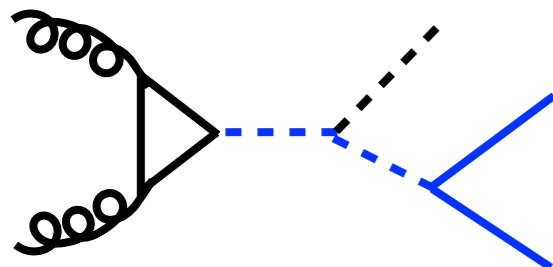
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Wide variety

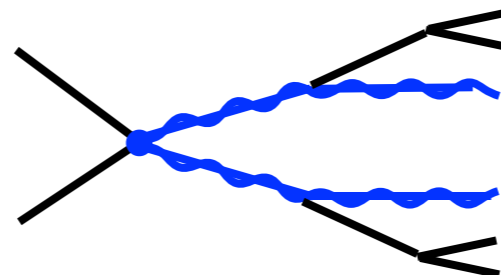
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SUSY scenarios

Cases with wino or higgsino-like LSP can give good DM candidates



Often simplified for practicality

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Long-lived particles

Not a model; rather, a class of signatures emerging from many of the others

Pros and cons of different benchmarks*



Simplified models

**Complete/
complex models**

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Ease of comparison between analyses and experiments

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No single answer. ATLAS & CMS lean on simplified models for comparisons; use complex models on analysis-by-analysis basis and for smaller comparison use cases

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How much should we care about ensuring benchmarks are **compatible with relic density**?

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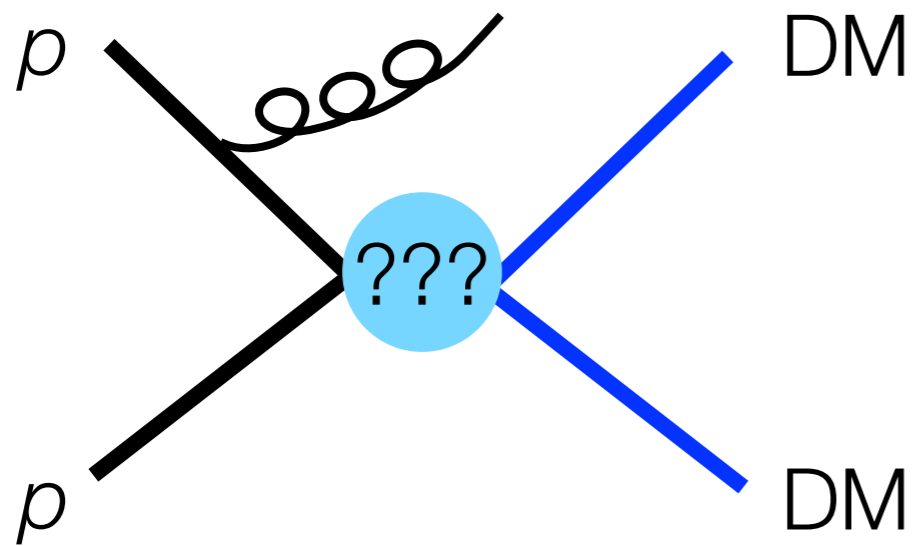
Could say a model is excluded once relic prediction reached

DM at the LHC



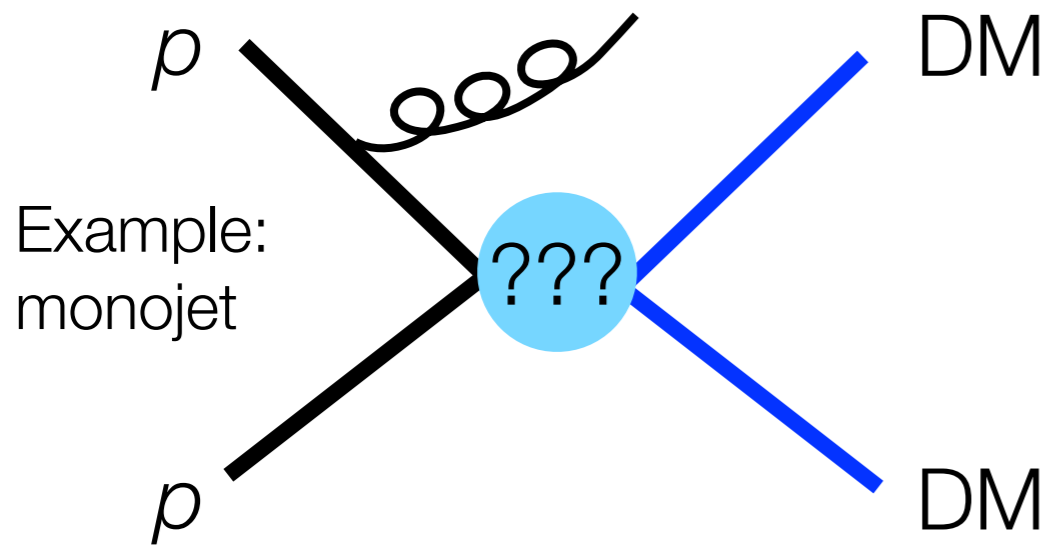
LHC signatures for DM searches

Most general: mono-X



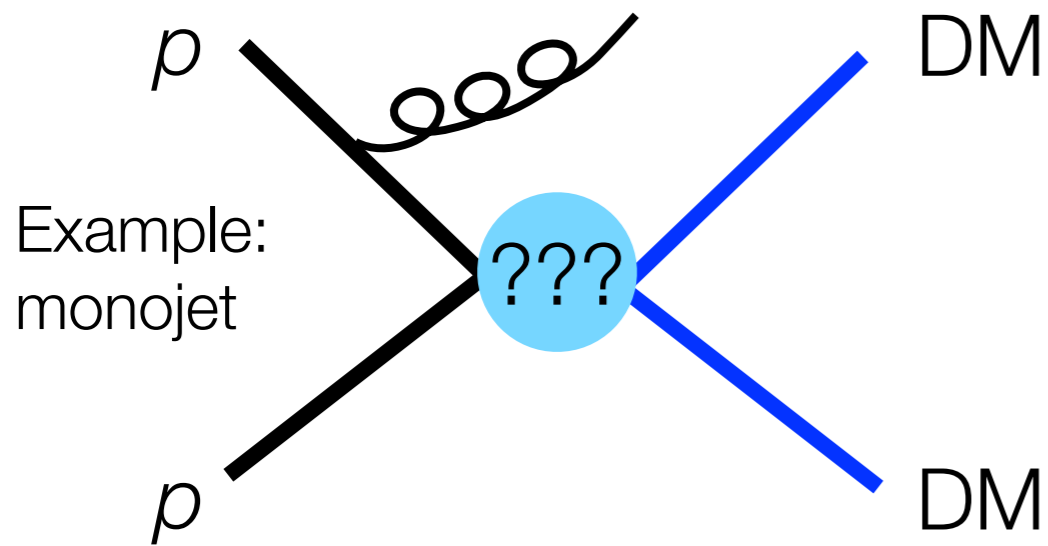
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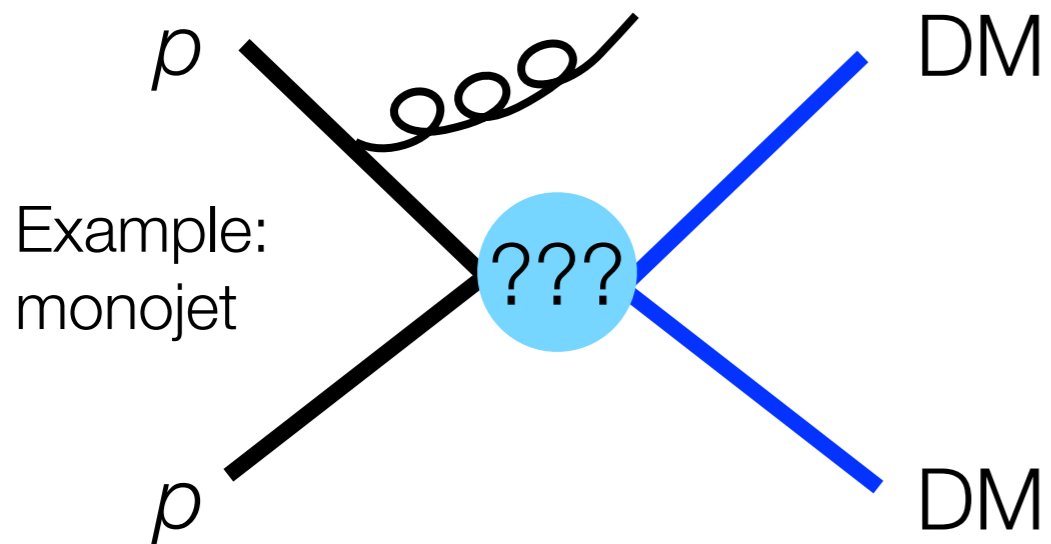
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Model-independent; high backgrounds.
ISR provides momentum, enabling
missing energy reconstruction

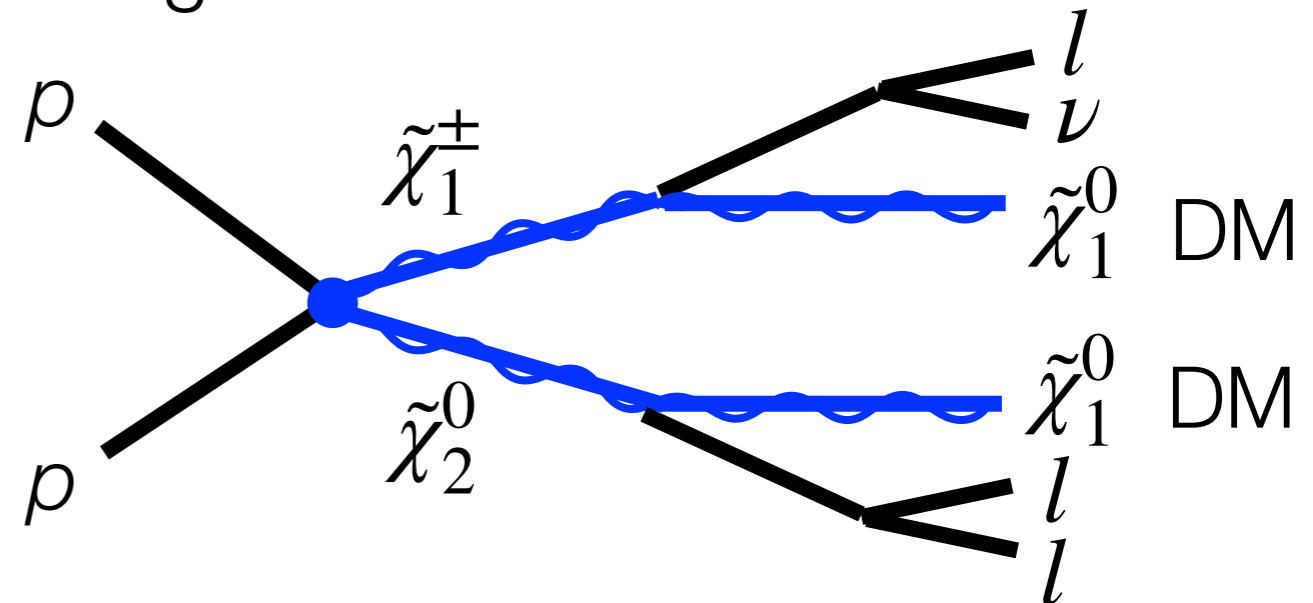
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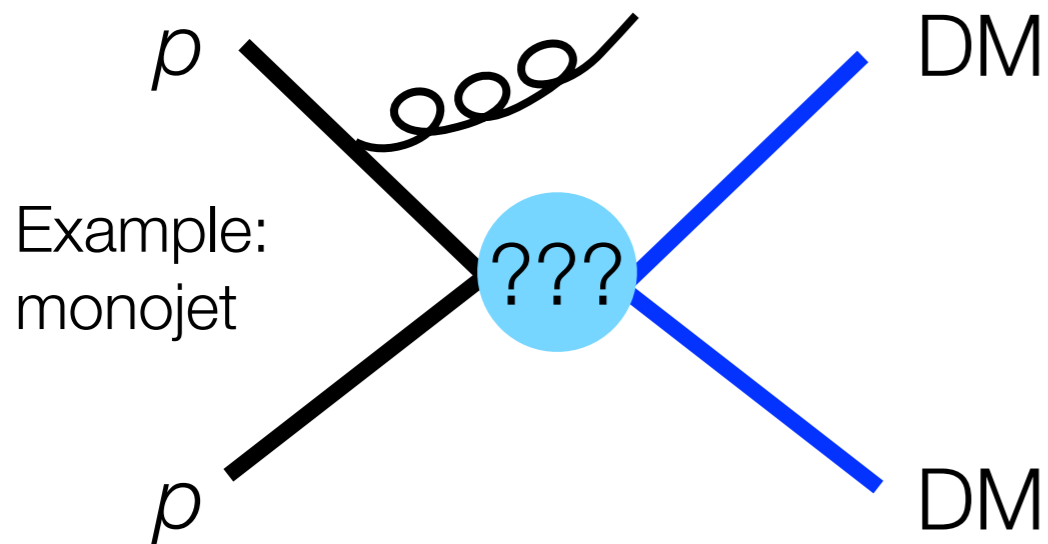
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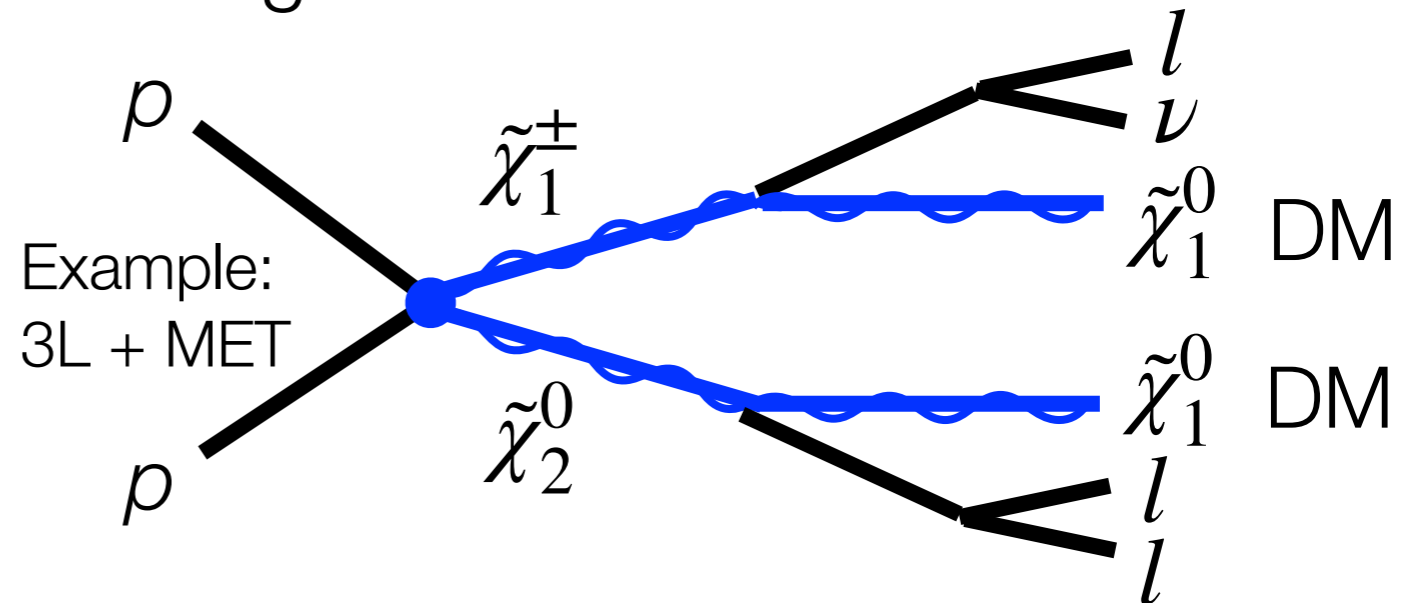
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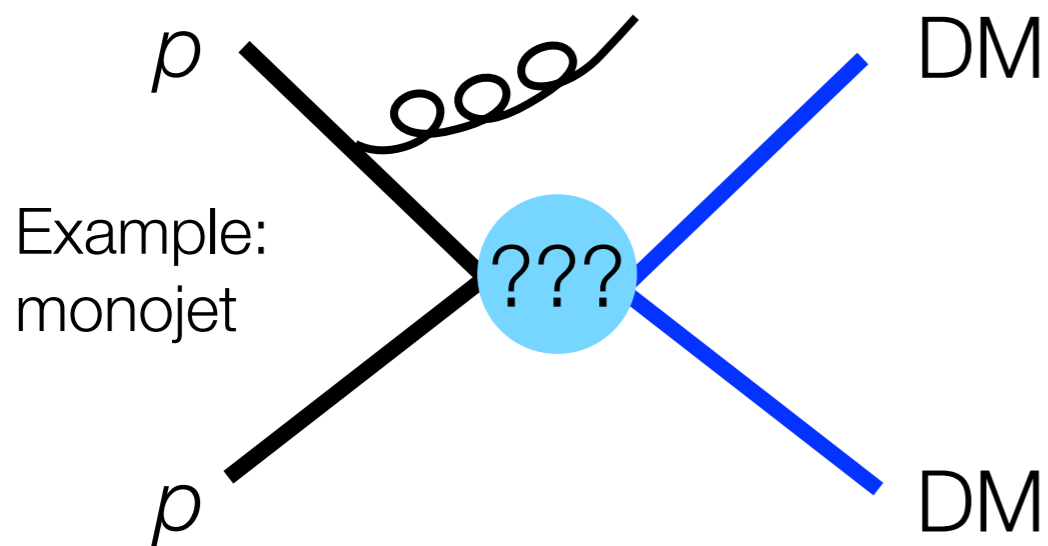
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3L + MET

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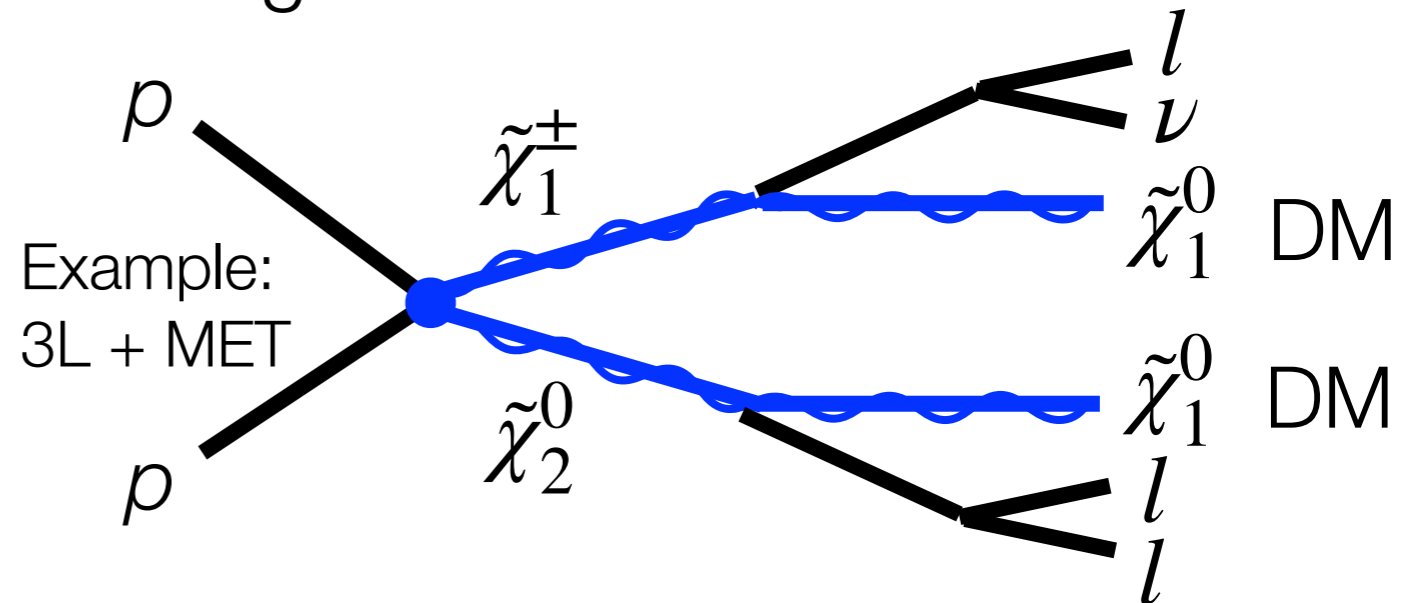
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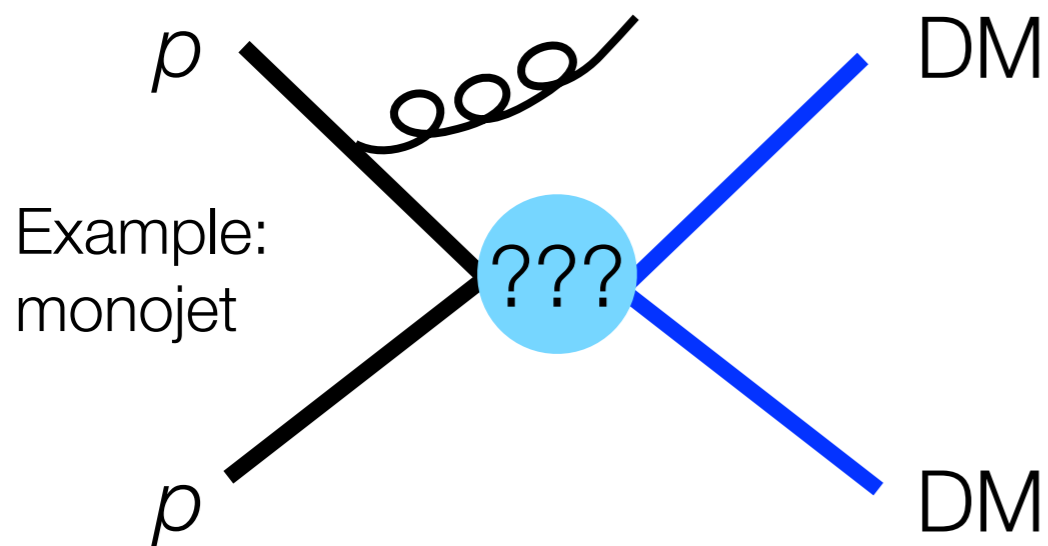


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3L + MET

Generally complex final state allowing
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MET remains key feature of selection

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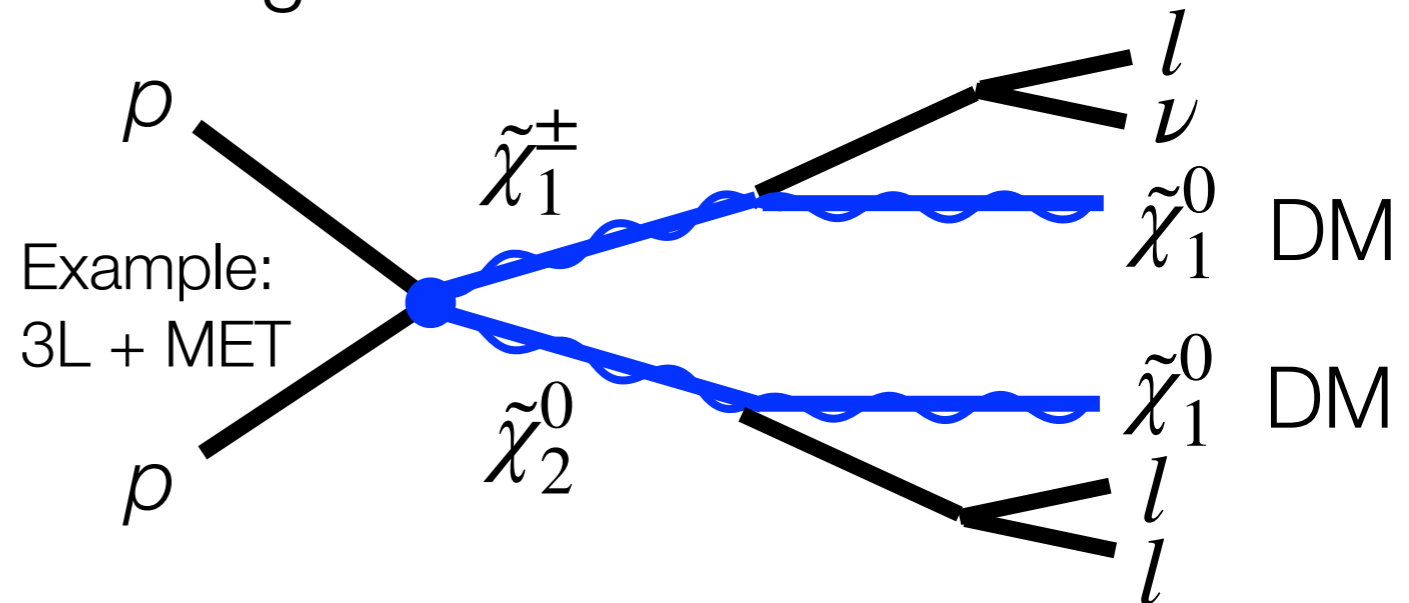
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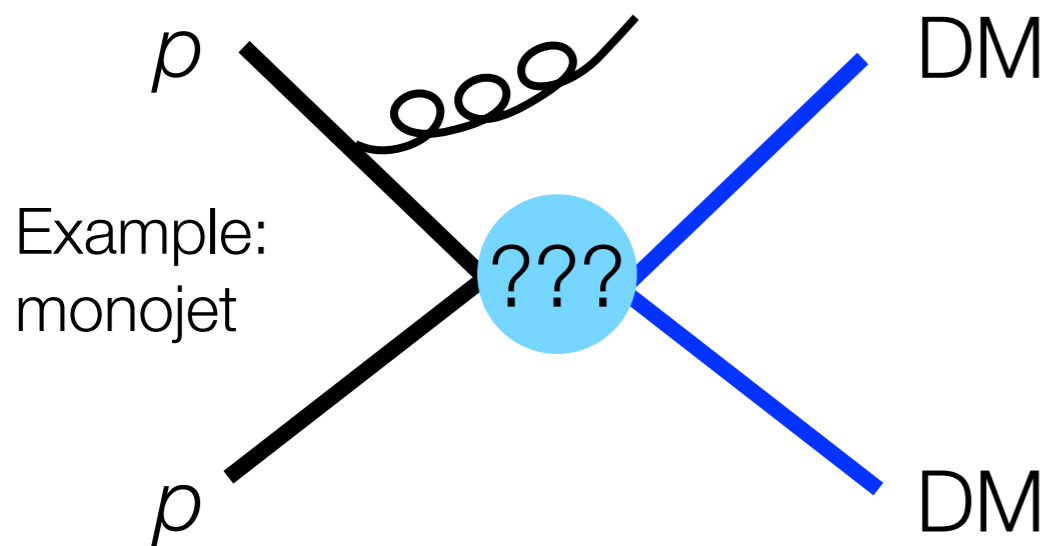
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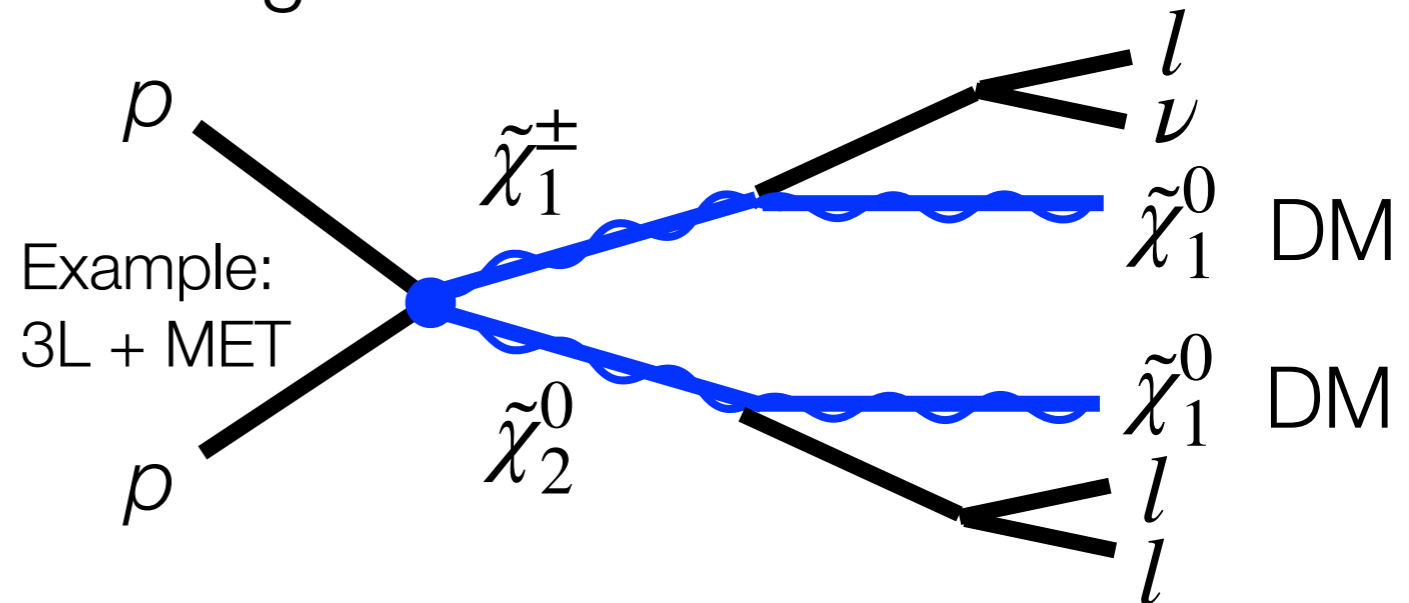
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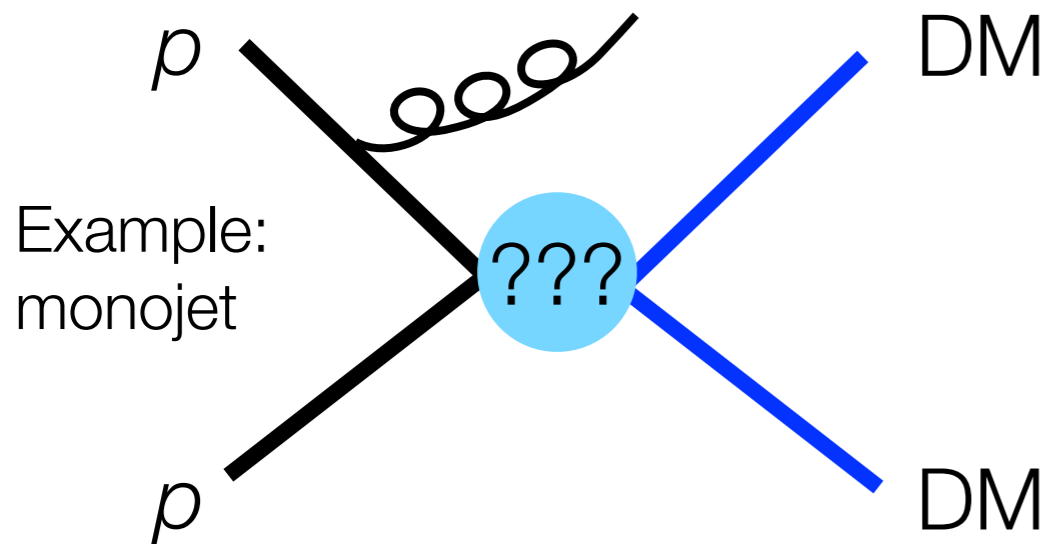
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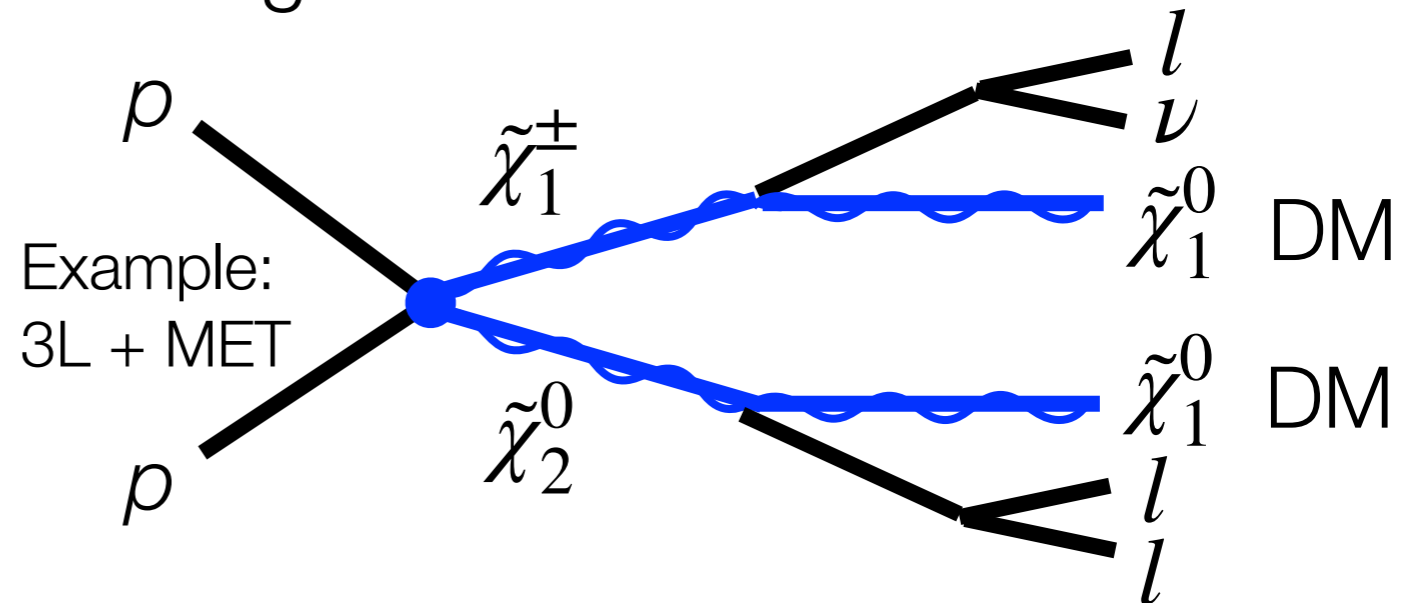
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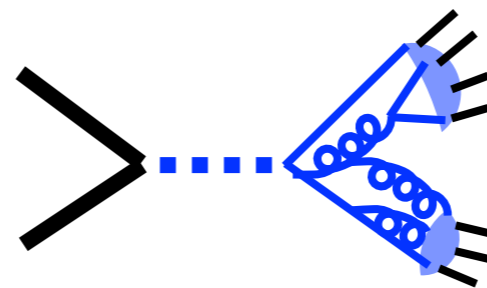
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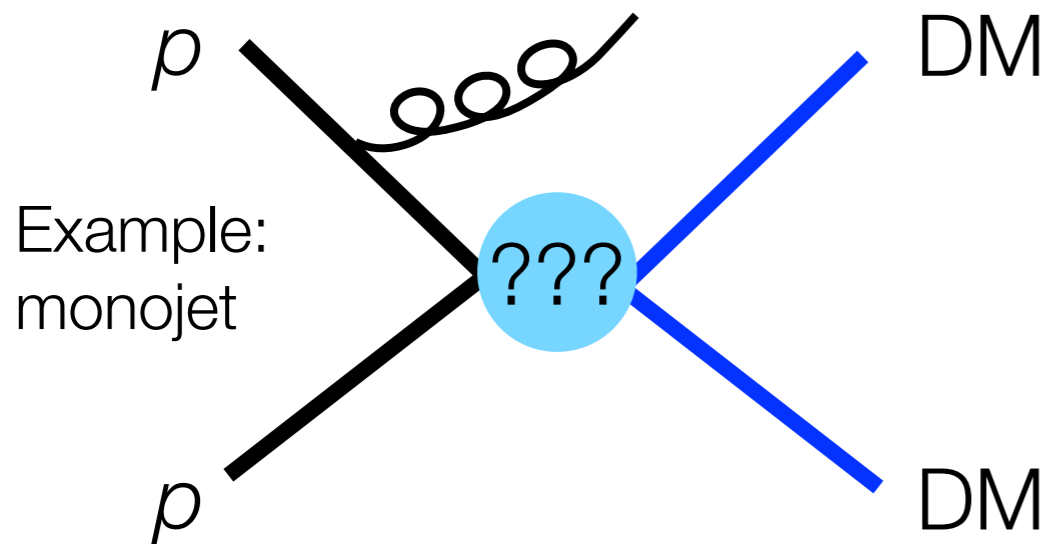
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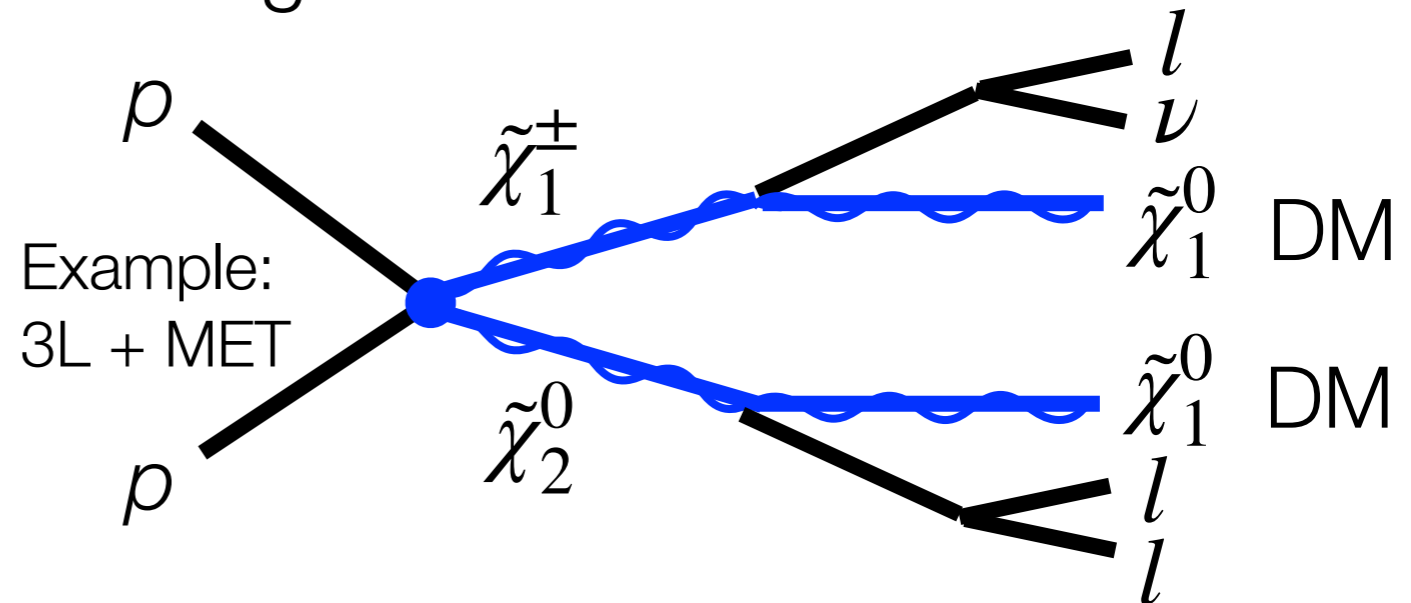
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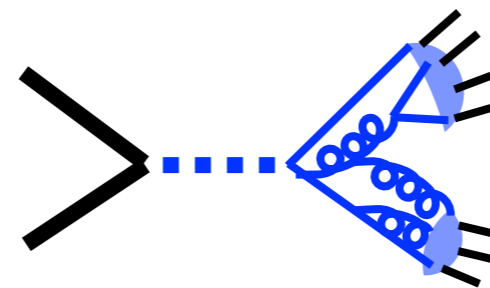
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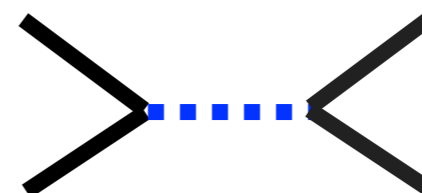
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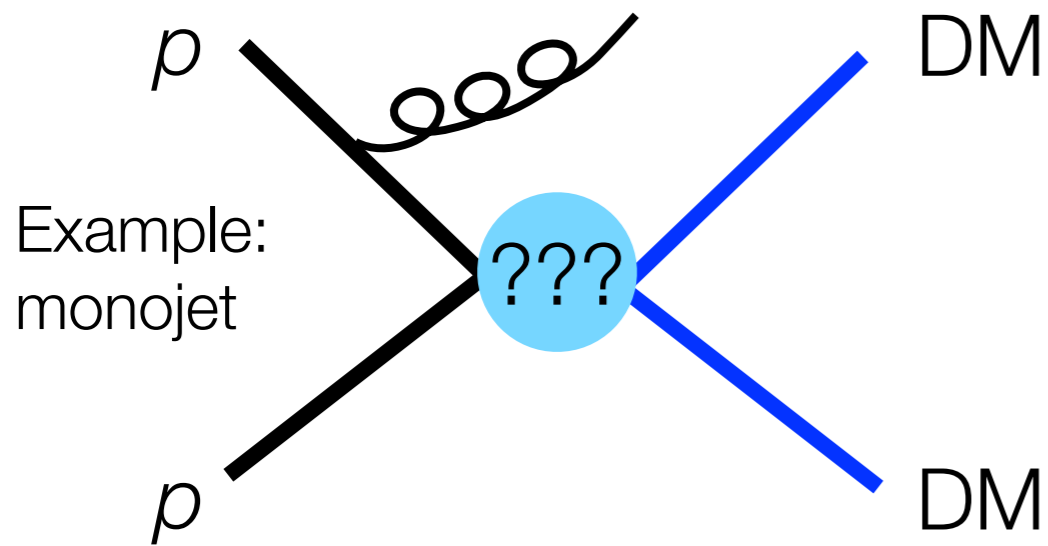
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SM decay of mediator



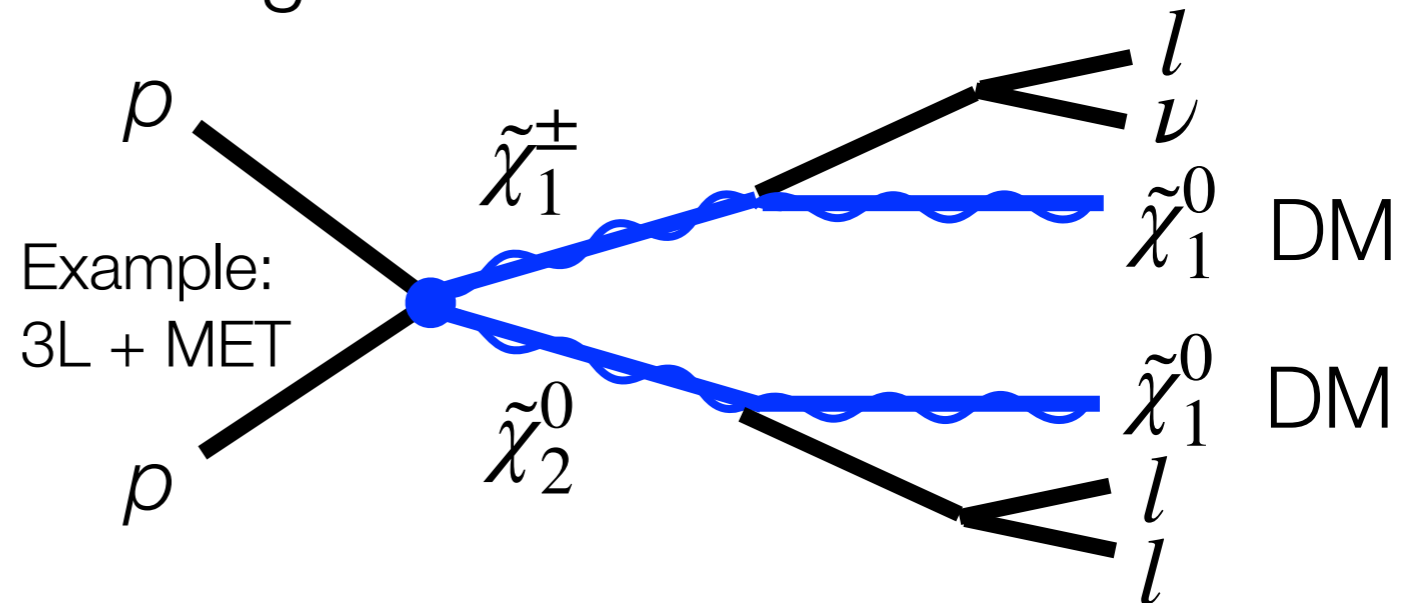
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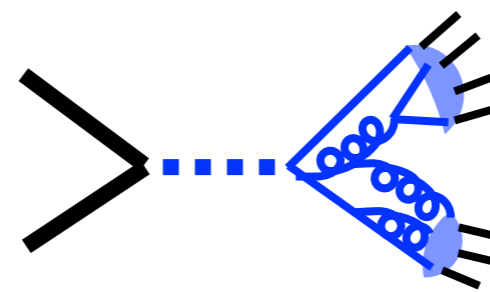
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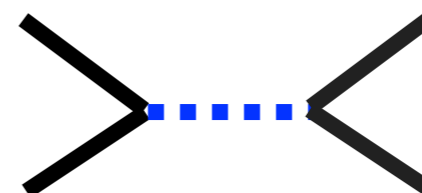
Non-MET-focused

Various searches target models with dark matter implications, but that do not rely on MET in final state. Extended dark sectors, direct mediator searches, LLPs

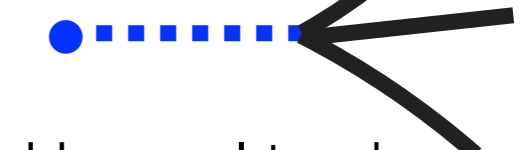


QCD final states with distinctive features

SM decay of mediator



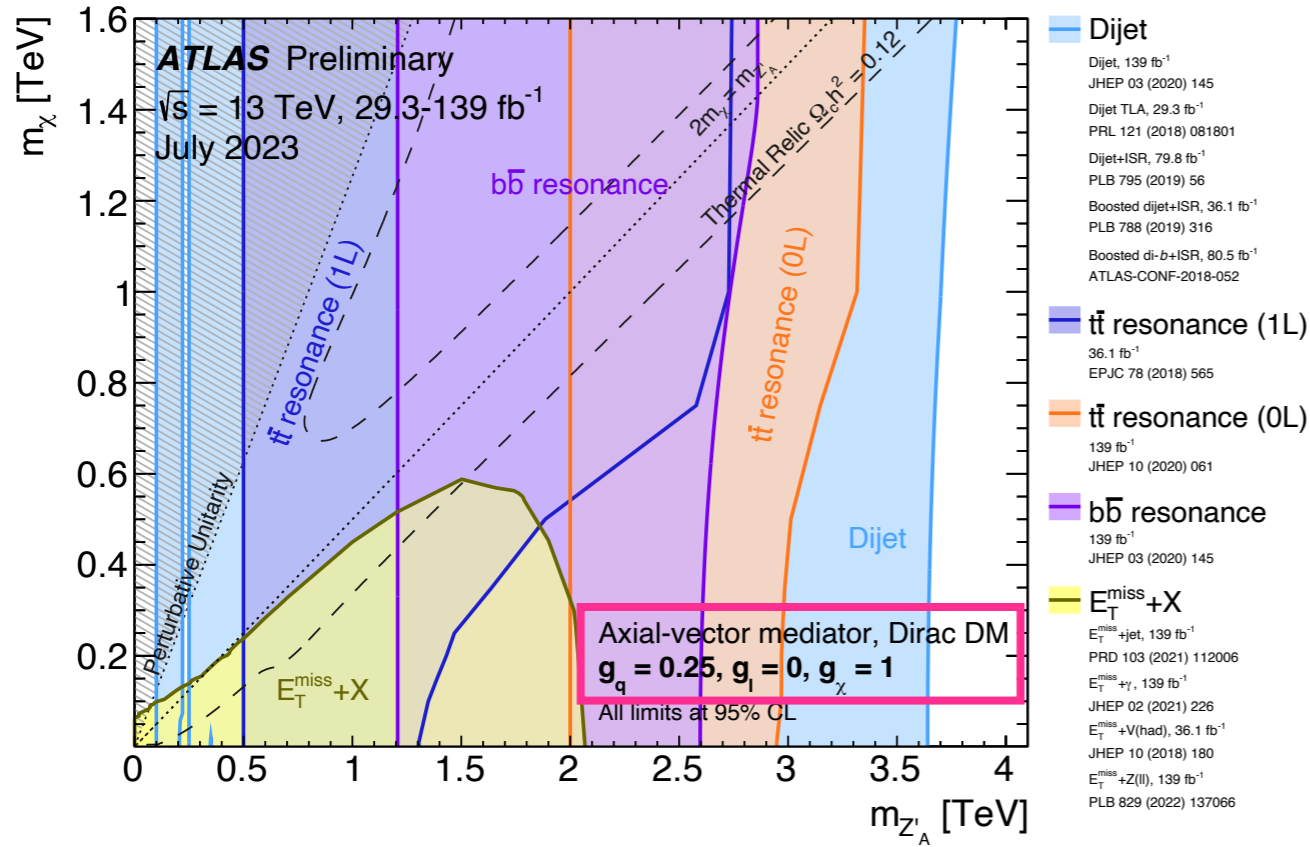
PV



Unusual tracks or displaced energy deposits

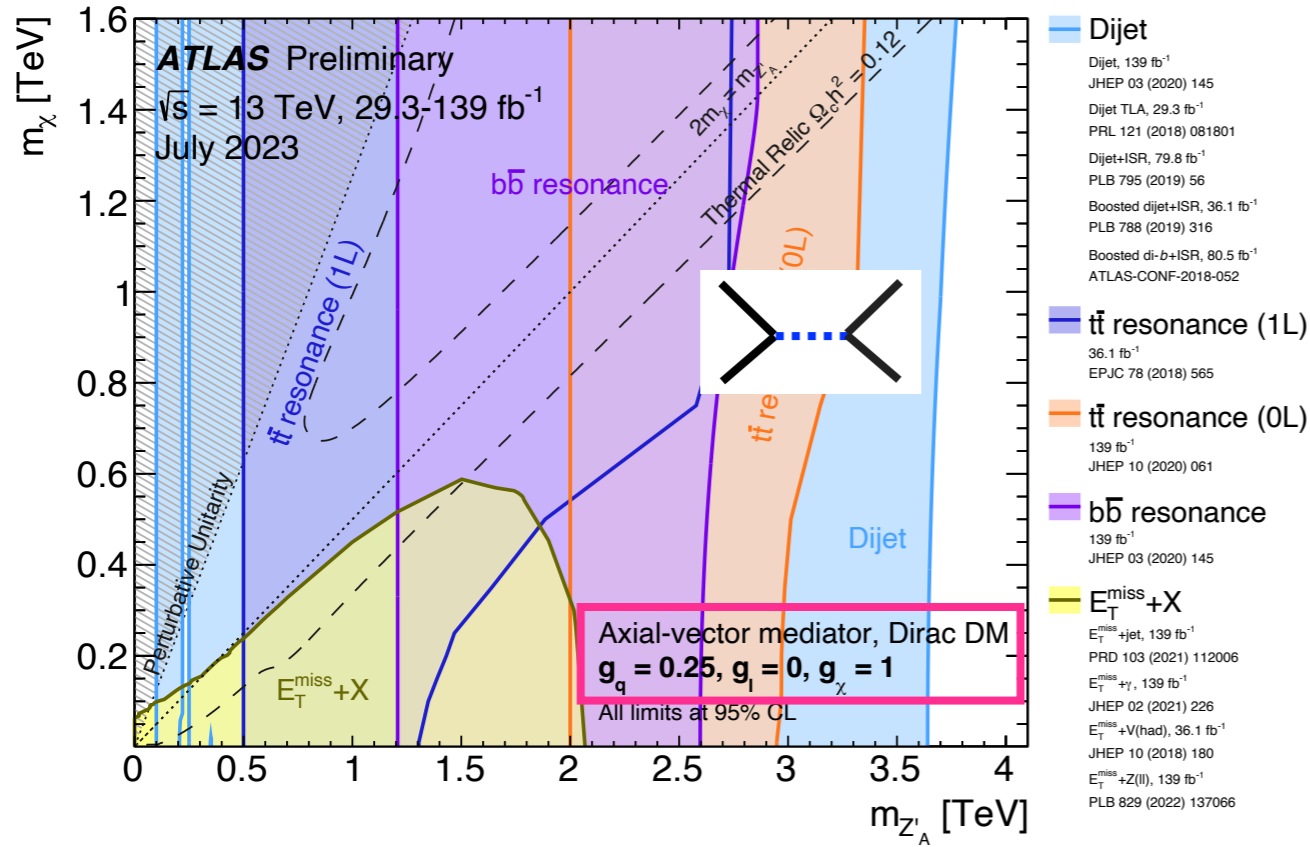
Current status of LHC spin-1 simplified models

ATL-PHYS-PUB-2023-018



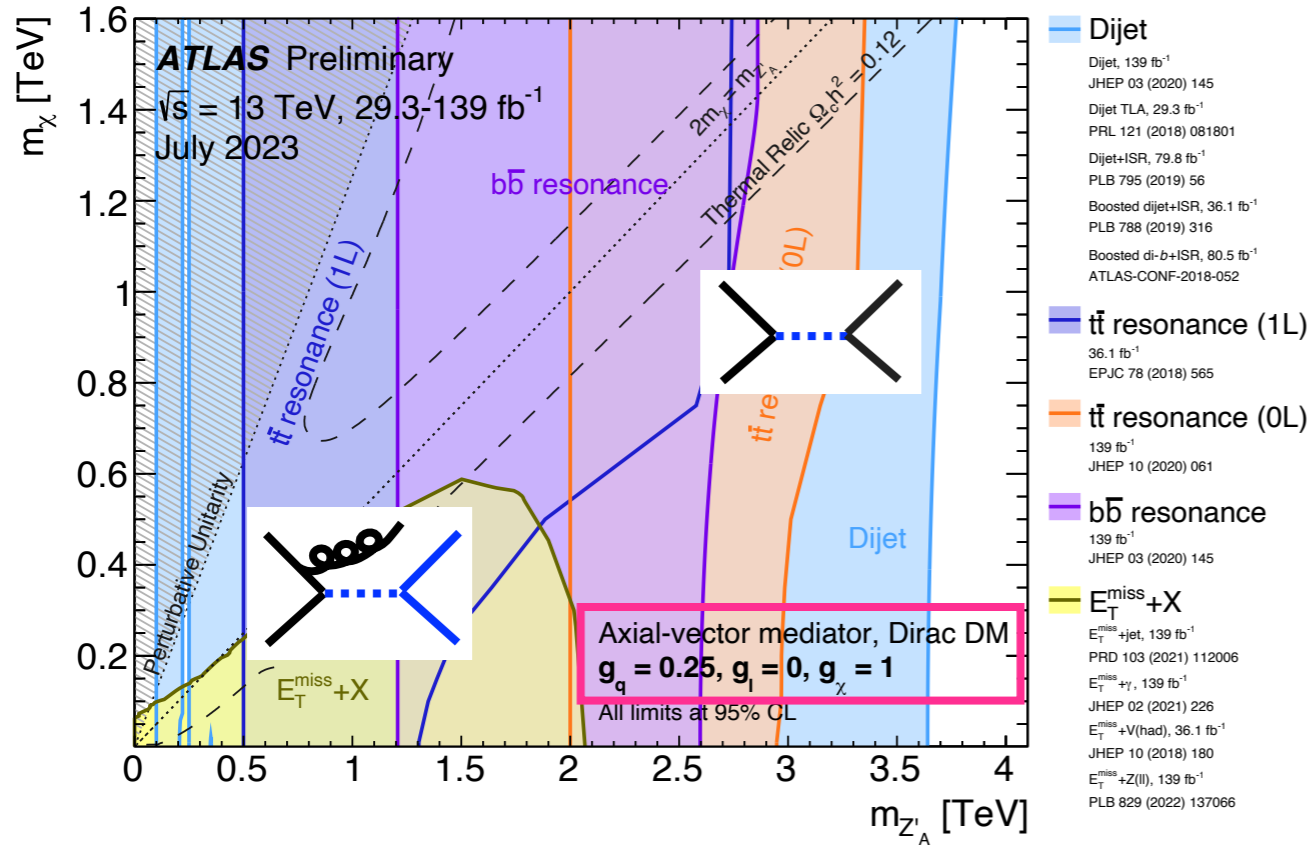
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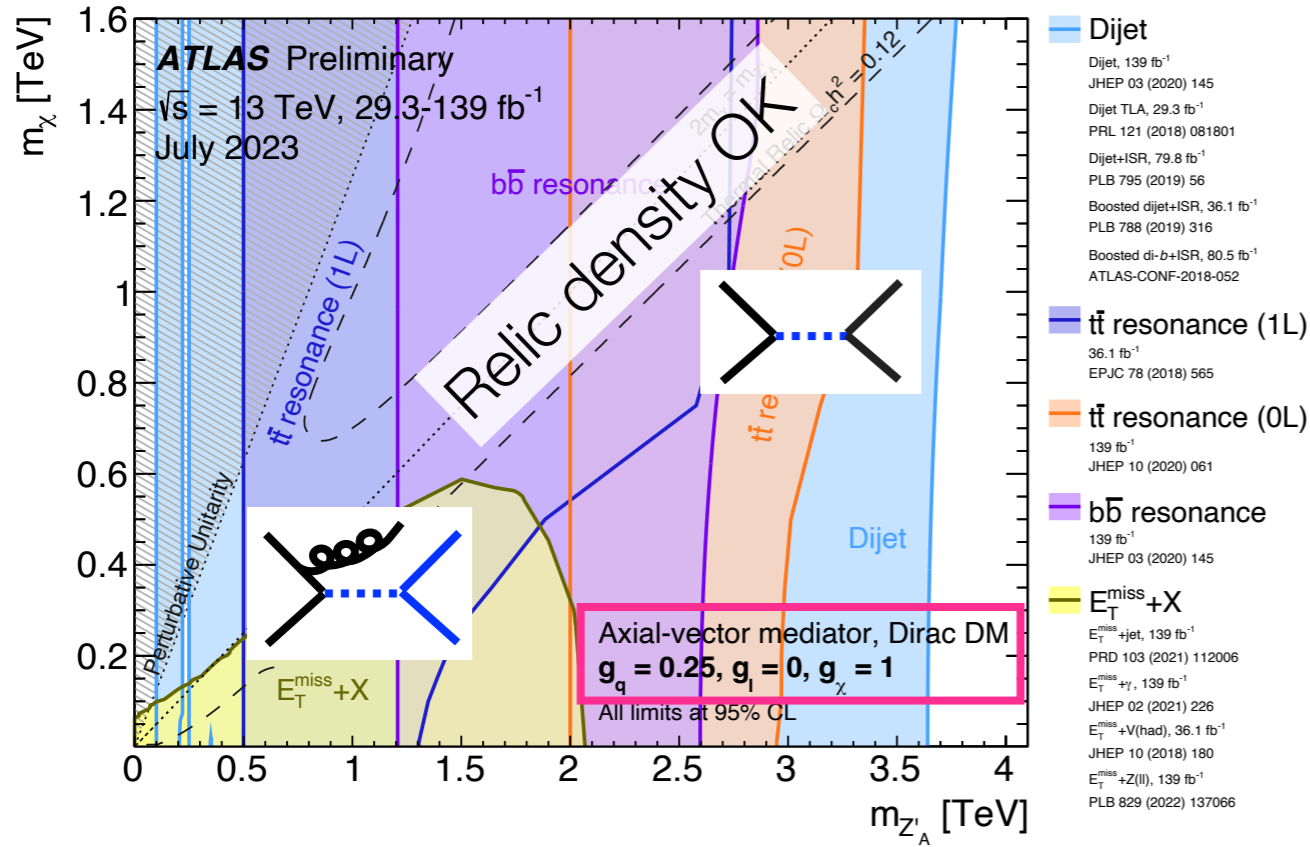
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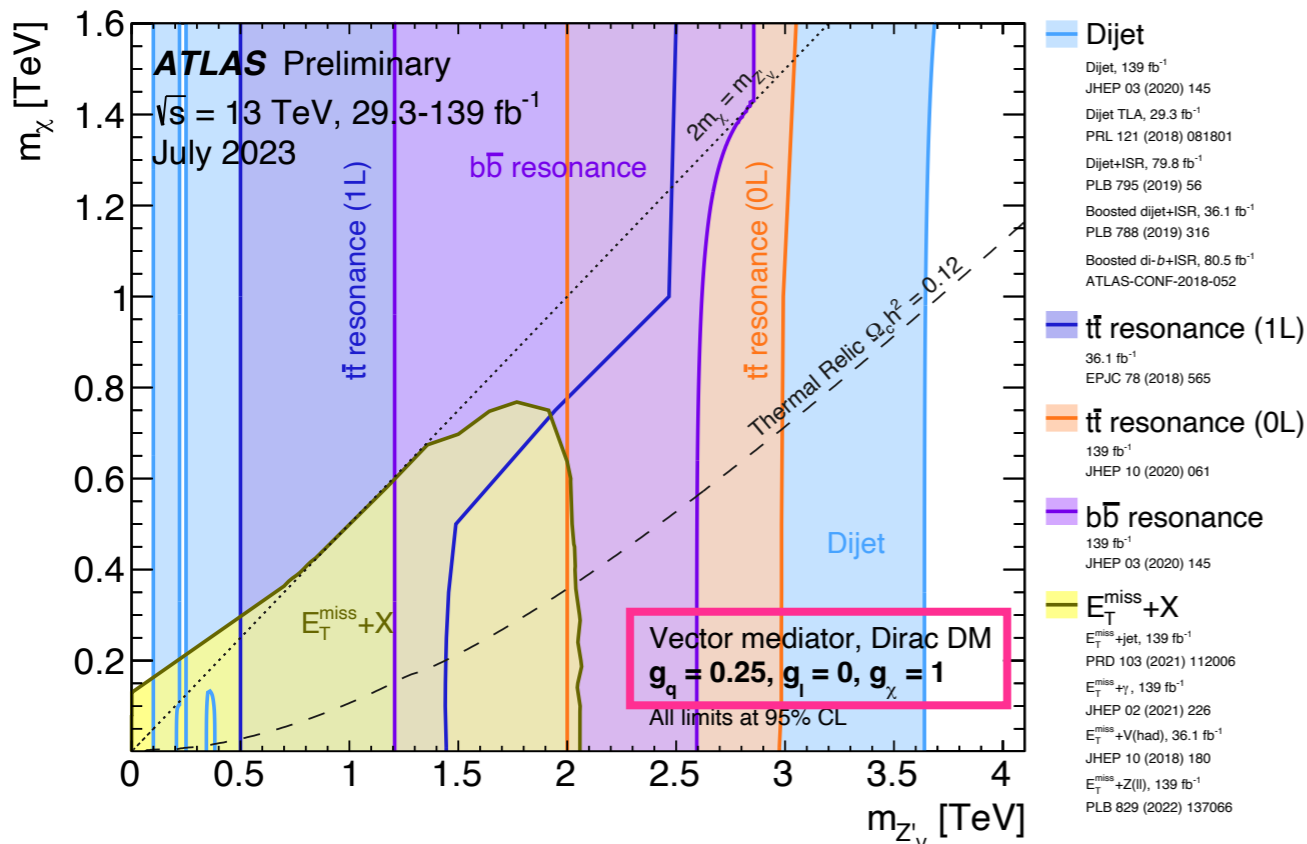
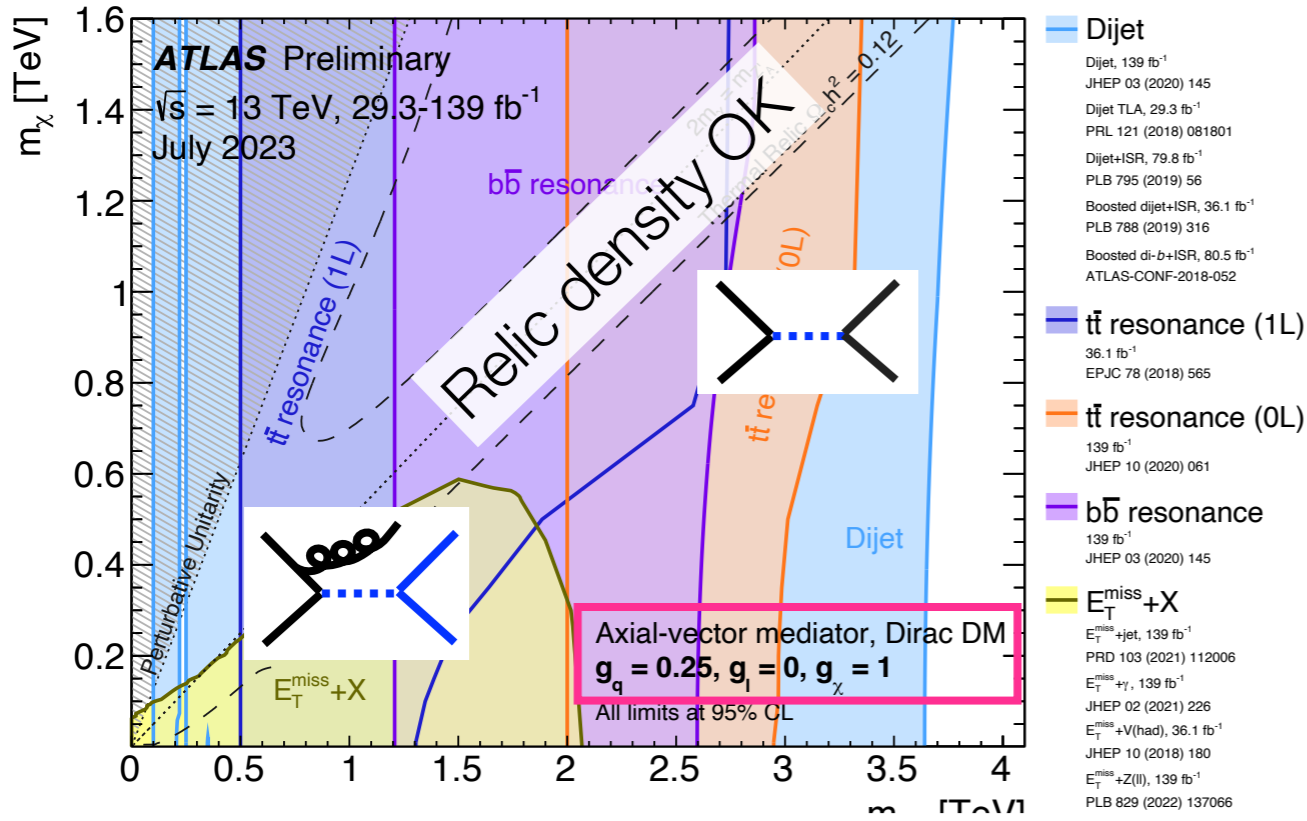
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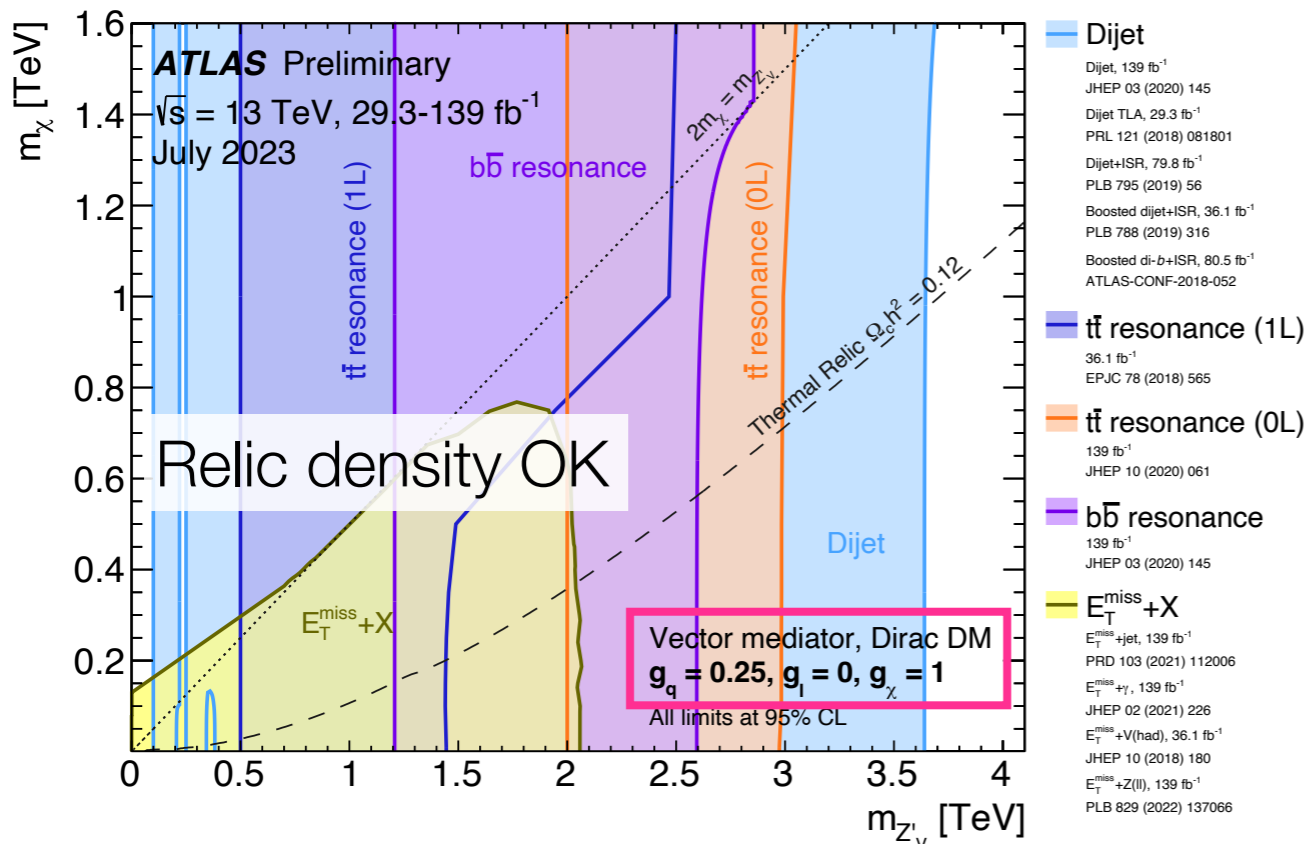
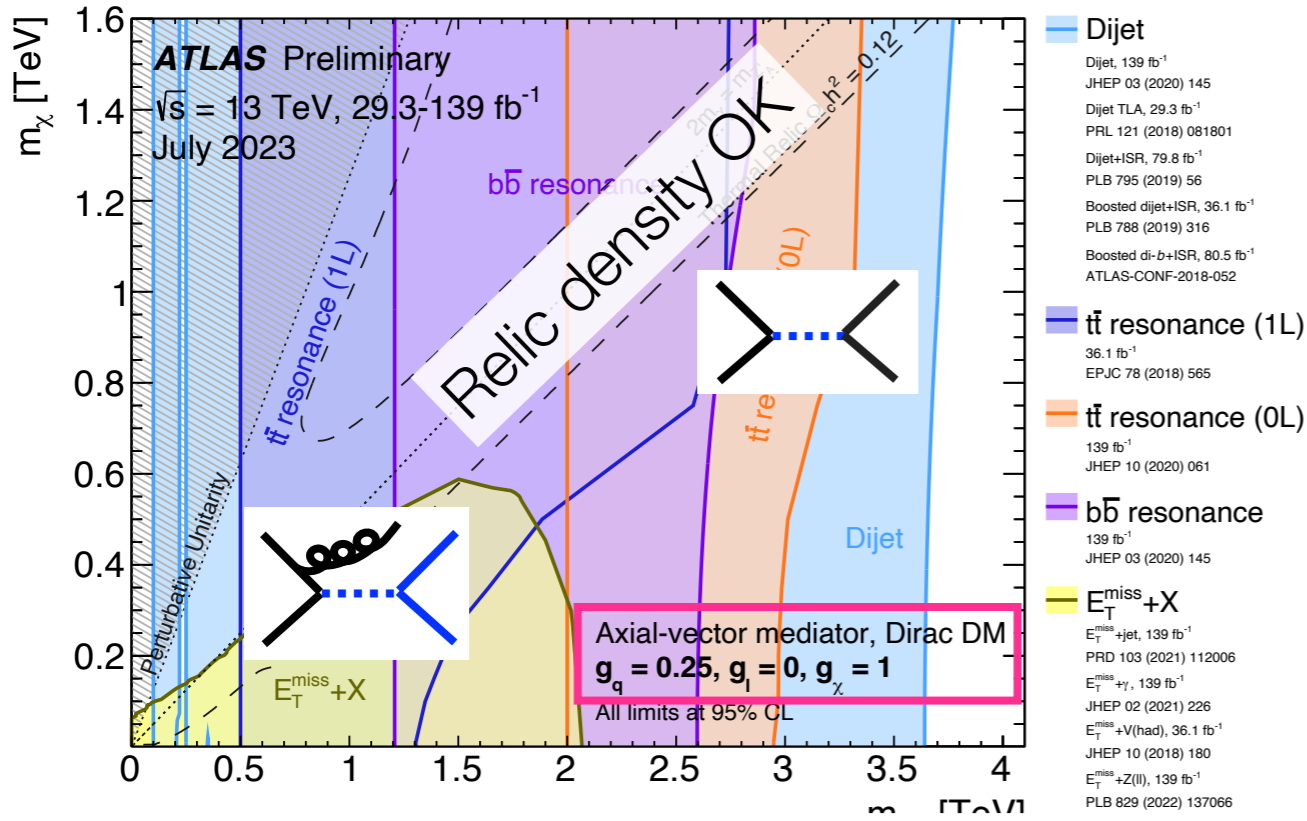
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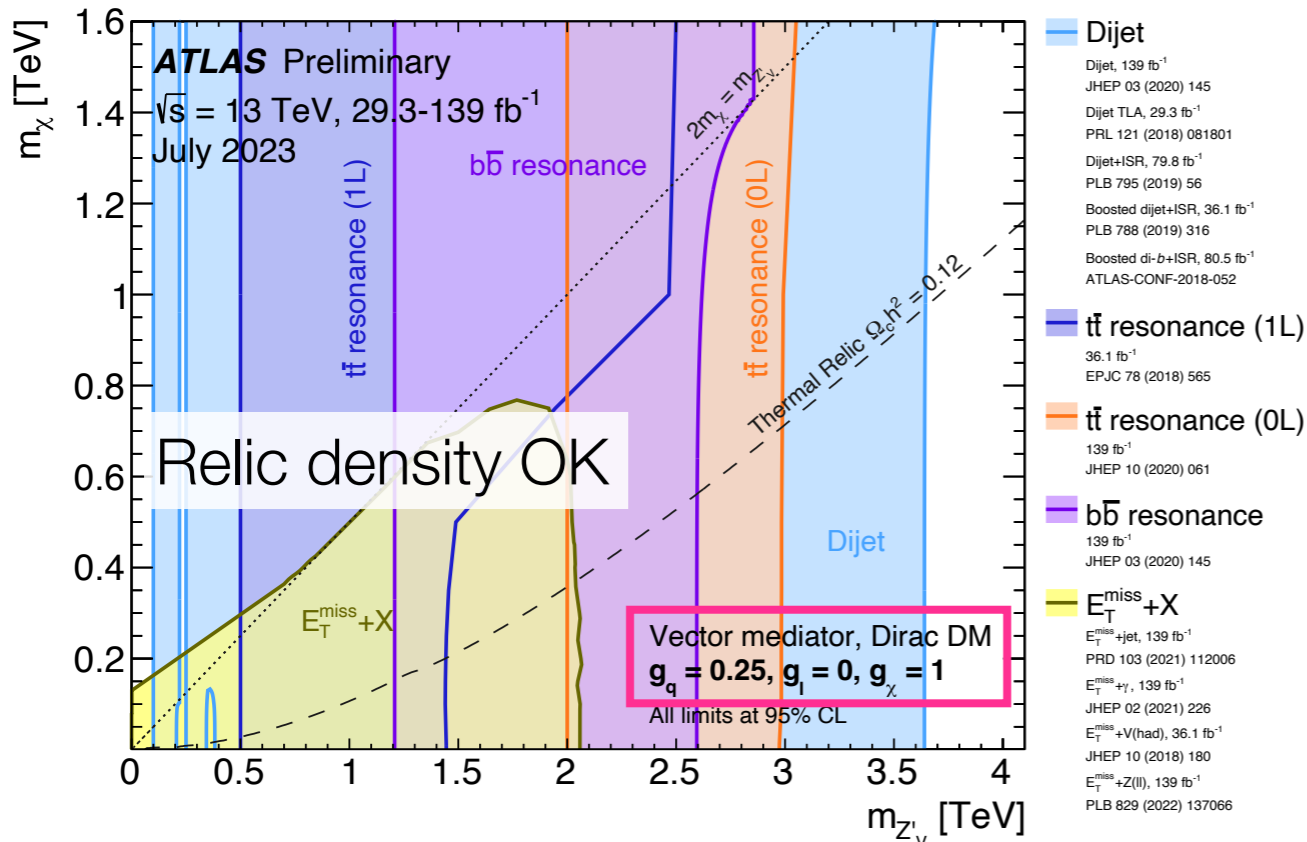
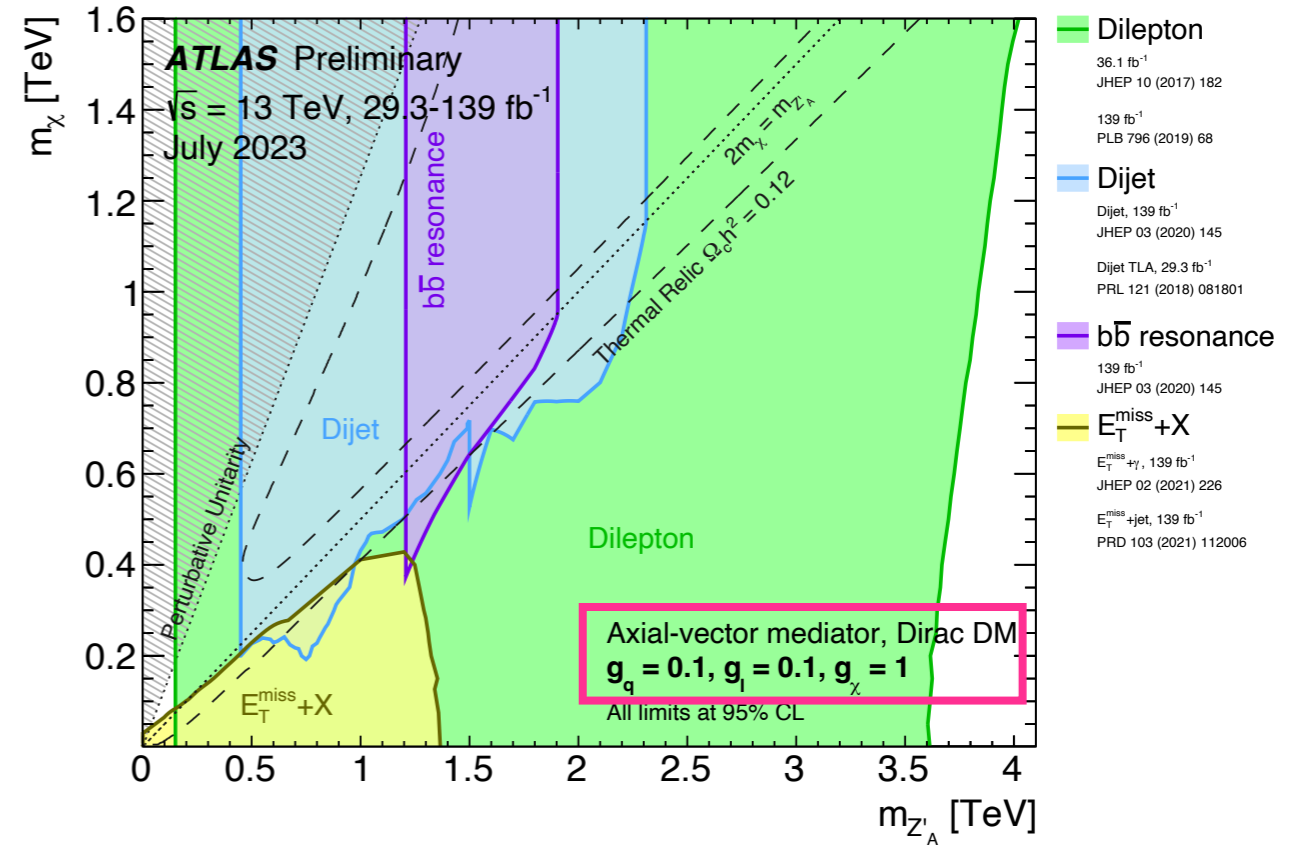
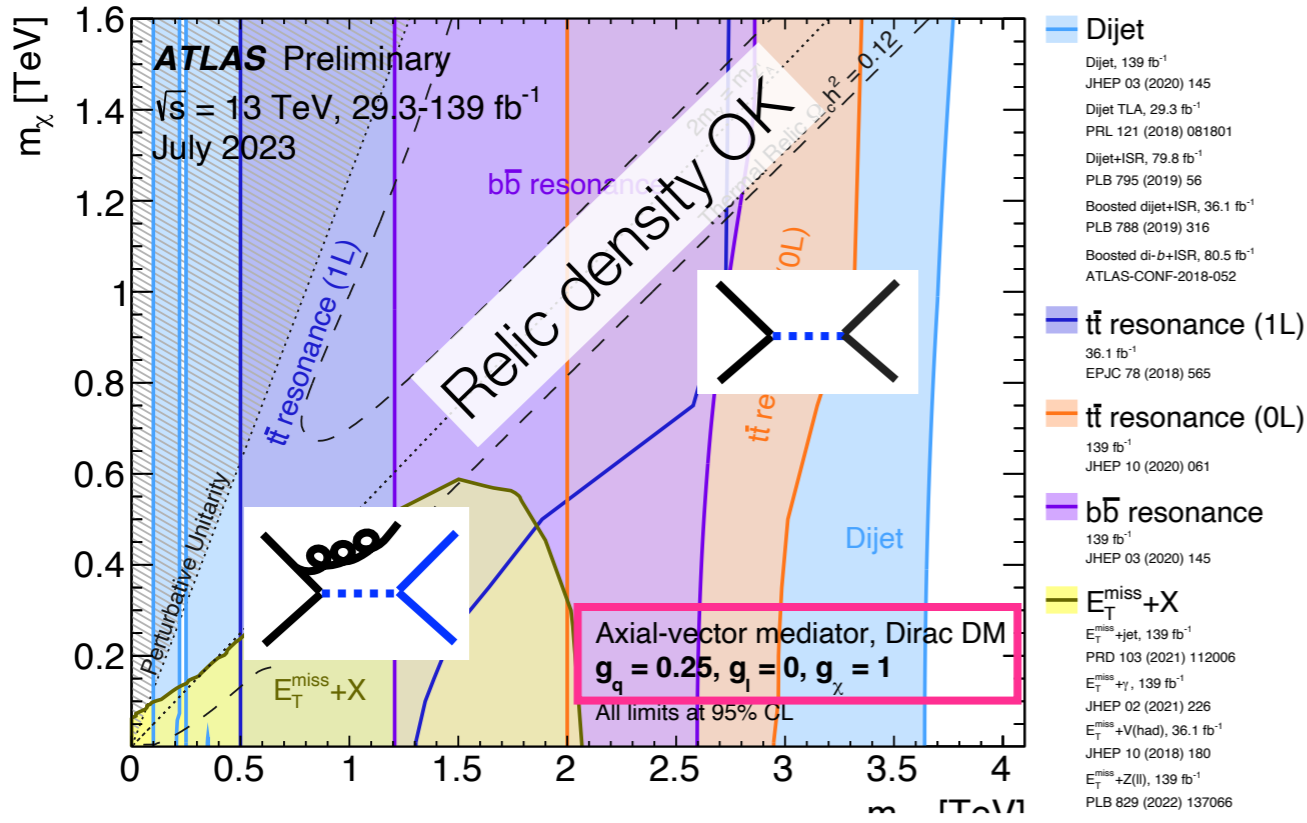
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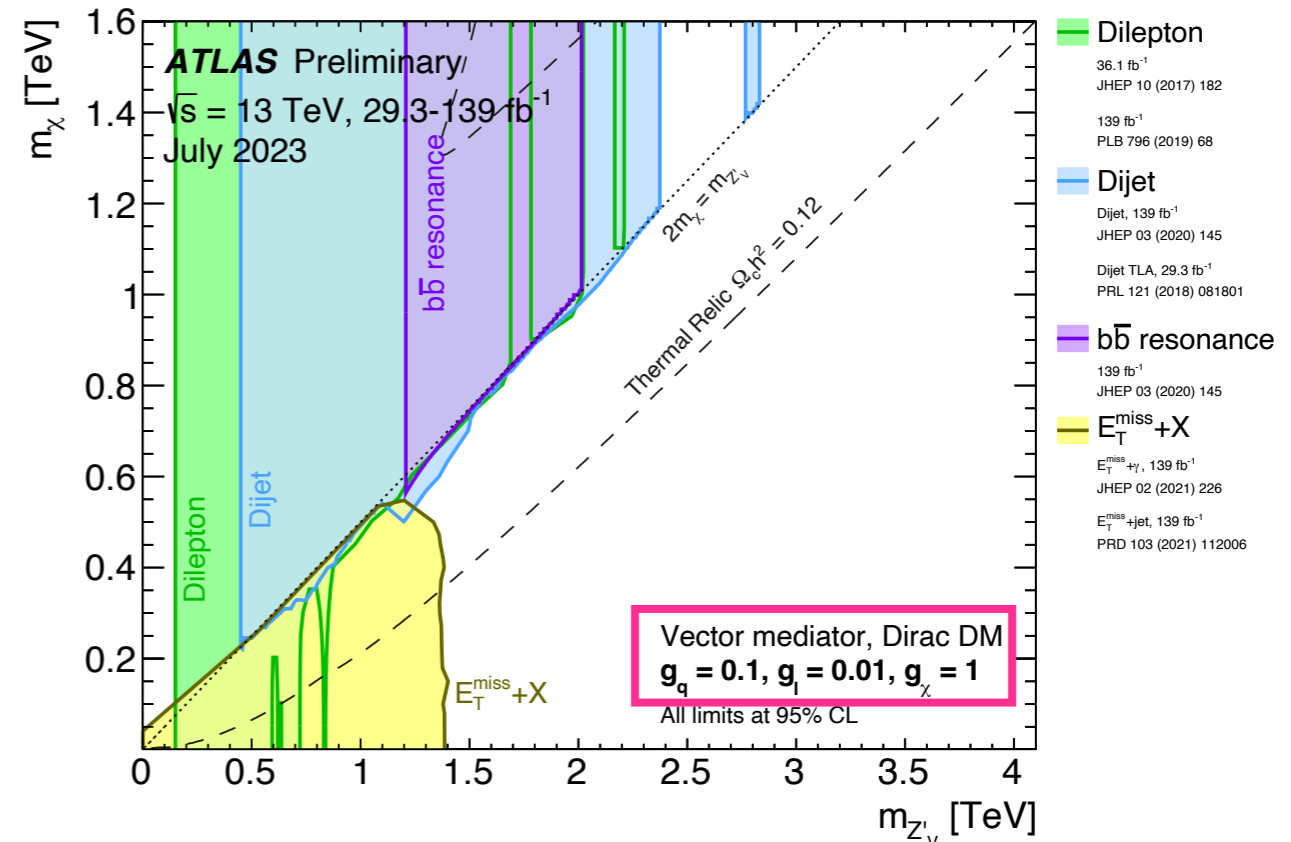
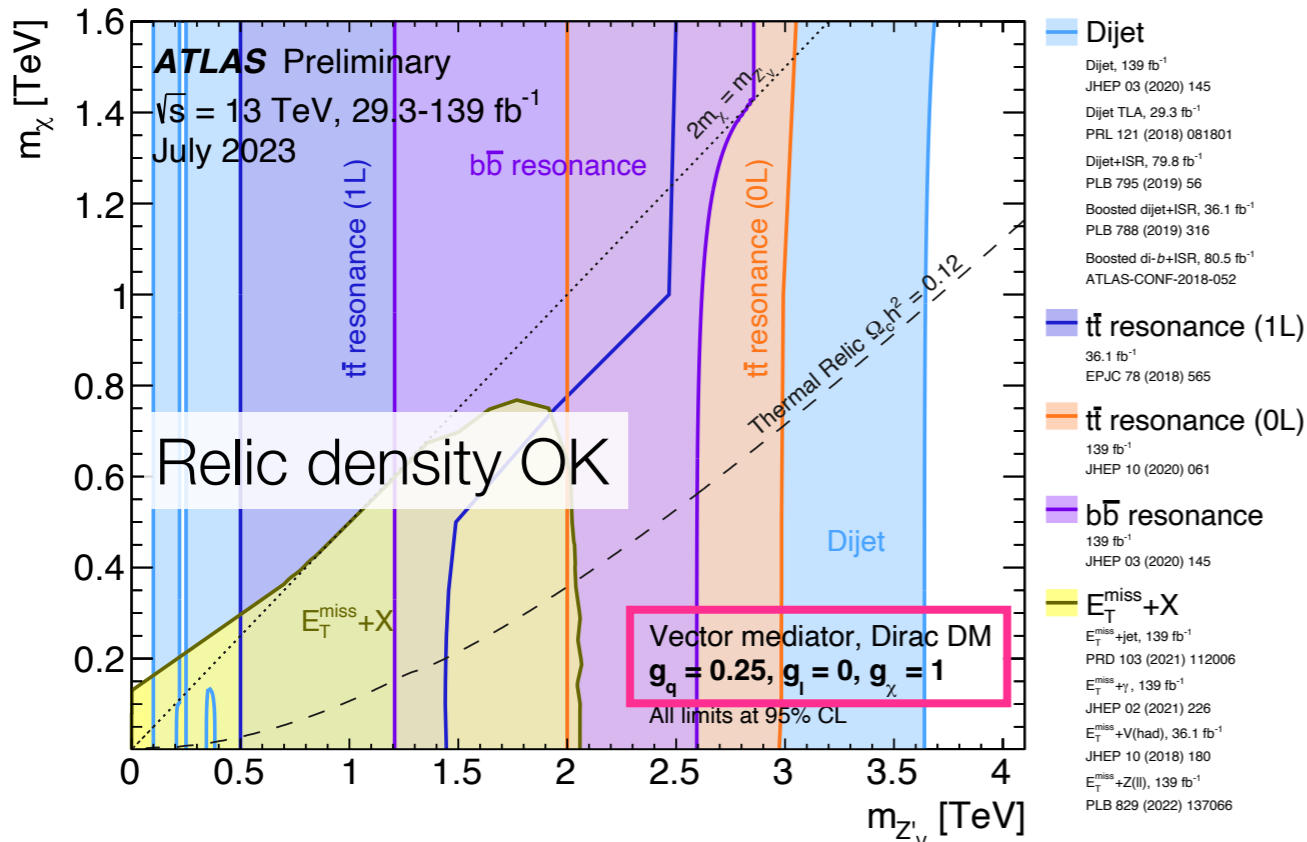
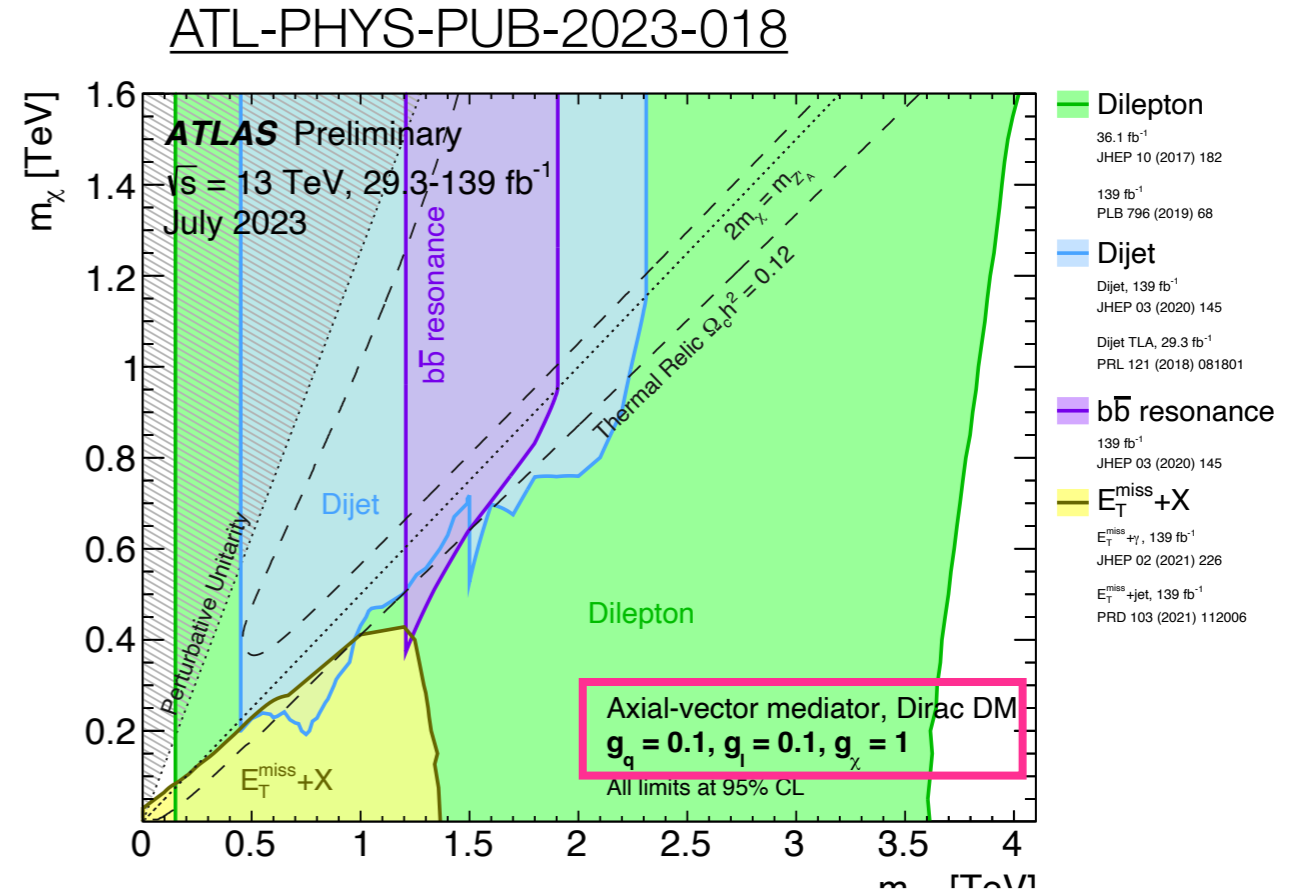
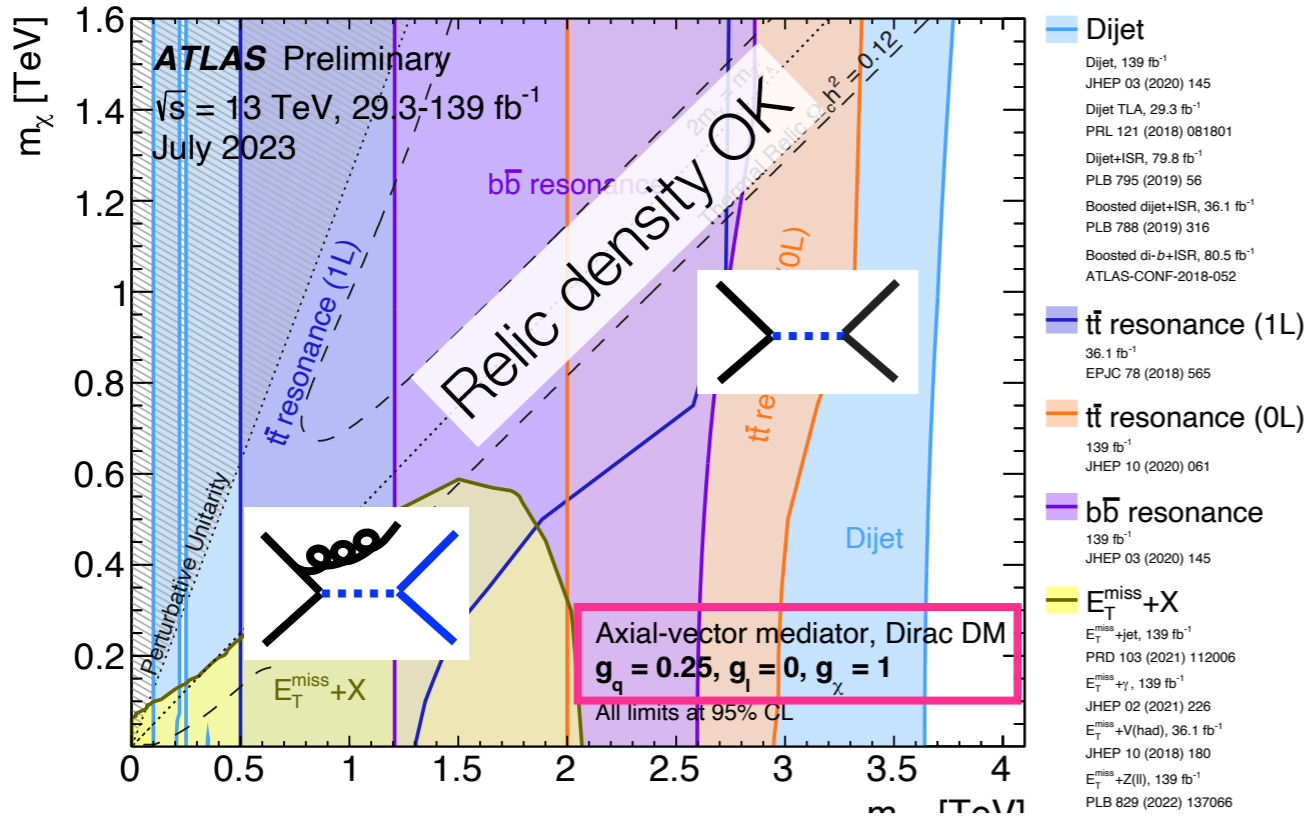


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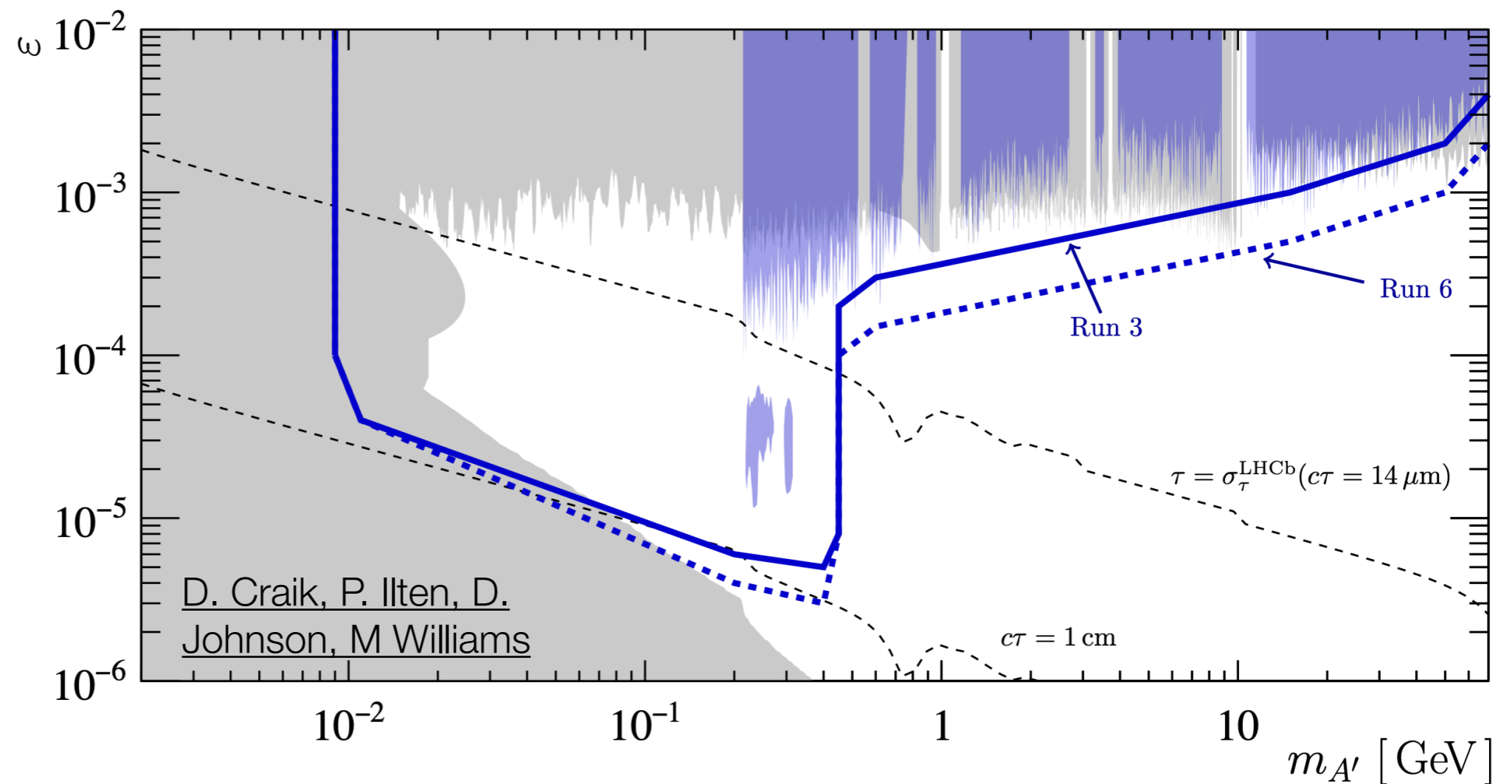
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Dark photons at the LHC

Very popular spin-1 vector benchmark, especially with intensity frontier and physics beyond colliders community

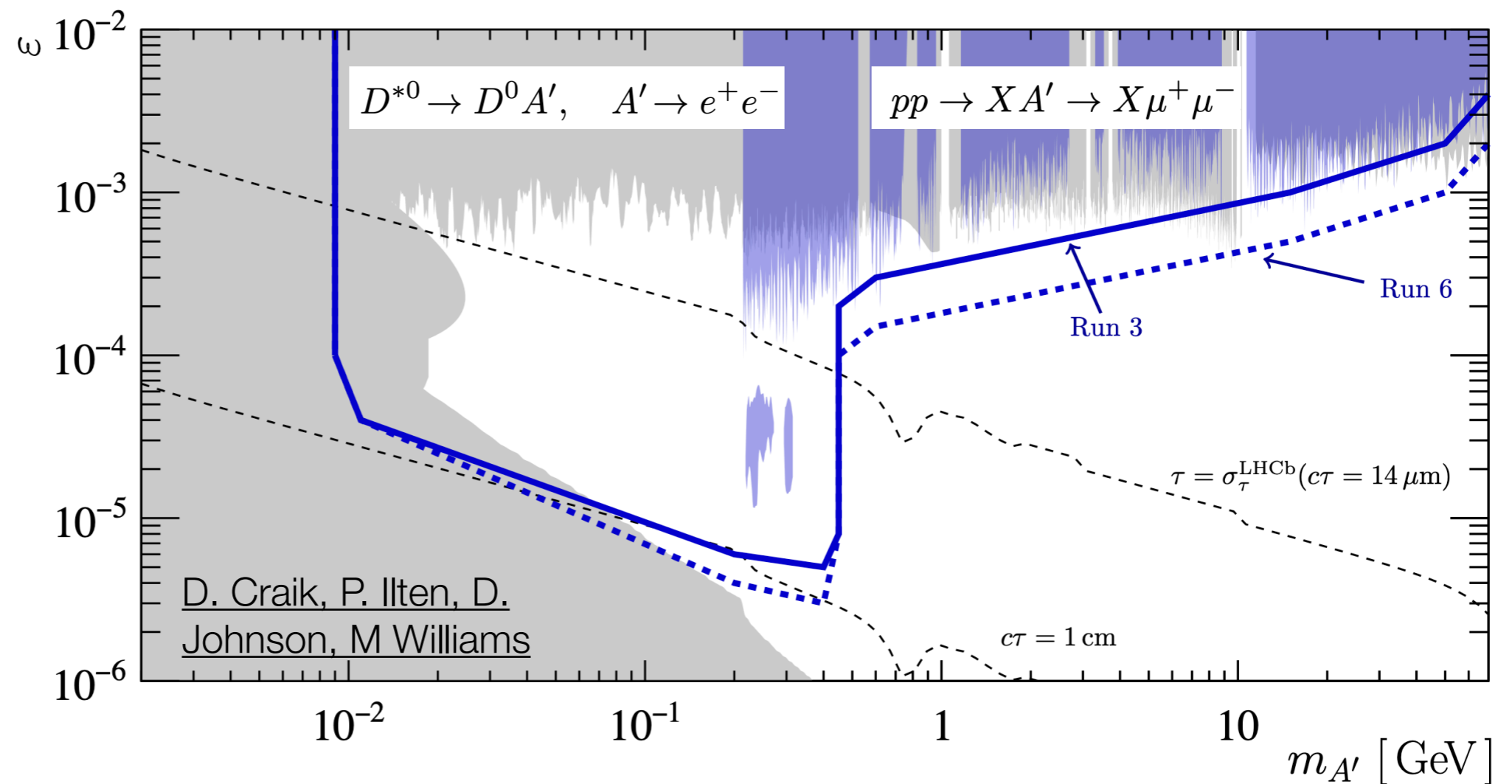
LHCb is a powerhouse with Run 3 triggerless readout



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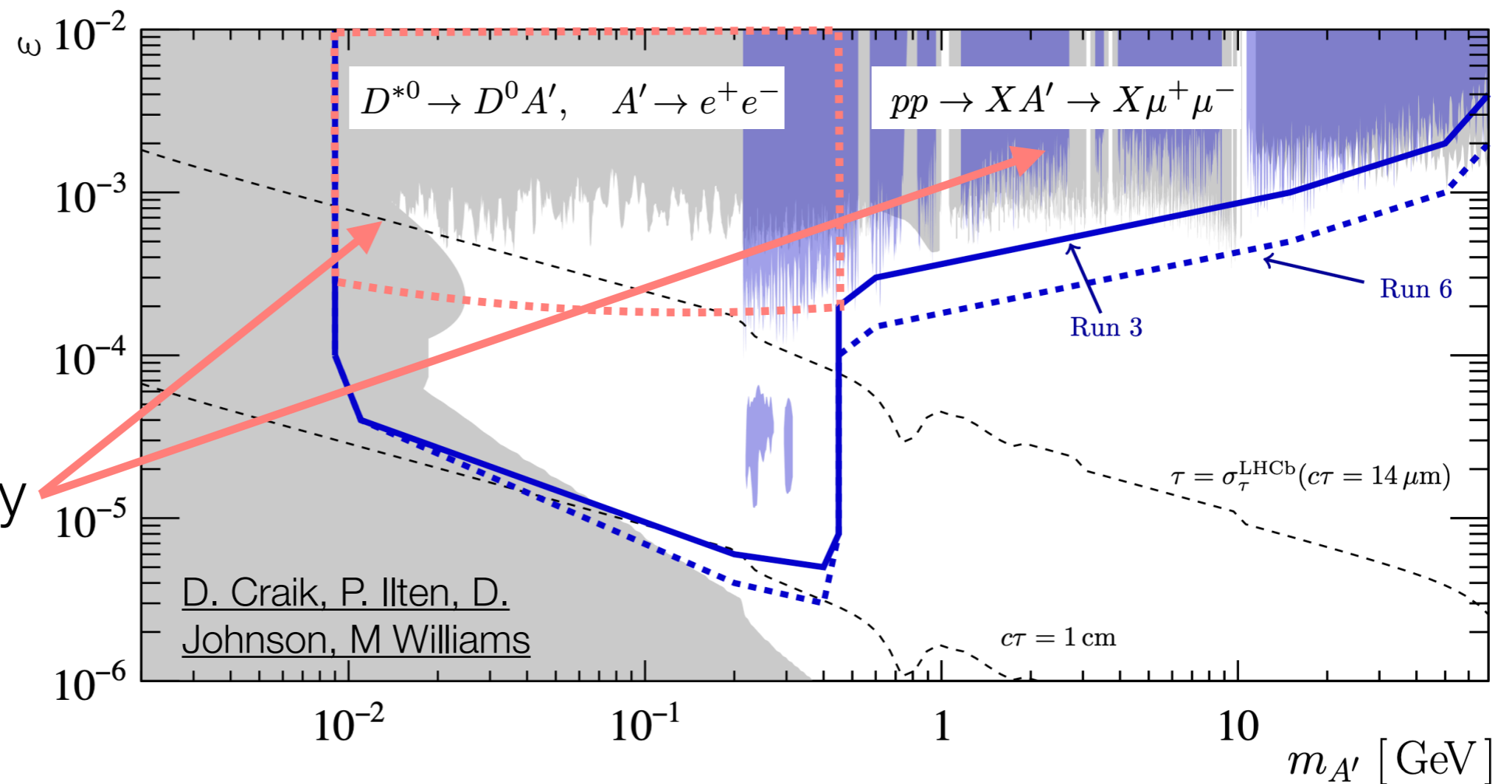


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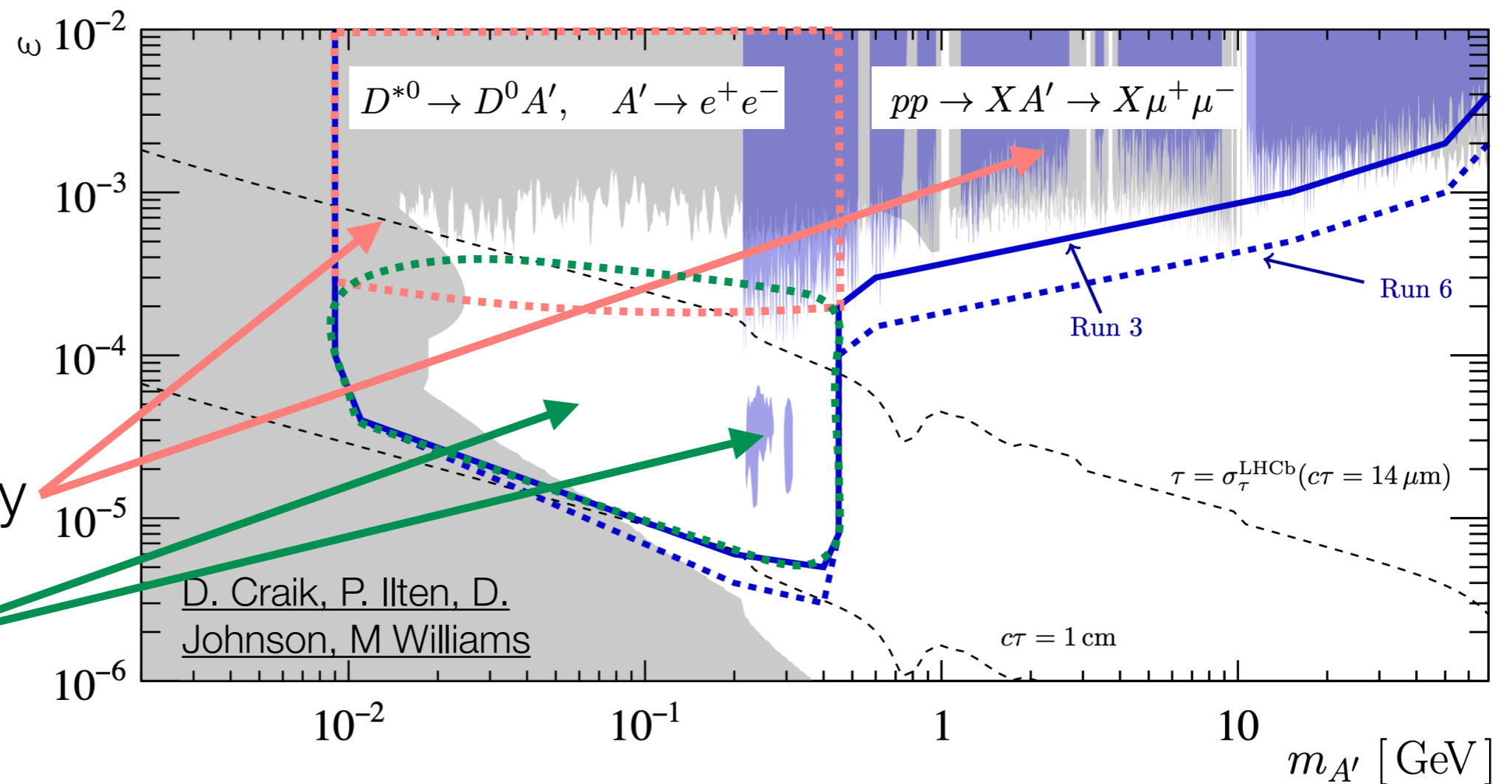
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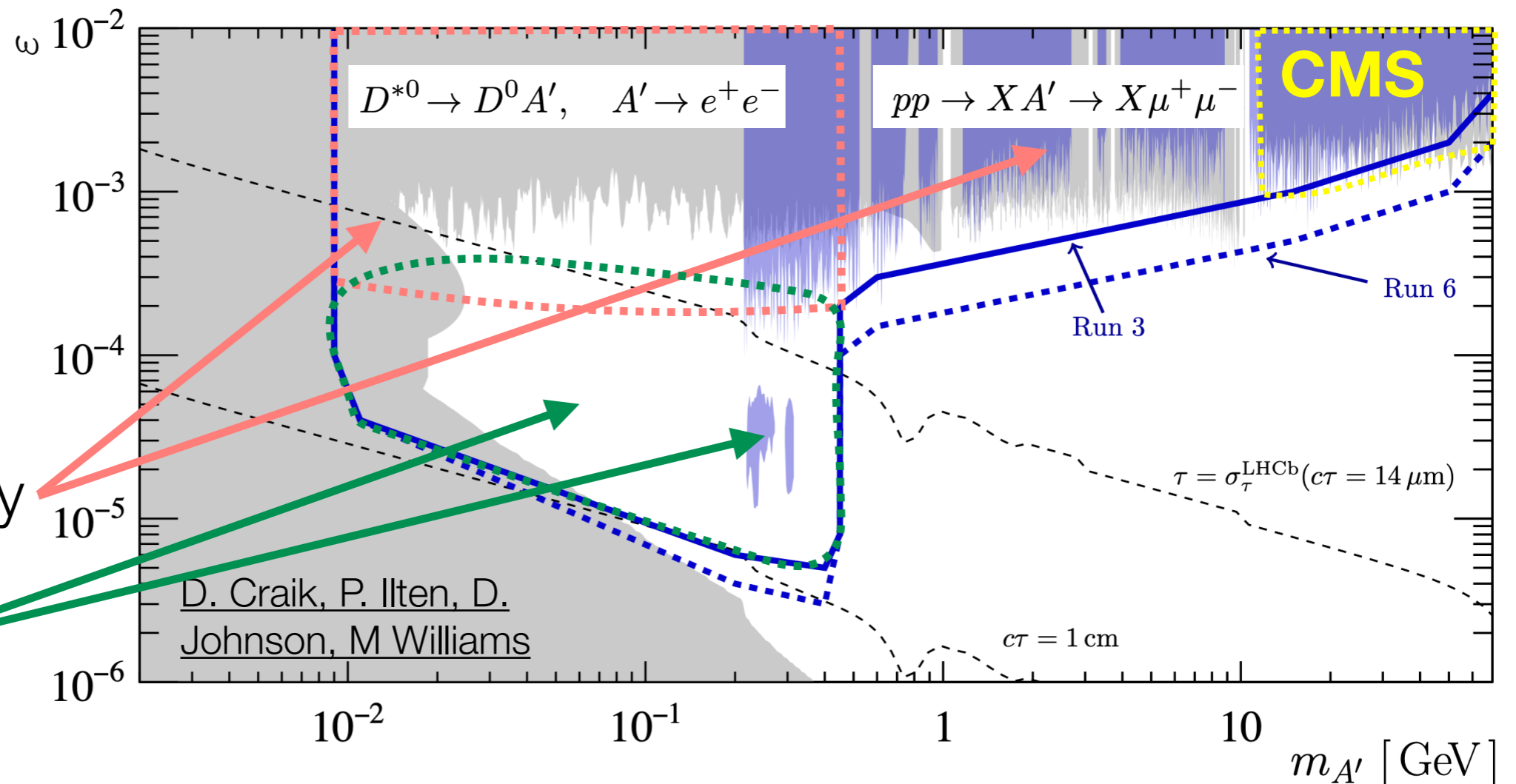
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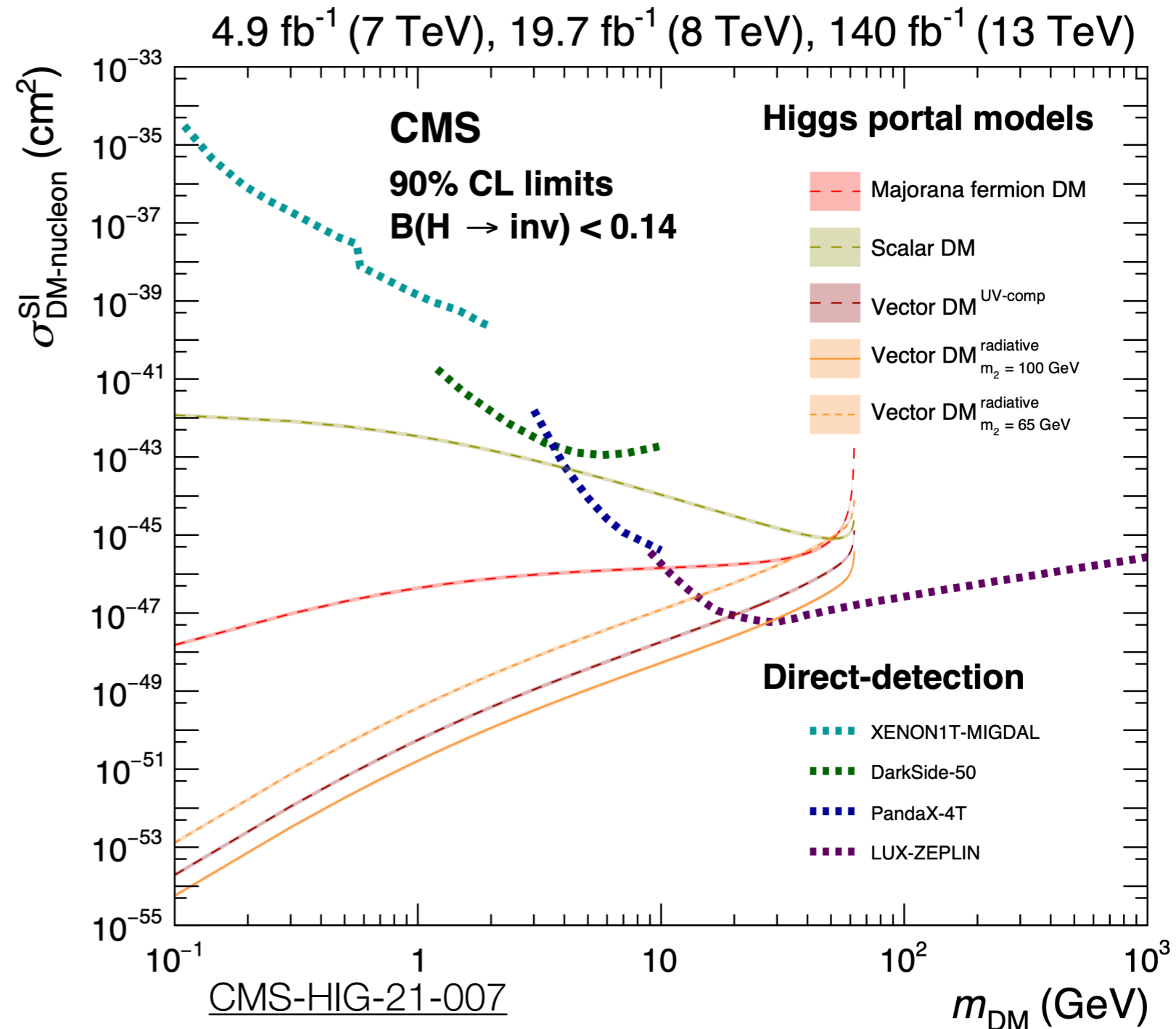
Displaced vertex



ATLAS & CMS can contribute at higher masses. Trigger poses a challenge. Simplified spin-1 limits translate fairly directly, but this is not currently a standard interpretation.

Higgs decays to dark matter

In Higgs portal models, the Higgs decays to DM, creating a MET signature

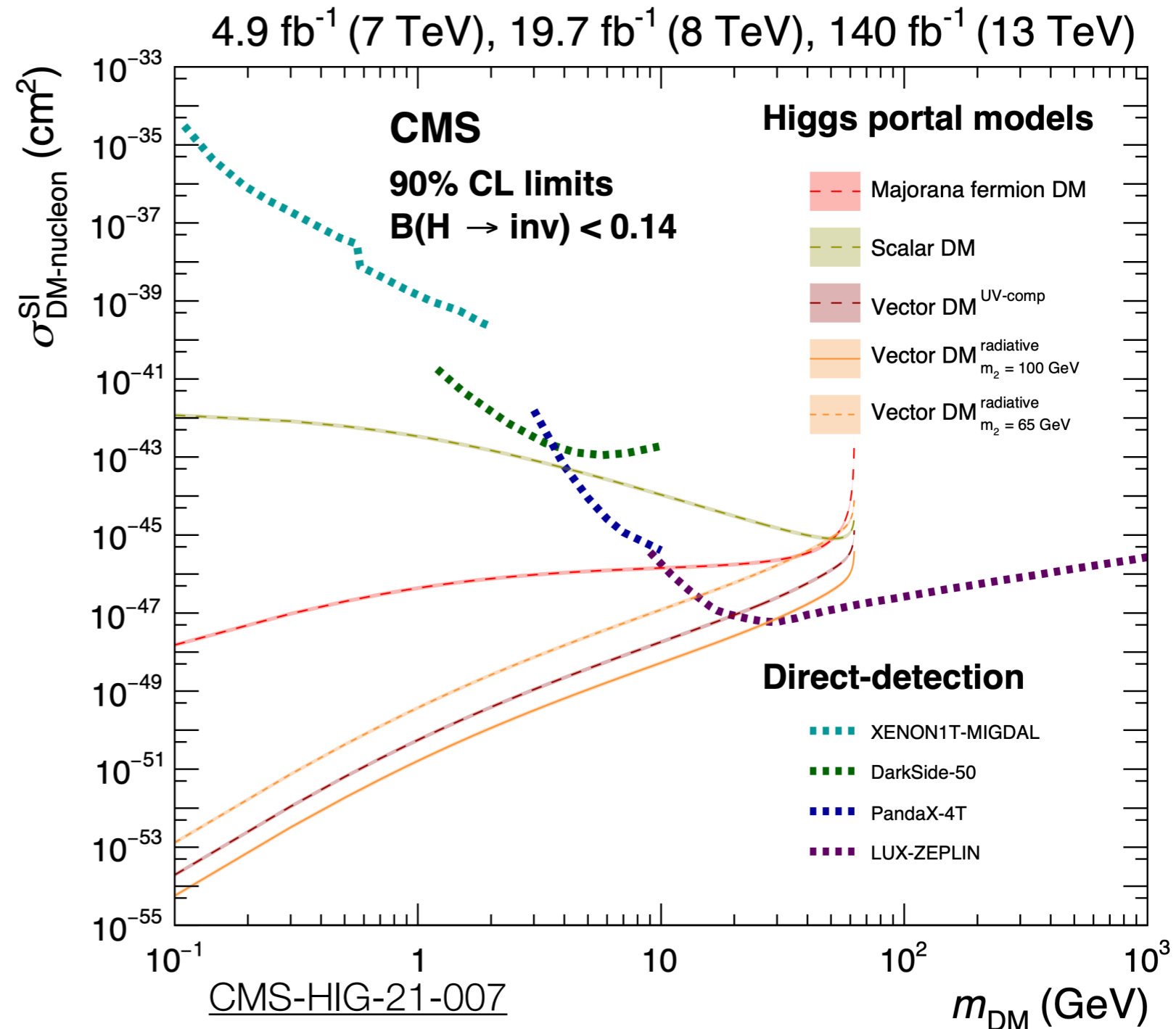


Model motivation from [Arcadi, Djouadi, and Kado](#)

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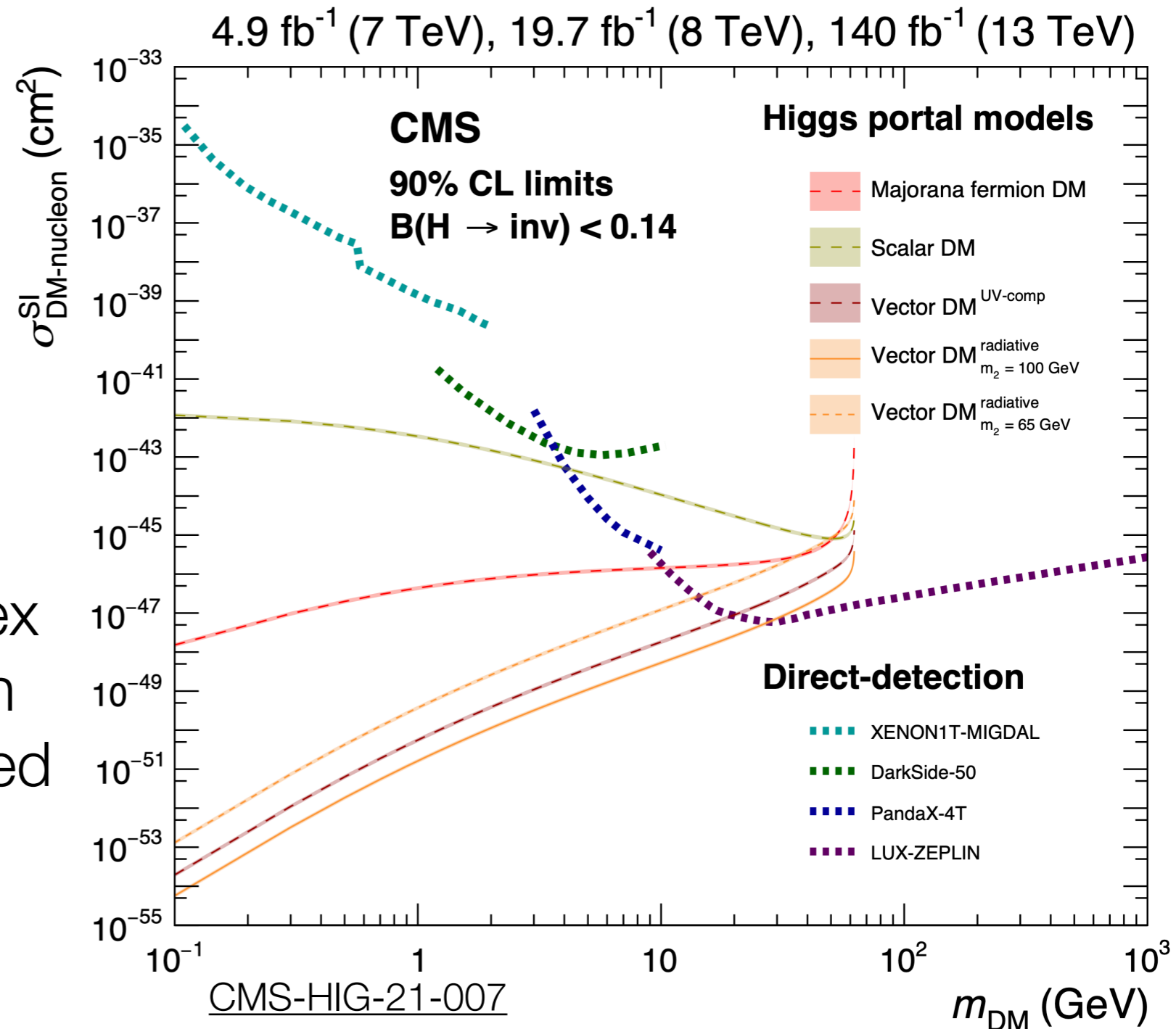
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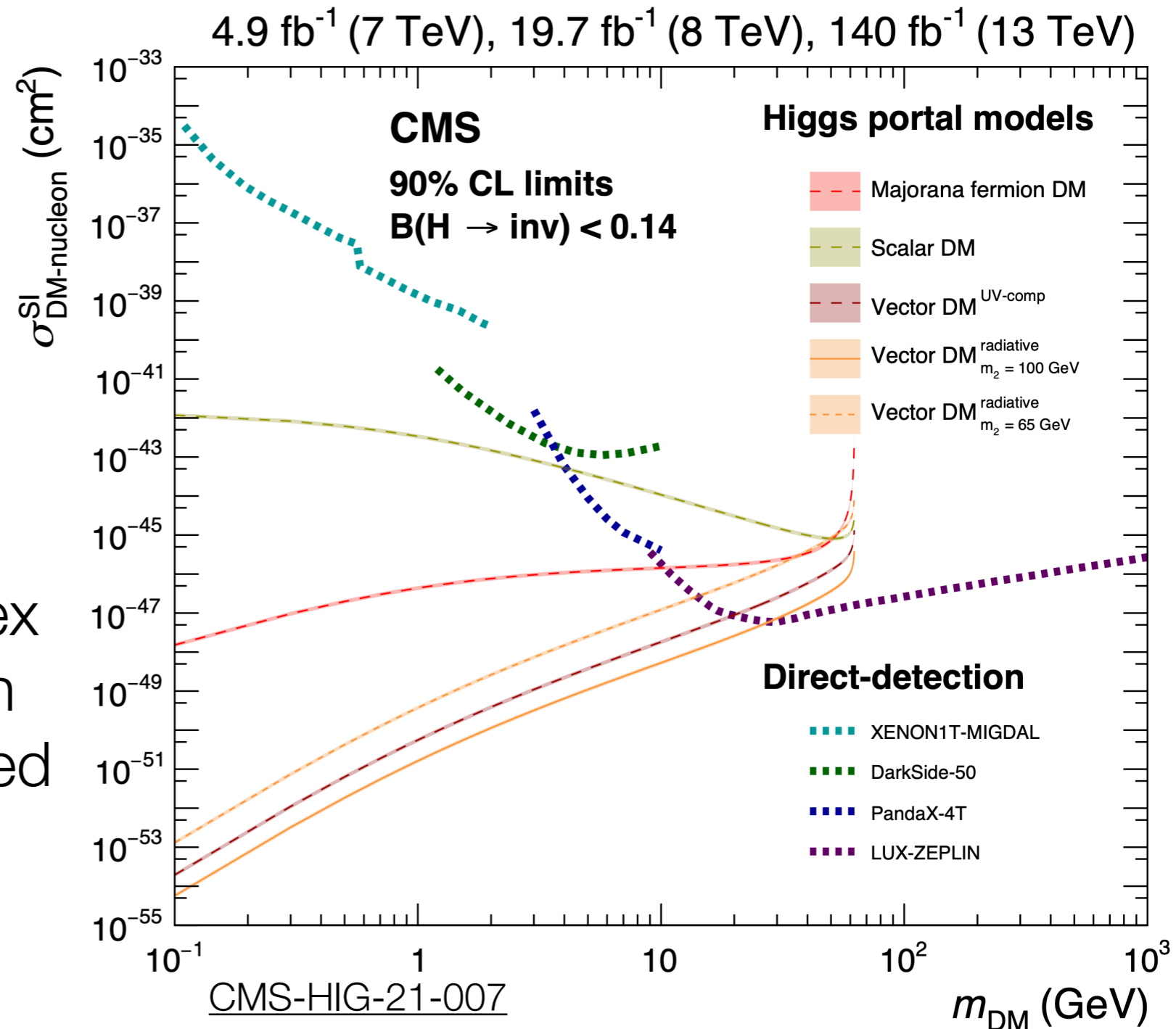
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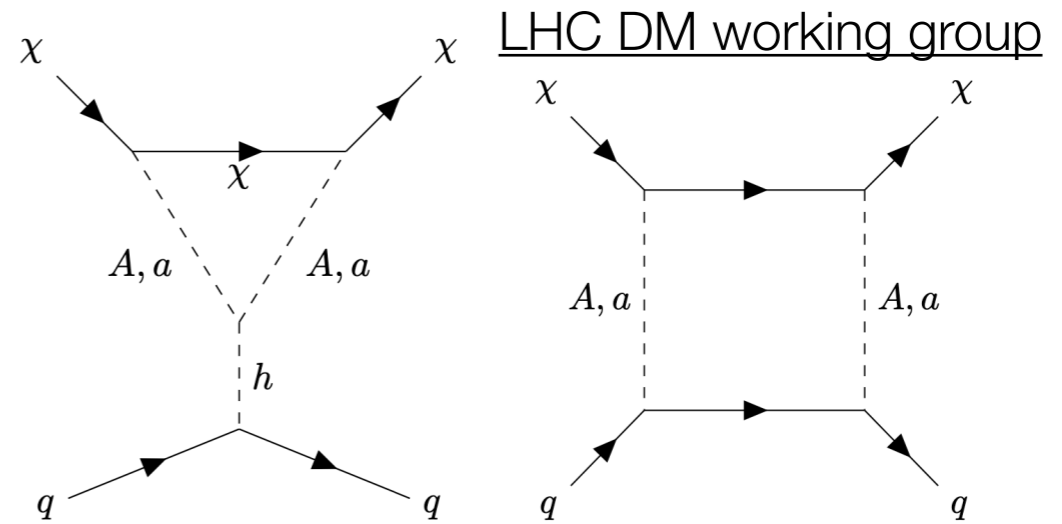
Current upper limits on $BR(h \rightarrow \text{inv}) \sim 0.11$ ([ATLAS](#))



Model motivation from [Arcadi, Djouadi, and Kado](#)

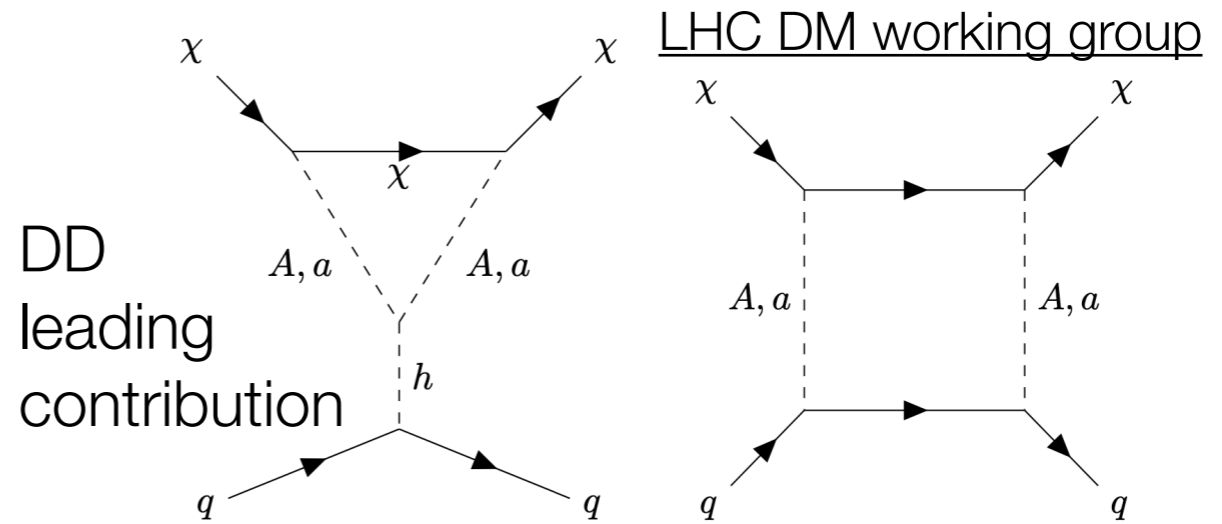
2HDM+a motivation and limits

DM with pseudoscalar mediator is a key LHC target because direct detection interactions are suppressed at tree level



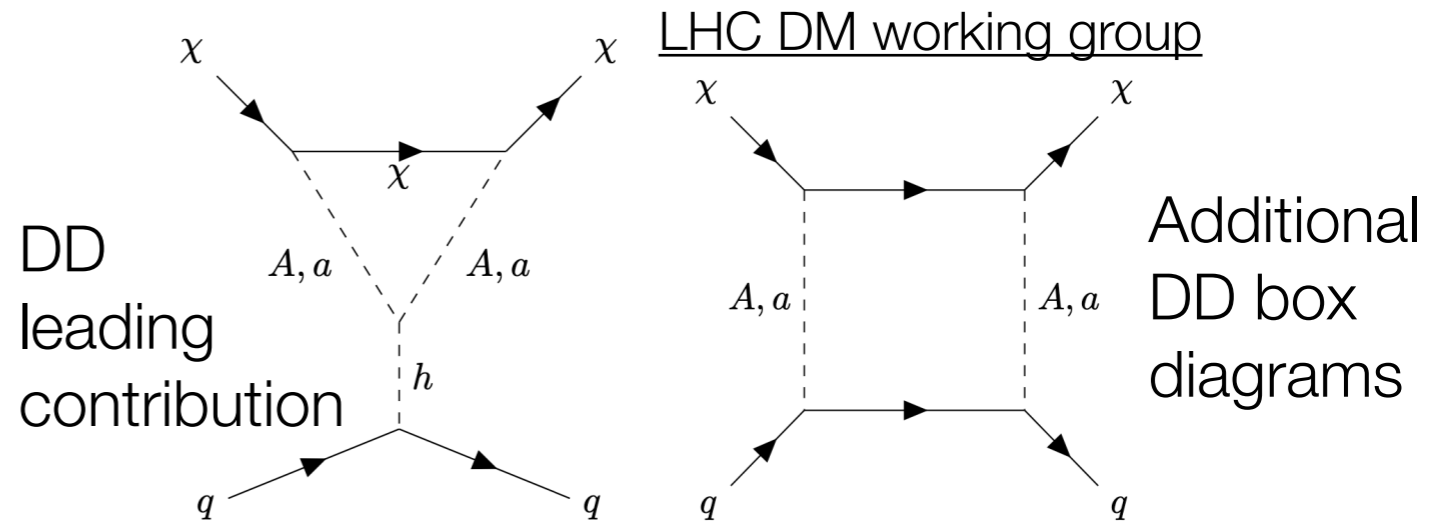
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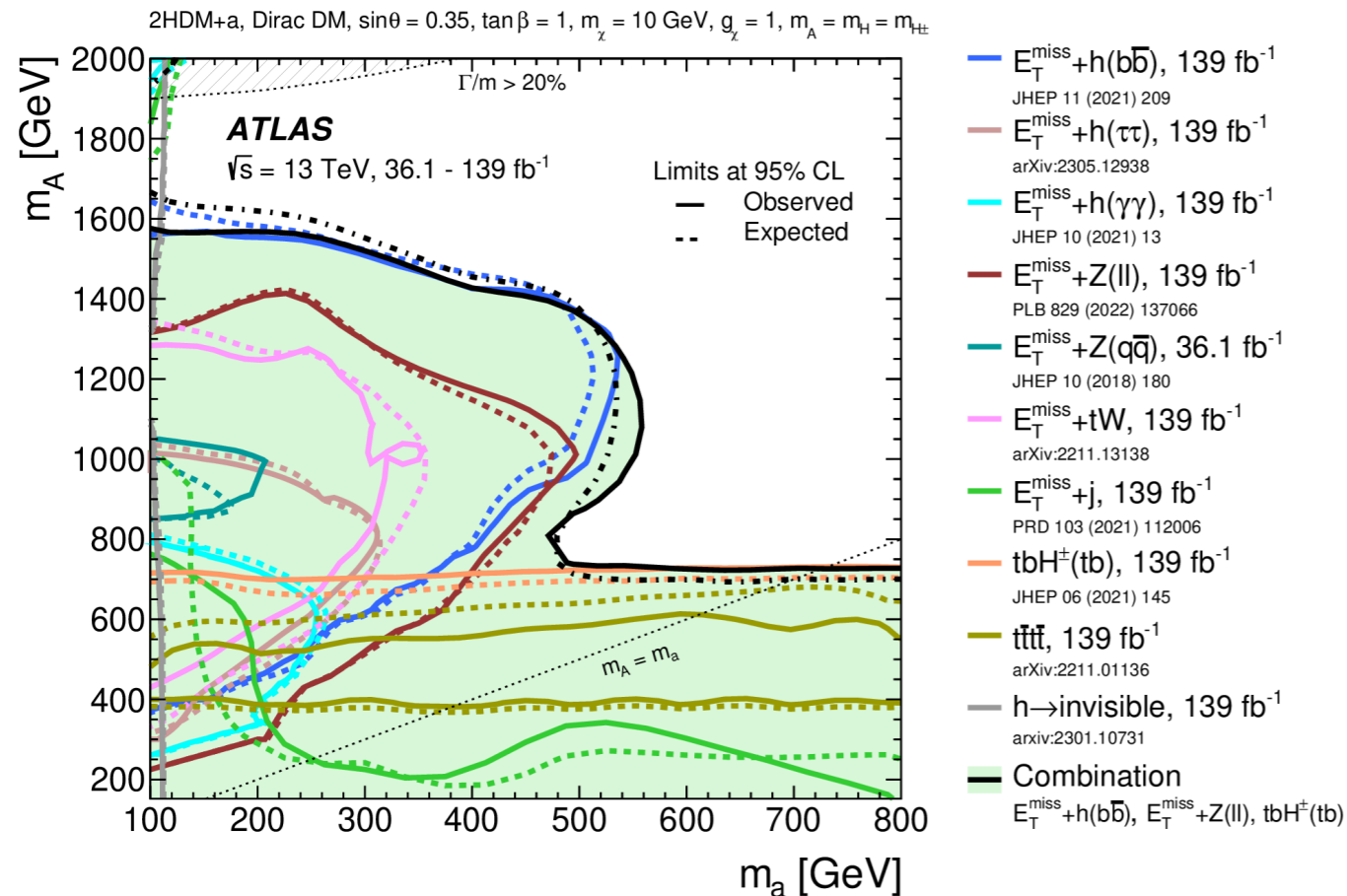
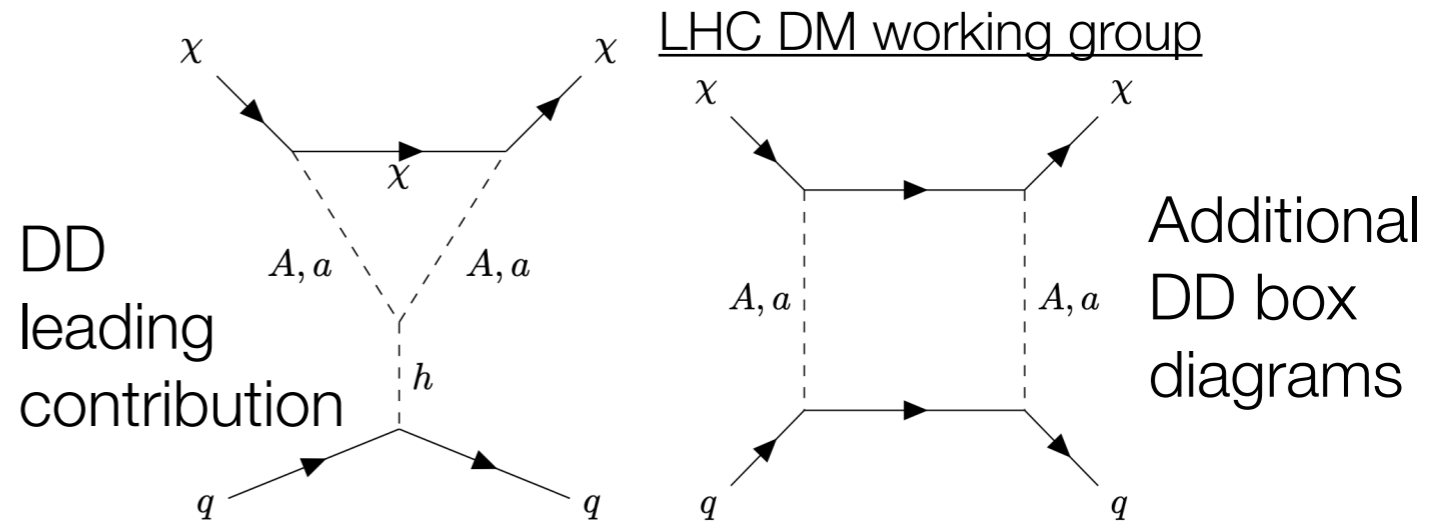
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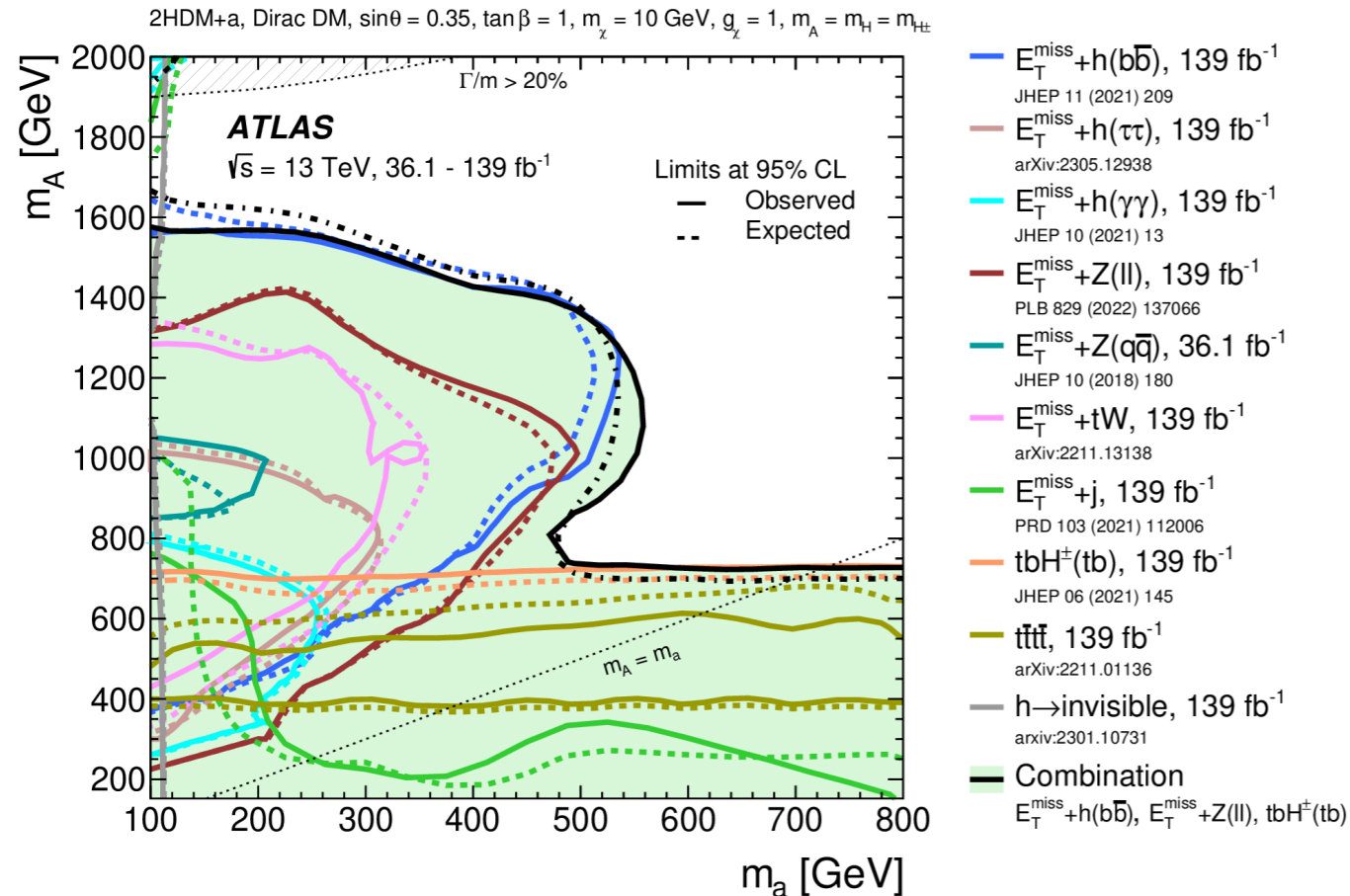
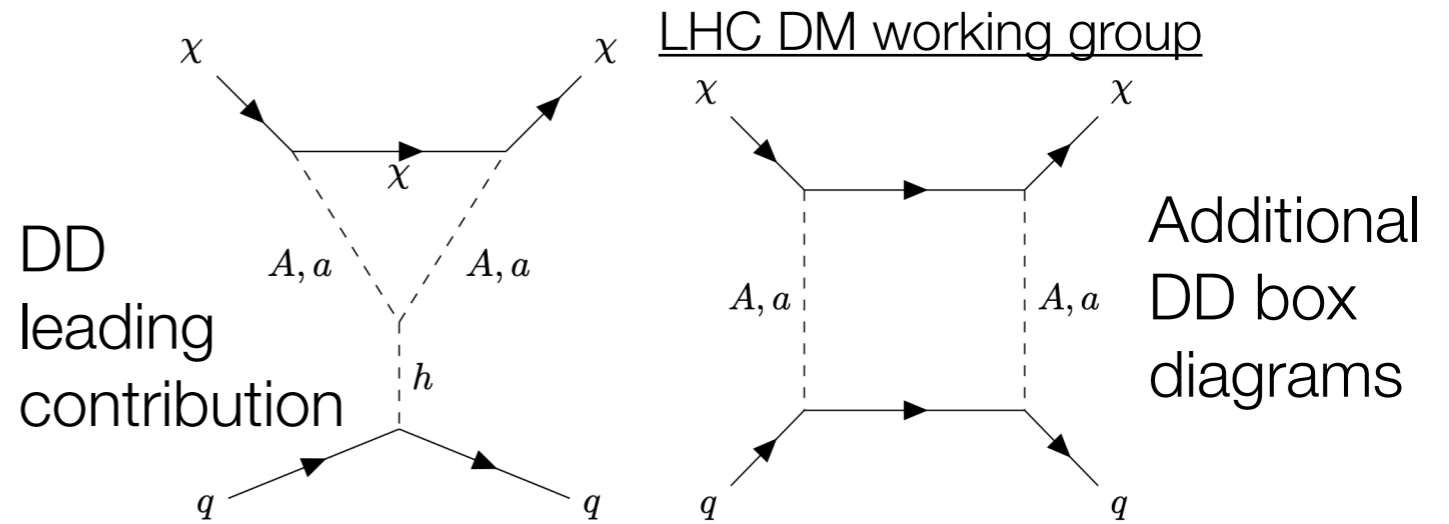
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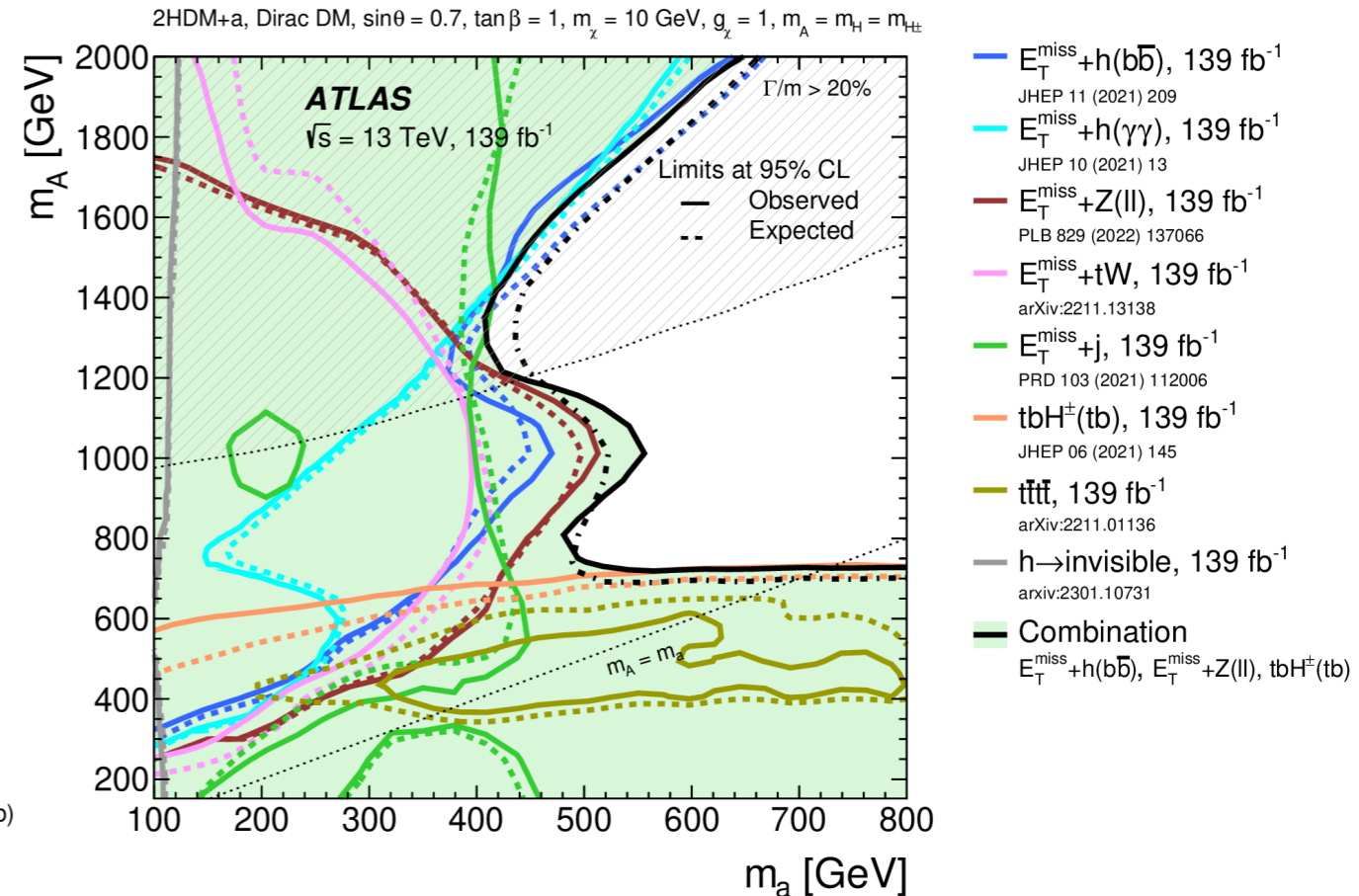


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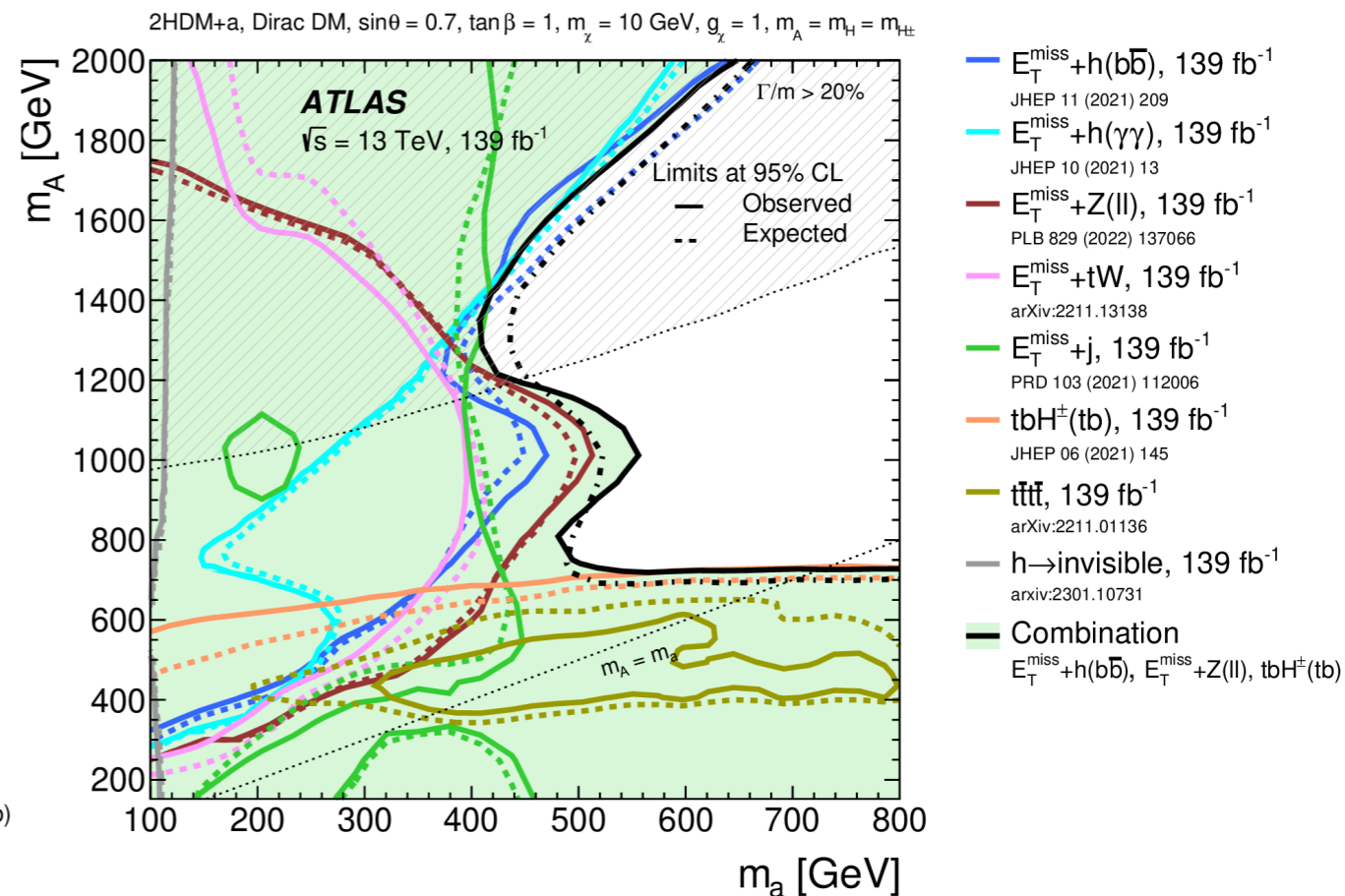
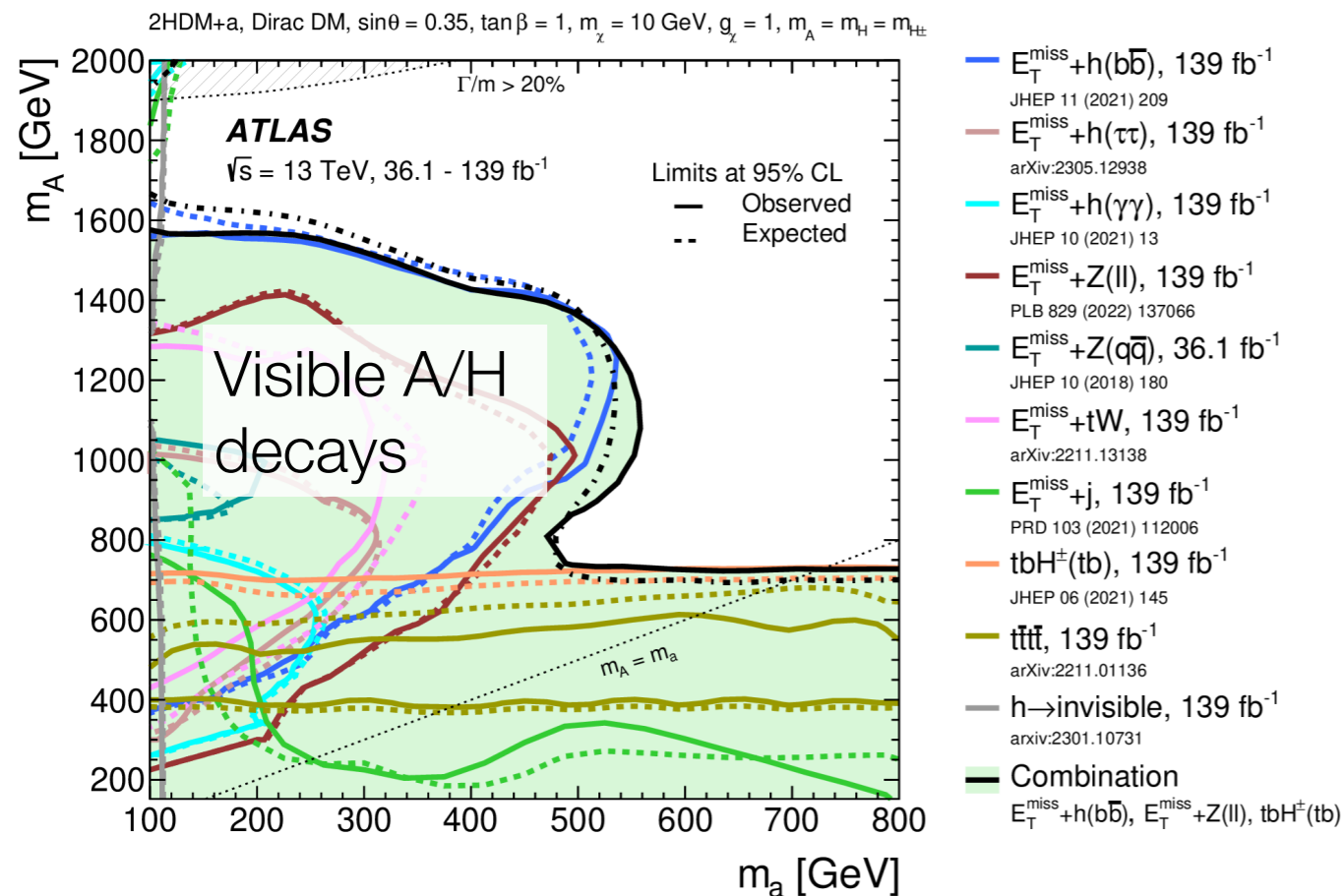
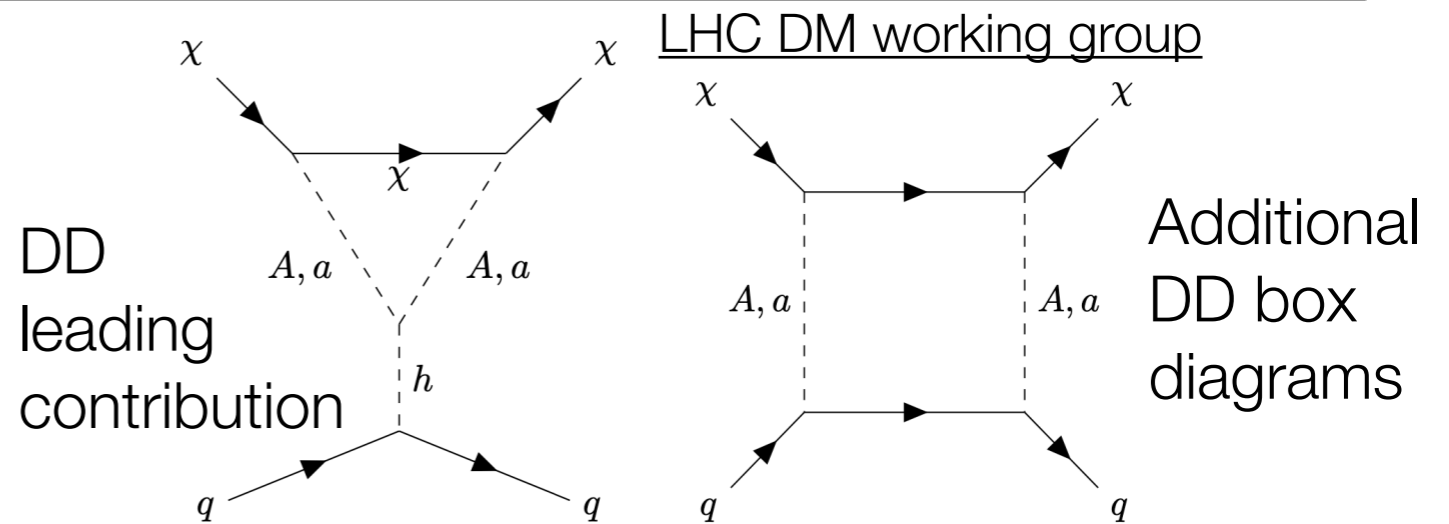


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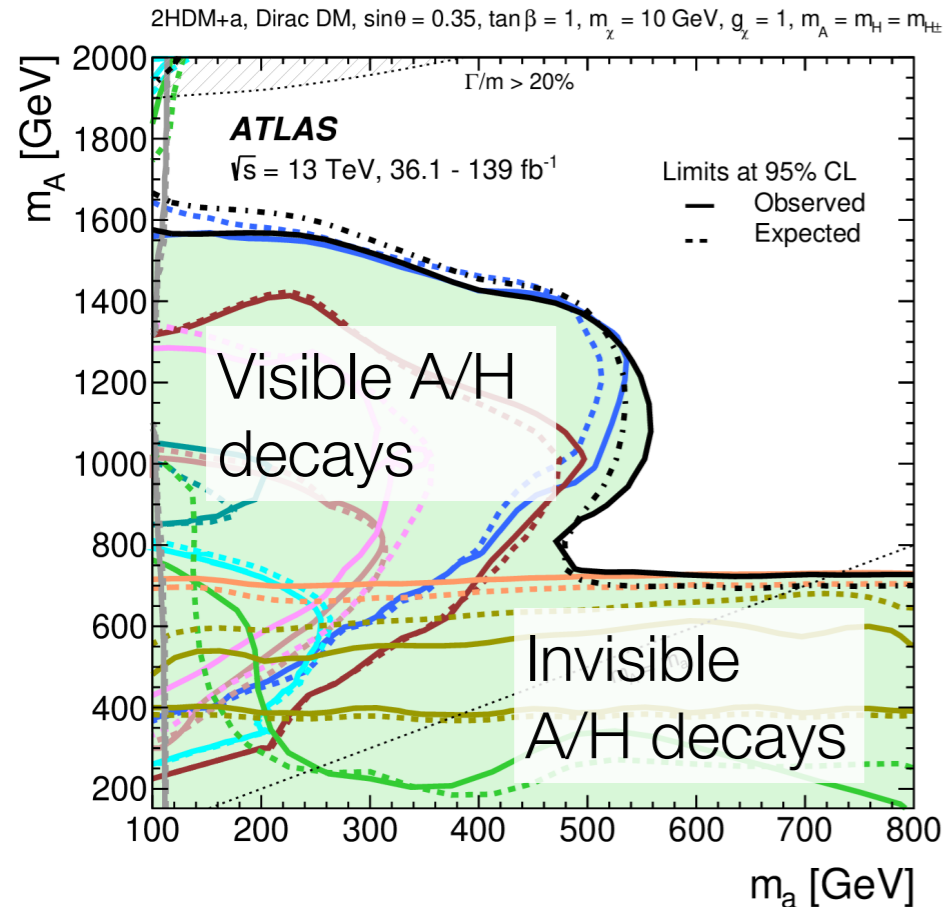
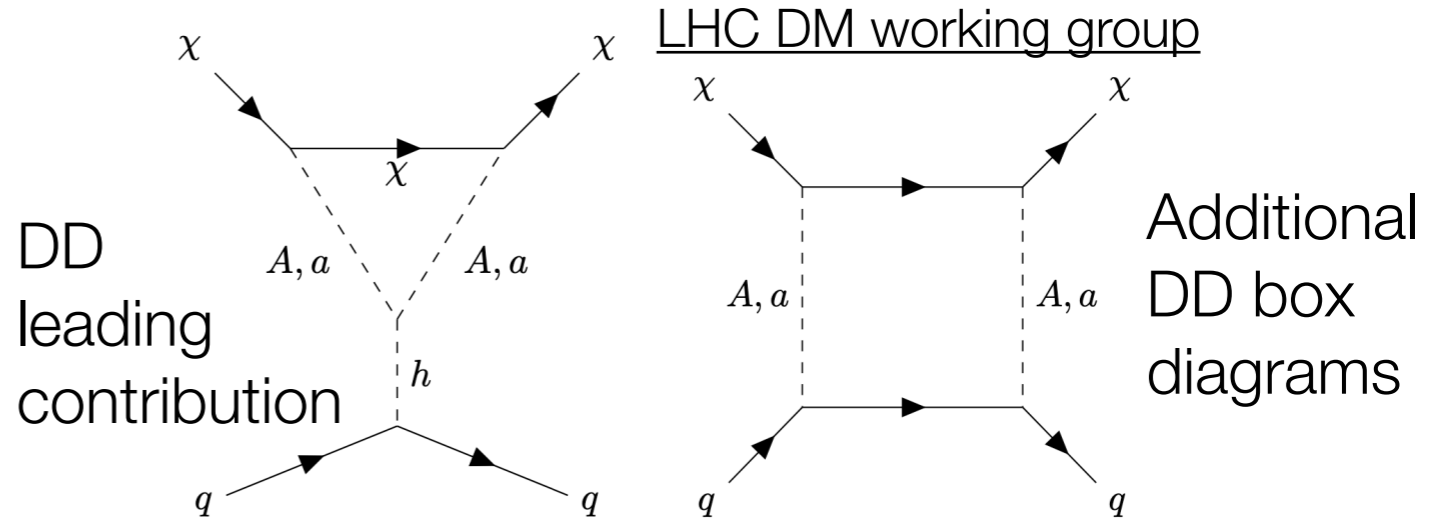
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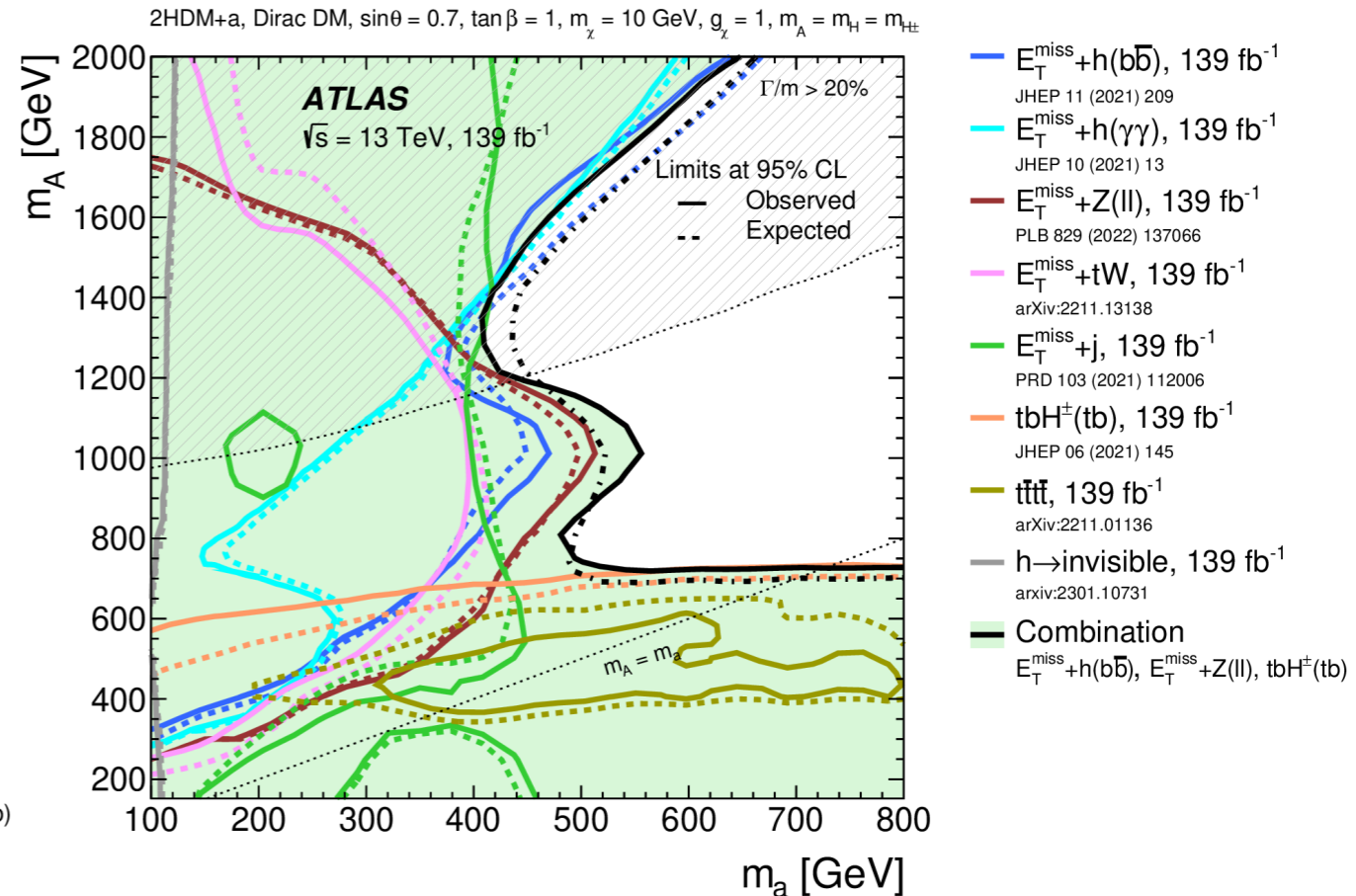


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- $E_T^{\text{miss}} + h(b\bar{b})$, 139 fb $^{-1}$
JHEP 11 (2021) 209
- $E_T^{\text{miss}} + h(\tau\tau)$, 139 fb $^{-1}$
arXiv:2305.12938
- $E_T^{\text{miss}} + h(\gamma\gamma)$, 139 fb $^{-1}$
JHEP 10 (2021) 13
- $E_T^{\text{miss}} + Z(\ell\ell)$, 139 fb $^{-1}$
PLB 829 (2022) 137066
- $E_T^{\text{miss}} + Z(q\bar{q})$, 36.1 fb $^{-1}$
JHEP 10 (2018) 180
- $E_T^{\text{miss}} + tW$, 139 fb $^{-1}$
arXiv:2211.13138
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The state of SUSY dark matter

Let's look at pMSSM scan of DM candidates

ATLAS CERN-EP-2024-021

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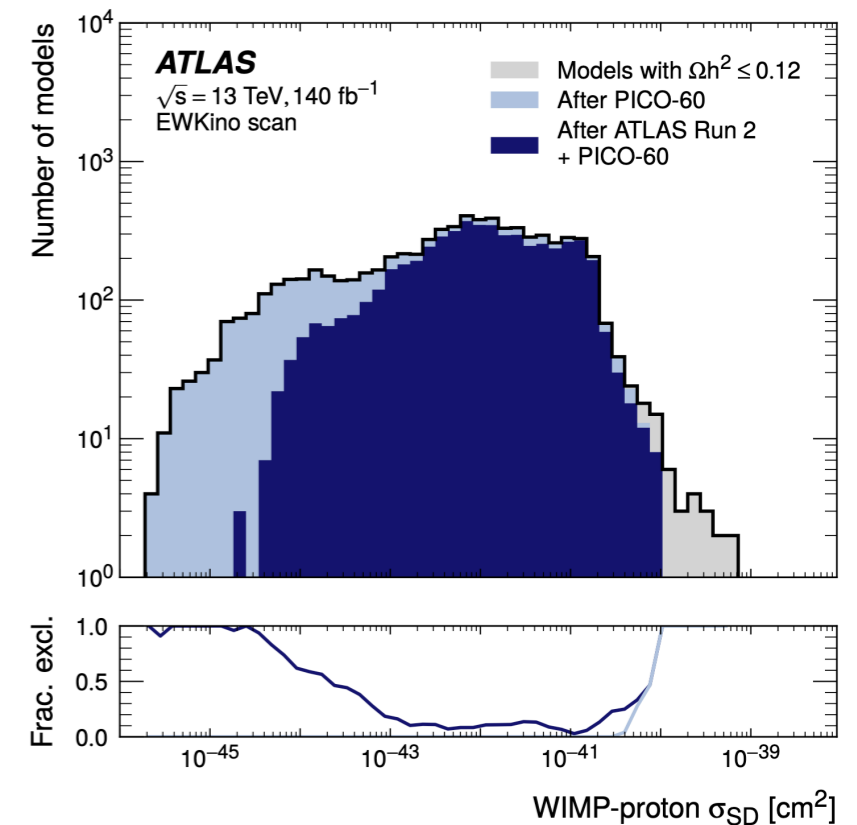
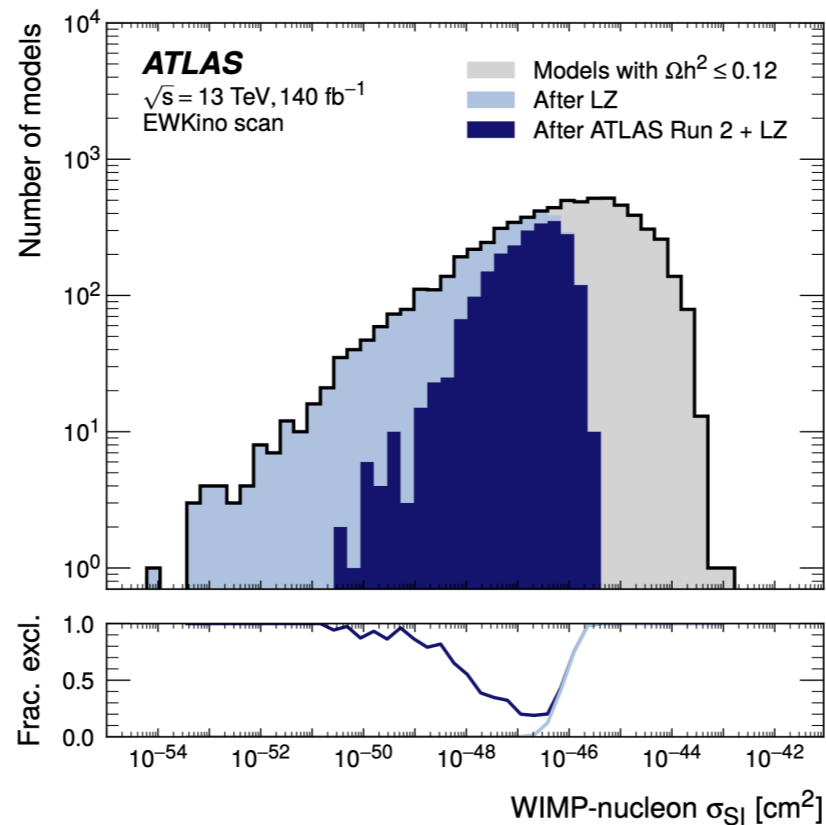
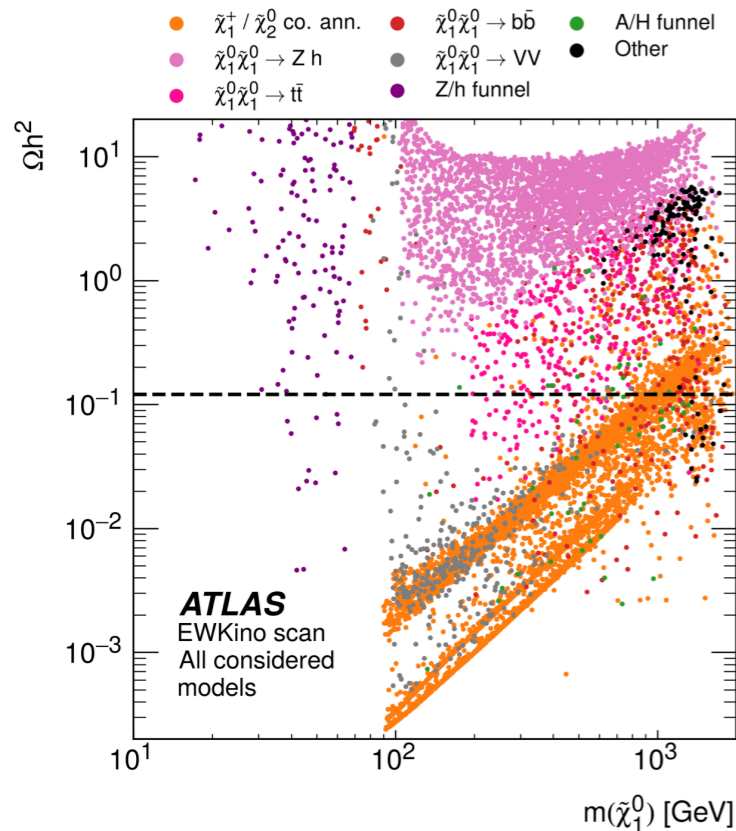
Co-annihilation with small mass splitting from wino/higgsino-like $\tilde{\chi}_1^\pm$ and $\tilde{\chi}_2^0$ to LSP gives most of the viable candidates explored here

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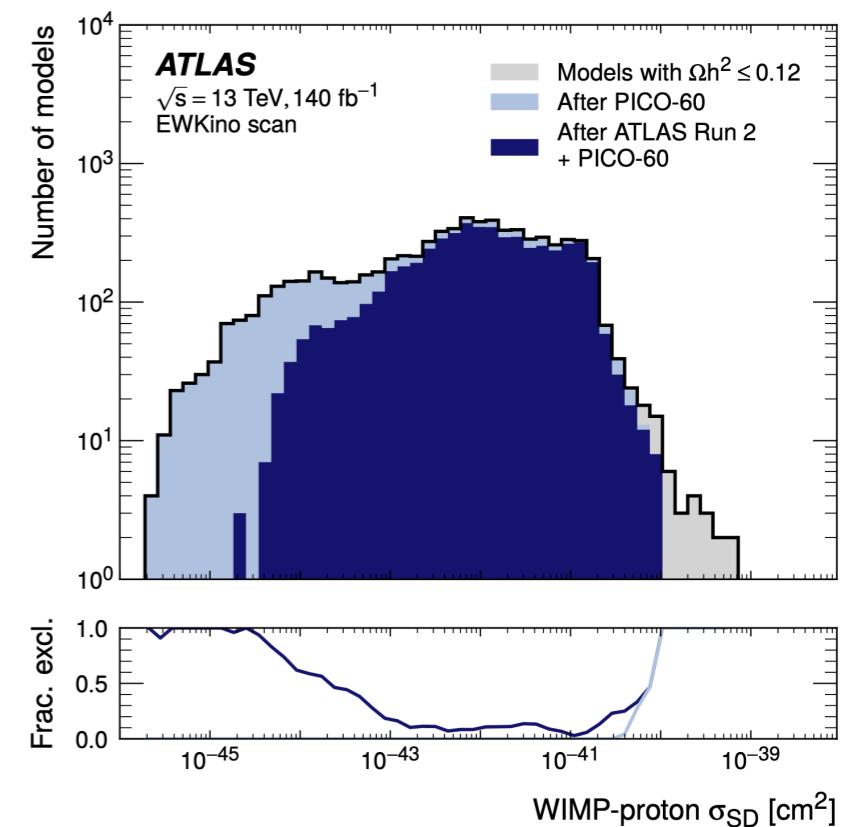
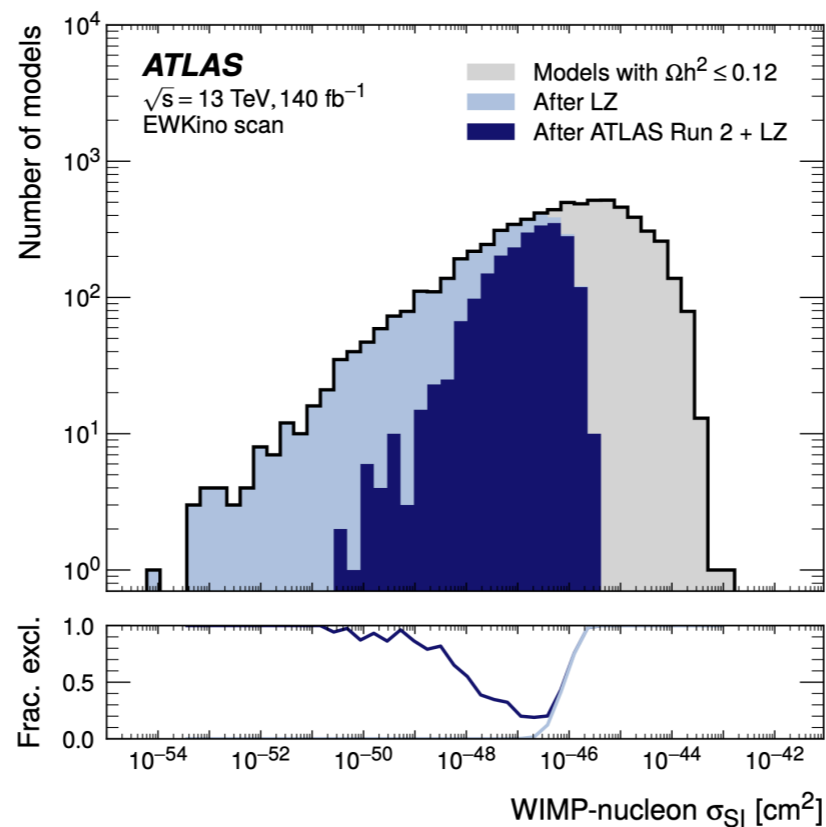
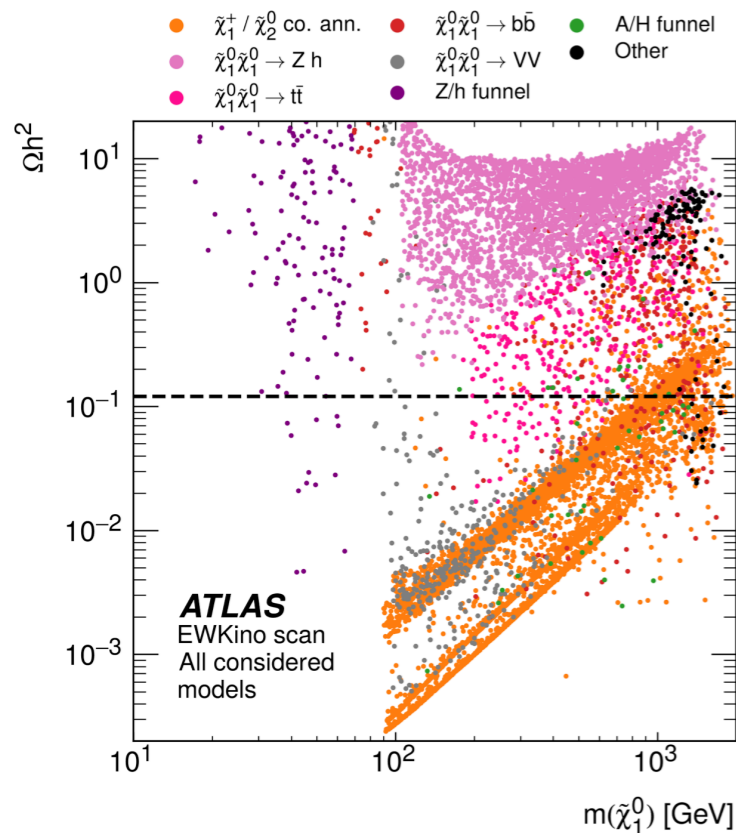


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ATLAS CERN-EP-2024-021

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Can see 1) there is considerable space left for SUSY DM candidates in hard-to-reach electroweak signatures, and 2) there is good complementarity between LHC and direct detection reach

Extended dark sectors: growing area of interest

Assume numerous additional particles, one of which could provide stable DM candidate

Extended dark sectors: growing area of interest

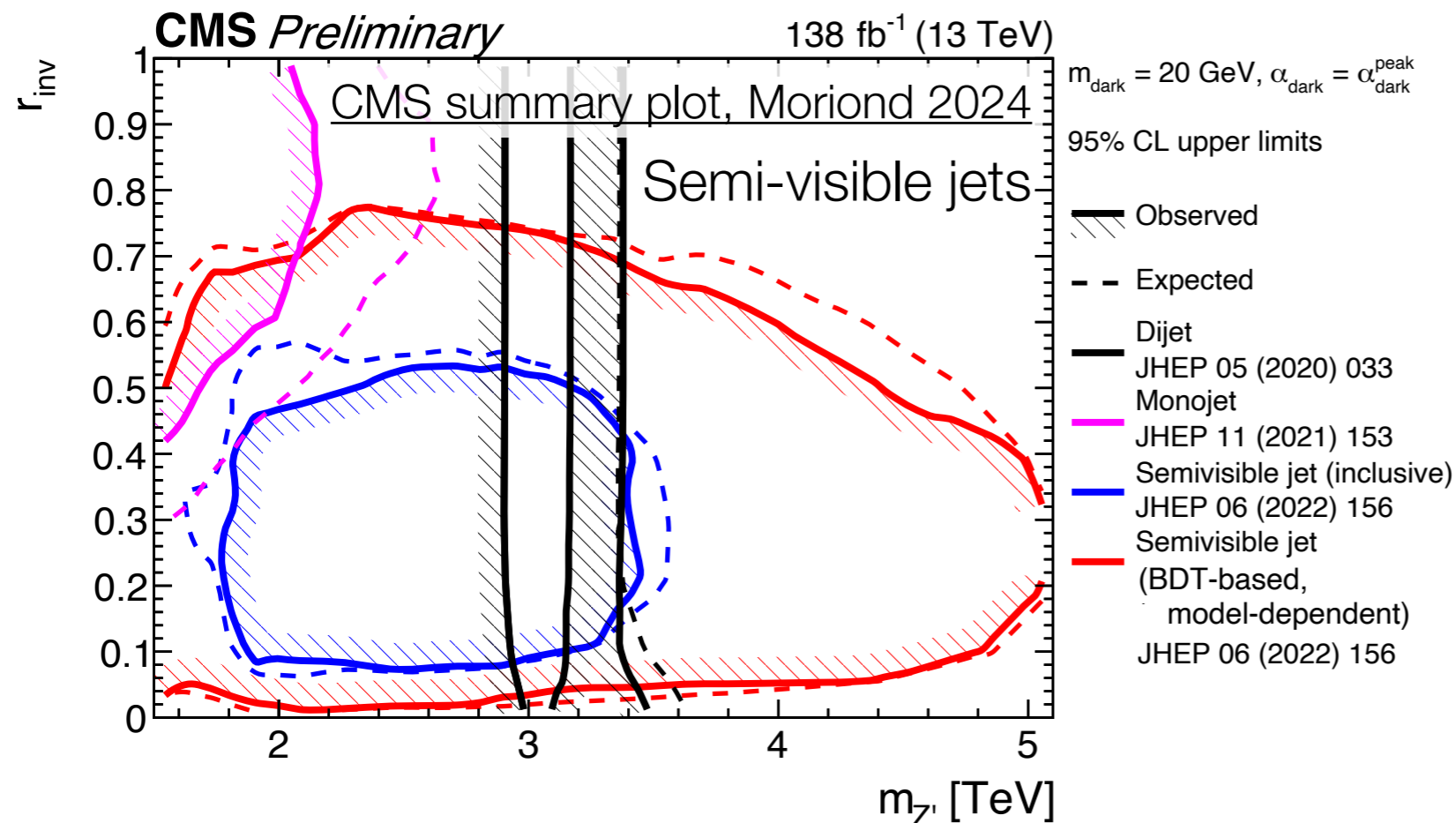
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Dark QCD & related give signatures with “weird jets”: containing displaced vertices, high fraction of invisible particles, etc depending on model details. Other cases give no jets at all (e.g. SUEPs)

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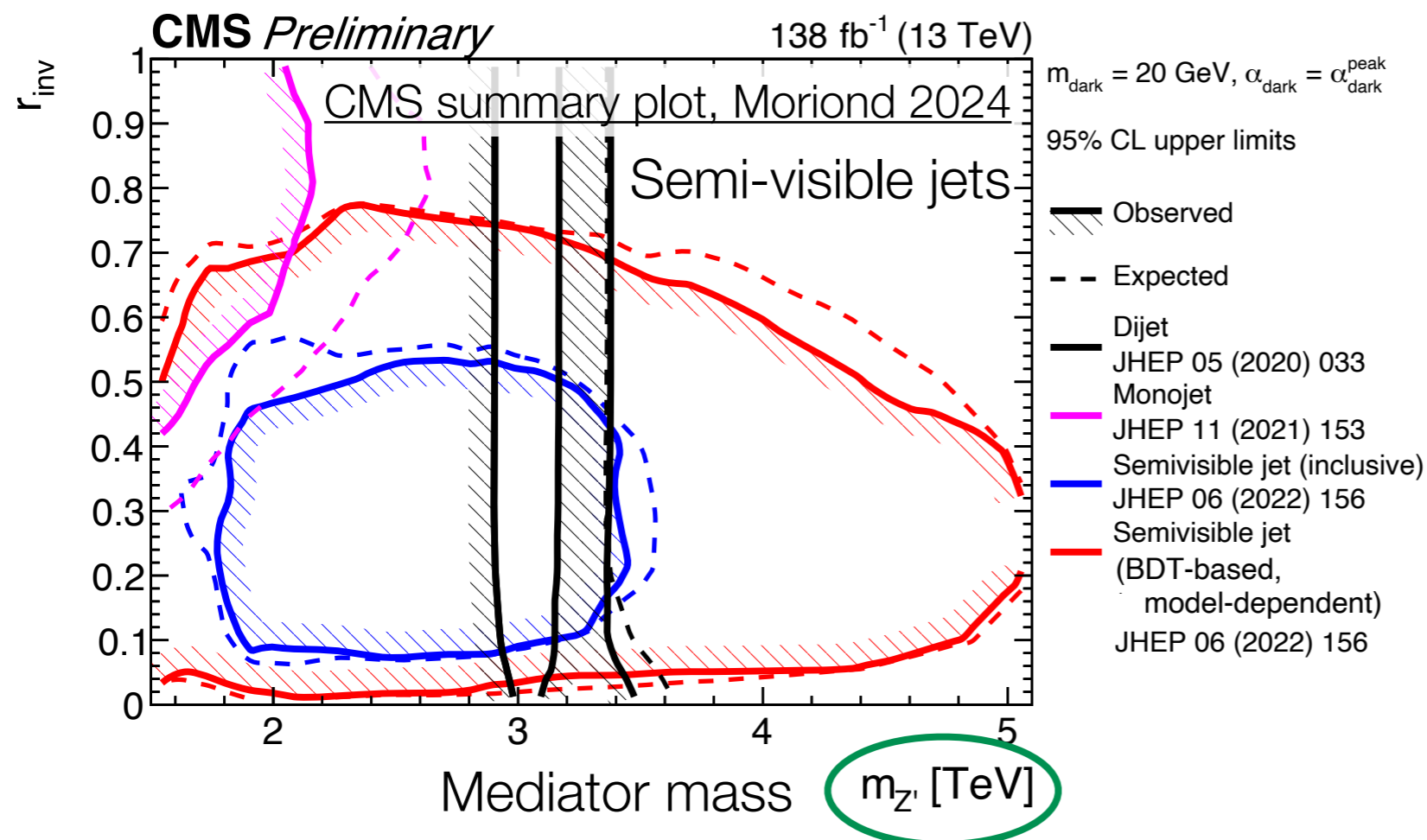
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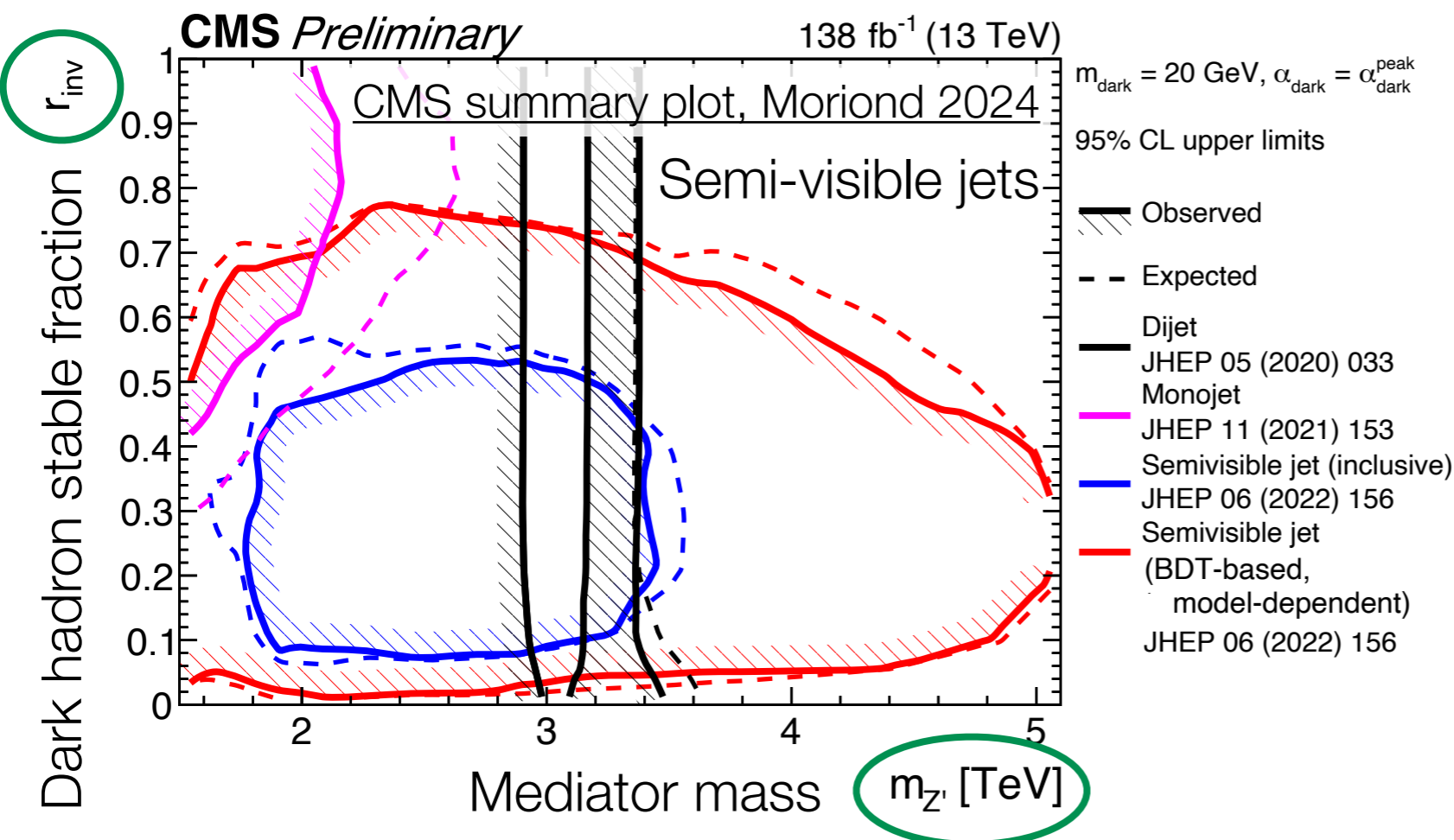
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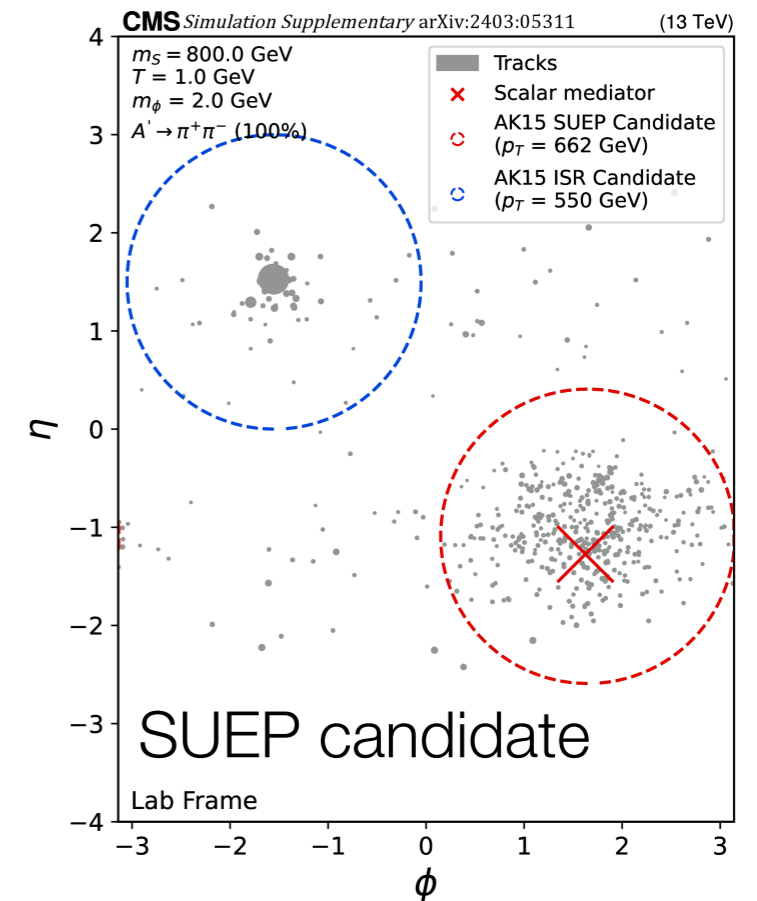
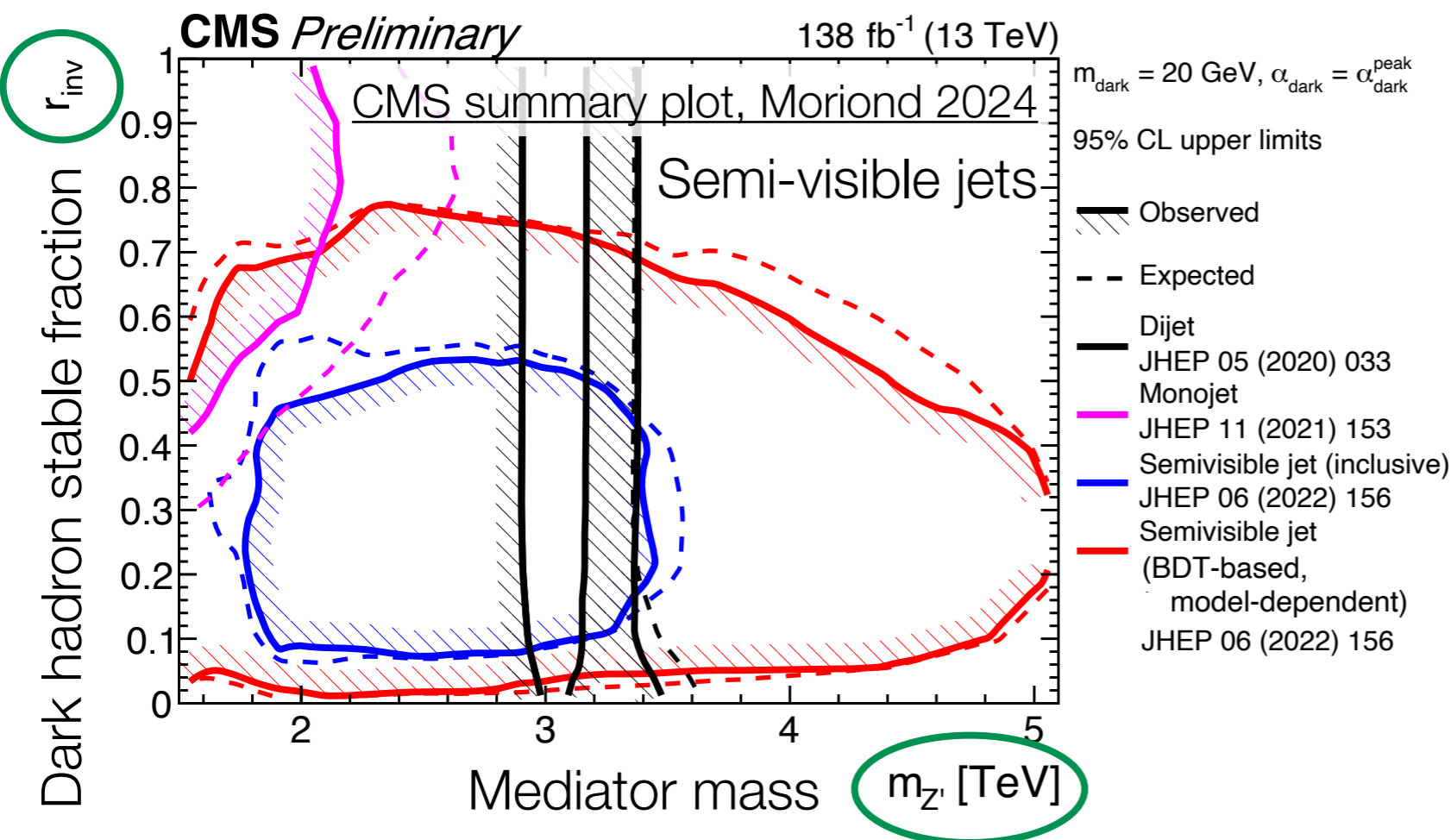
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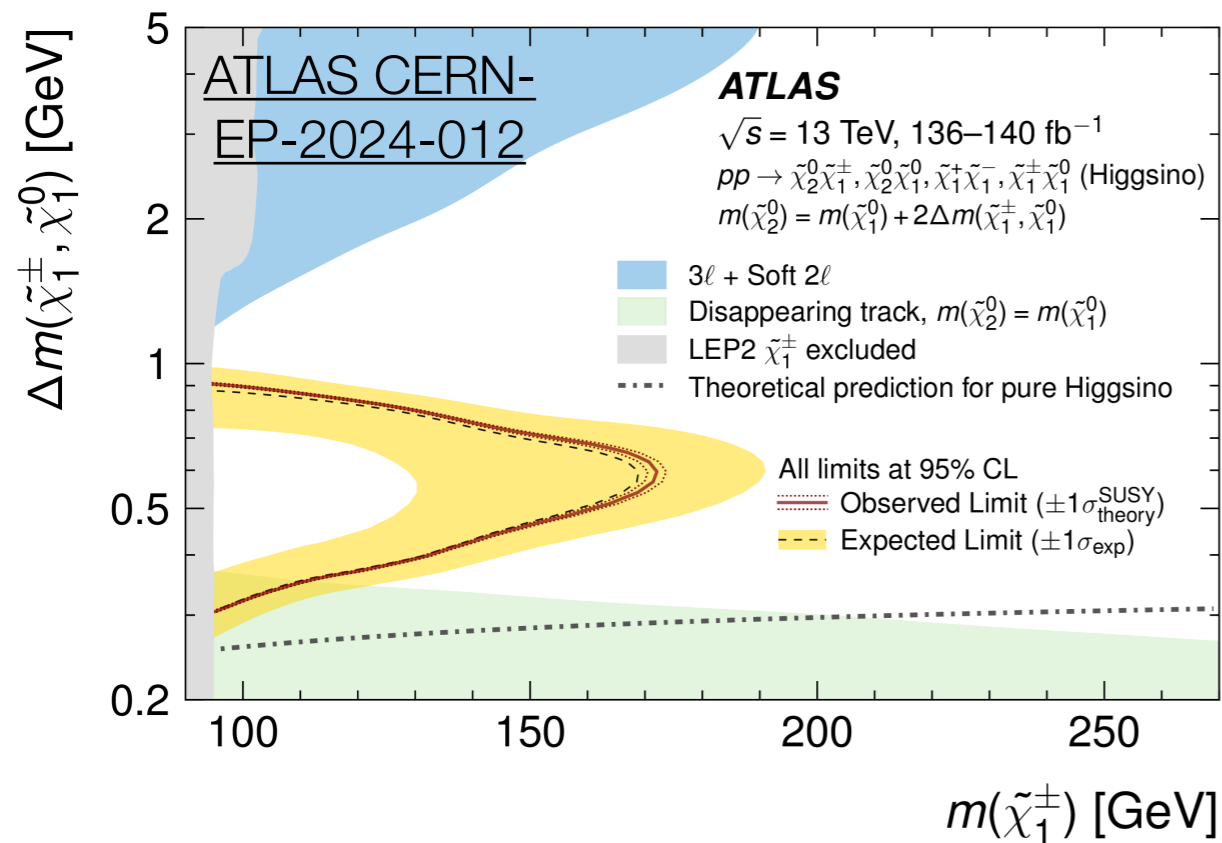
Long-lived particle searches

Saw one case already: displaced decays in dark photons with small ε . Other important examples:

Long-lived particle searches

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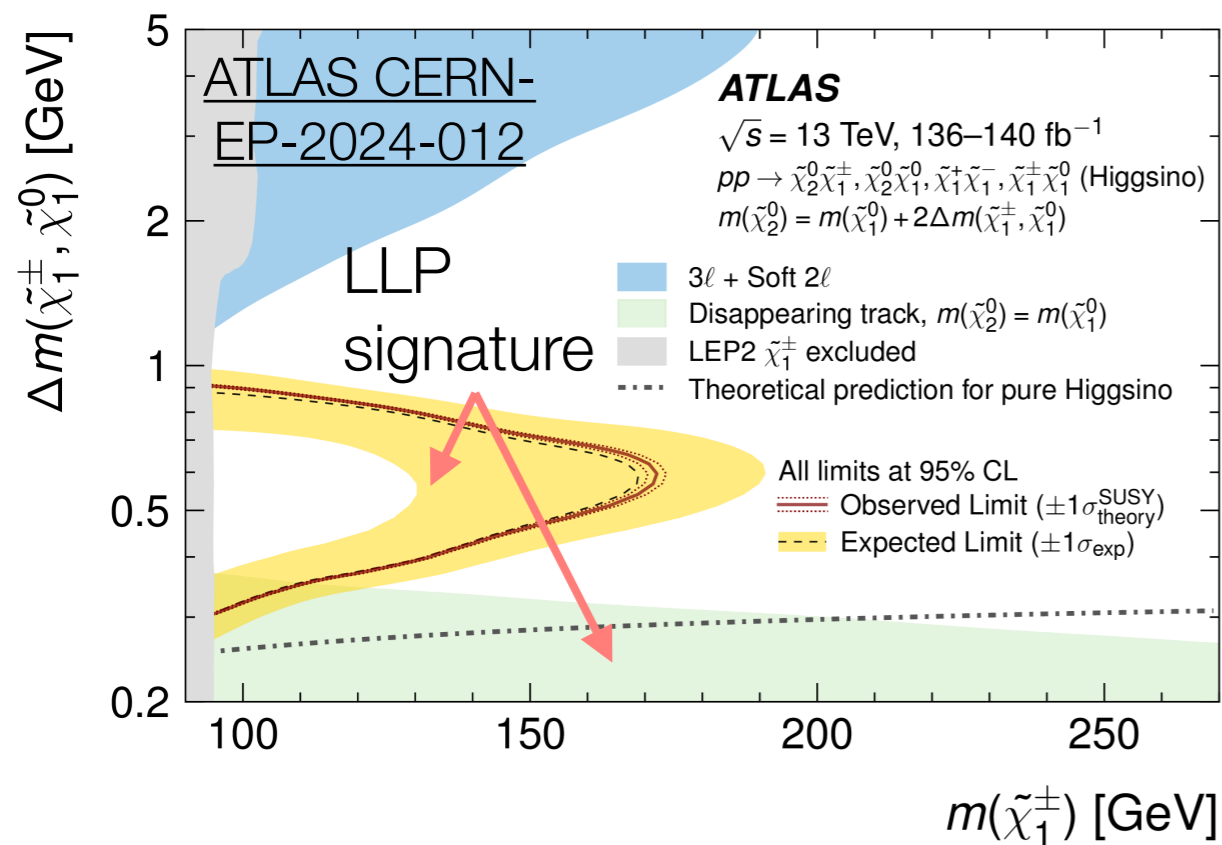
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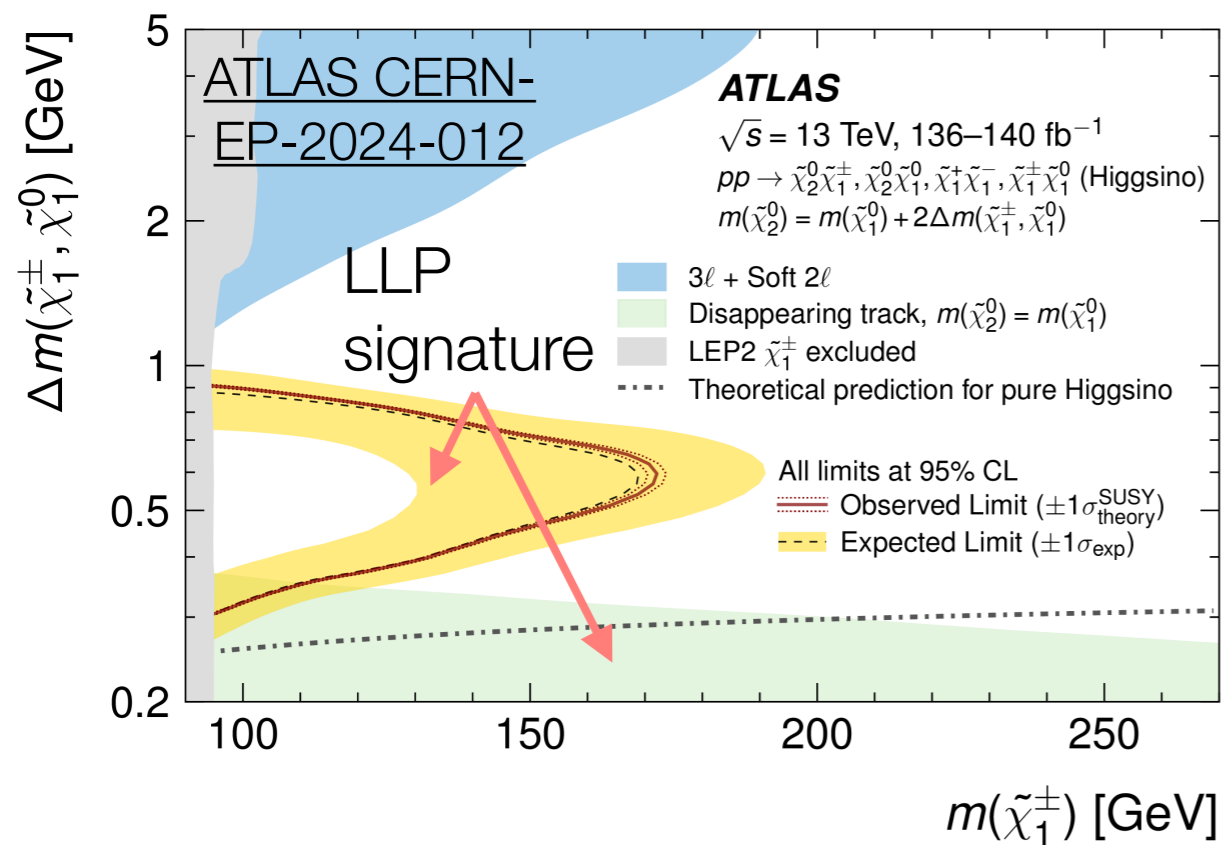
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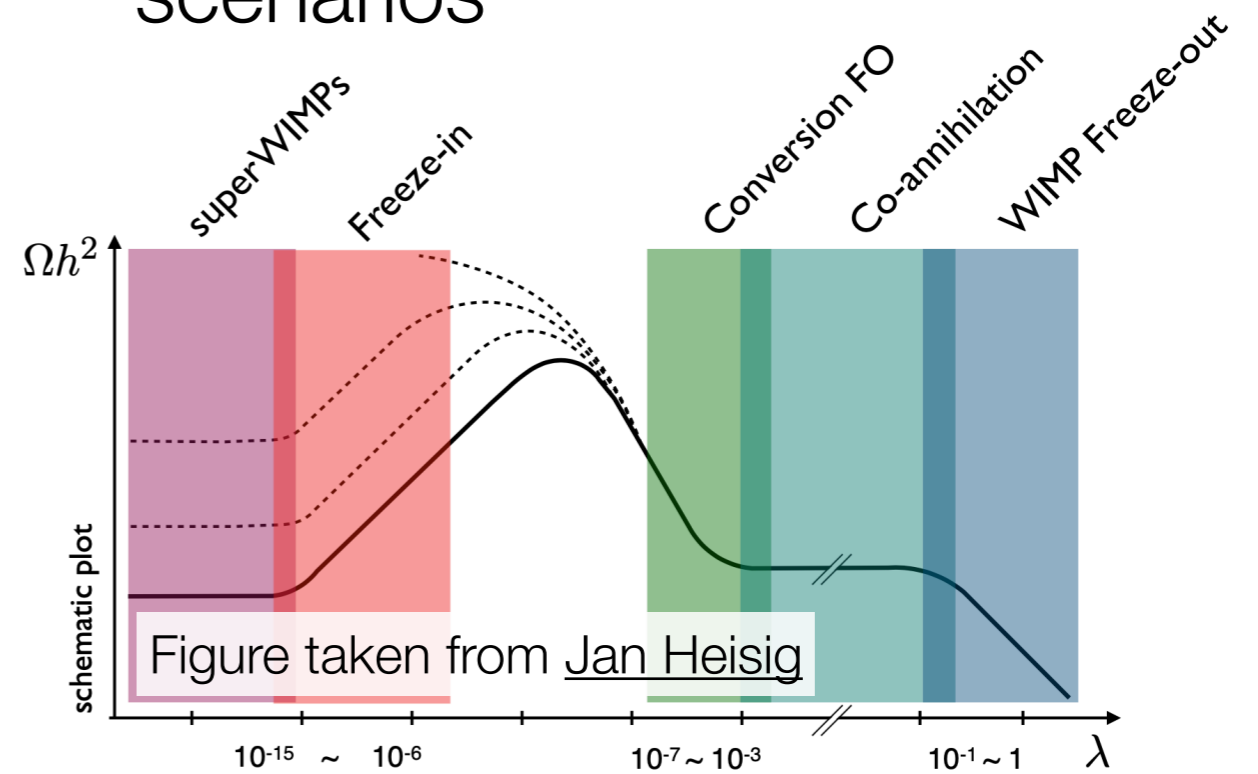
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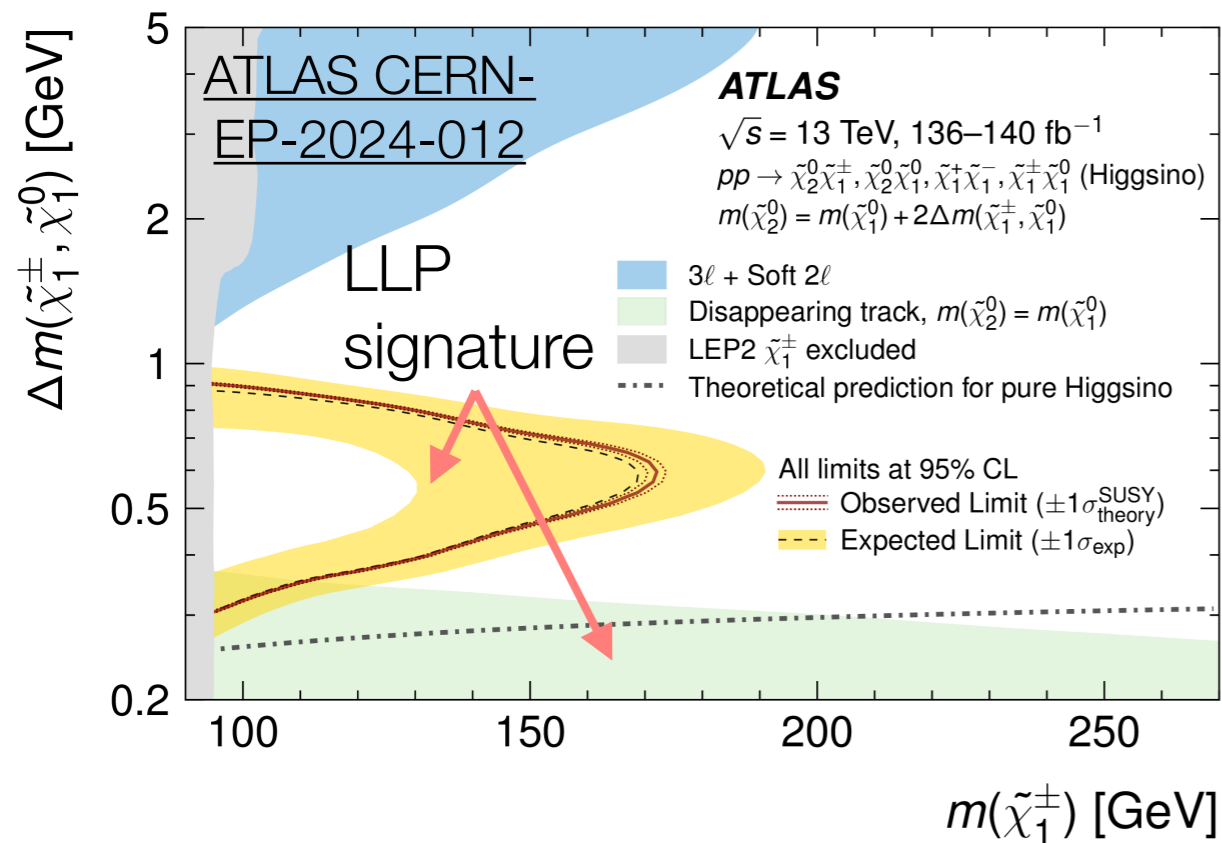
Freeze-in dark matter scenarios



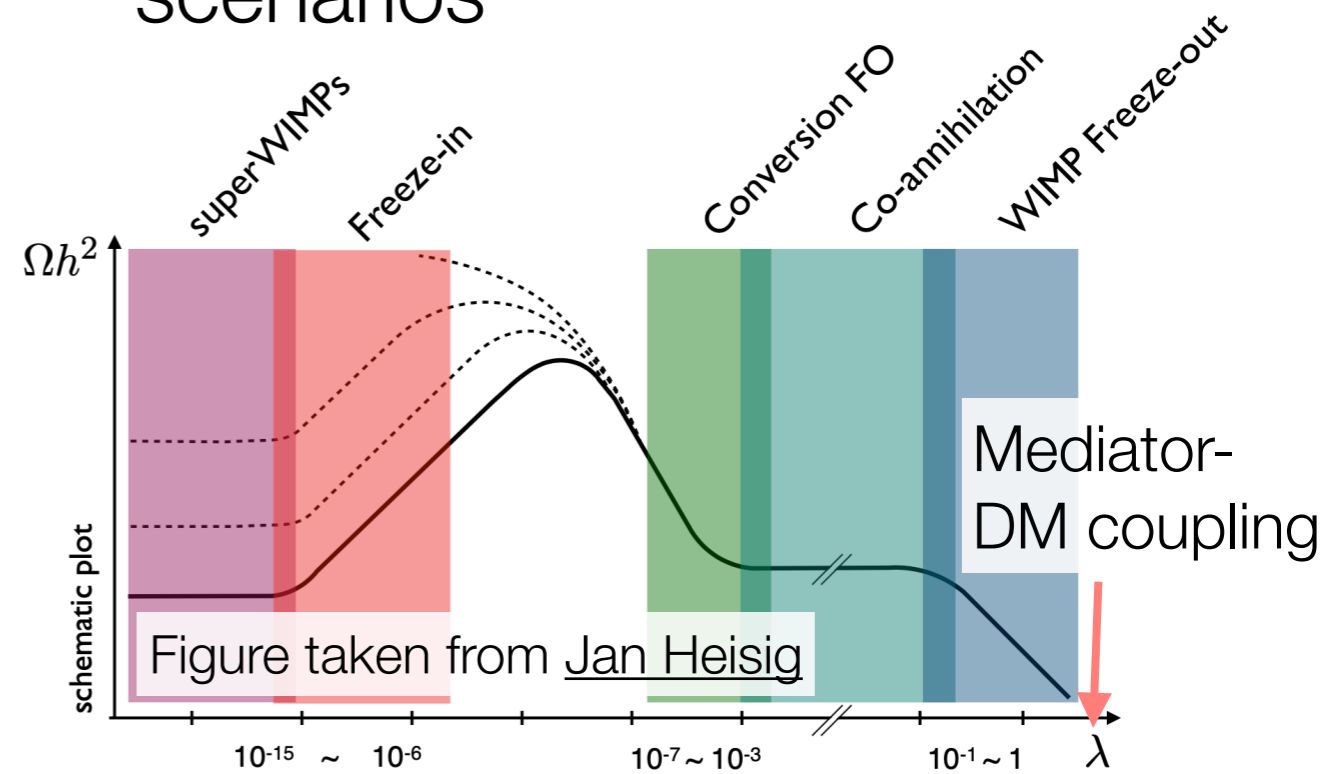
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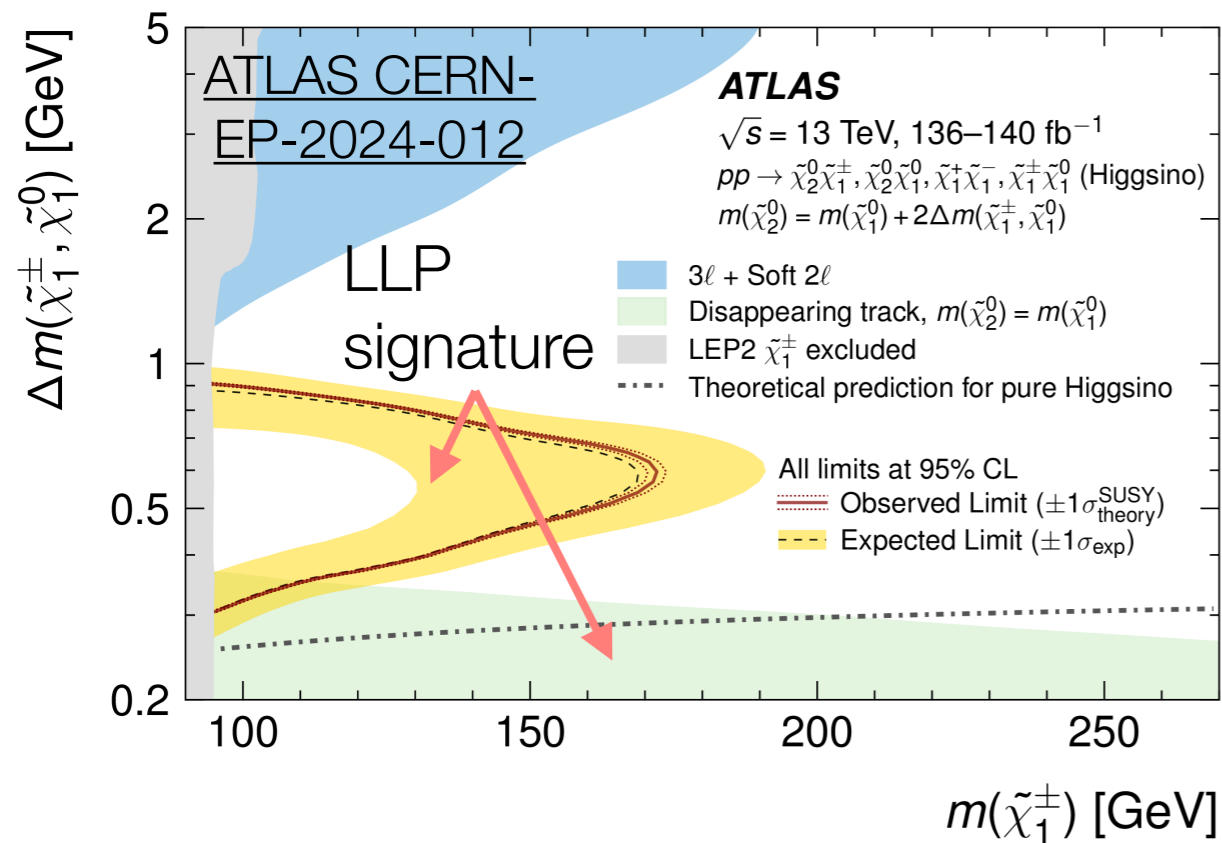
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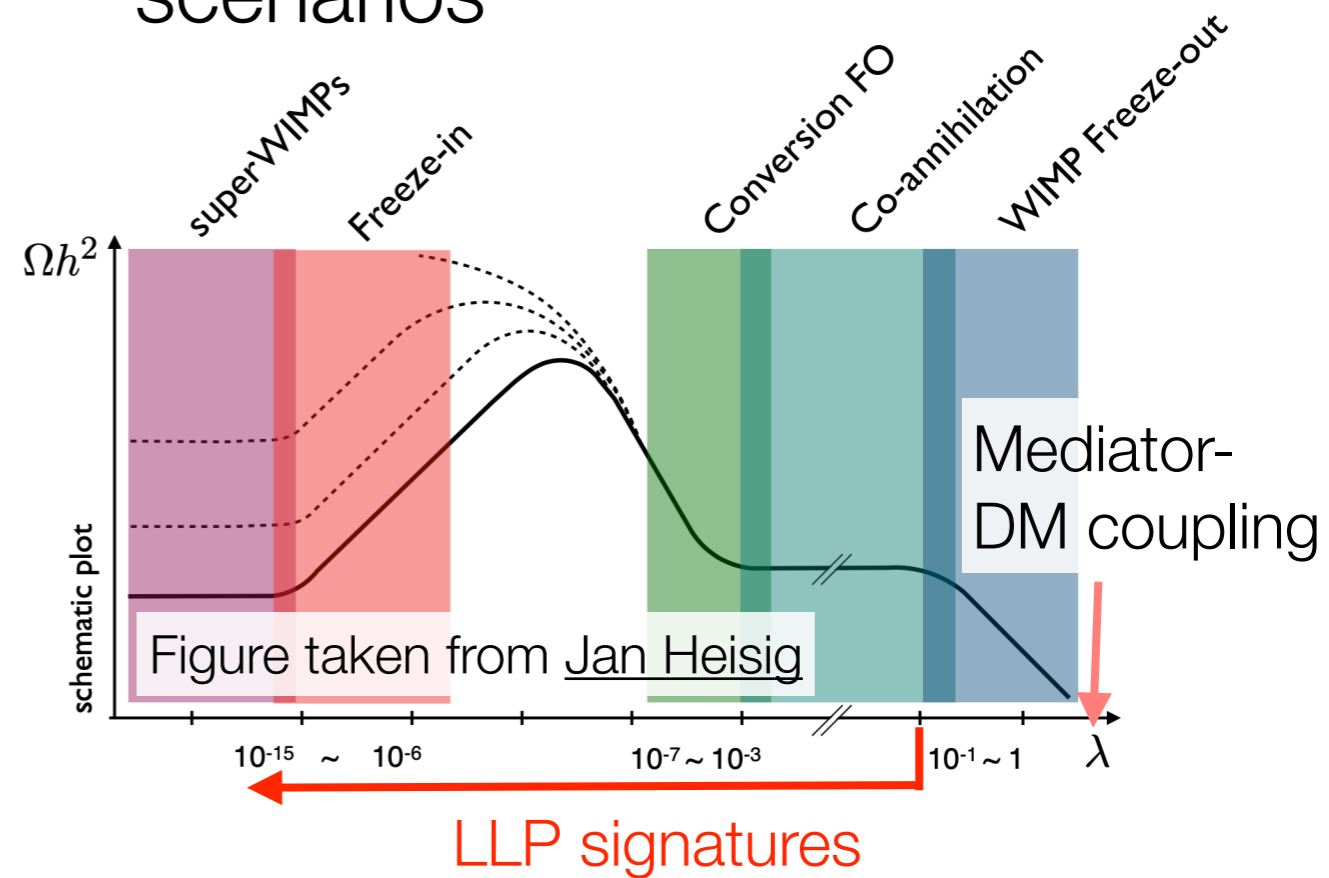
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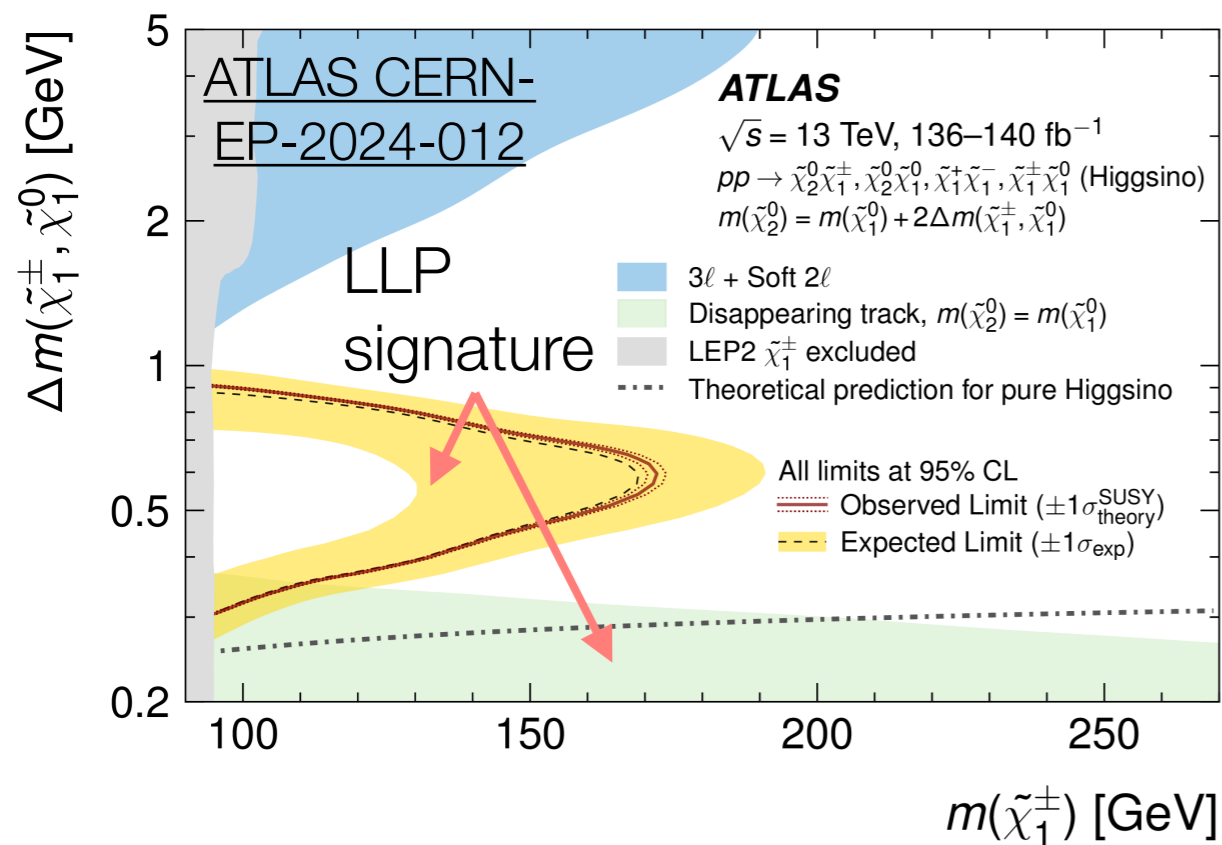
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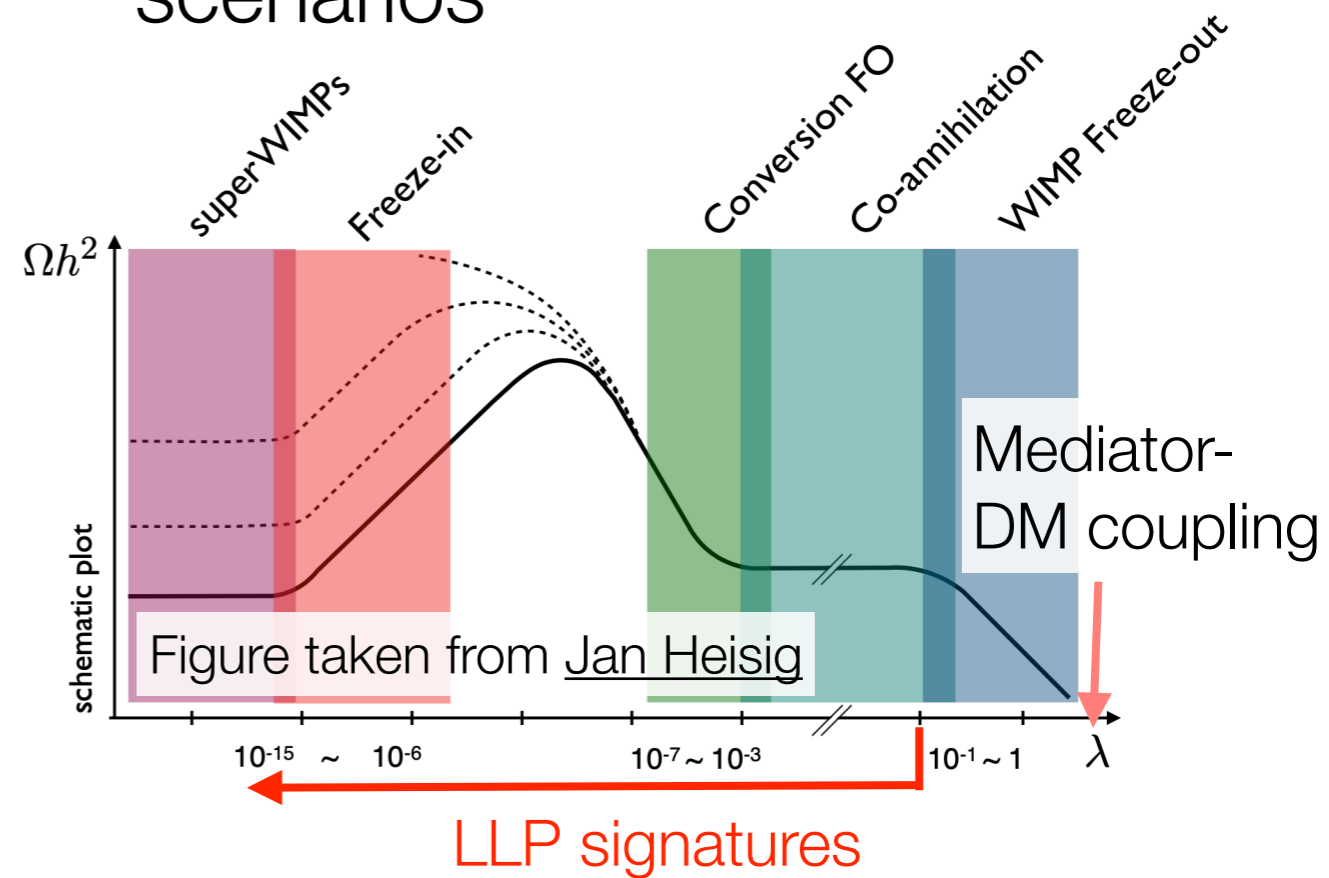
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Freeze-in dark matter scenarios



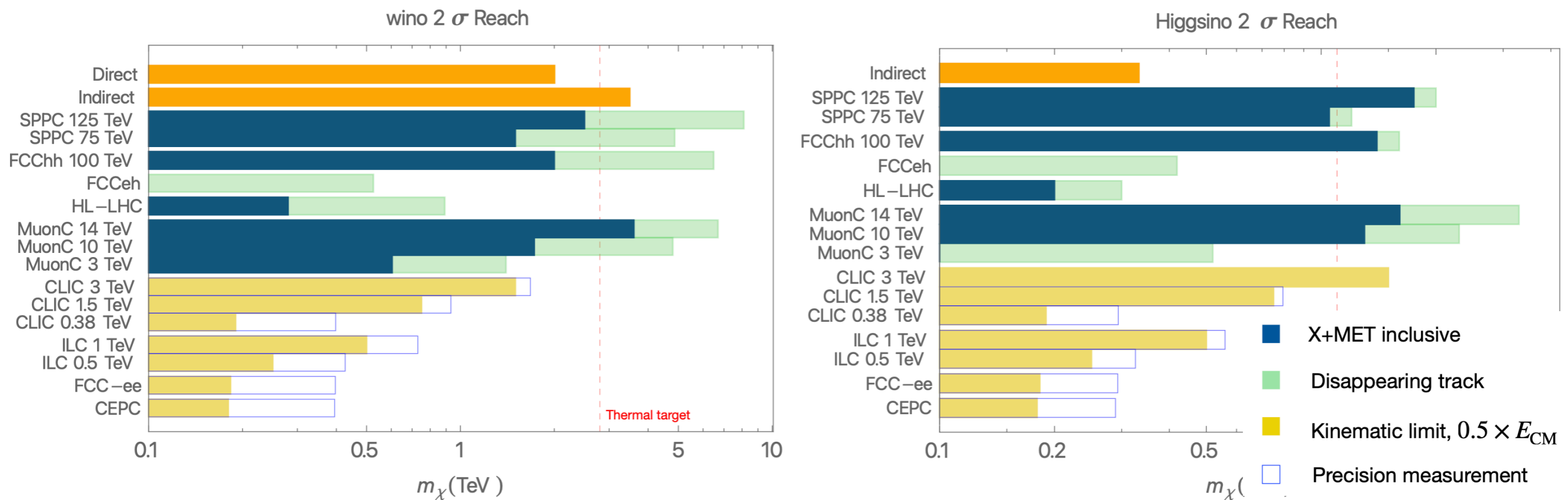
Can get LLPs from small mass splittings or small couplings, and turn up frequently in asymmetric, freeze-in, & SUSY DM



DM at HL-LHC and future colliders

Opportunities at future colliders: SUSY DM

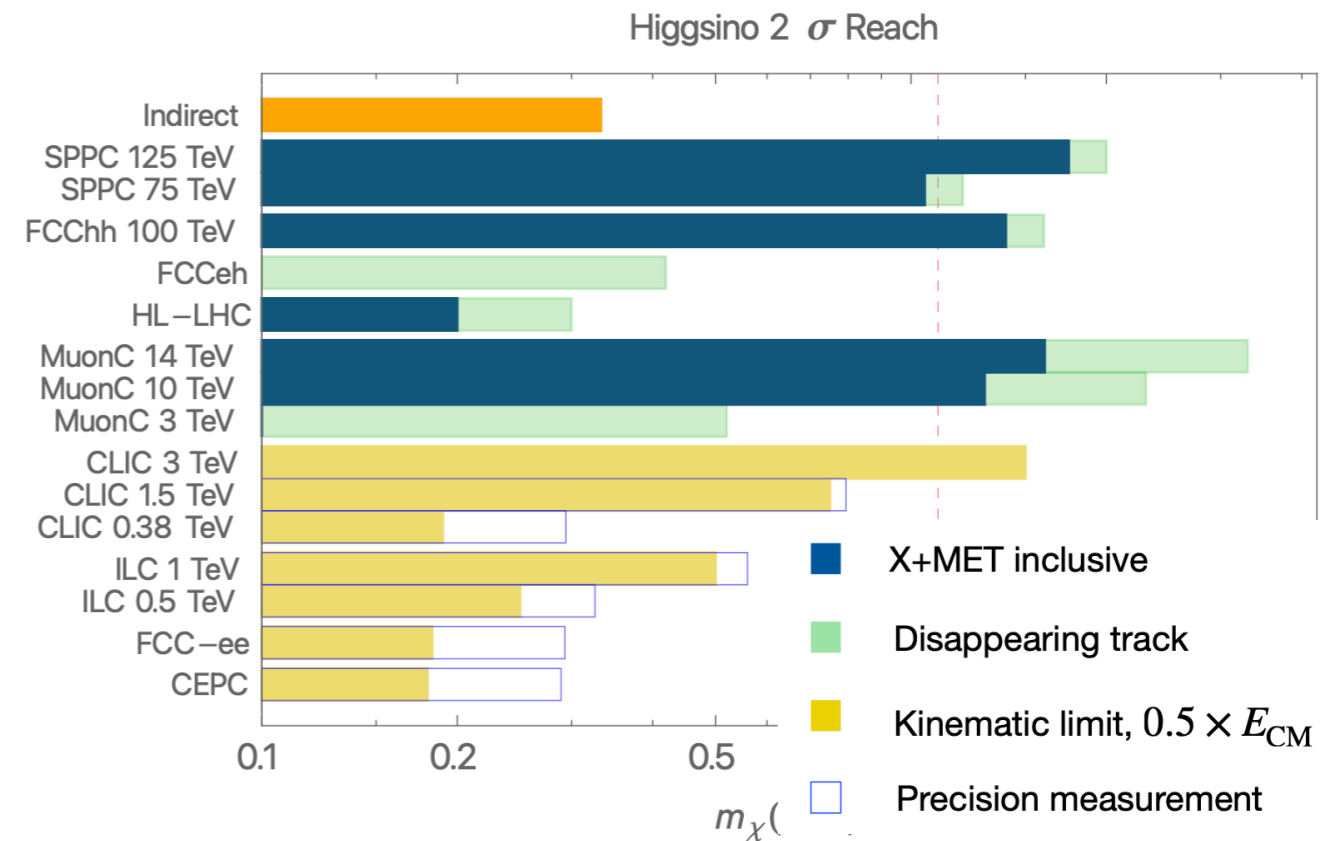
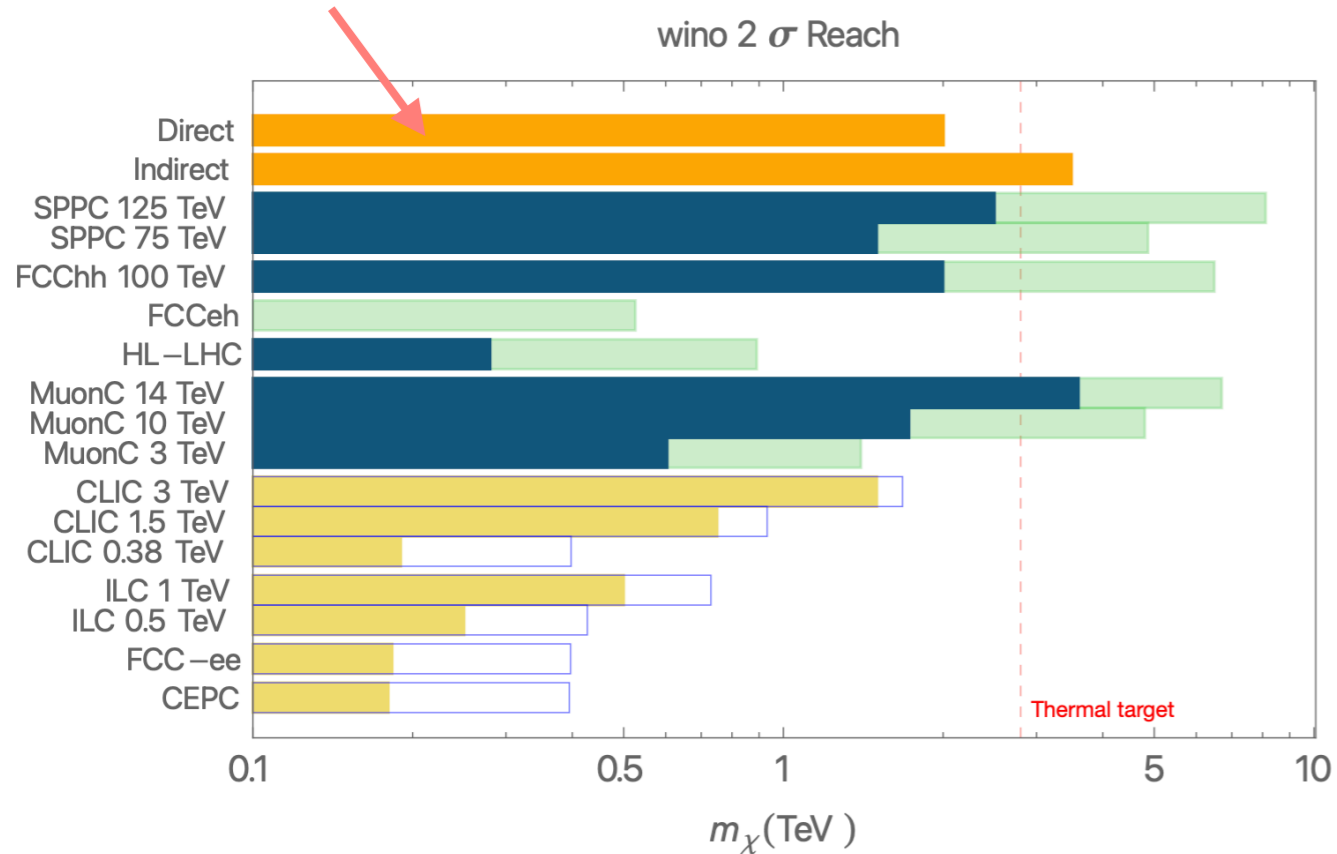
Minimal EW multiplet scenario: SM gauge couplings fix interactions so mass is only free parameter and thermal DM predictions simple.



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DARWIN (50T) projection

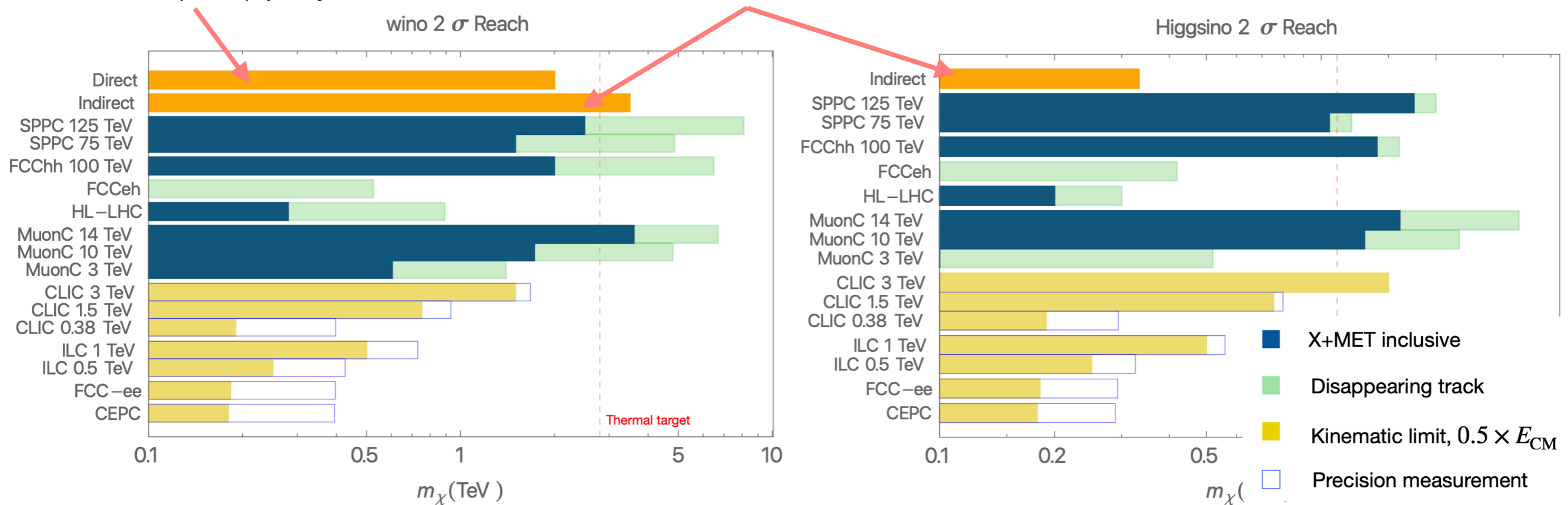


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FERMI & H.E.S.S.



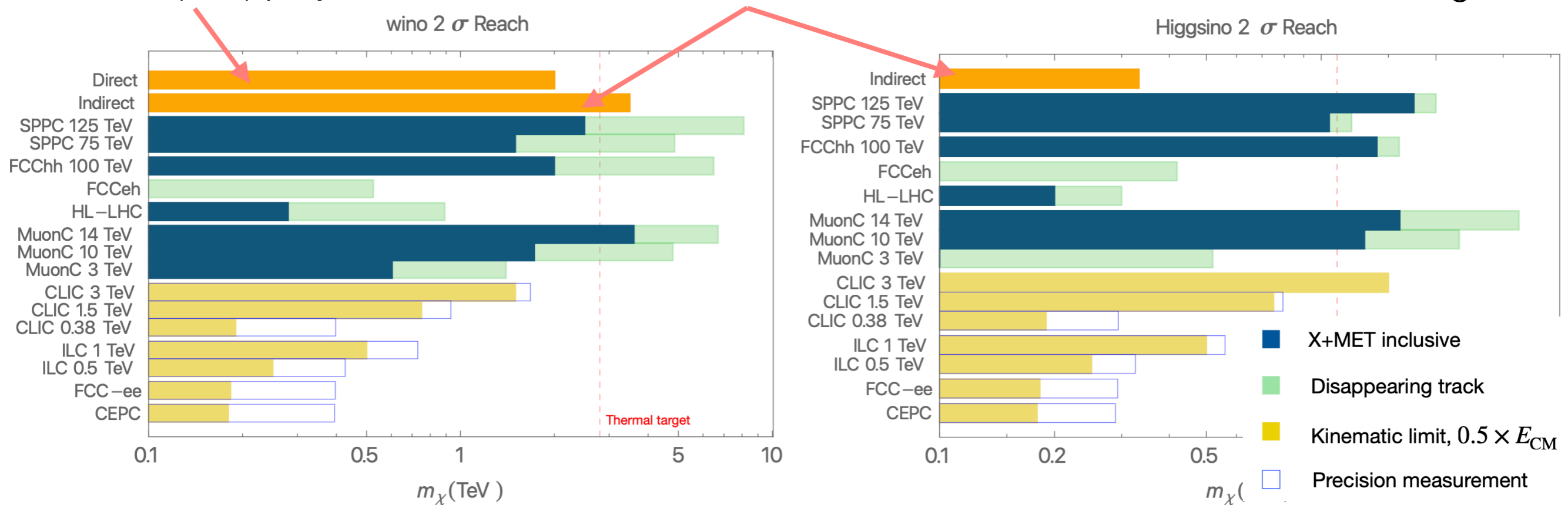
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DD below neutrino fog



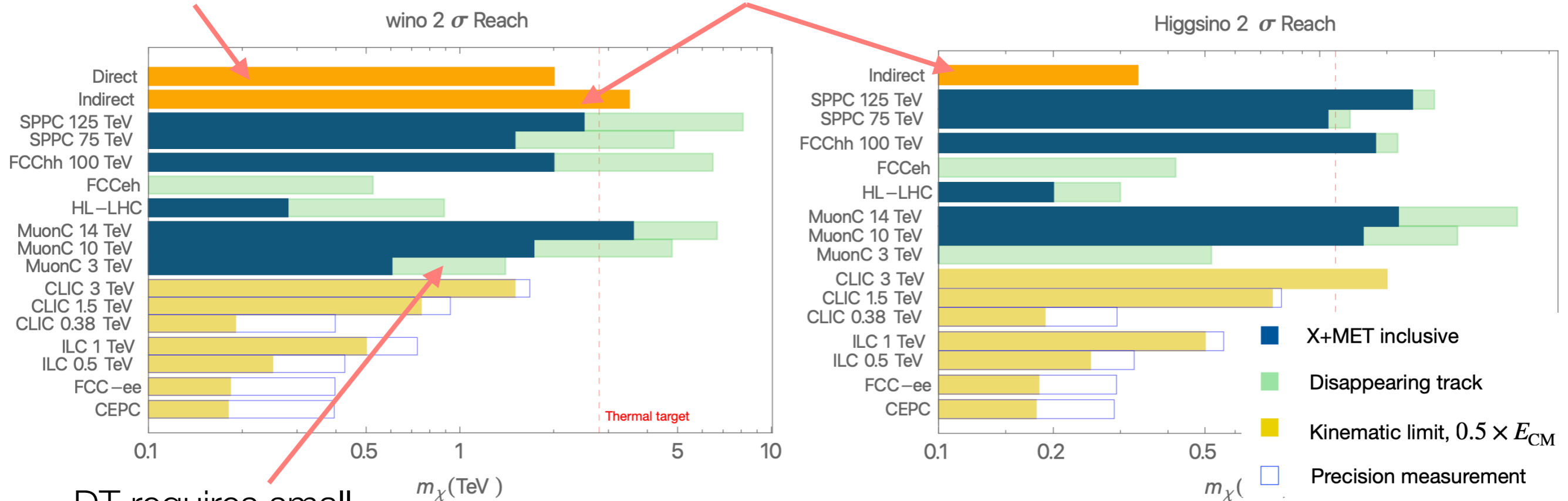
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DT requires small mass splittings

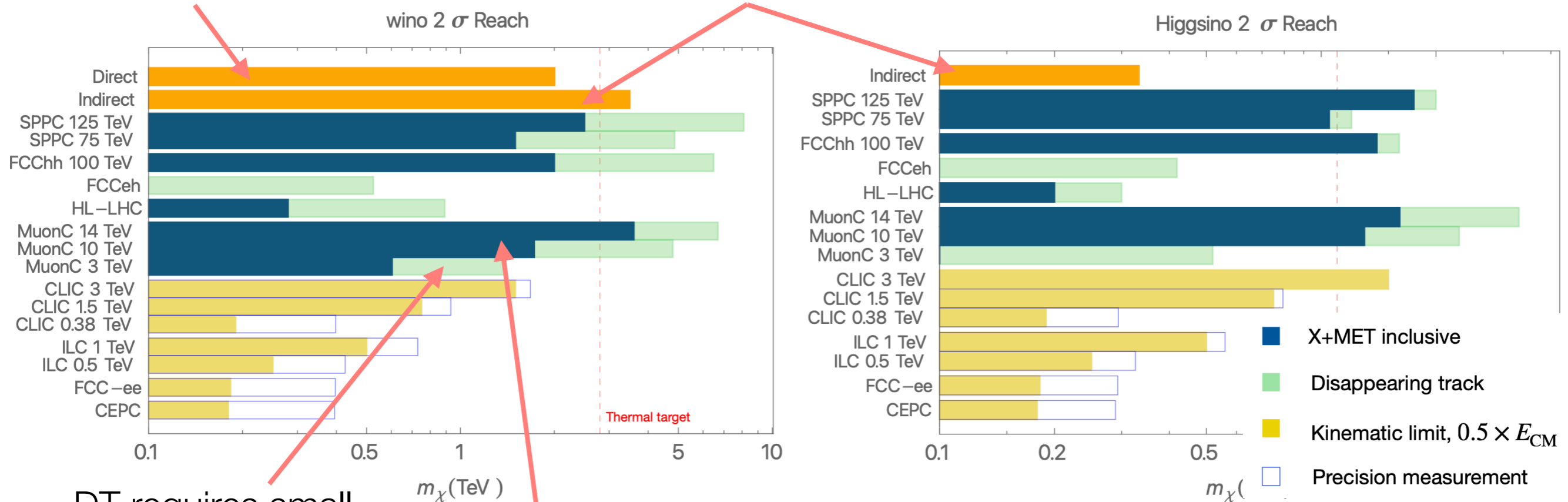
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DD below neutrino fog



DT requires small mass splittings

More general for other scenarios

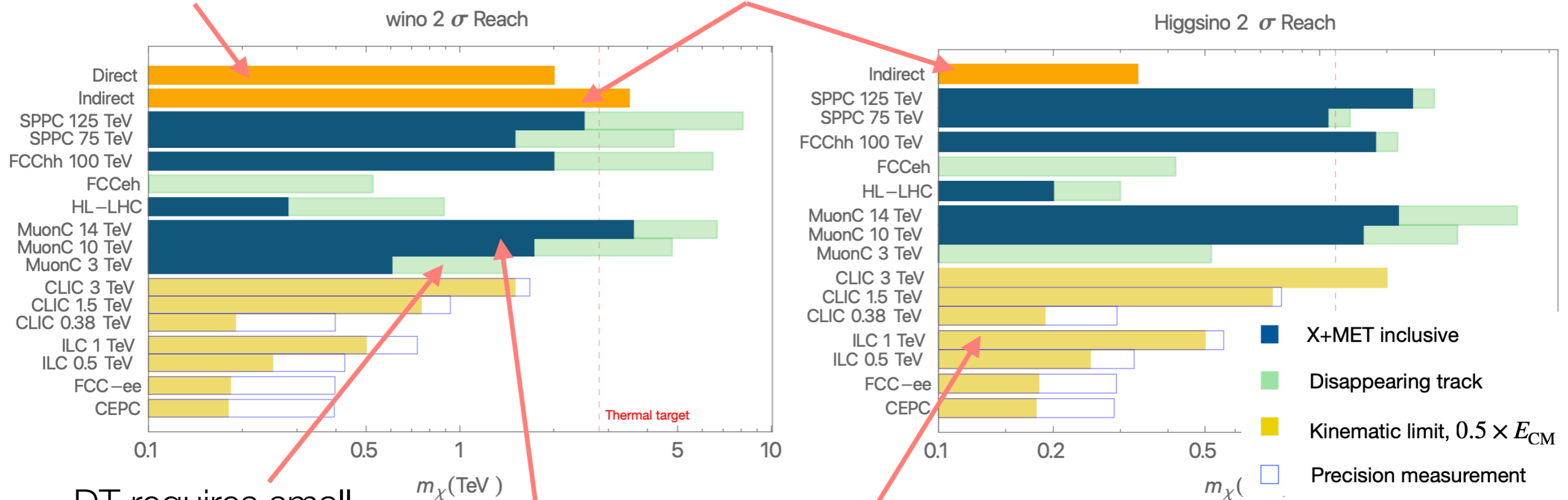
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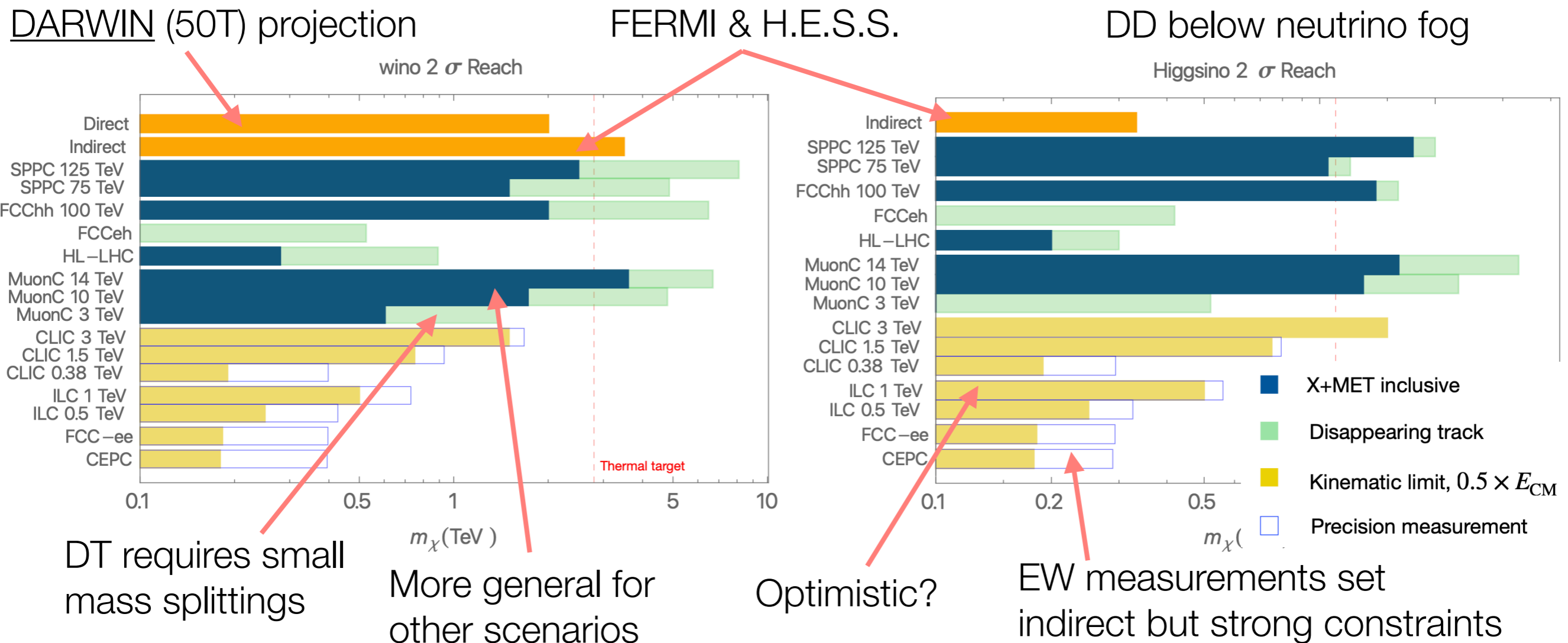
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More general for other scenarios

Optimistic?

Opportunities at future colliders: SUSY DM

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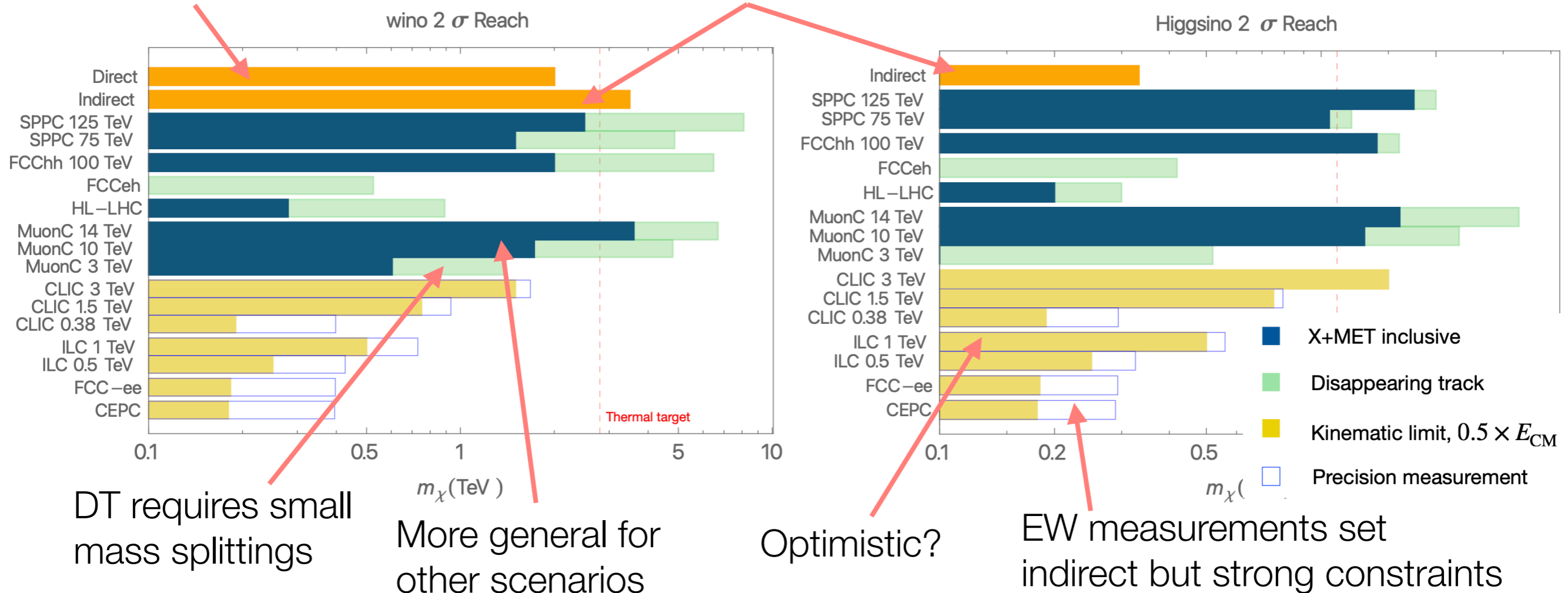
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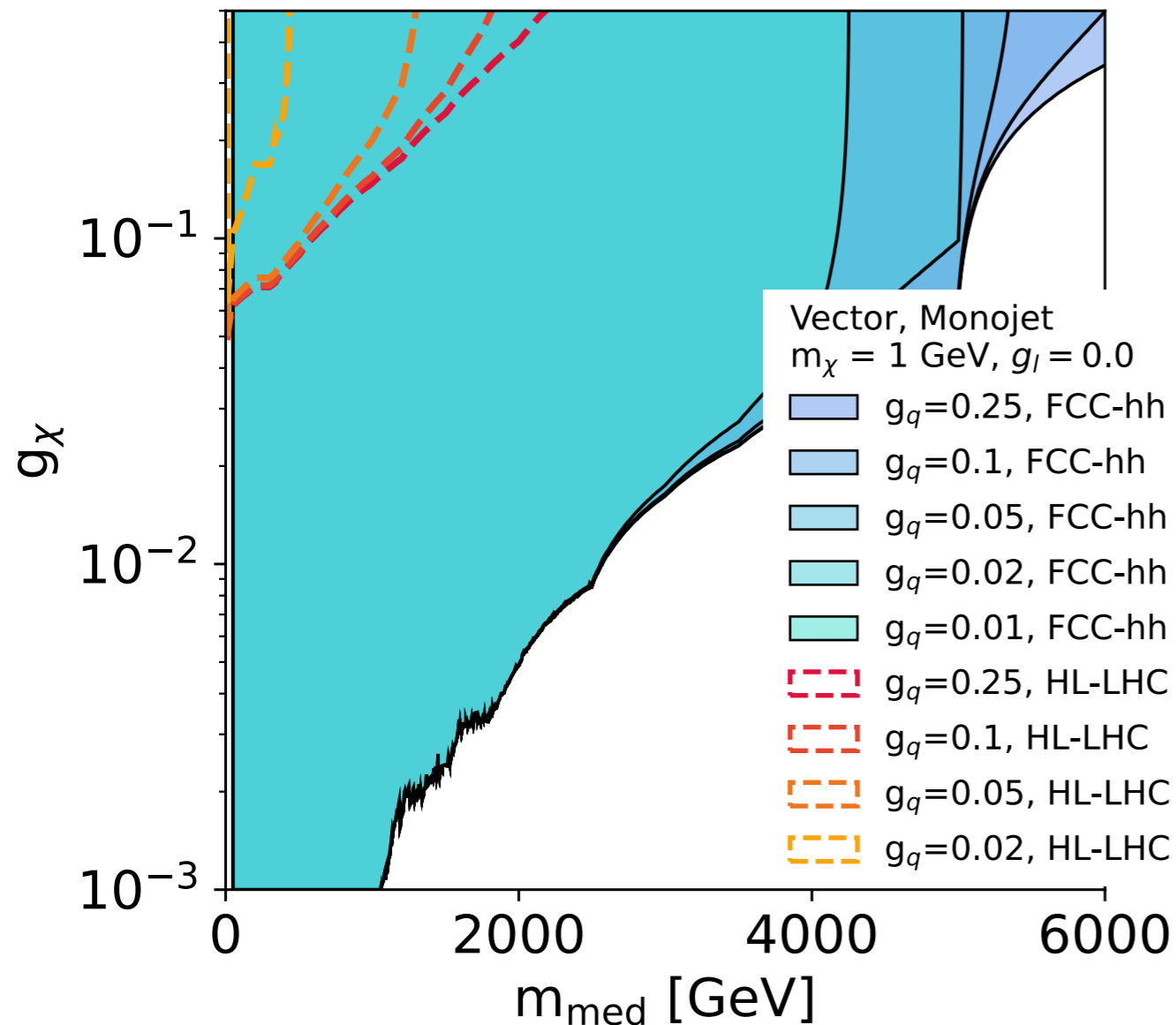
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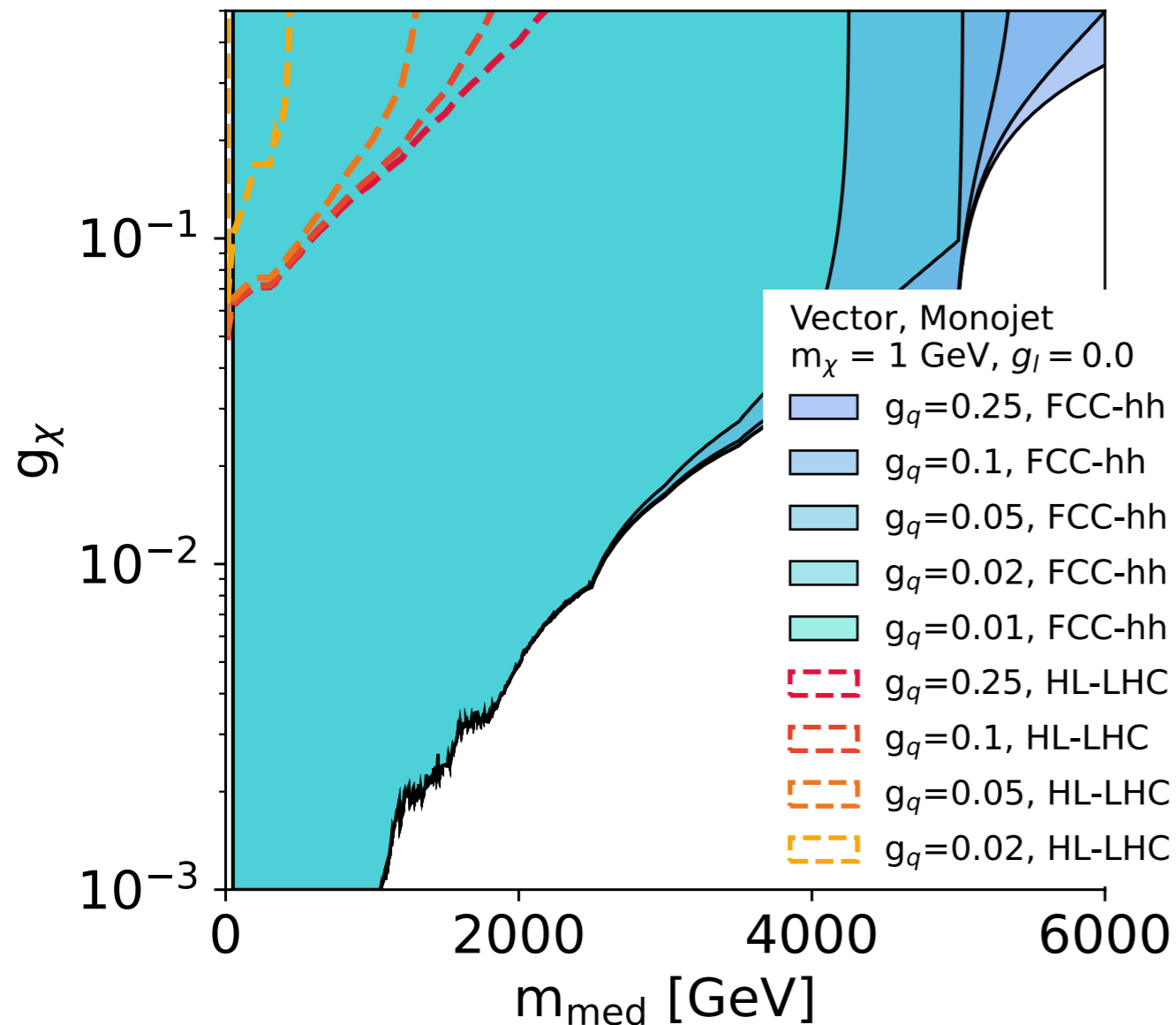
Reaching thermal target is not easy, but possible at some colliders

Opportunities at future colliders: non-SUSY DM

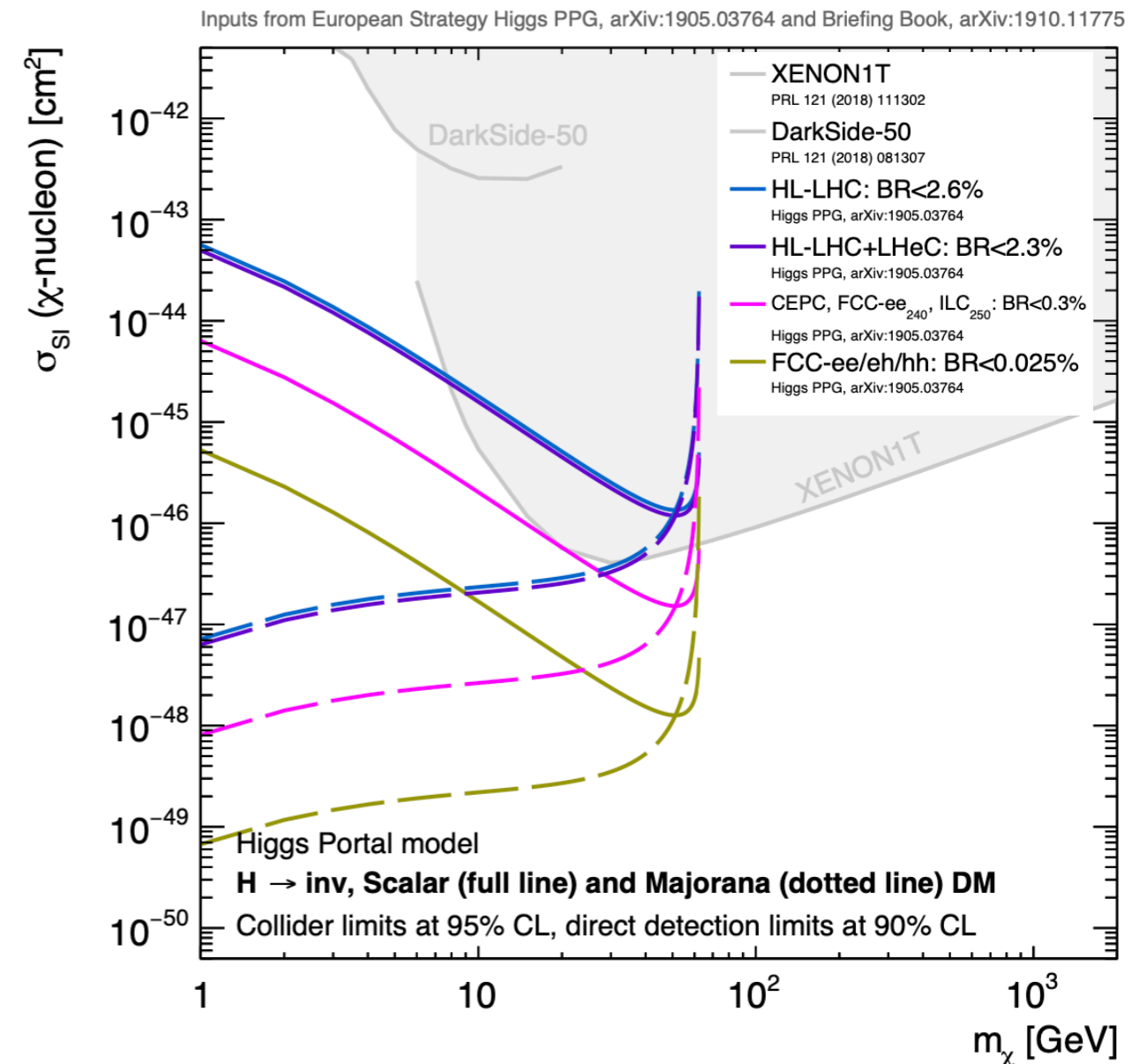


Spin-1 vector mediator: monojet sensitivity to DM coupling

Opportunities at future colliders: non-SUSY DM



Spin-1 vector mediator: monojet sensitivity to DM coupling



Higgs portal: $H \rightarrow \text{inv}$ sensitivity compared to current DD

Parasitic experiments at future colliders

Future colliders have possibilities beyond collision point detectors

Parasitic experiments at future colliders

Future colliders have possibilities beyond collision point detectors

Dedicated LLP experiments

Valuable when LLP signature is trigger limited

Limited use at e^+e^- machines but useful at hadron & probably muon machines

Different signatures can favour forward (FASER-esque) vs off-axis far detectors

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Beam dump experiments

Missing energy/mass experiments not possible at EF machines

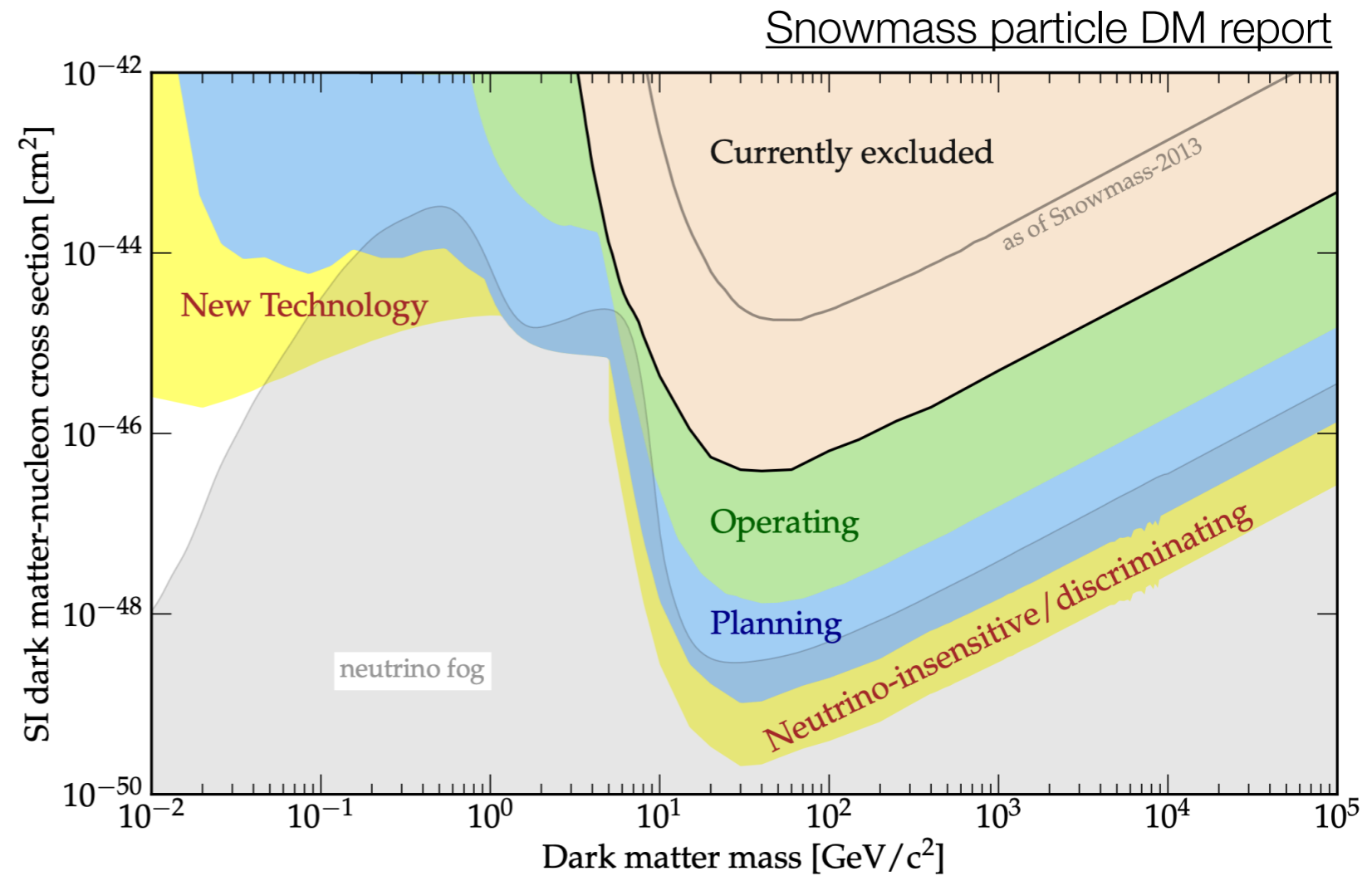
Could probably do a re-scattering experiment here but I've not seen it talked about

Visible decay searches are well suited and could be added to future colliders (examples 1, 2)

A hand-drawn diagram on a dark background. It features several intersecting lines, some solid and some dashed. There are several small circles, some of which are spirals. There are also several small crosses and L-shaped symbols scattered throughout the diagram. The overall appearance is that of a technical or scientific sketch.

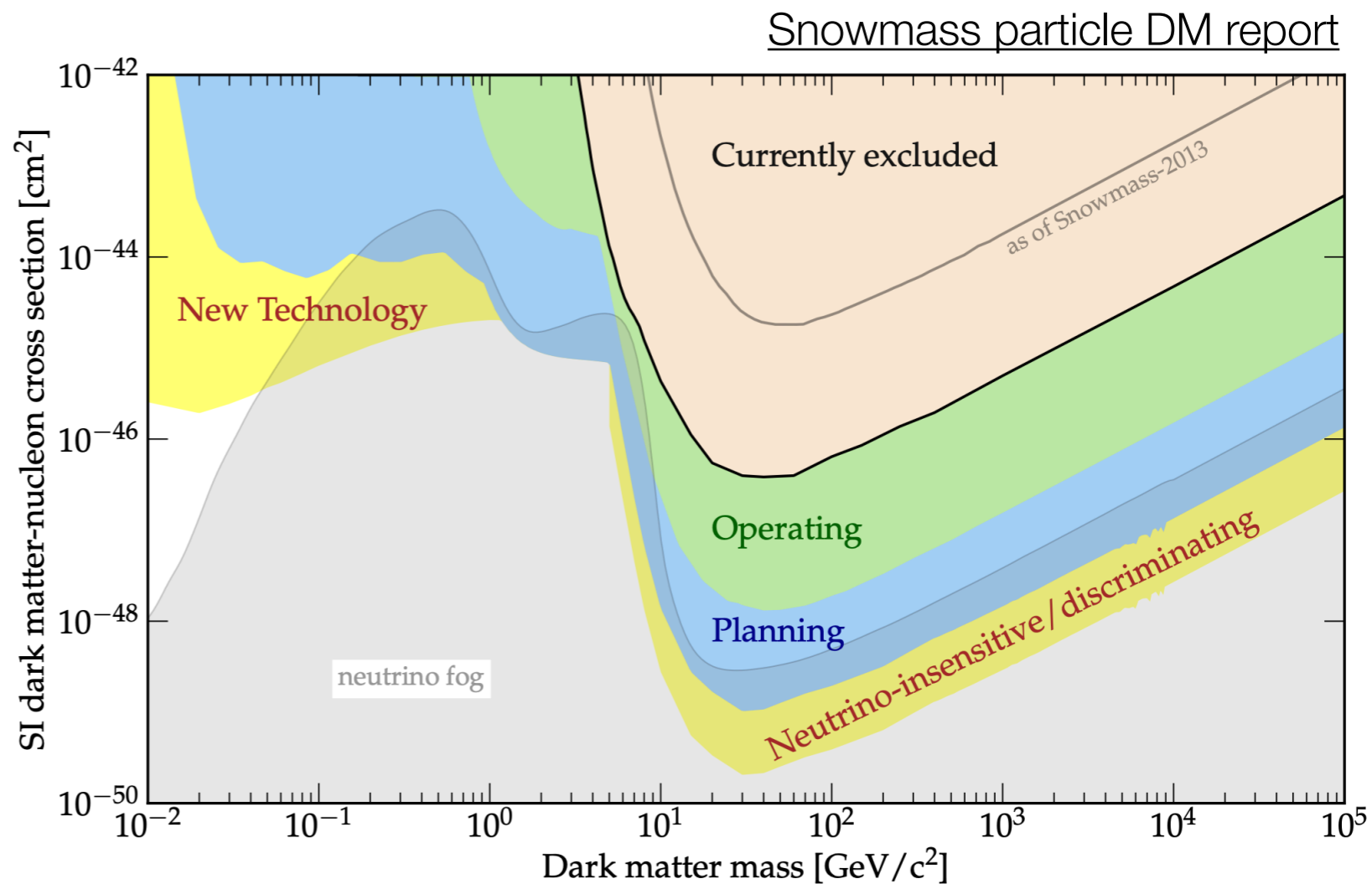
Discussing complementarity

Mentioned earlier that we need to highlight complementary areas of strength between DD, ID, and future colliders



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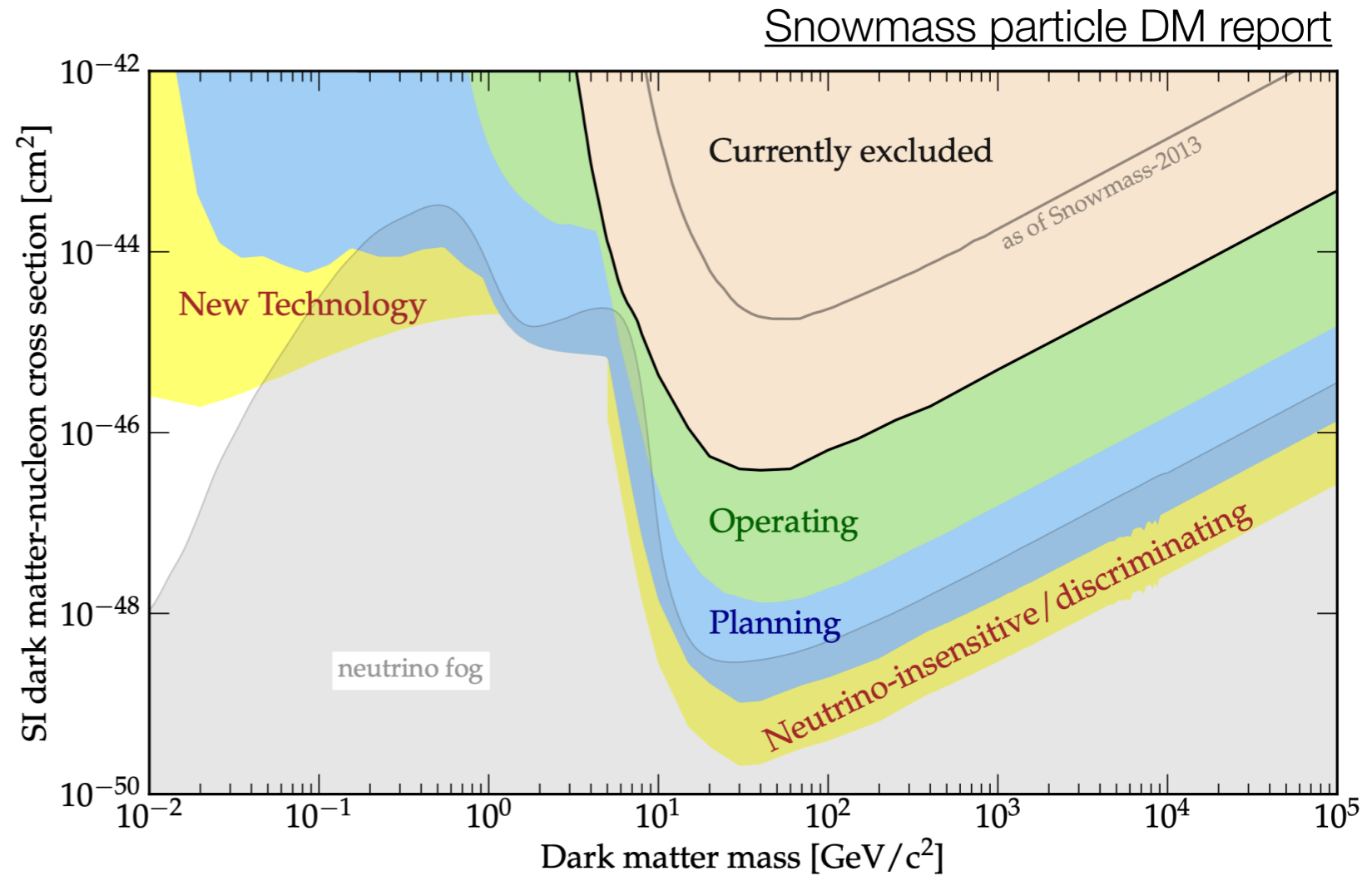
This will be key to building the field we want to see



Mentioned earlier that we need to highlight complementary areas of strength between DD, ID, and future colliders

This will be key to building the field we want to see

Often easier said than done.

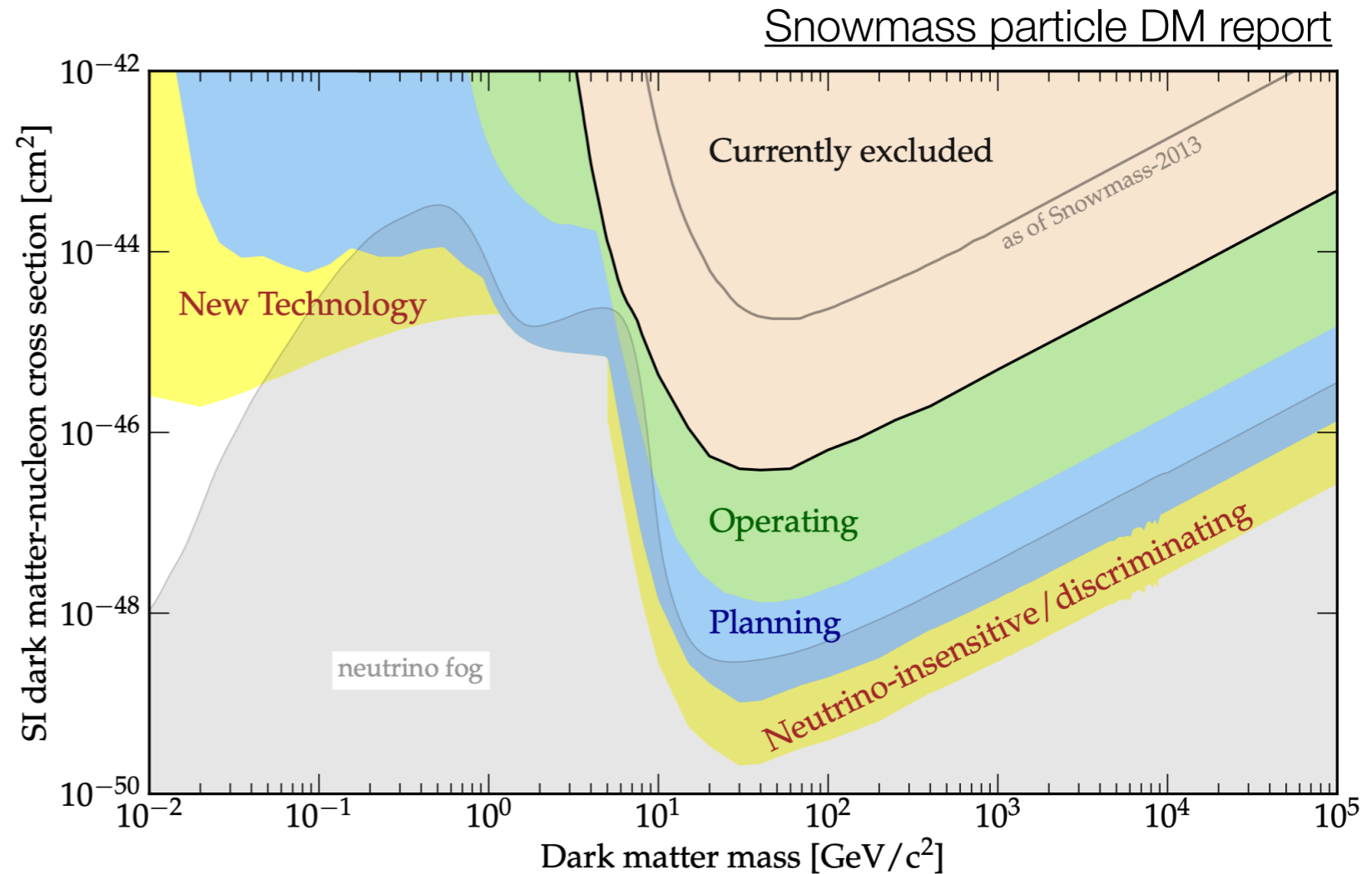


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DD limits can use EFT;
EF searches require model assumptions.
Reducing problem dimensions to 2D plane usually needs extra assumptions

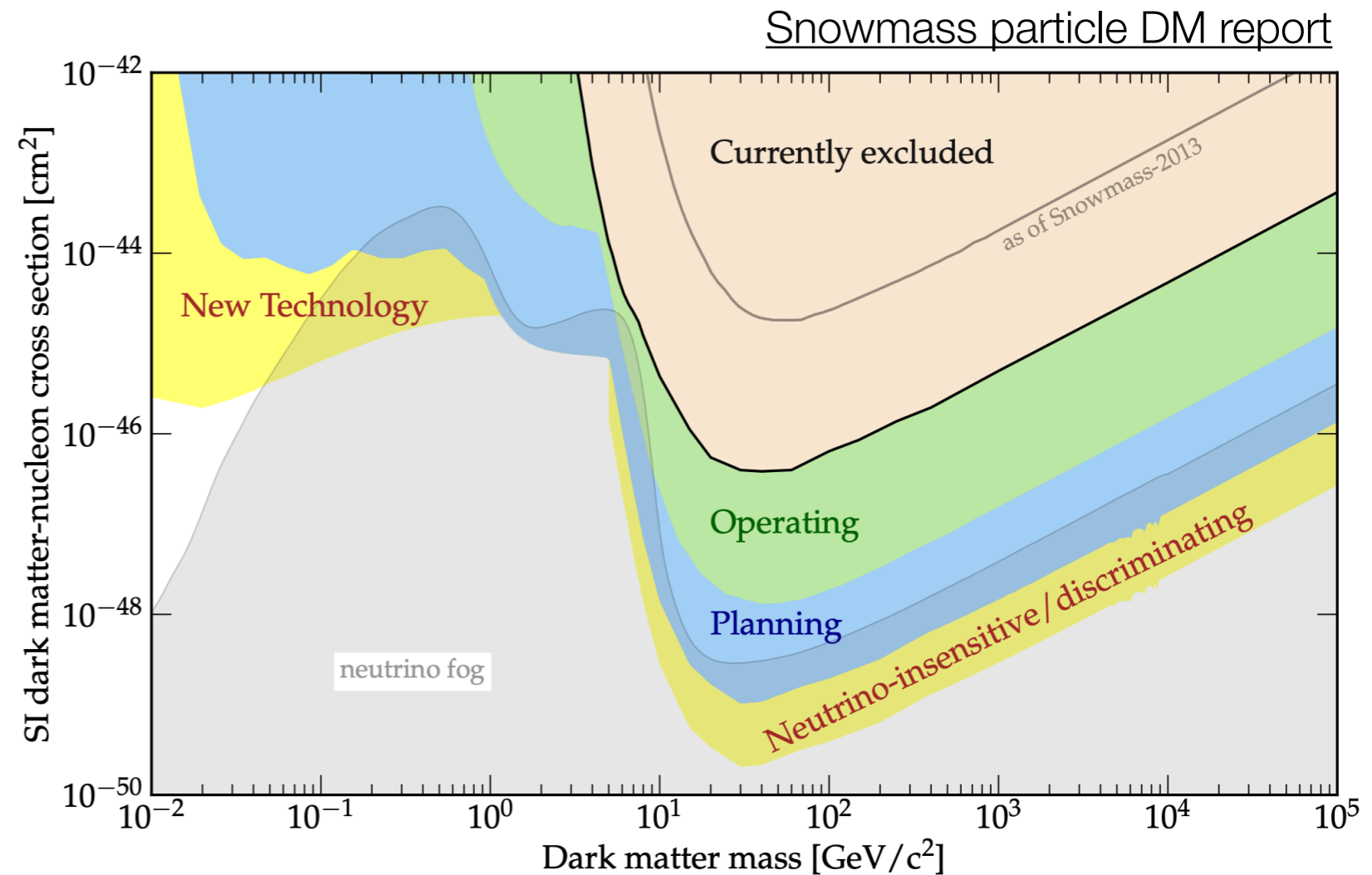


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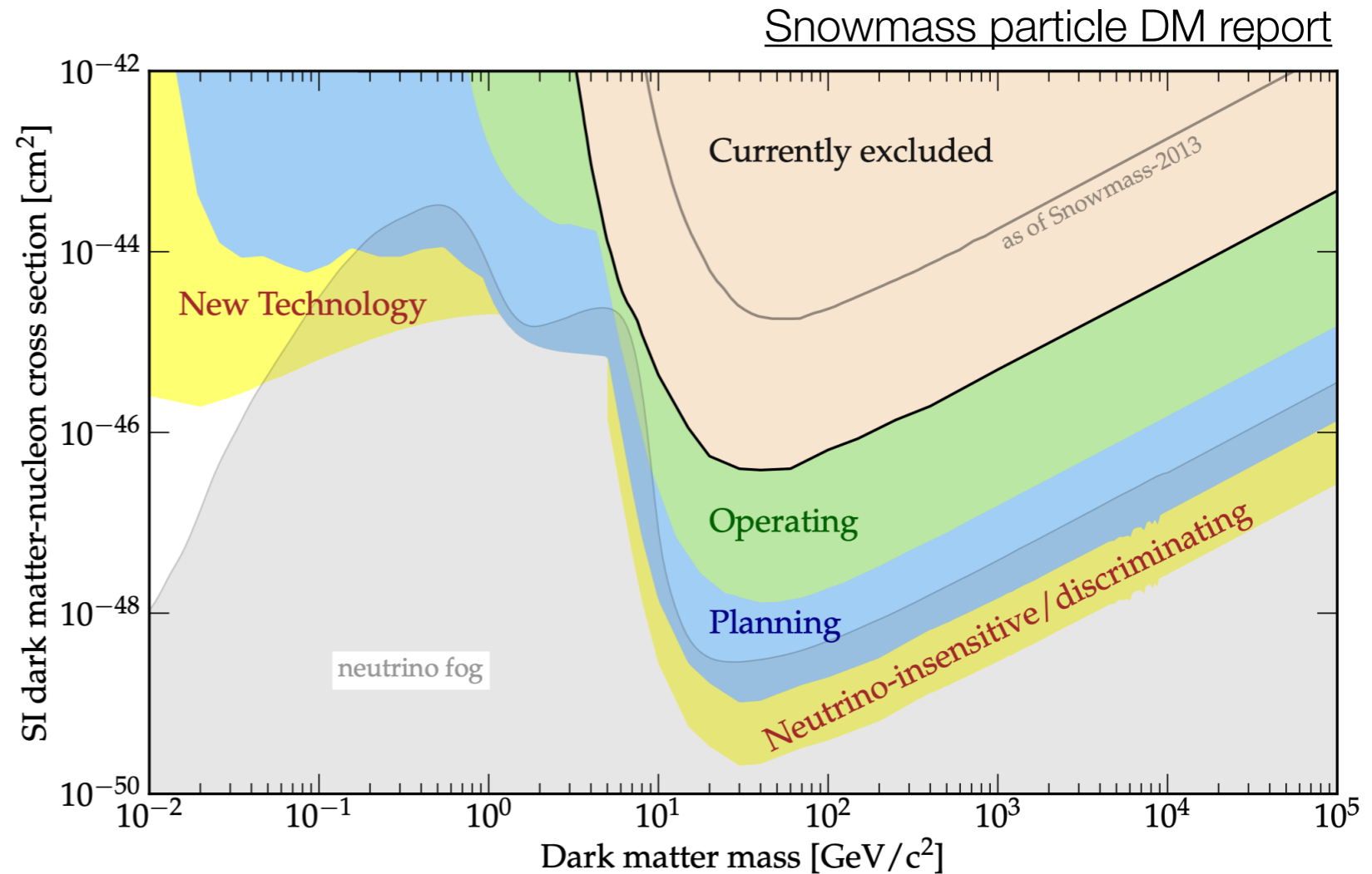
Show example I know best: LHC DMWG spin-1 simplified model

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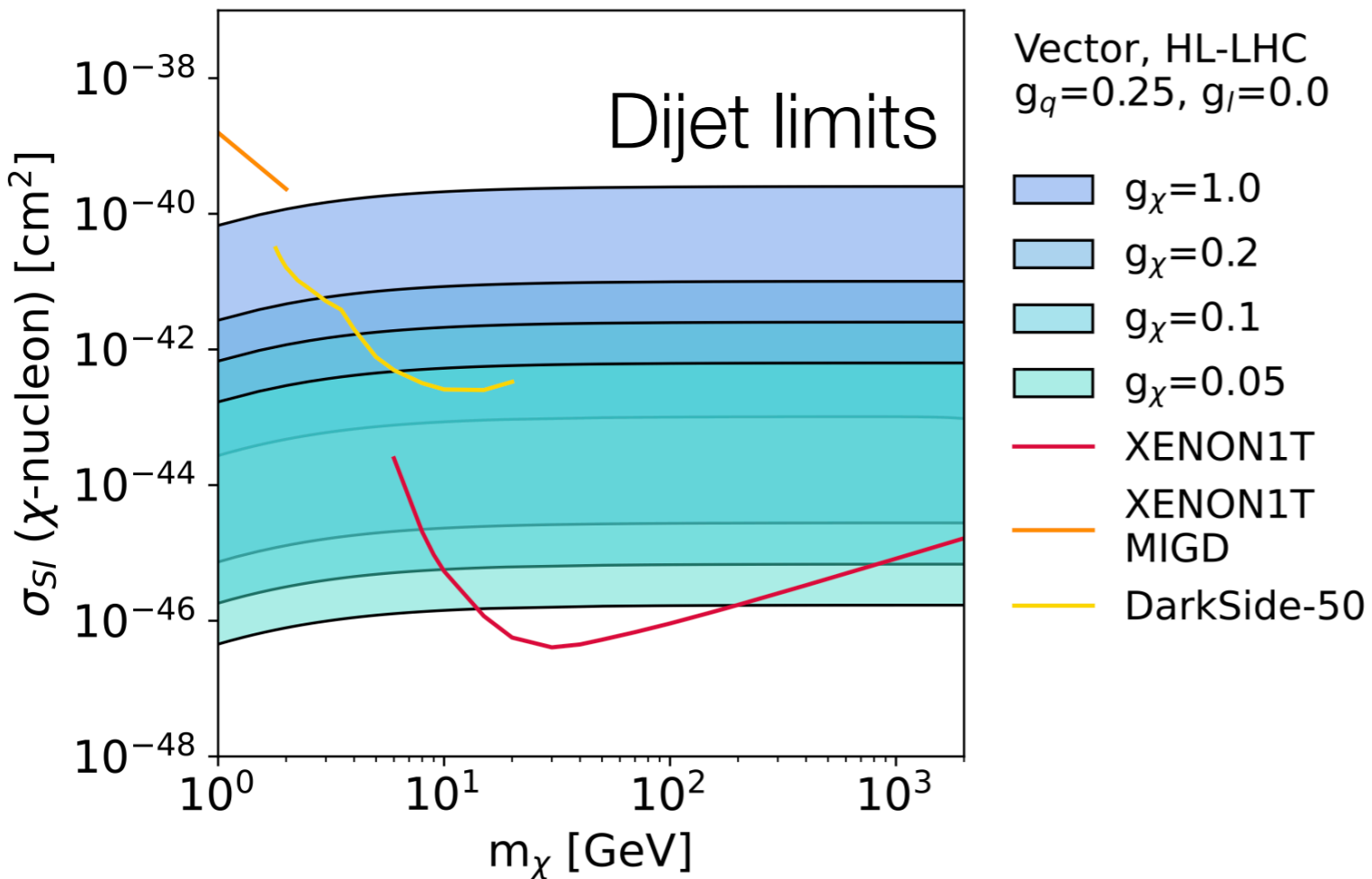
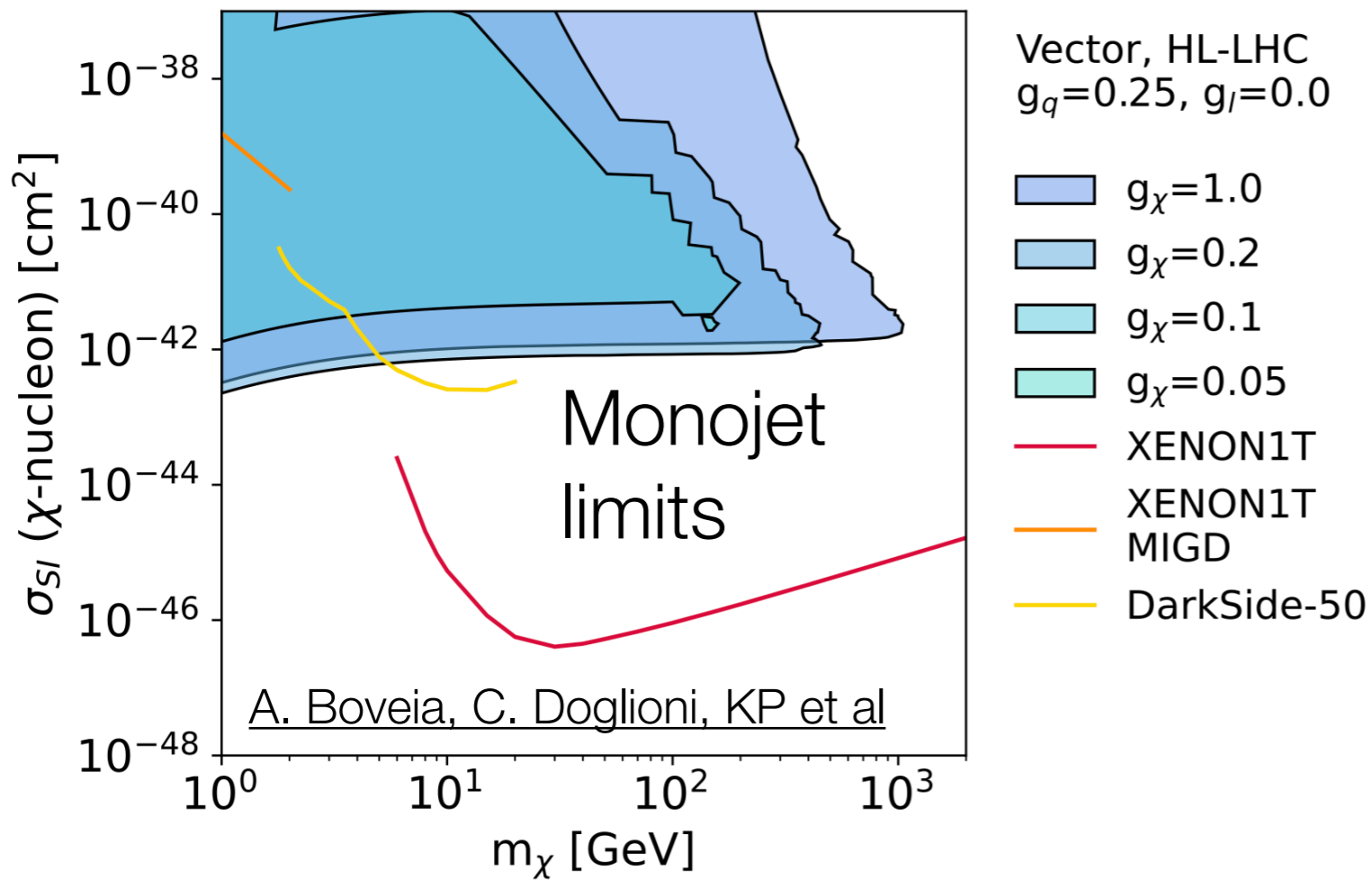
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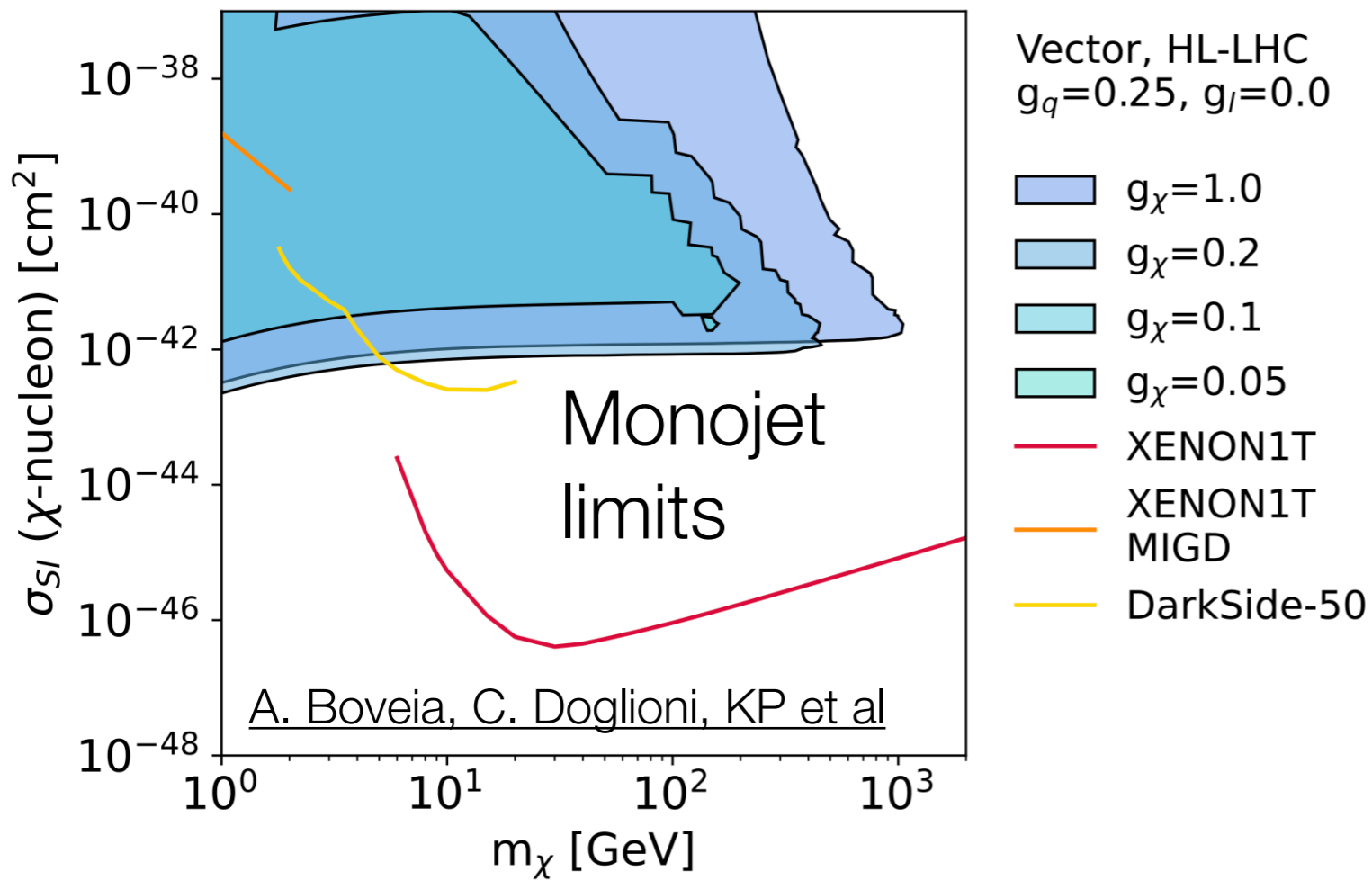


Show example I know best: LHC DMWG spin-1 simplified model

Must reduce 4-5 free parameters ($m_{\text{med}}, m_\chi, g_{SM}, g_\chi$) to 2

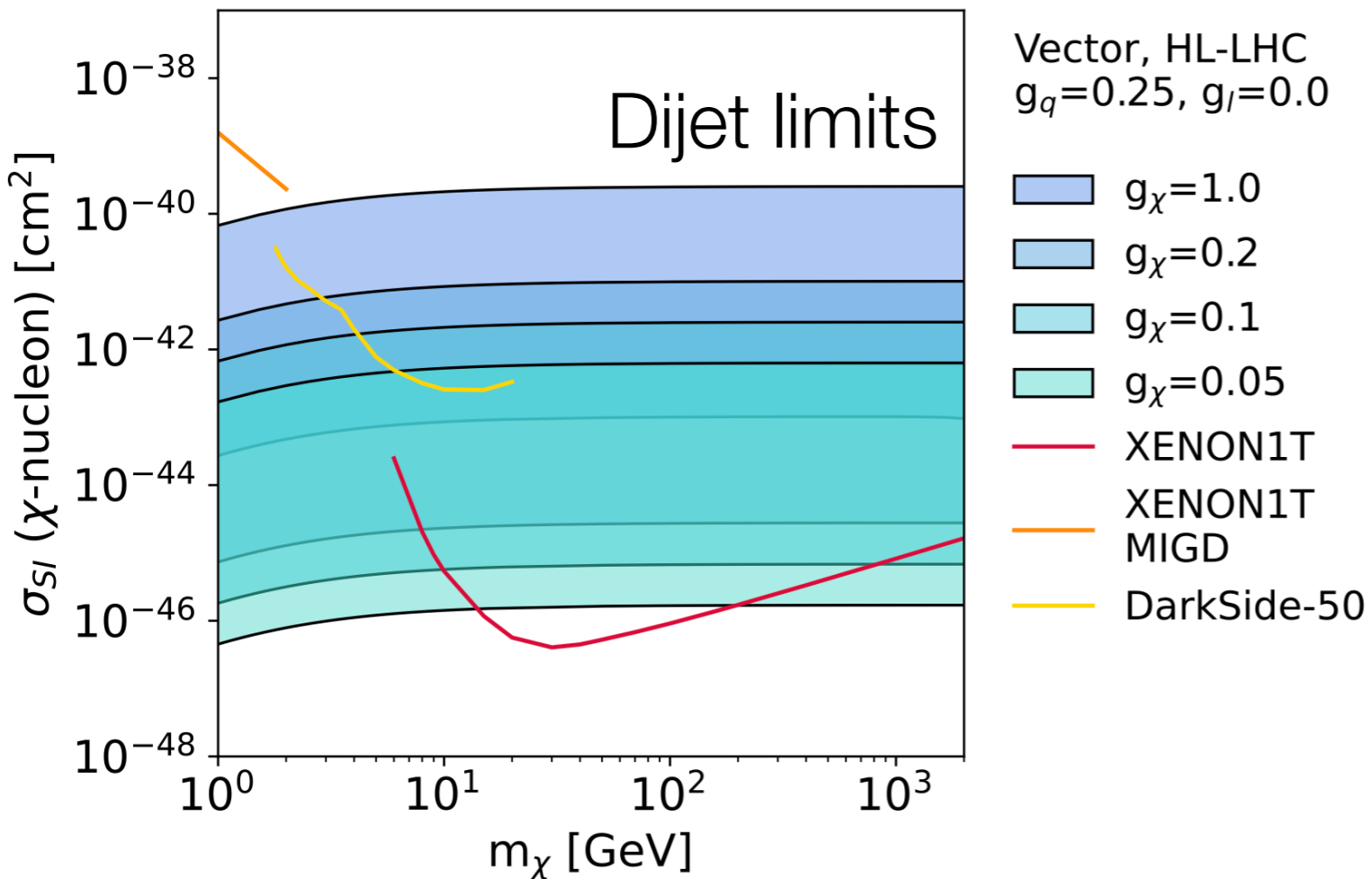


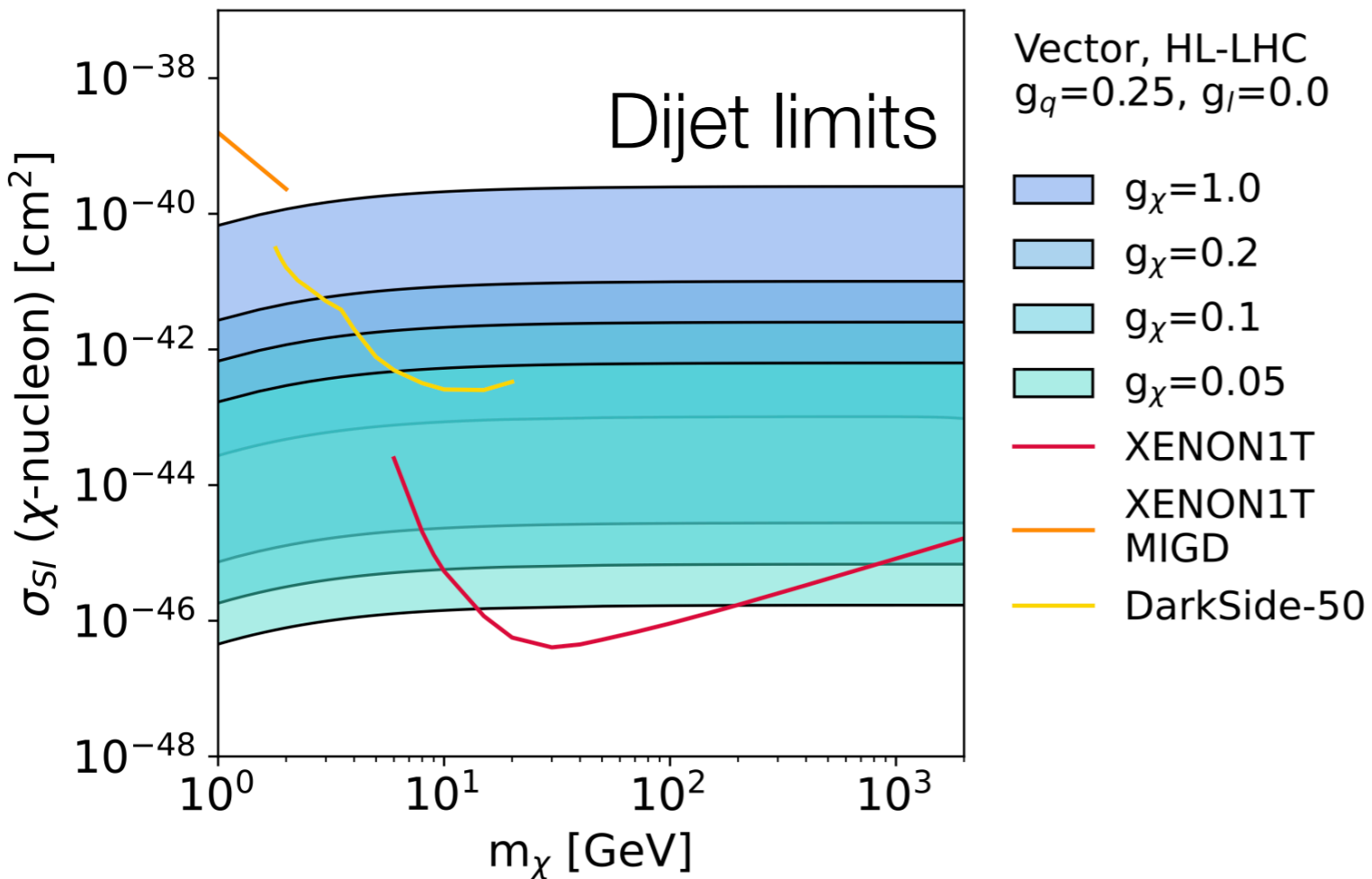
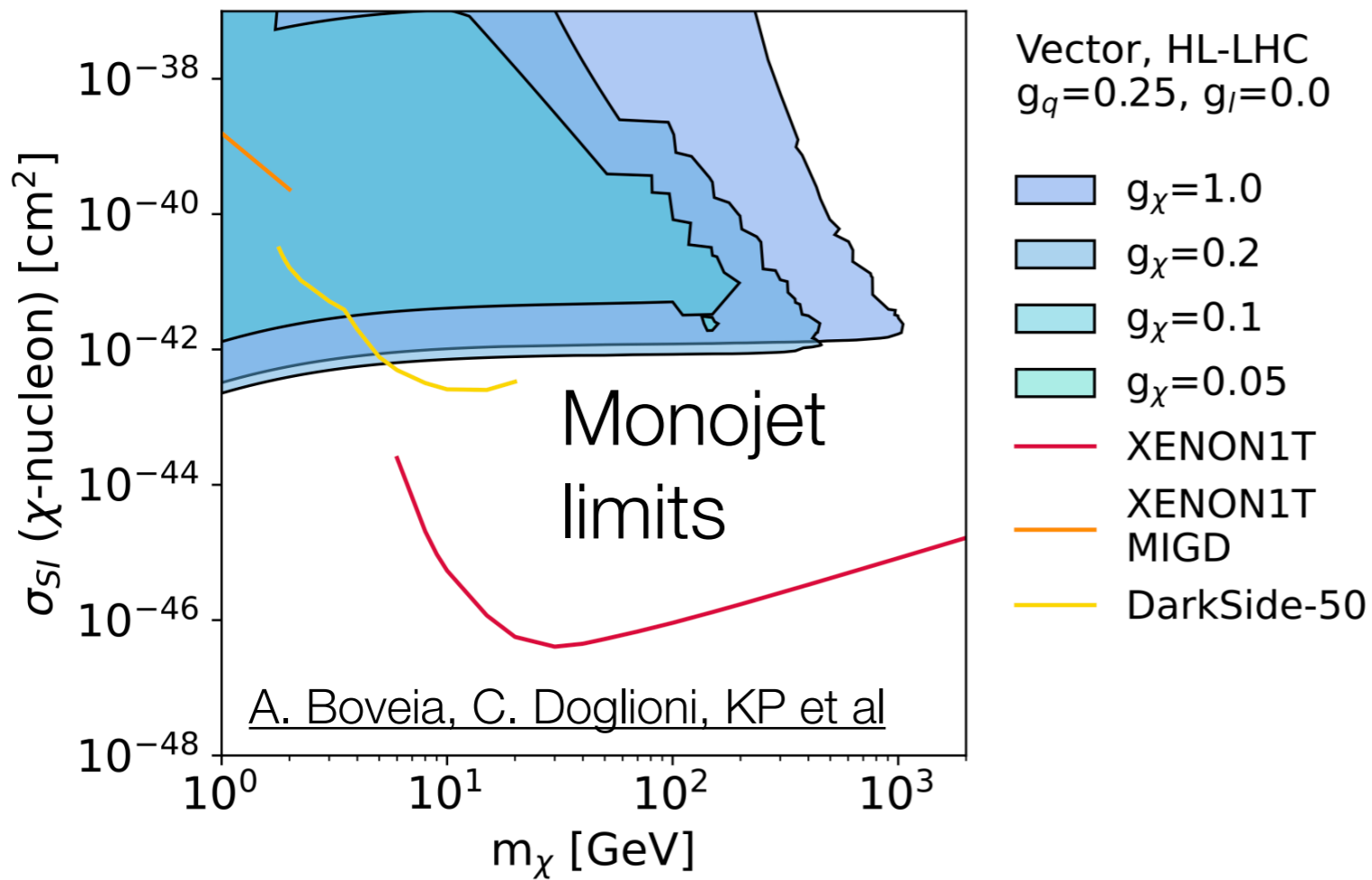
These are the type of projections we usually show from ATLAS and CMS



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Couplings take explicit values

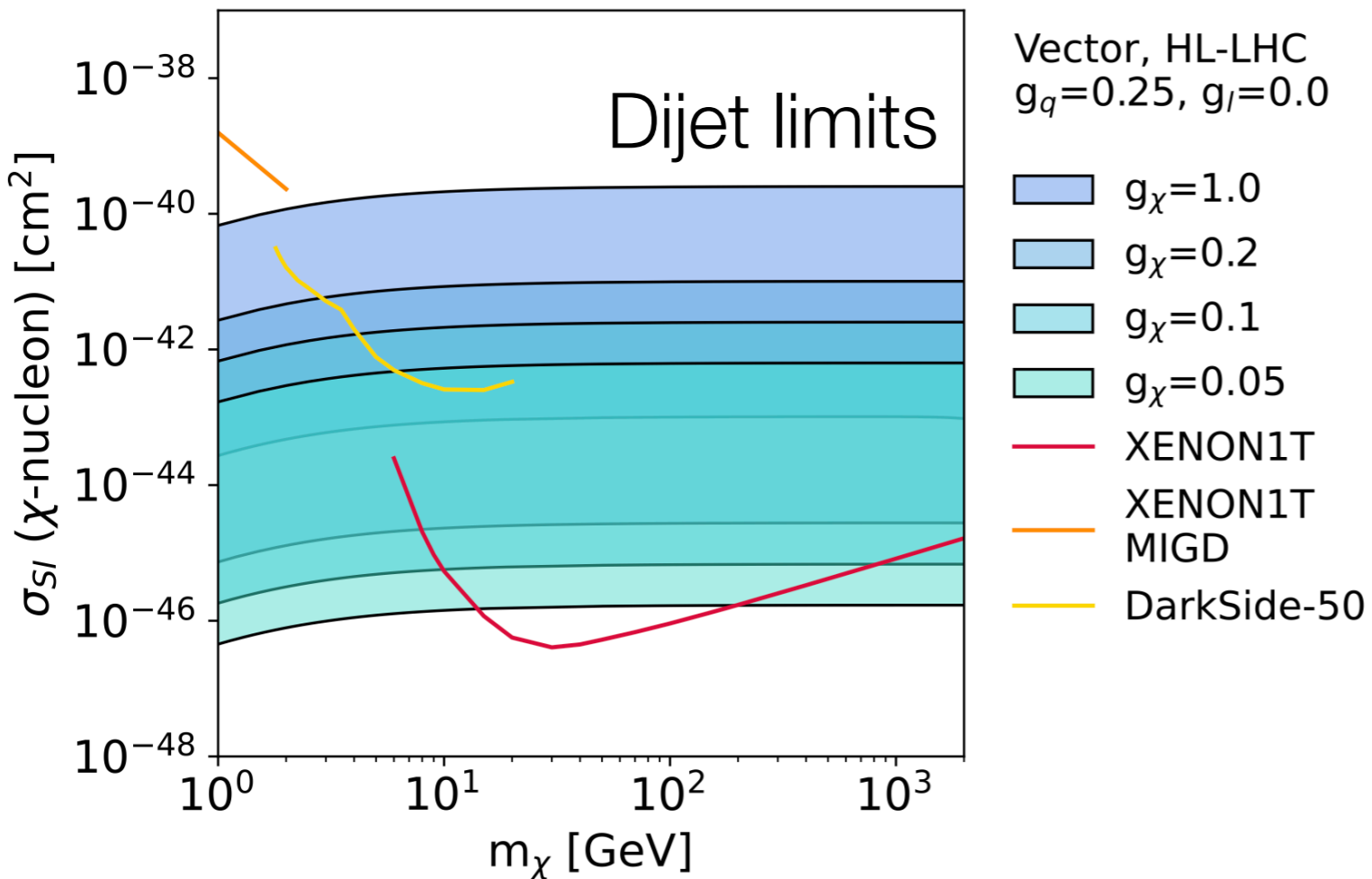
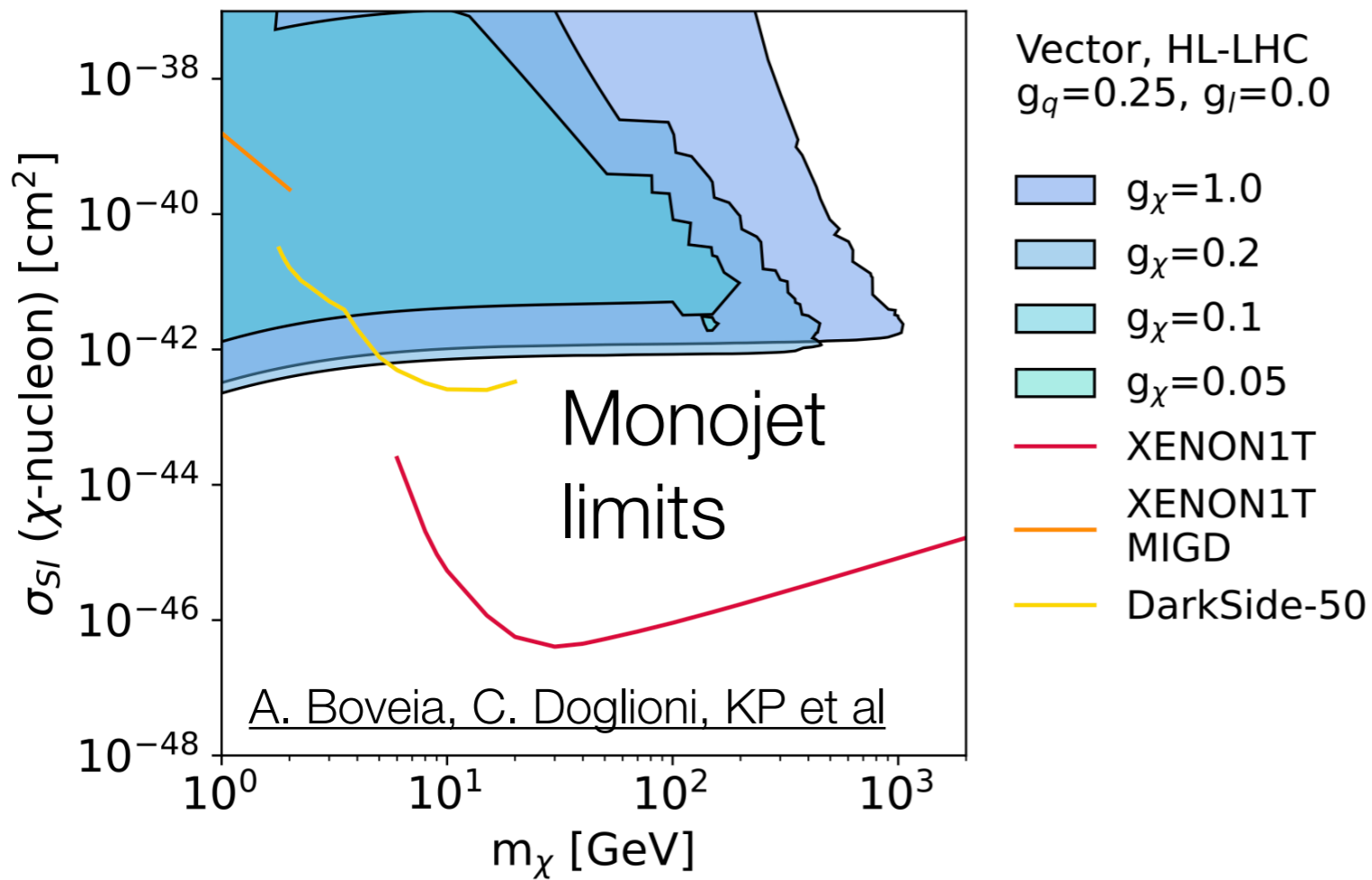




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Couplings take explicit values

Mediator mass absorbed into y axis variable

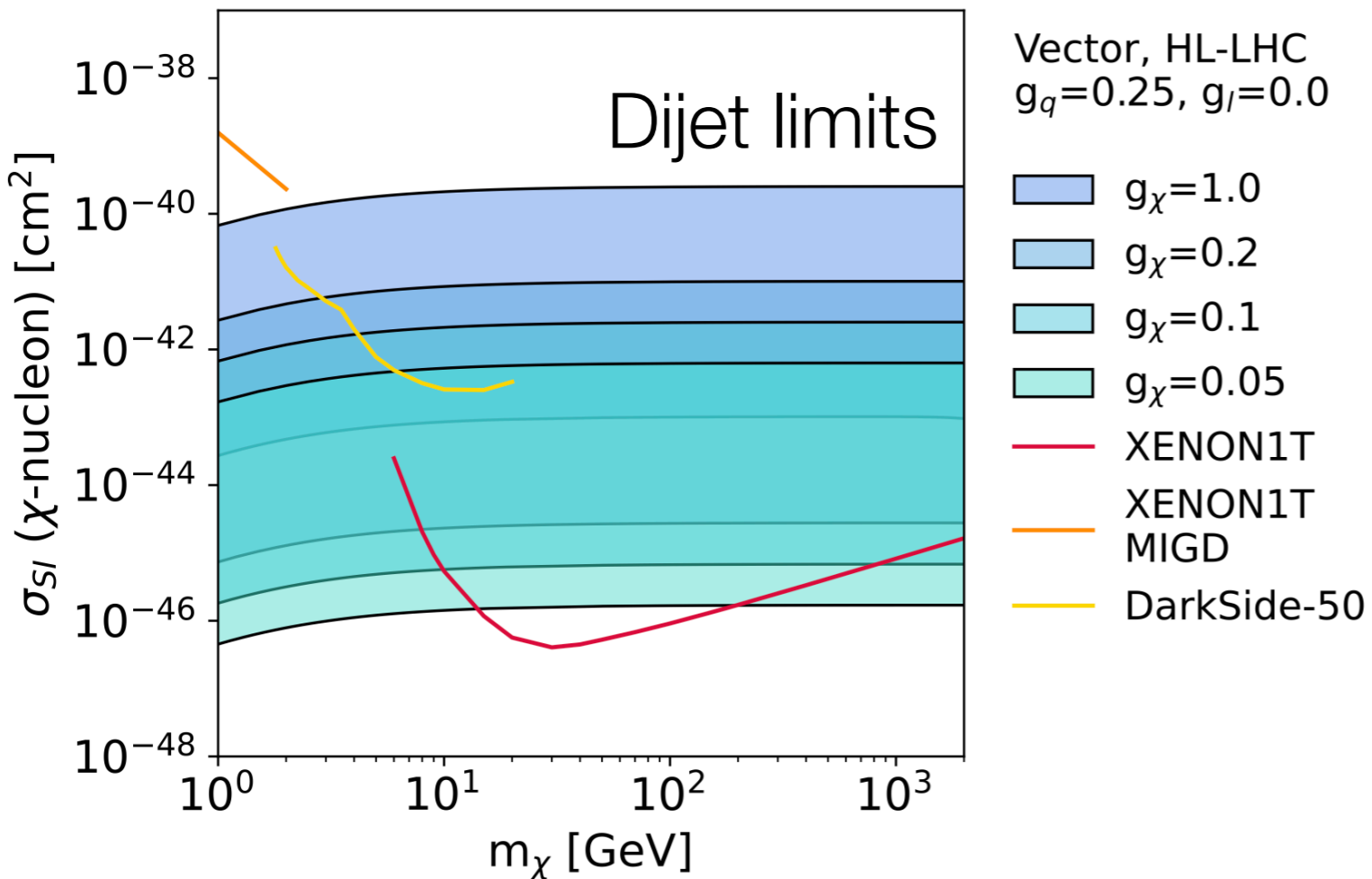
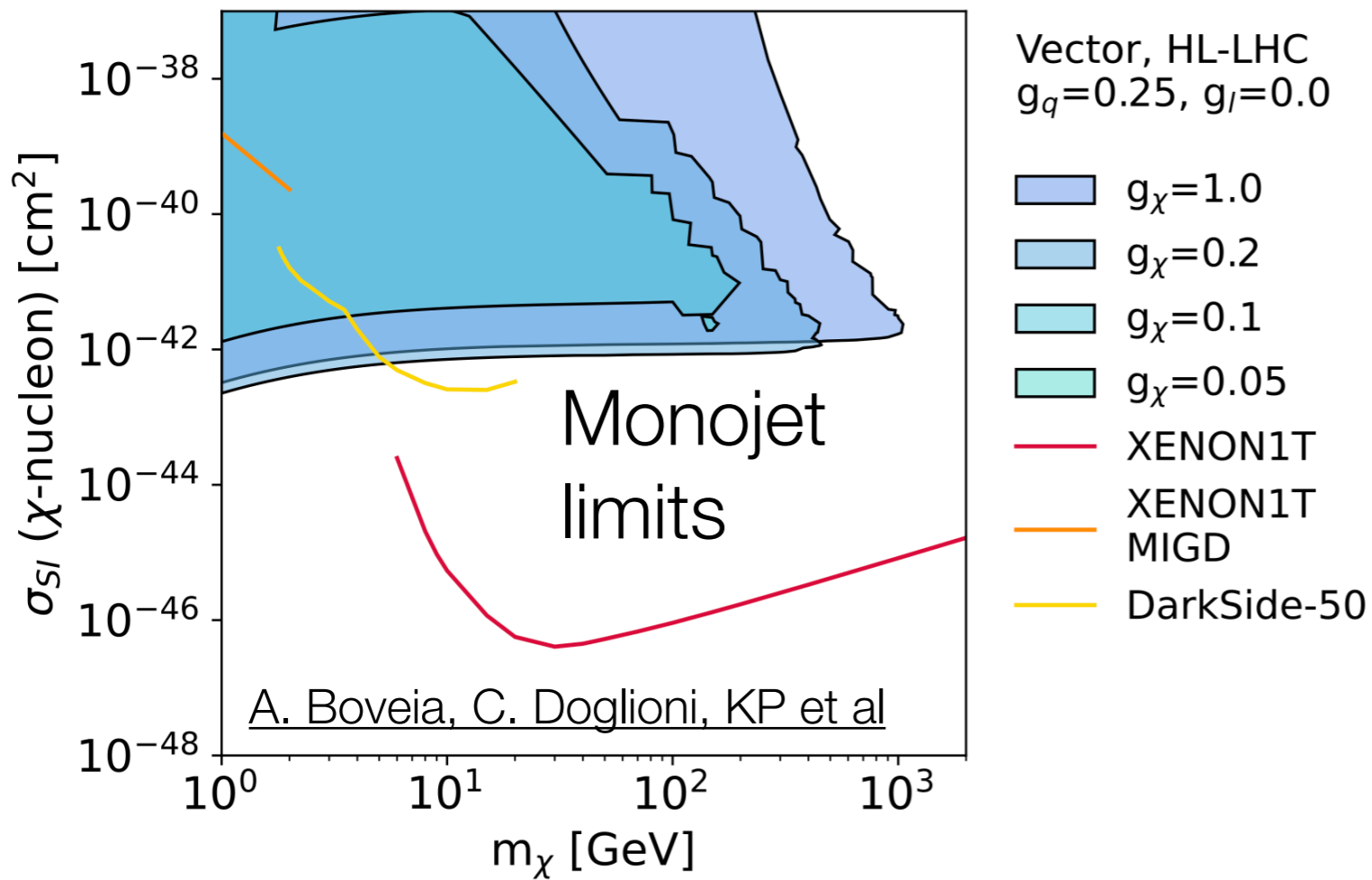


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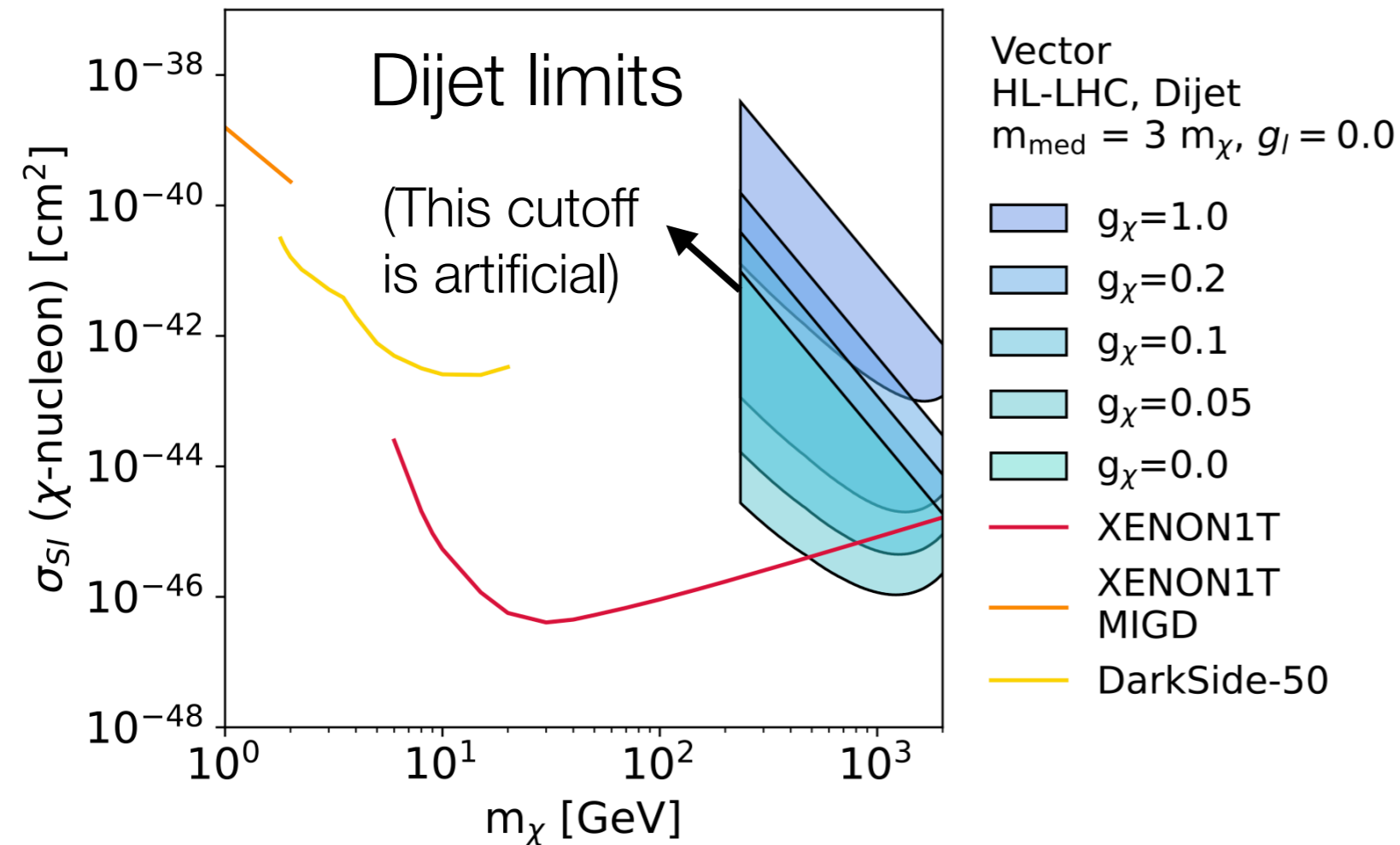
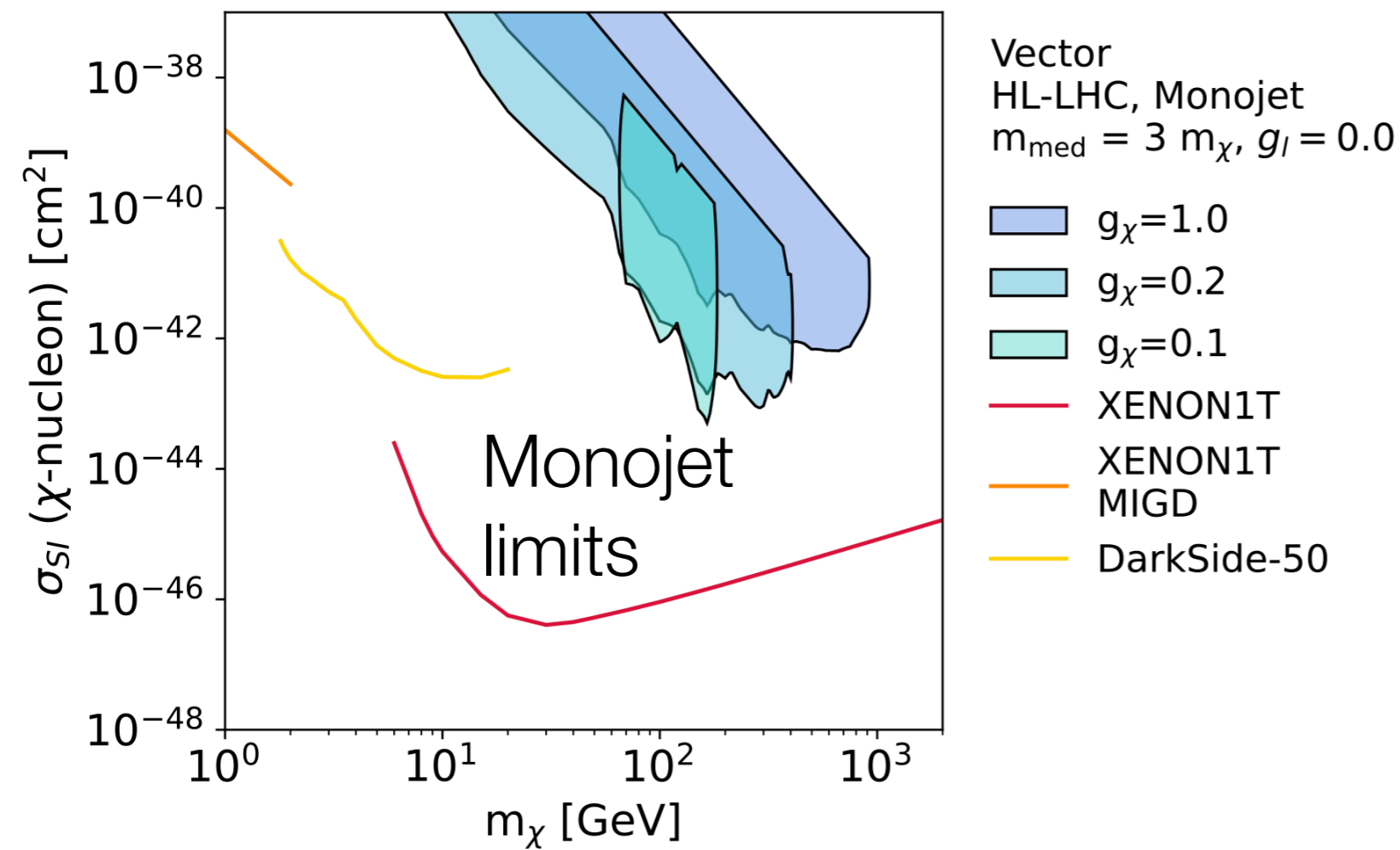
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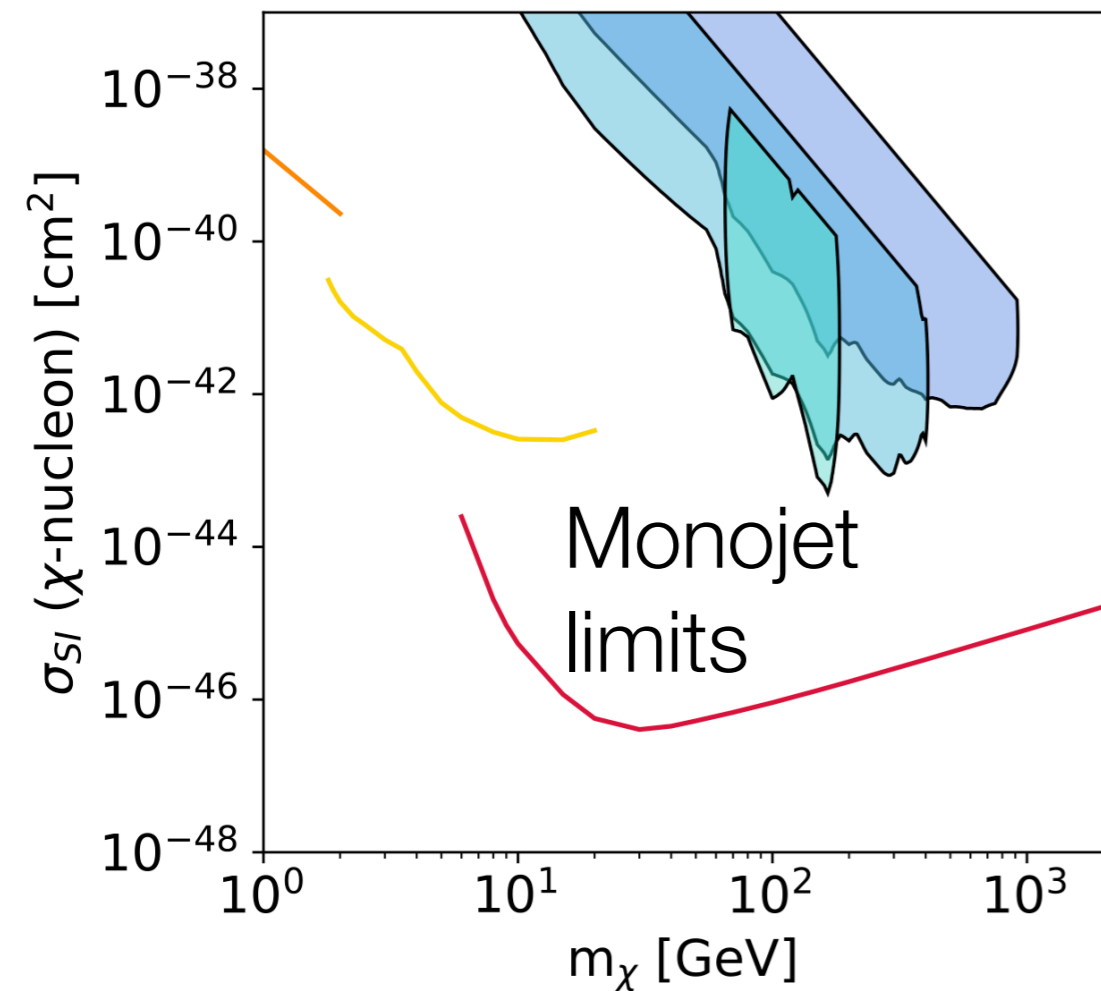
Mediator mass absorbed into y axis variable

Implication: no constraint on mediator mass

Points with strong collider limits have high mediator mass to DM mass ratio

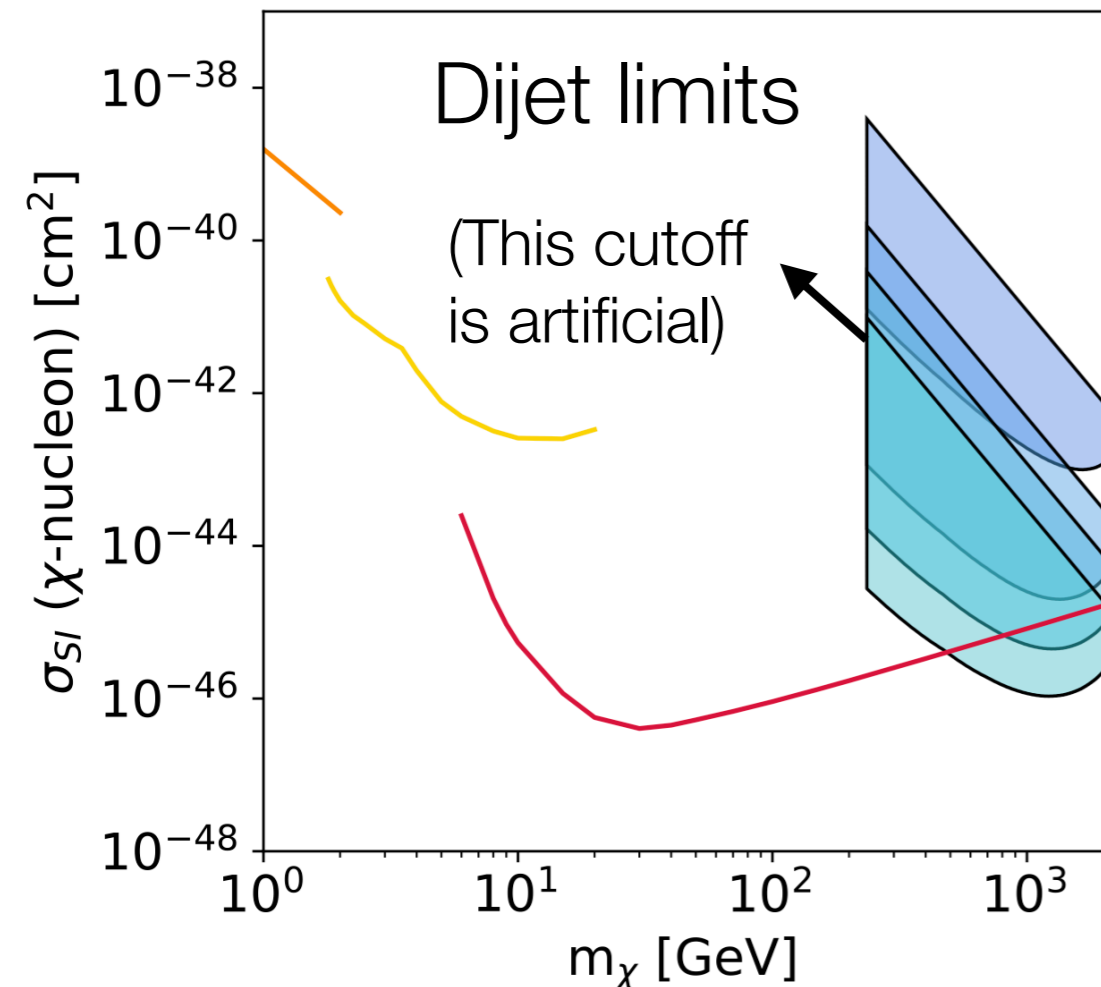
Same concept,
different projection
into two dimensions

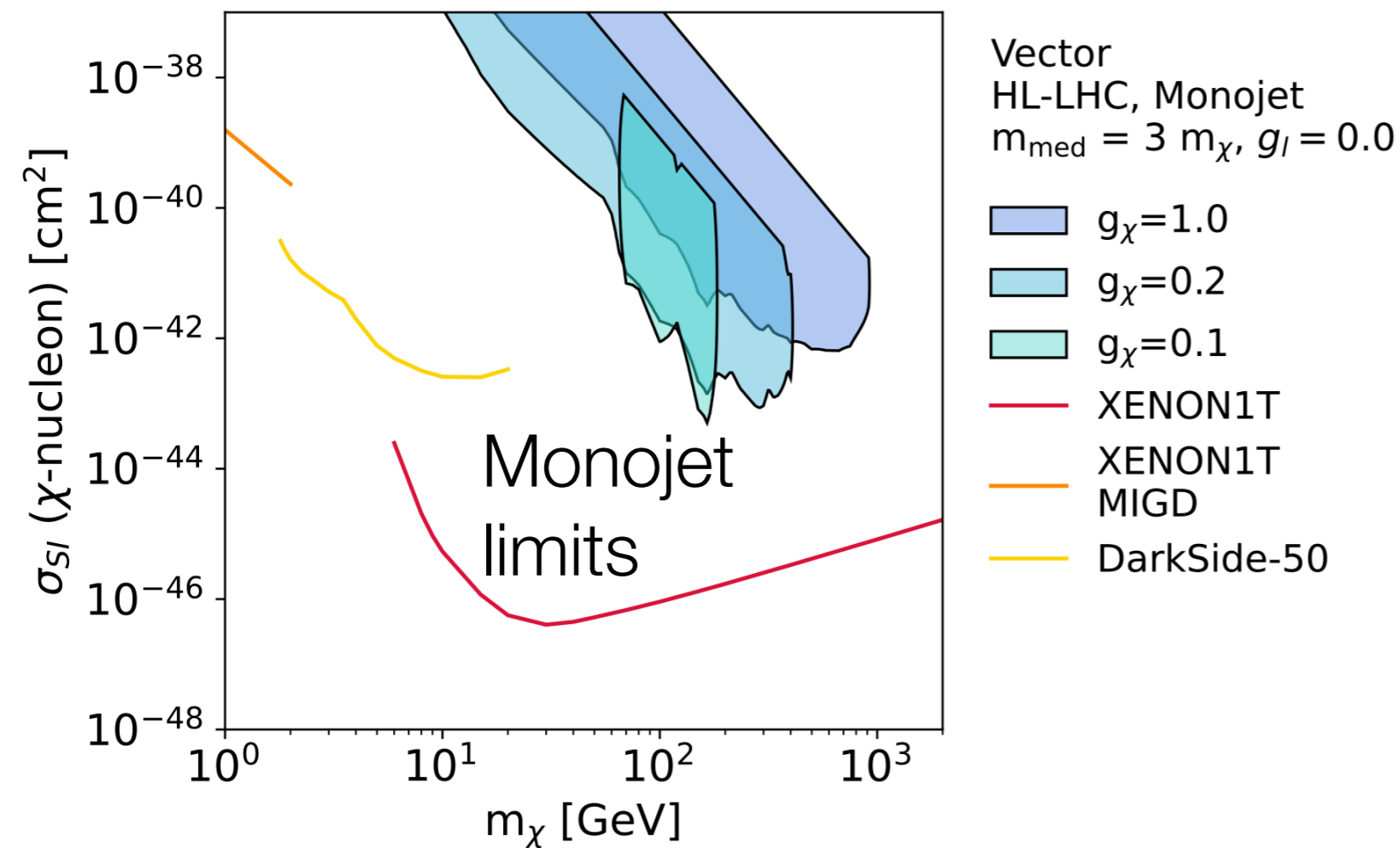




Same concept,
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Now ratio between
mediators is fixed and g_q
is absorbed into y axis

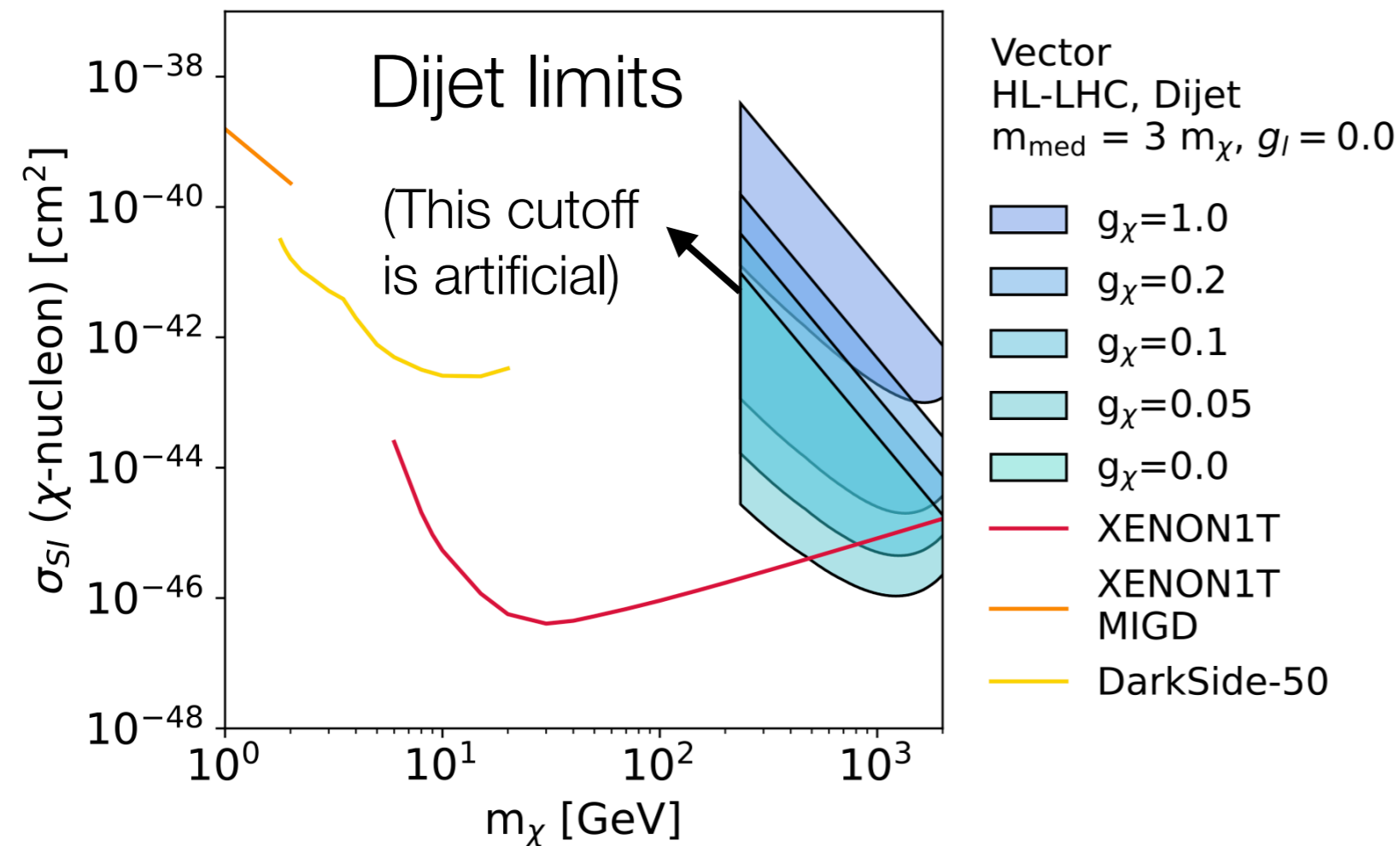


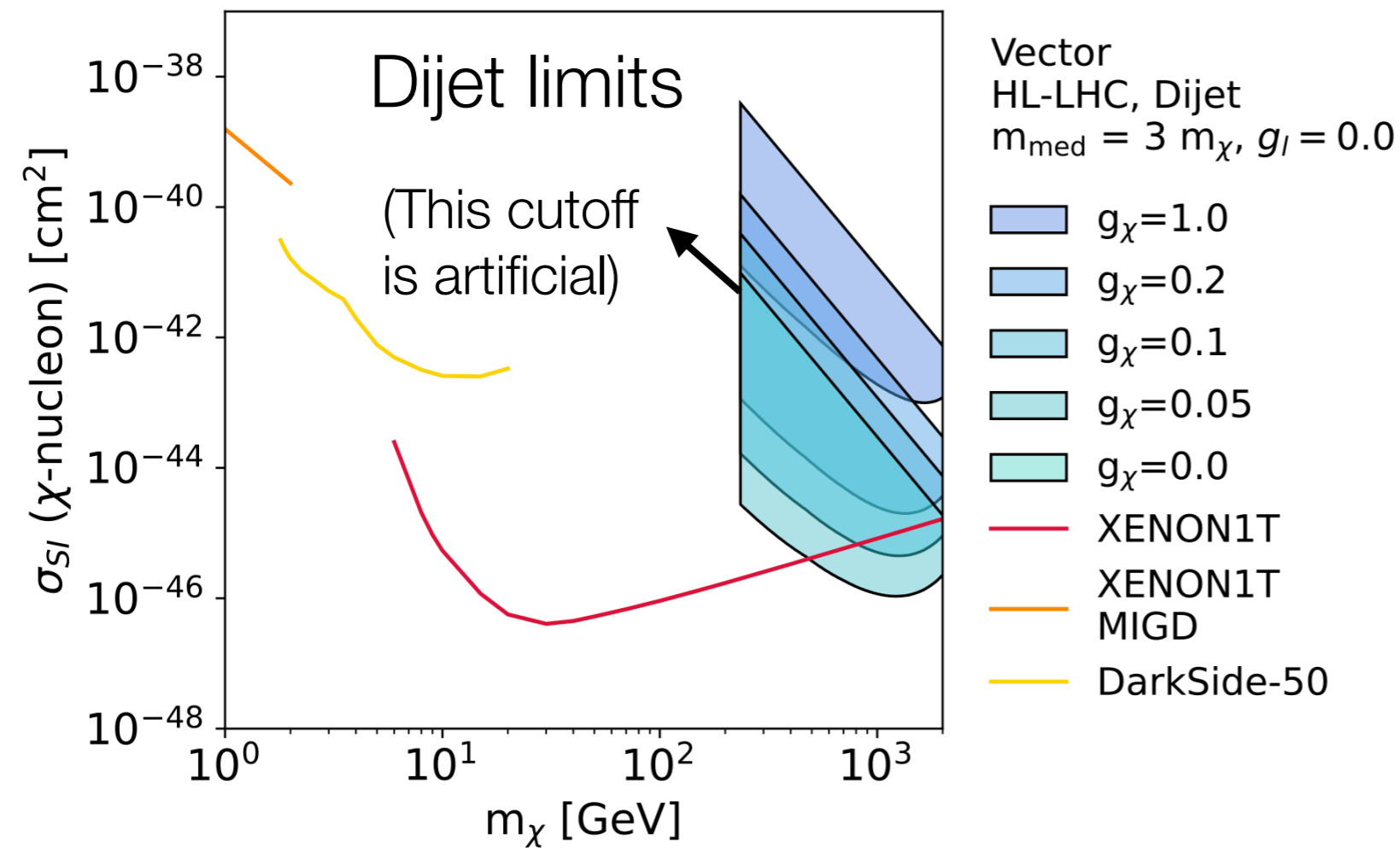
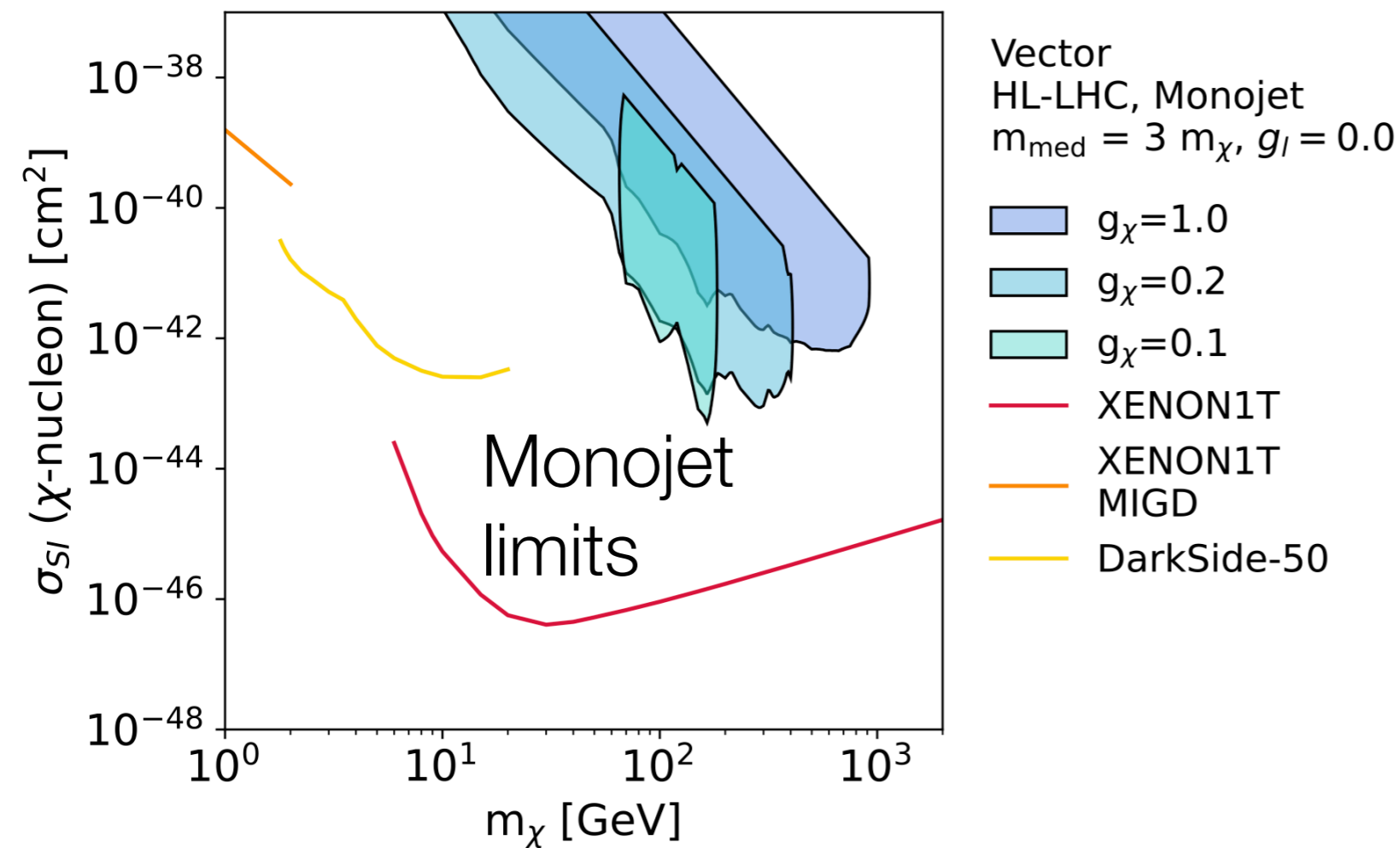


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Now ratio between
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Colliders have unique
strengths in accessing
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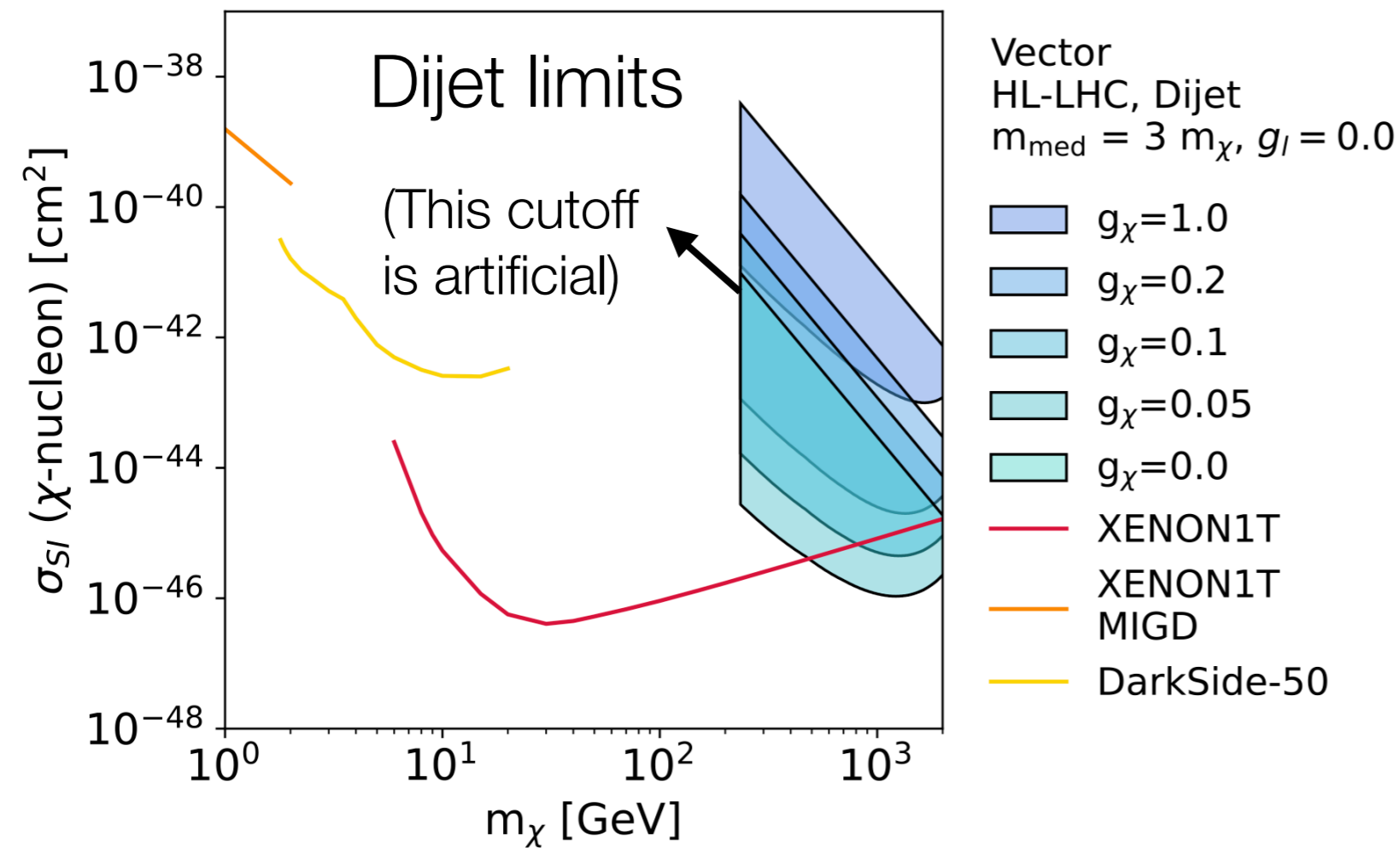
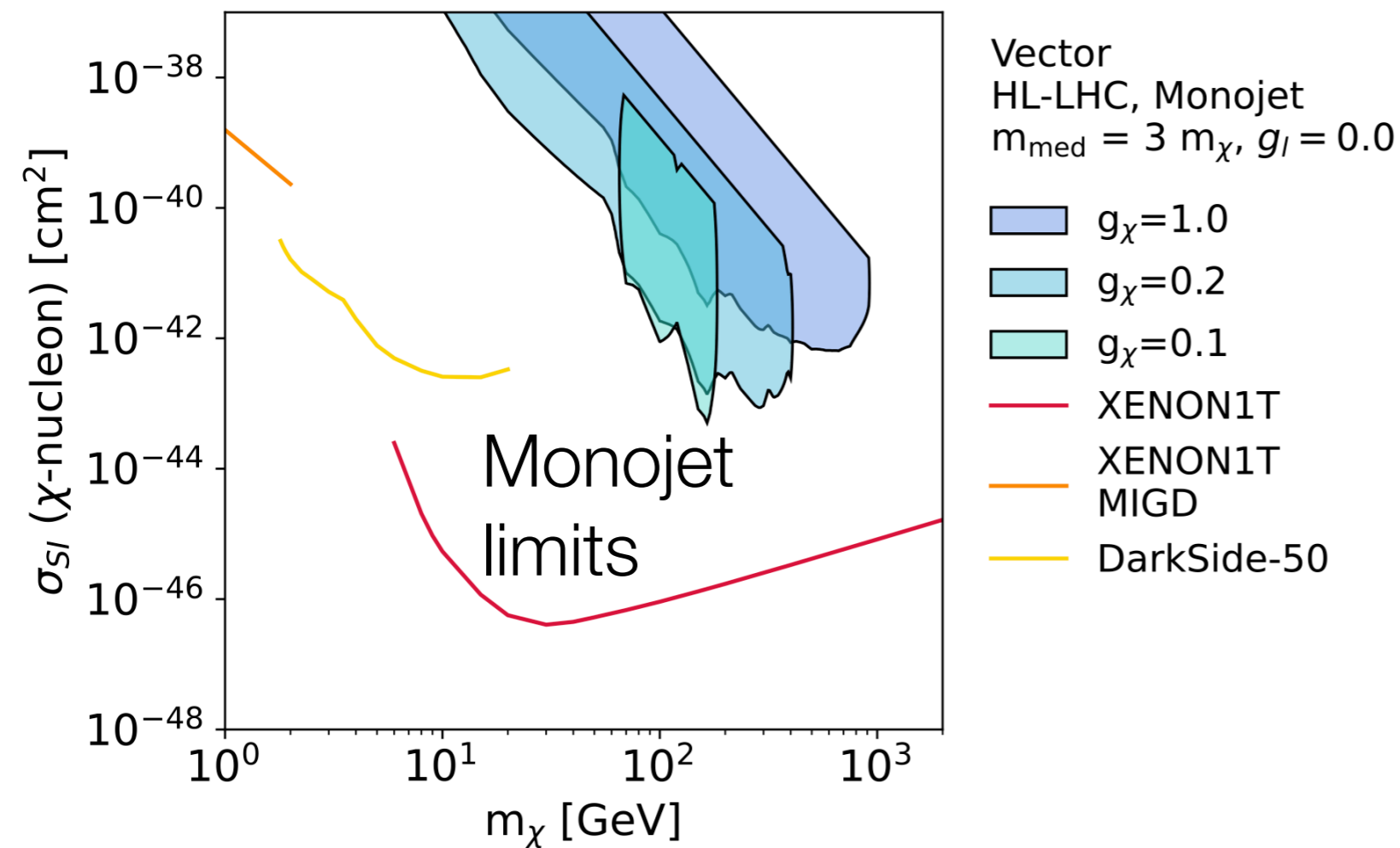


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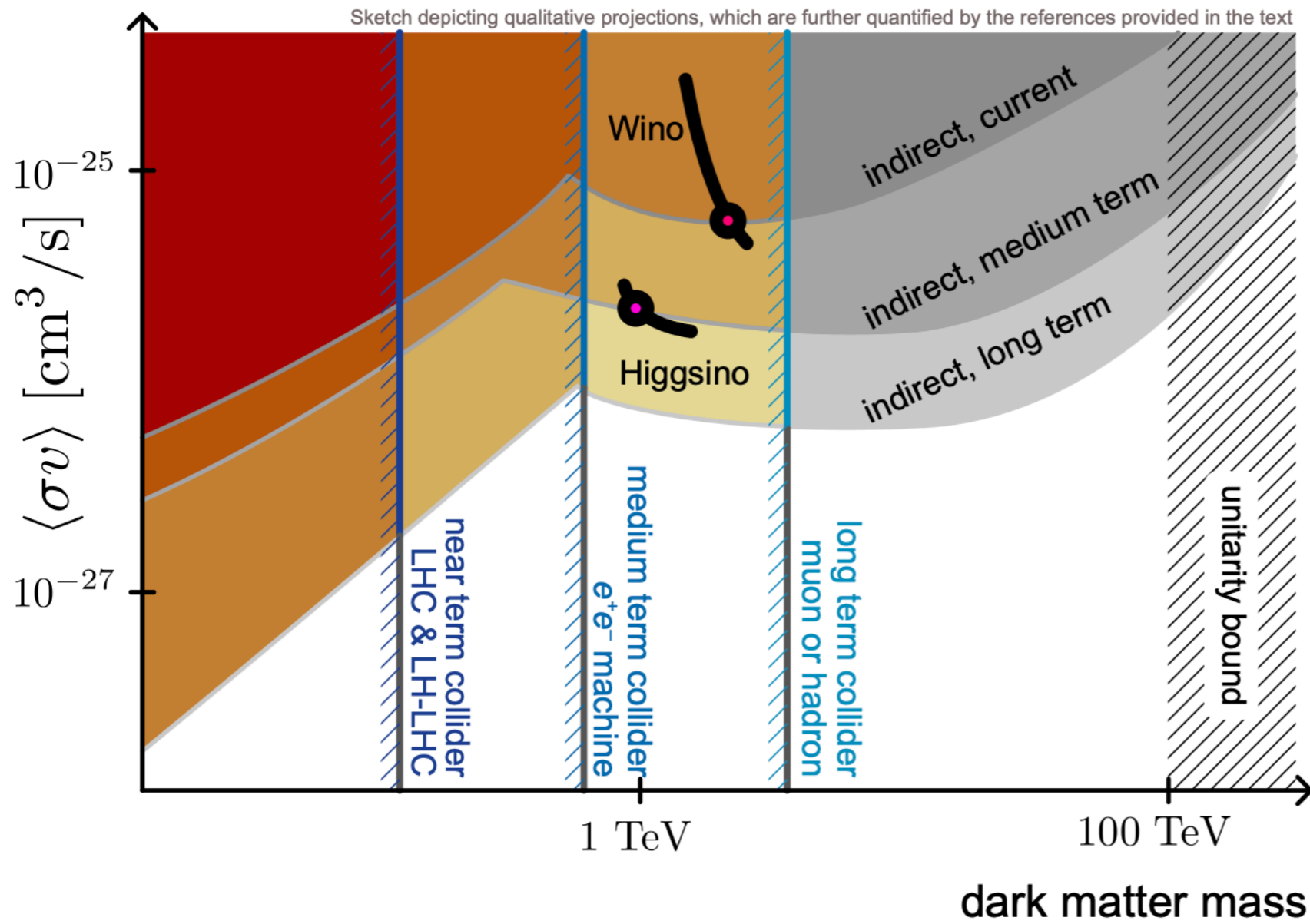
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Must present both for
complete picture

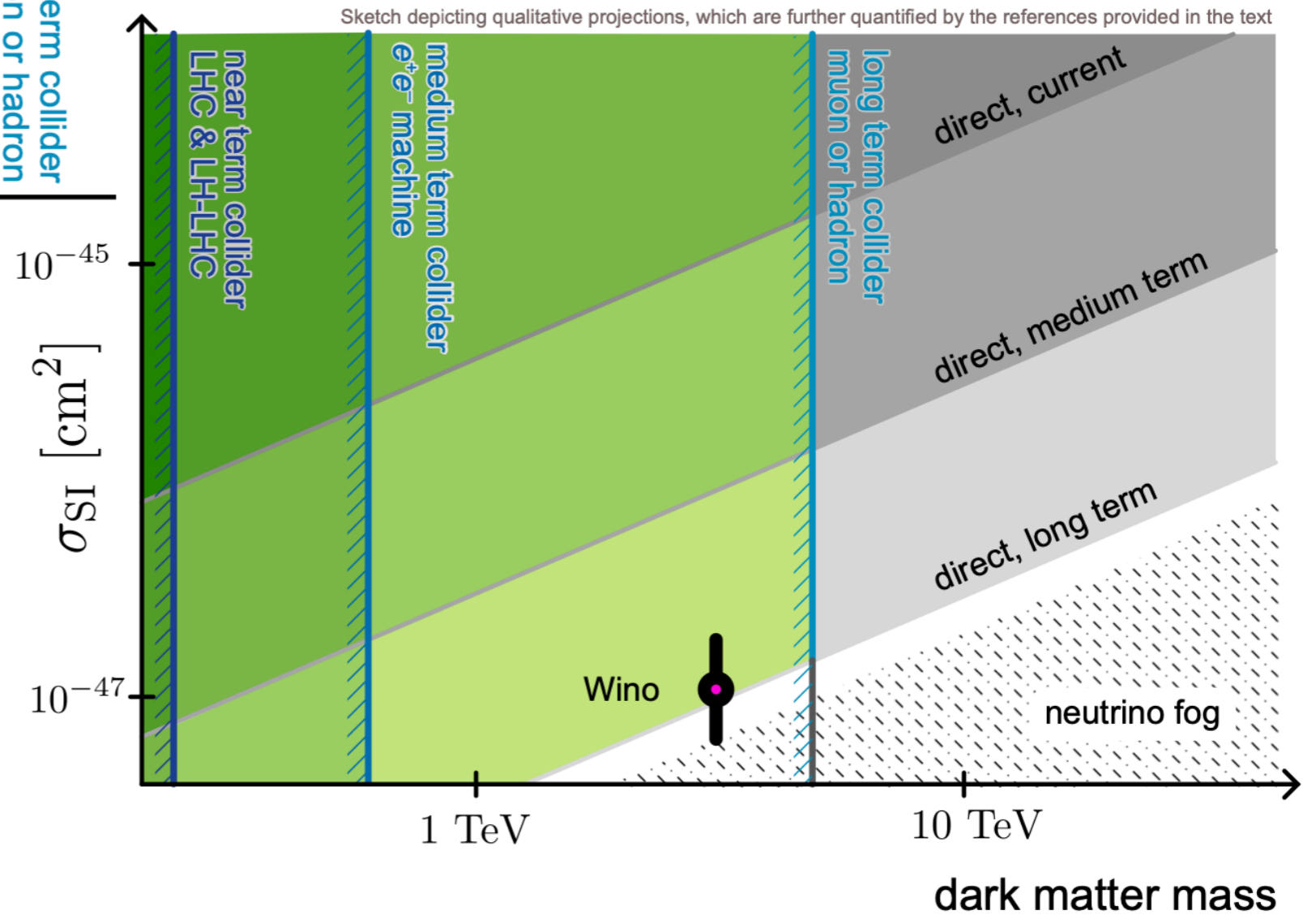
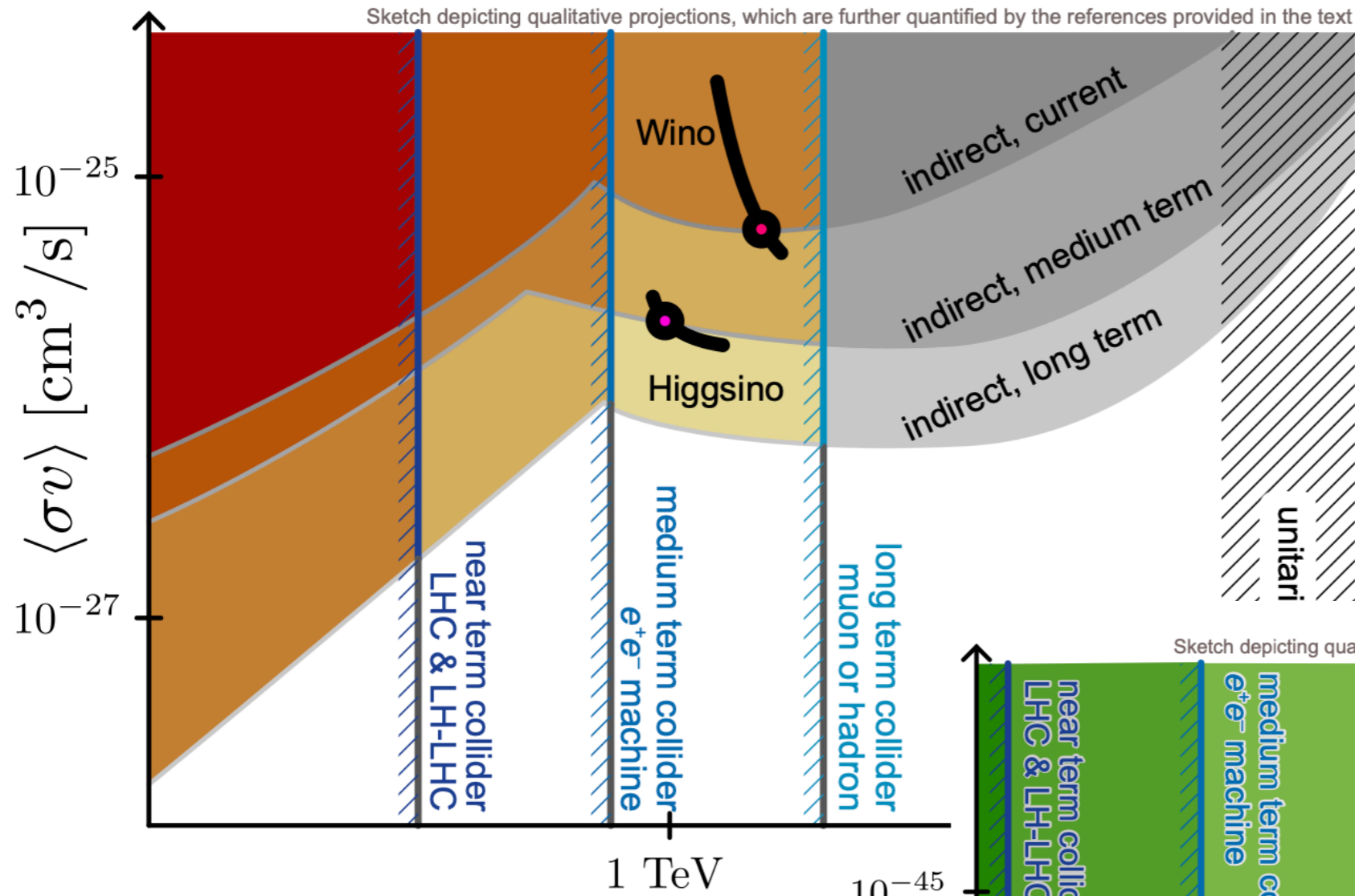


A few sketches from
Snowmass dark matter
complementarity report

[arXiv:2210.01770](https://arxiv.org/abs/2210.01770)

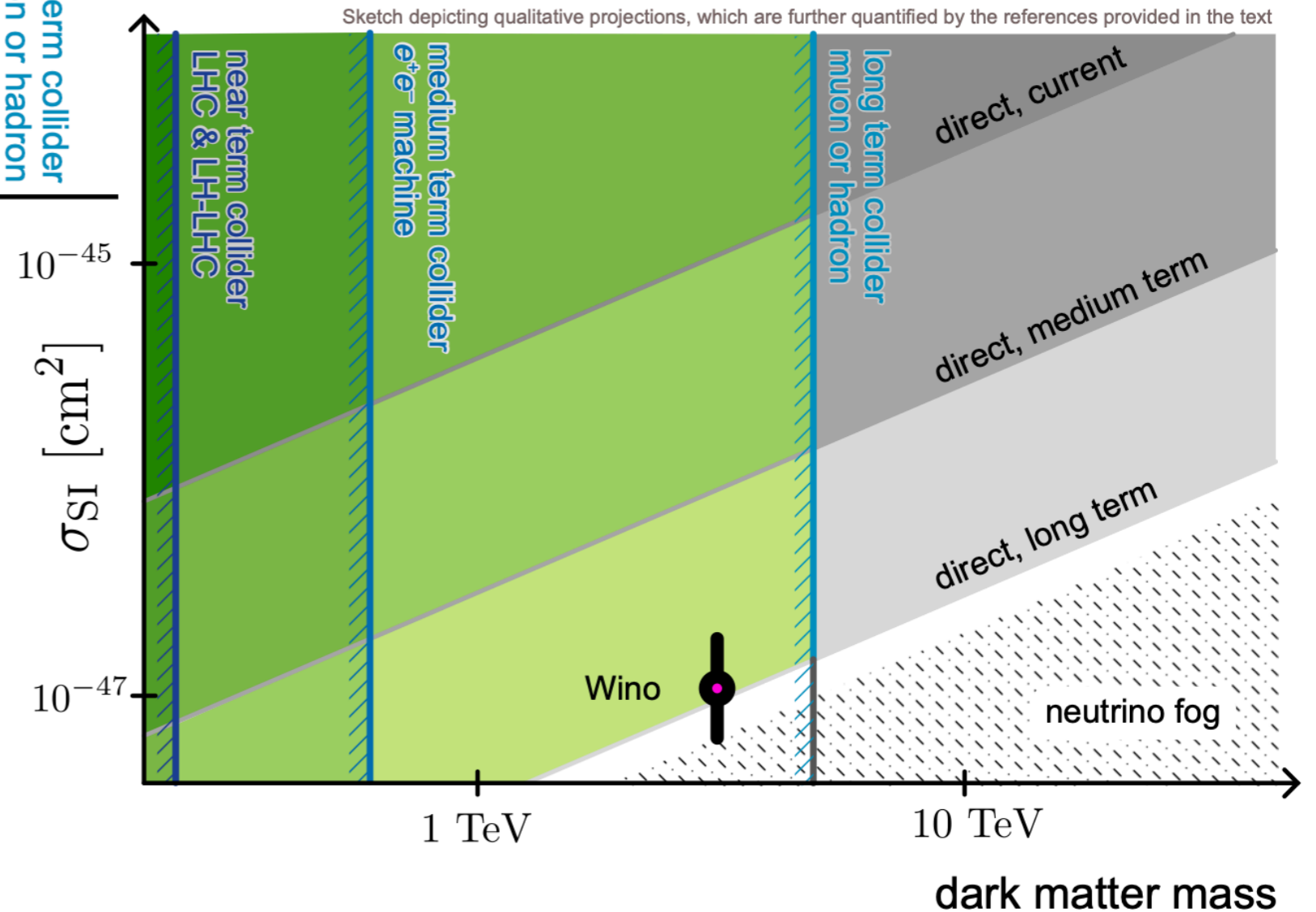
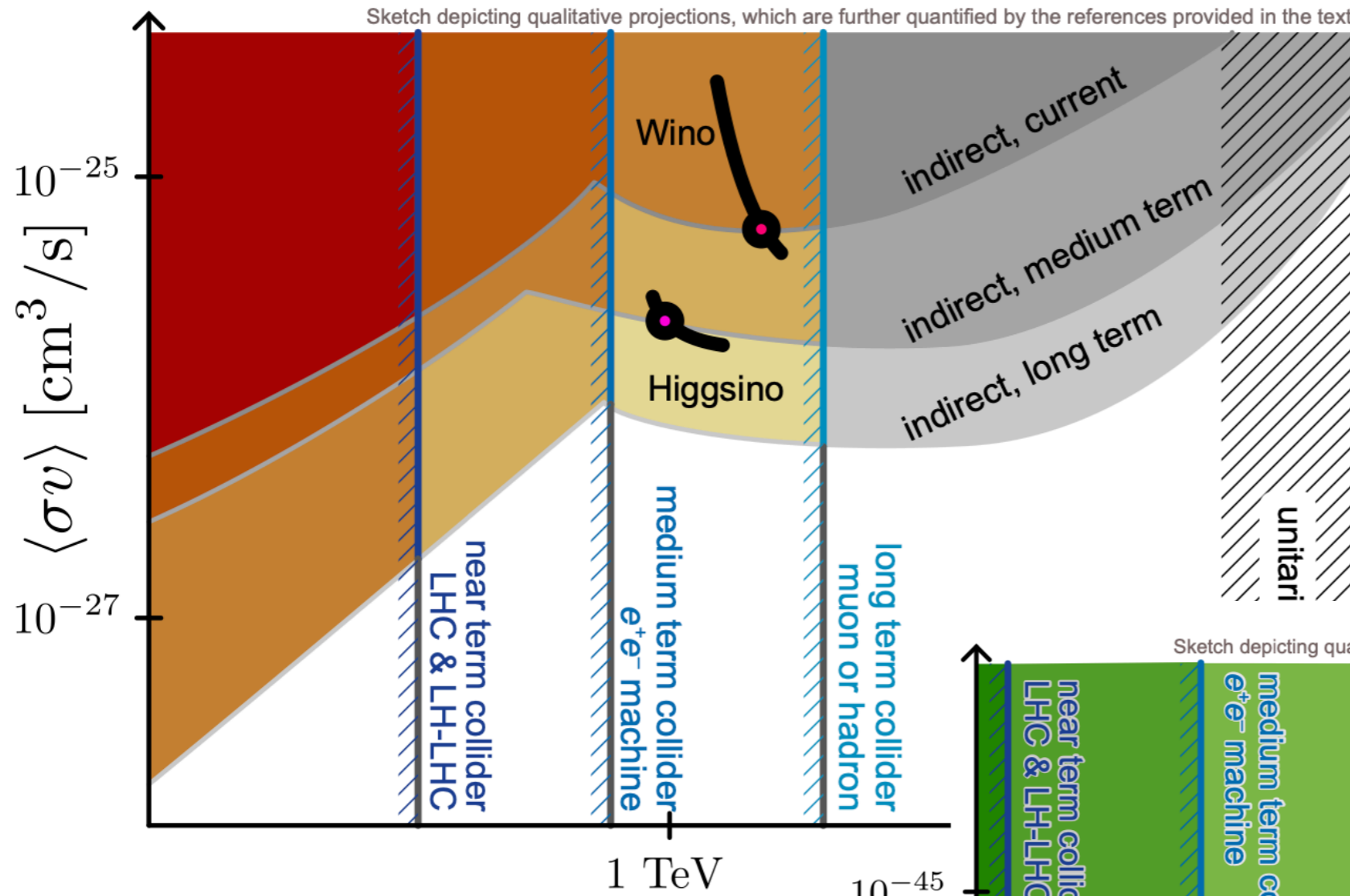
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Wino & Higgsino DM
candidate sensitivity vs
mass for indirect and
direct detection and
future colliders

dark matter mass

A hand-drawn diagram on a grid background. The diagram consists of several intersecting lines, including a prominent vertical line on the left and several diagonal lines. There are several small circles, some of which are spirals, and several small crosses scattered throughout the drawing. A semi-transparent black box is overlaid on the left side of the image, containing the word "Conclusion" in white text.

Conclusion

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Conclusion

Dark matter searches at colliders are complicated, take many forms, and are still not fully explored

We rely on theory community to help us guide this work

There remains plenty of non-excluded space for cosmologically motivated particle dark matter above the \sim GeV scale

There are also areas of DM phase space that only colliders can probe, just as there are areas that only direct or indirect detection experiments can probe

Complementarity, DM discovery potential, and the potential to exclude values aligning with cosmological observations should be thoroughly understood and included in future collider proposals

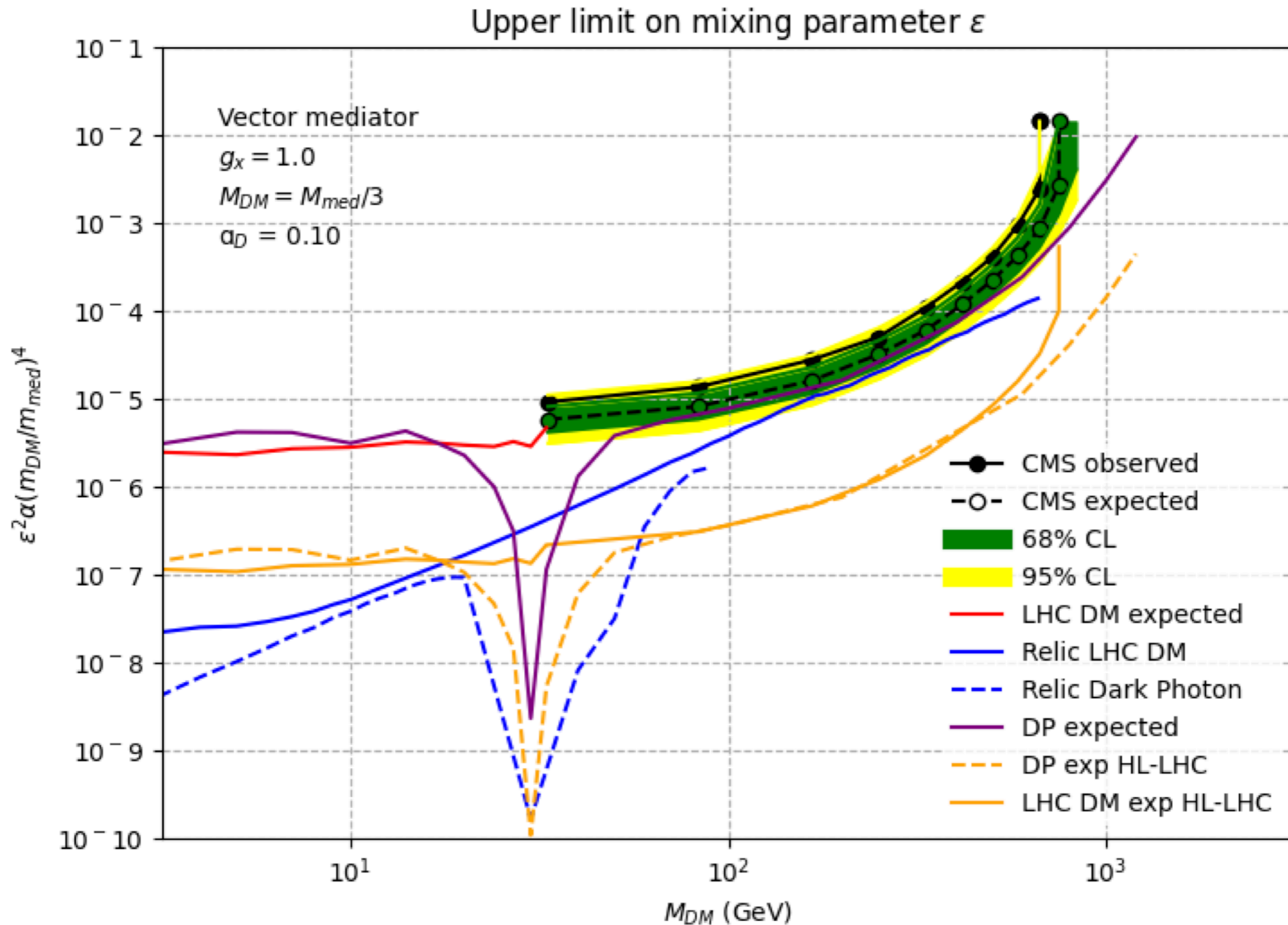


Additional materials

References

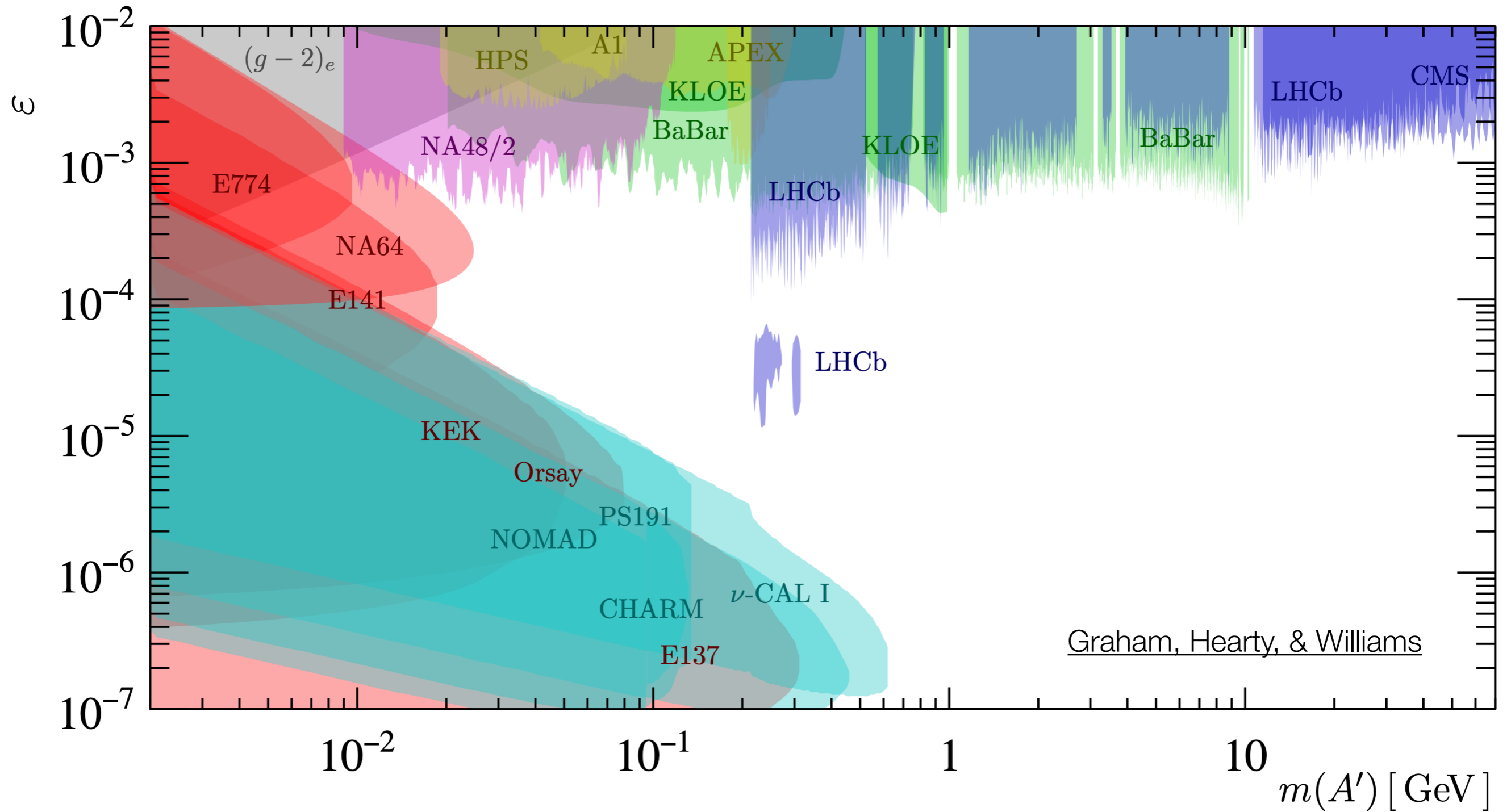
- LHC simplified models (s-channel mediators) [arXiv:1507.00966](#)
- LHC 2HDM+a model: [arXiv:1810.09420](#)
- Notes on Higgs portal: [arXiv:2001.10750](#), [arXiv:1903.03616](#)
- Snowmass BSM topical group report [arXiv:2209.13128](#)
- Snowmass particle dark matter topical group report [arXiv:2209.07426](#)
- Snowmass DM complementarity report: [arXiv:2210.01770](#)
- Spin-1 projection comparisons for HL-LHC and FCC [arXiv:2206.03456](#)
- European Strategy briefing document: [cds link](#)

Comparison between true dark photon model and LHC simplified Z' mediator model, demonstrating good agreement above Z peak



A. Boveia, C. Doglioni, P. Harris, KP, et al

Current limits on visible dark photon decays, by experiment



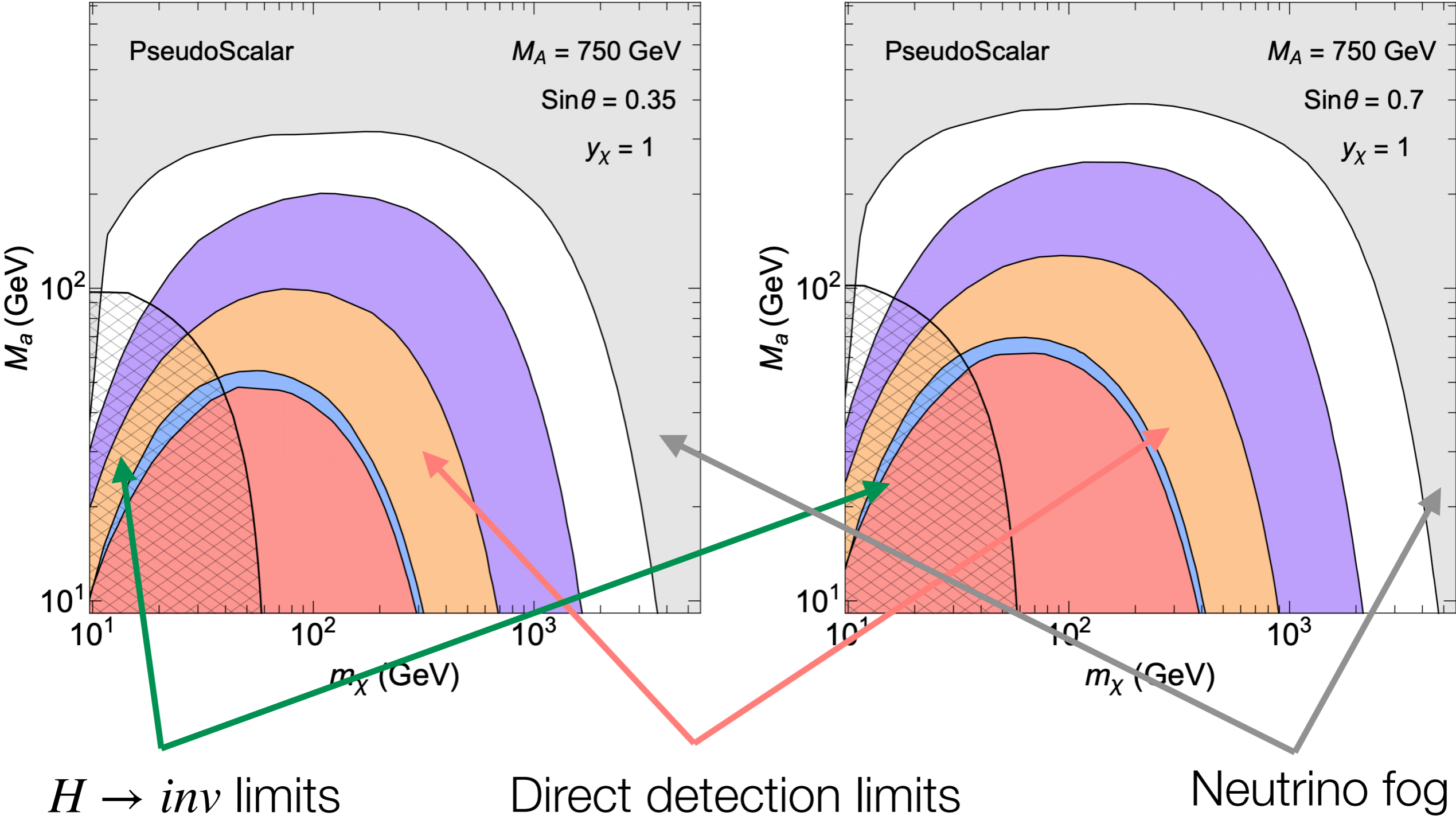
2HDM+a model and parameter choice description

The model considered here is the 2HDM+a model suggested by the LHC DM Working Group, which is the simplest gauge-invariant and renormalizable ultraviolet completion of the simplified pseudoscalar model initially recommended by the LHC DM Forum, which only contained the DM candidate and the mediator. This model is a type-II two-Higgs-doublet (2HDM) model to which an additional pseudoscalar a and a fermionic DM candidate χ are added. After electroweak symmetry breaking, the 2HDM contains five Higgs bosons: a lighter CP-even boson, h , a heavier CP-even boson, H , a CP-odd boson, A , and two charged bosons, H^\pm . While the phenomenology of the model would be determined by 14 free parameters, some benchmark choices are made in order to match h with the observed SM Higgs boson, to ensure the stability of the Higgs potential, or to evade electroweak precision measurement constraints. In the end, the benchmarks are defined by five parameters: the mass of the heavy Higgs bosons, which are taken to be degenerate, $m_A = m_H = m_{H^\pm}$; the mass of the pseudoscalar mediator, m_a ; the mass of the DM particle, m_χ ; the mixing angle θ between the two CP-odd states a and A ; and the ratio of the vacuum expectation values of the two Higgs doublets, $\tan \beta$.

ATLAS EXOT-2023-14

Shape of direct detection exclusions in 2HDM+a model, M_a vs m_χ plane. Requires fixing of other three parameters

LHC Dark Matter Working Group



How spin-1 simplified model to DD plane conversion works

For details, see [this talk](#)

$$\sigma_{SI} \simeq 6.9 \times 10^{-41} \text{ cm}^2 \cdot \left(\frac{g_q g_{\text{DM}}}{0.25} \right)^2 \left(\frac{1 \text{ TeV}}{M_{\text{med}}} \right)^4 \left(\frac{\mu_{n\chi}}{1 \text{ GeV}} \right)^2$$

1 variable

3 variables

Fix two and the other one becomes the thing that changes as σ_{SI} changes.

Implications and consequences can be very different, but can also be somewhat opaque when just looking at final 2D plot.